

Thematic Interiorization Project

Compact Groups of Galaxies in the Multidark Cosmological Simulation

Sebastián Ortiz Gómez December 14th, 2023 Computational Tools for Astroinformatics

O1 Previously on...

The idea



Objectives:

- To comprehend basic aspects of working with simulations (hydrodynamics or SAM).
- Deepen and apply knowledge of Python in relation to what is required.
- To relate the analysis to what I do with my thesis.

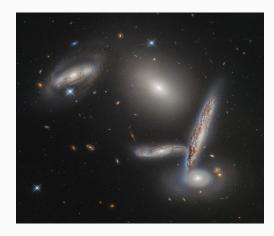
Galaxy-galaxy interaction

Physical and Morphological Changes

What is happening to the star-formation?

Candidates for galaxy pre-processing

How do the dark matter halos behave?



HCG 40, WFC3/HST



Procedure:

- To determine which simulation would let me explore and accomplish the objectives. \checkmark
- To analyze the tables and the information available, to establish the feasibility of applying observational criteria to classify compact groups of galaxies as such:

Richness and magnitude

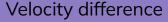
$$3 \le N(14.00 \le r \le 17.77) \le 10$$

Isolation

$$\theta_N \geq 3\theta_G$$

Compactness

$$\mu_r \le 26.0 \text{ mag arcsec}^{-2}$$



$$|V - V_i| \le 1000 \text{ km/s}$$

- Establish/apply a method that allows finding compact groups in these simulations, and visualizing their properties.

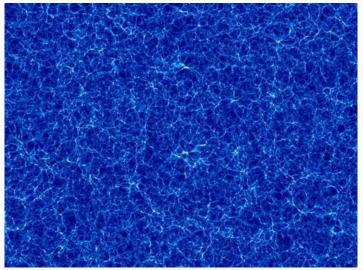


02

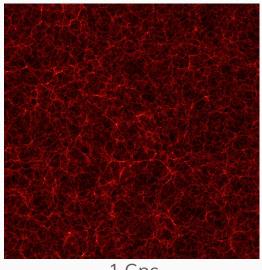
The Data

Which simulation?

Multidark-Galaxy Simulation



Box size	1 Gpc/h	side length of the cosmological cube
Number of particles	3840 ³	total number of dark matter particles
Mass resolution	1.51*10 ⁹ M _{sun} /h	mass of one dark matter particle



1 Gpc

https://www.cosmosim.org/cms/images-and-movies/



Datalinks for MDPL2_sag_snapdir

Data Links

access_url	description	content_length
https://www.cosmosim.org/data/catalogs/NewMD_3840_Planck1/SAG/snapshot_125/gal_125_SAG-7.128_BOX_001.hdf5	MDPL2: SAG - Snapshots	437 MB
https://www.cosmosim.org/data/catalogs/NewMD_3840_Planck1/SAG/snapshot_125/gal_125_SAG-7.128_BOX_002.hdf5	MDPL2: SAG - Snapshots	437 MB
https://www.cosmosim.org/data/catalogs/NewMD_3840_Planck1/SAG/snapshot_125/gal_125_SAG-7.128_BOX_003.hdf5	MDPL2: SAG - Snapshots	437 MB
https://www.cosmosim.org/data/catalogs/NewMD_3840_Planck1/SAG/snapshot_125/gal_125_SAG-7.128_BOX_004.hdf5	MDPL2: SAG - Snapshots	437 MB
https://www.cosmosim.org/data/catalogs/NewMD_3840_Planck1/SAG/snapshot_125/gal_125_SAG-7.128_BOX_005.hdf5	MDPL2: SAG - Snapshots	437 MB
https://www.cosmosim.org/data/catalogs/NewMD_3840_Planck1/SAG/snapshot_125/gal_125_SAG-7.128_BOX_006.hdf5	MDPL2: SAG - Snapshots	437 MB
https://www.cosmosim.org/data/catalogs/NewMD_3840_Planck1/SAG/snapshot_125/gal_125_SAG-7.128_BOX_007.hdf5	MDPL2: SAG - Snapshots	437 MB
https://www.cosmosim.org/data/catalogs/NewMD_3840_Planck1/SAG/snapshot_125/gal_125_SAG-7.128_BOX_008.hdf5	MDPL2: SAG - Snapshots	437 MB
https://www.cosmosim.org/data/catalogs/NewMD_3840_Planck1/SAG/snapshot_125/gal_125_SAG-7.128_BOX_009.hdf5	MDPL2: SAG - Snapshots	437 MB
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https://www.cosmosim.org/data/catalogs/NewMD_3840_Planck1/SAG/snapshot_125/gal_125_SAG-7.128_BOX_011.hdf5	MDPL2: SAG - Snapshots	437 MB



HDF5 files!



VX Vy Vz X Y

Multidark-Galaxy Simulation

Identifiers

Main Halo ID, Galaxy Type

Galaxy properties

X, Y, Z position; Vx, Vy, Vz velocities

Absolute magnitudes

SDSS r filter

Halo properties

R200 Main Halo

03 +

Pipeline



How was it done?



1) Hickson criteria to select Compact Groups

Richness and magnitude

$$3 \le N(14.00 \le r \le 17.77) \le 10$$

Isolation

 $\theta_N \ge 3\theta_G$

HCG 79 Legacy Survey DR10

Compactness

 $\mu_r \le 26.0 \text{ mag arcsec}^{-2}$

Velocity difference

 $|V - V_i| \le 1000 \text{ km/s}$





2) Adaptability to data from simulation

Richness and magnitude

 $3 \le N(14.00 \le r \le 17.77) \le 10$

Apparent Magnitude

Distances

Angular sizes

Compactness

 $\mu_r \le 26.0 \text{ mag arcsec}^{-2}$

Isolation

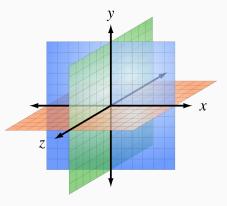
$$\theta_N \ge 3\theta_G$$

Velocity difference

$$|V - V_i| \le 1000 \text{ km/s}$$

Using 2xR200 of the Main Halo as the radial size of the group

Difference from the velocities of galaxies with respect to the host





3) Application of the criteria to the sample

Trying to do it efficiently...

Richness and magnitude

$$3 \le N(14.00 \le r \le 17.77) \le 10$$



Velocity difference

$$|V - V_i| \le 1000 \text{ km/s}$$



Compactness

$$\mu_r \leq 26.0 \text{ mag arcsec}^{-2}$$



Isolation

$$\theta_N \ge 3\theta_G$$

 \Rightarrow

04 ♦

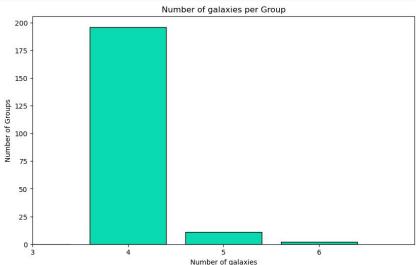
Results



What was it found? (or at least I think it was)

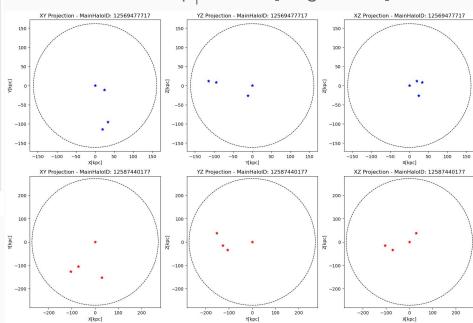


Applying the code to the simulation...



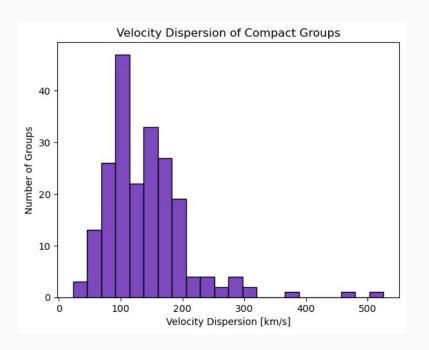
196 groups \rightarrow 4 members 11 groups \rightarrow 5 members 2 group \rightarrow 6 members

Group sizes from 72 to 682 [kpc] $20.14 \le \mu_r \le 23.46$ [mag/arcsec²]





Applying the code to the simulation...



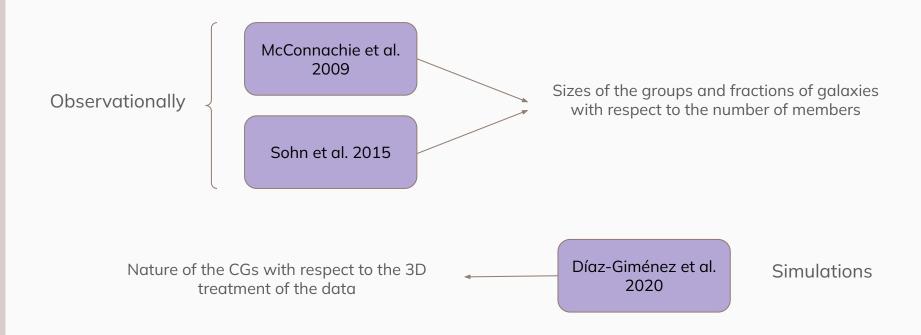
Velocity dispersion ranging between 22.57 to 525.70 km/s

Mean velocity dispersion ~ 139.92 km/s





Results in general agree with literature...

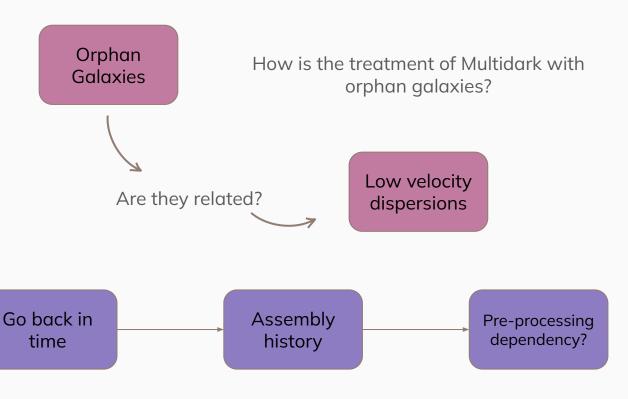


05 \(\phi\) Conclusions \(\phi\)

What can be done?



Things to keep in mind...







Thanks!

Do you have any questions?

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References

- [1] McConnachie, A. W., Patton, D. R., Ellison, S. L., and Simard, L., "Compact groups in theory and practice III. Compact groups of galaxies in the Sixth Data Release of the Sloan Digital Sky Survey", Monthly Notices of the Royal Astronomical Society, vol. 395, no. 1, pp. 255–268, 2009. doi:10.1111/j.1365-2966.2008.14340.x.
- [2] Sohn, J., "Compact Groups of Galaxies with Complete Spectroscopic Redshifts in the Local Universe", Journal of Korean Astronomical Society, vol. 48, no. 6, pp. 381–398, 2015. doi:10.5303/JKAS.2015.48.6.381.
- [3] Díaz-Giménez, E., Taverna, A., Zandivarez, A., and Mamon, G. A., "Compact groups from semi-analytical models of galaxy formation I. A comparative study of frequency and nature", Monthly Notices of the Royal Astronomical Society, vol. 492, no. 2, pp. 2588–2605, 2020. doi:10.1093/mnras/stz3356.
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- [5] Knebe, A., "MULTIDARK-GALAXIES: data release and first results", *Monthly Notices of the Royal Astronomical Society*, vol. 474, no. 4, pp. 5206–5231, 2018. doi:10.1093/mnras/stx2662.