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$$g(N, U, V) = f(N) V^N U^{3N/2}$$

From g we can find the entropy

$$S = k_B \ln g = k_B \ln (f(N) V^N U^{3N/2})$$

$$= k_B \ln (f(N) V^N) + k_B \ln U^{3N/2}$$

$$= k_B \ln (f(N) V^N) + \underbrace{\frac{3Nk_B}{2} \ln U}_{\text{function of } U}$$

From S can find the temperature

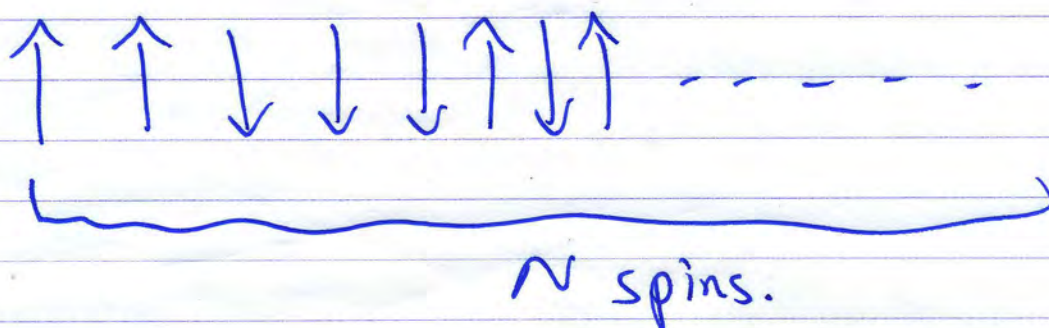
$$\frac{1}{T} = \left(\frac{\partial S}{\partial U} \right)_{N, V} = \left(\frac{\partial \left(\frac{3Nk_B}{2} \ln U \right)}{\partial U} \right)_{N, V}$$

$$= \frac{3Nk_B}{2U} \Rightarrow \boxed{U = \frac{3}{2} N k_B T}$$

2. Paramagnetism

a) We know the multiplicity of a chain of N spins is:

$$g(N, N_{\uparrow}, N_{\downarrow}) = \frac{N!}{N_{\uparrow}! N_{\downarrow}!}$$



Hence the entropy is given by the formula:

$$\begin{aligned} S &= k_B \ln(g(N, N_{\uparrow}, N_{\downarrow})) \\ &= k_B \ln\left(\frac{N!}{N_{\uparrow}! N_{\downarrow}!}\right) \end{aligned}$$

$$= k_B \ln(N!) - k_B \ln(N_{\uparrow}!) - k_B \ln(N_{\downarrow}!)$$

As $N, N_{\uparrow}, N_{\downarrow} \gg 1$ individually we may use:

$$\ln(N!) \approx N \ln(N) - N \quad (\text{Stirling Approx.})$$

b) By definition the temperature is:

$$\frac{1}{T} = \left(\frac{\partial S}{\partial U} \right)_N$$

Now, we have $S = S(N, s)$ and $s = s(U)$.
Hence we may write,

$$\frac{1}{T} = \frac{\partial S}{\partial s} \frac{\partial s}{\partial U} \quad \text{by chain rule.}$$

Firstly, noting the total energy of the chain

$$U = -2smB$$

$$\Rightarrow s = -\frac{U}{2mB}$$

$$\therefore \frac{\partial s}{\partial U} = -\frac{1}{2mB}$$

Then, using our expression from a)

$$\frac{\partial S}{\partial s} = \frac{\partial}{\partial s} \left(-k_B \left(\frac{1}{2}N + s \right) \ln \left(\frac{1}{2}N + s \right) - k_B \left(\frac{1}{2}N - s \right) \ln \left(\frac{1}{2}N - s \right) \right)$$

$$= -k_B \ln \left(\frac{1}{2}N + s \right) - k_B \frac{\left(\frac{1}{2}N + s \right)}{\left(\frac{1}{2}N + s \right)}$$

$$+ k_B \ln \left(\frac{1}{2}N - s \right) - k_B \frac{\left(\frac{1}{2}N - s \right)}{-\left(\frac{1}{2}N - s \right)}$$

$$= -k_B \left[\ln\left(\frac{1}{2}N+s\right) - \ln\left(\frac{1}{2}N-s\right) \right]$$

Thus,

$$\frac{\partial S}{\partial U} = \frac{k_B}{2mB} \left[\ln\left(\frac{1}{2}N+s\right) - \ln\left(\frac{1}{2}N-s\right) \right] = \frac{k_B}{2mB} \left[\ln\left(\frac{\frac{1}{2}N+s}{\frac{1}{2}N-s}\right) \right]$$

$$\Rightarrow T = \frac{2mB}{k_B} \left[\ln\left(\frac{\frac{1}{2}N+s}{\frac{1}{2}N-s}\right) \right]^{-1}$$

$$= \frac{2mB}{k_B} \left[\ln\left(\frac{N - U/mB}{N + U/mB}\right) \right]^{-1}$$

where in the last line we use $s = -\frac{U}{2mB}$.

c) Rearranging from above:

$$\frac{2mB}{k_B T} = \ln\left(\frac{N - U/mB}{N + U/mB}\right)$$

$$e^{2mB/k_B T} = \frac{N - U/mB}{N + U/mB}$$

$$\Rightarrow (N + U/mB) e^{2mB/k_B T} = N - U/mB$$

pls

$$U = N m B \left[\frac{1 - e^{2mB/k_B T}}{1 + e^{2mB/k_B T}} \right]$$

$$= N m B \left[\frac{e^{-mB/k_B T} - e^{mB/k_B T}}{e^{-mB/k_B T} + e^{mB/k_B T}} \right] \cdot \frac{e^{mB/k_B T}}{e^{mB/k_B T}}$$

$$= N m B \cdot \frac{-\sinh\left(\frac{mB}{k_B T}\right)}{\cosh\left(\frac{mB}{k_B T}\right)} \quad \left(\text{see definition of } \cosh(x), \sinh(x) \right)$$

$$= -N m B \tanh\left(\frac{mB}{k_B T}\right)$$

d) For large T we have $k_B T \gg mB$

$$\text{ie. } \frac{mB}{k_B T} \ll 1$$

The Taylor expansion for $\tanh(x)$ for $x \ll 1$ is:

$$\tanh(x) \approx x$$

Hence we may write

$$U \approx -N m B \left(\frac{mB}{k_B T} \right) \quad \text{for } k_B T \gg mB$$

$$= -N \frac{m^2 B^2}{k_B T}$$

Using ~~$\langle U \rangle$~~ $\langle U \rangle = -2\langle s \rangle mB$ we can express the fractional magnetization as:

$$-2\langle s \rangle mB = - \frac{Nm^2 B^2}{k_B T}$$

$$\Rightarrow \frac{2\langle s \rangle}{N} \approx \frac{mB}{k_B T}$$

Thus we indeed get a scaling of $1/T$ for the fractional magnetization, consistent with Curie's law.

e) For $T \rightarrow 0$, $\frac{mB}{k_B T} \rightarrow \infty$.

For such a case we note,

$$\tanh(x) \rightarrow 1 \quad \text{as } x \rightarrow \infty$$

Hence we rewrite the energy as:

$$U \approx -NmB.$$

This is consistent as it implies all spins align at low temperature, minimising the energy, i.e.

$$\langle s \rangle = \frac{N}{2}$$

Question 3: Paramagnetism (continued).

a) see attached plots.

b) By definition the heat capacity is given by:

$$C_B = \left(\frac{\partial \langle U \rangle}{\partial T} \right)_{N, B}$$

Using $\langle U \rangle$ derived in Question 4, this becomes:

$$C_B = -NmB \cdot \frac{\partial}{\partial T} \left(\tanh\left(\frac{mB}{k_B T}\right) \right)$$

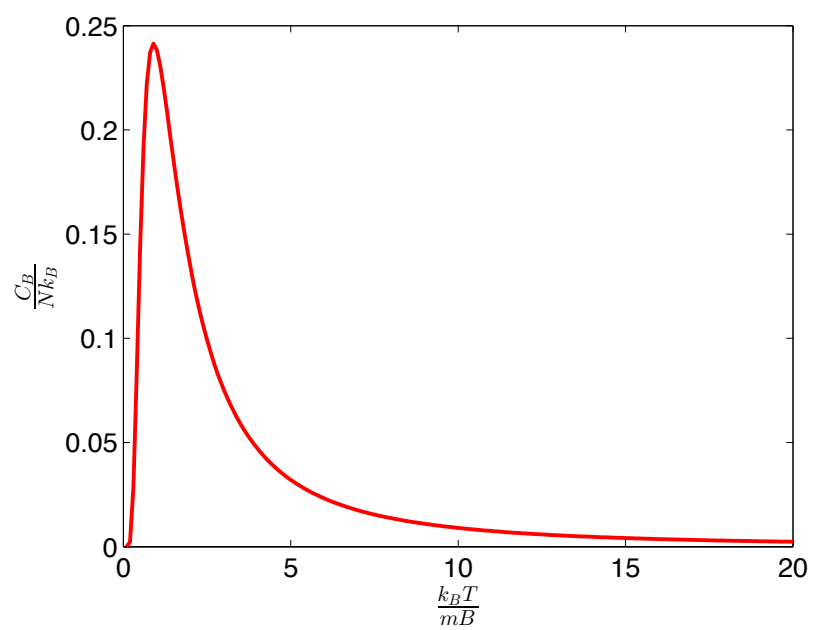
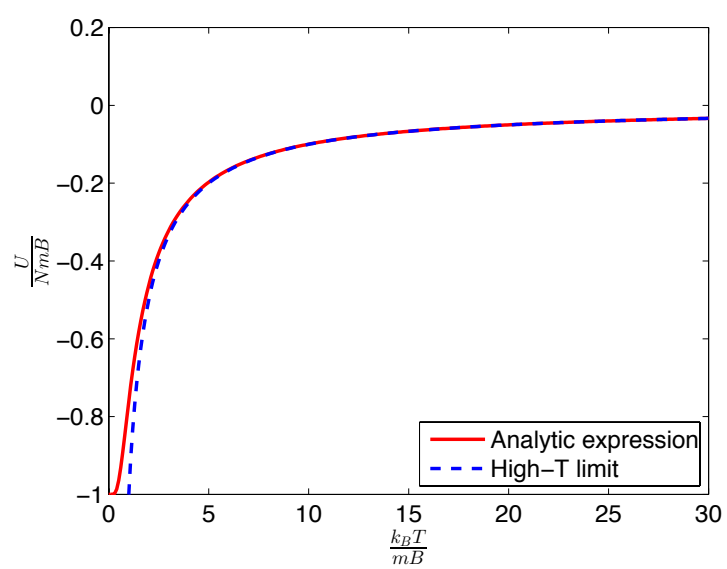
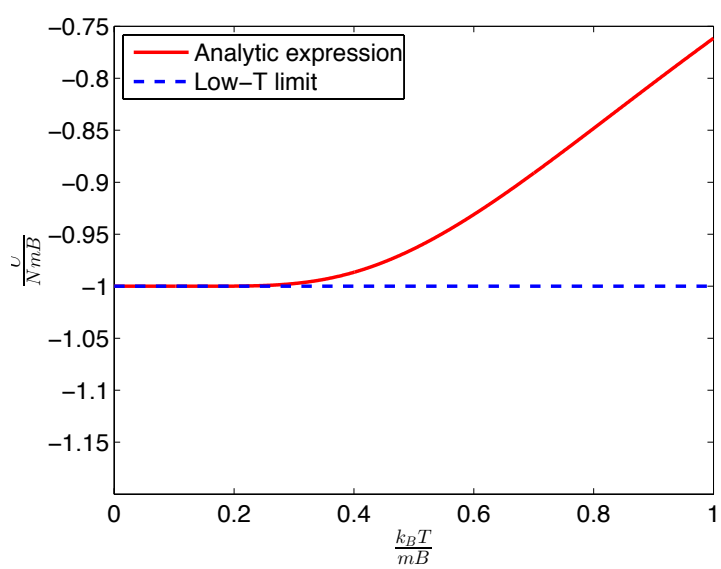
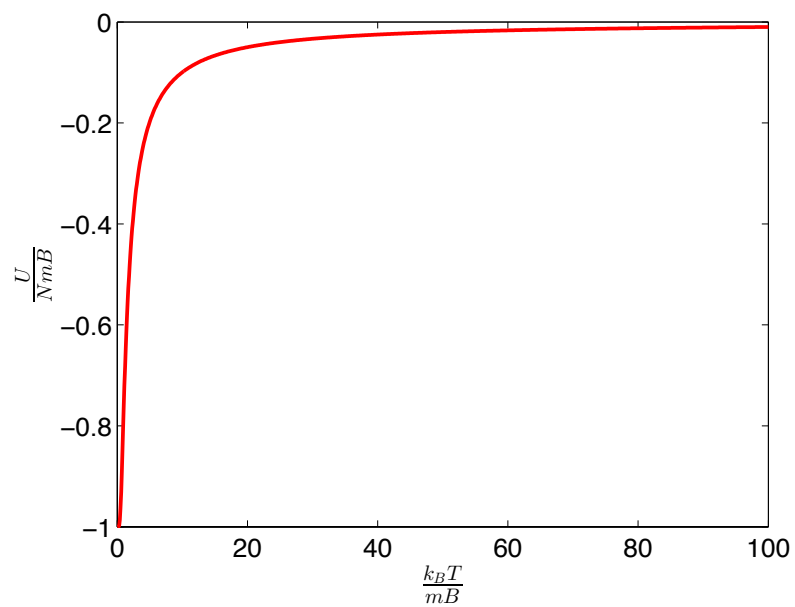
$$= \frac{Nm^2 B^2}{k_B T^2} \cdot \frac{1}{\cosh^2\left(\frac{mB}{k_B T}\right)}$$

For $T \rightarrow 0$, ~~we have~~ $k_B T \ll mB$

$$\text{ie. } \cosh\left(\frac{mB}{k_B T}\right) \approx \frac{e^{mB/k_B T}}{2} \quad \left[\text{see definition of } \cosh \right]$$

$$\Rightarrow C_B \approx \frac{Nm^2 B^2}{k_B T^2} \cdot \frac{4}{e^{2mB/k_B T}}$$

$$= \frac{4Nm^2 B^2}{k_B} \cdot \frac{1}{T^2} \cdot e^{-2mB/k_B T}$$



Thus for $T \rightarrow 0$, we expect the heat capacity to exponentially ~~decr~~ decay.

For $T \rightarrow \infty$, $k_B T \gg mB$.

ie. $\cosh\left(\frac{mB}{k_B T}\right) = 1 + \frac{1}{2}\left(\frac{mB}{k_B T}\right)^2 + \dots$ (Taylor series)

Hence,

$$C_B = \frac{Nm^2 B^2}{k_B T^2} \cdot \frac{1}{\left(1 + \frac{1}{2}\left(\frac{mB}{k_B T}\right)^2\right)^2}$$

$$\approx \frac{Nm^2 B^2}{k_B T^2} \cdot \left[1 - \left(\frac{mB}{k_B T}\right)^2\right] *$$

$$\approx Nk_B \cdot \left(\frac{mB}{k_B T}\right)^4 \quad \text{for } T \rightarrow \infty$$

* Note here we use: $(1+ax)^n \approx 1+nax$
for $ax \ll 1$

Hence for $T \rightarrow \infty$ we expect the heat capacity to go to zero as $\frac{1}{T^4}$.

c) We assume the multiplicity of the chain can be approximated as:

$$g(N, s) = g(N, 0) e^{-2s^2/N}$$

which is valid for: $|s| \ll N$ + $N \gg 1$.

Then the entropy is given by:

$$\begin{aligned} S &= k_B \log(g) \\ &= k_B \log(g(N, 0)) - \frac{2k_B s^2}{N} \end{aligned}$$

Trivially we then calculate the equilibrium temperature as:

$$\frac{1}{T} = \left(\frac{\partial S}{\partial U} \right)_N = \frac{\partial S}{\partial s} \frac{\partial s}{\partial U}$$

Noting $U = -2smB$ from Q4,

$$\Rightarrow \frac{\partial s}{\partial U} = \frac{\partial}{\partial U} \left(\frac{U}{-2mB} \right) = \frac{-1}{2mB}$$

$$\begin{aligned} \Rightarrow \frac{1}{T} &= -\frac{4k_B s}{N} \cdot \frac{-1}{2mB} \\ &= \frac{2k_B s}{NmB} \end{aligned}$$

Rewriting S in terms of U ,

$$T = - \frac{Nm^2 B^2}{k_B U}$$

which can be inverted to give the thermal average energy:

$$U = - \frac{Nm^2 B^2}{k_B T}$$

which agrees with the high temperature limit derived in Q4.

Note that our method does not recover the low-temperature behaviour as the initial approximations are invalid. Namely as $T \rightarrow 0$ we intuitively expect all spins to align, which would imply $|S| \sim N$, explicitly contradicting the approximation $|S| \ll N$.

Question 4: The Einstein Model of a Solid

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a) We treat each atom in the lattice as 3 harmonic oscillators (one each to correspond to motion in the x, y, z directions).

Each oscillator has the same frequency, ω , and for $\frac{N}{3}$ atoms, there exist N oscillators.

If we assume a total of q_{tot} quanta to be shared then,

$$U = q_{\text{tot}} kT = NkT$$

Using the sticks and ~~boxes~~ balls method, (see lecture notes) it can easily be shown

$$g(N, q) = \frac{(q + N - 1)!}{q!(N-1)!}$$

is the multiplicity where $q = q_{\text{tot}}$.

Following the derivation of Q in problem set 1 we may calculate the entropy, equilibrium temperature etc. We refer the student to the solutions of that problem set for the full results. In the following we only require the result:

$$U = \frac{NkT}{e^{kT/k_B T} - 1}$$

which is the thermal equilibrium total energy for a given temperature T .

b) We use the expression for U found in a) and consider the limits of high & low temperature.

For $T \rightarrow \infty$, $k_B T \gg \hbar \omega$, we may expand

$$e^{\hbar \omega / k_B T} \approx 1 + \frac{\hbar \omega}{k_B T} + \dots \quad (\text{Taylor series})$$

We then ~~rewrite~~ rewrite the total energy as:

$$U \approx \frac{N \hbar \omega}{1 + \frac{\hbar \omega}{k_B T} - 1} = N \hbar \omega \cdot \frac{k_B T}{\hbar \omega} \approx N k_B T. \quad (\propto T)$$

This result is consistent with the equipartition theorem for classical particles. Each oscillator has energy $k_B T$ as the Hamiltonian is quadratic in both position & momentum. Thus for N atoms, with

For $T \rightarrow 0$, $k_B T \ll \hbar \omega$, we note that

$$e^{\hbar \omega / k_B T} \gg 1$$

Hence we write the total energy as:

$$\cancel{U = \frac{N \hbar \omega}{2}} \quad U = N \hbar \omega e^{-\hbar \omega / k_B T}$$

ie. the energy decays exponentially as $T \rightarrow 0$

\Rightarrow See attached plots for U vs. T .

c) Using the expression for U , the heat capacity is:

$$C_V = \left(\frac{\partial U}{\partial T} \right)_{N,V}$$

$$= \frac{\partial}{\partial T} \left[\frac{N \hbar \omega}{e^{\hbar \omega / k_B T} - 1} \right]$$

$$= \cancel{\frac{N (\hbar \omega)^2}{k_B T^2}} \quad N k_B \cdot \left(\frac{\hbar \omega}{k_B T} \right)^2 \cdot \frac{e^{\hbar \omega / k_B T}}{(e^{\hbar \omega / k_B T} - 1)^2}$$

d) For $T \rightarrow \infty$, $k_B T \gg \hbar \omega$ and we again use:

$$e^{\hbar \omega / k_B T} \approx 1 + \frac{\hbar \omega}{k_B T}$$

Hence,

$$\begin{aligned} C_v &\approx \cancel{Nk_B} Nk_B \left(\frac{\hbar \omega}{k_B T} \right)^2 \cdot \frac{1 + \hbar \omega / k_B T}{\left(\frac{\hbar \omega}{k_B T} \right)^2} \\ &\approx Nk_B \left(1 + \hbar \omega / k_B T \right) \\ &\approx Nk_B \quad \text{for} \quad \frac{\hbar \omega}{k_B T} \ll 1 \end{aligned}$$

Again, this is the expected result for a classical system due to the equipartition theorem.

For $T \rightarrow 0$, $k_B T \ll \hbar \omega$, we use:

$$e^{\hbar \omega / k_B T} \gg 1 \quad \text{and} \quad \cancel{e^{\hbar \omega / k_B T}} \quad \cancel{e^{2\hbar \omega / k_B T}}$$

Hence,

$$\begin{aligned} C_v &\approx Nk_B \left(\frac{\hbar \omega}{k_B T} \right)^2 \frac{e^{\hbar \omega / k_B T}}{e^{2\hbar \omega / k_B T}} \\ &\approx Nk_B \left(\frac{\hbar \omega}{k_B T} \right)^2 e^{-\hbar \omega / k_B T} \end{aligned}$$

\Rightarrow We expect an exponential decay in the heat capacity for low T .

