Section 3.8 Geodetic Measurements

Acronyms, Abbreviations and Symbols

DGPS	Differential	Global	Positioning	System

ECEF Earth Centered Earth Fixed coordinate system

GPS Global Positioning System

INS Inertial Navigation System

WGS84 World Geodetic System 1984

- a Earth's semi-major axis radius
- b Earth's semi-minor axis radius
- D Great circle distance between two points
- *e* eccentricity of the Earth square
- f Earth's flatness factor
- *h* geodetic height
- N radius of curvature in prime vertical
- P radius of curvature in prime vertical
- \rightarrow Vector from earth center extending to coordinates
- r Earth's radius
- X ECEF x coordinate
- Y ECEF y coordinate
- Z ECEF z coordinate
- Geodetic latitude
- φ Angle between the two \xrightarrow{P} vectors originating at the Earth's center and extending to their respective coordinates at the start and end points.
- λ Geodetic longitude
- Ψ Runway heading with respect to true North.

Earth Modeling

The Geodetic System defines the location of any point relative to the earth using latitude, longitude and height (Figure 3.8-1, point P). Longitude and latitude are expressed in degrees, minutes, seconds. Longitude lines extend \pm 180 degrees from the Prime Meridian, run north to south, and converge at the poles. Latitude lines are parallel to the equator and extend \pm 90°.

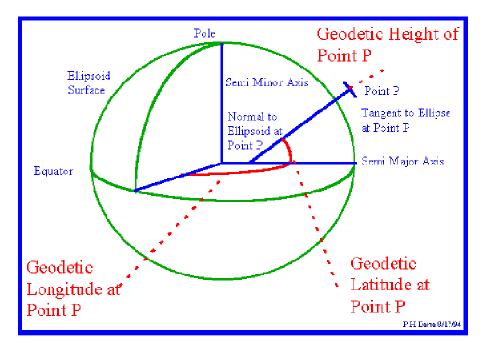


Figure 3.8-1 Geodetic Coordinate System

The 1984 world geodetic system, WGS84, models the earth's surface as an oblate spheroid - an ellipsoid rotated about its semi-minor axis. In this model, used by global positioning systems, the earth's semi-major axis, a is 6,378,137.0 meters and the semi-minor axis, b is 6,356,752.314 meters. The flatness factor (f) is defined as:

$$f = \frac{a - b}{a}$$

For the WGS84 model, f = 1/298.257223563

Because the earth is not perfectly spherical, there are various methods for defining latitude. Unlike the geocentric latitude which uses the earth's center for determining a point's latitude, the **geodetic latitude** (used herein) is the angle between the equatorial plane and a line *normal to the reference ellipsoid*. Figure 3.8-1 exaggerates this with a normal line being well offset from the earth's center. This definition leads to a degree of latitude being longer at the pole than at the equator: 111,694 m (60.3 nm) vs. 110,574 m (59.7 nm).

The **geodetic longitude** of a point is the angle between a reference plane and a plane passing through the point, both planes being perpendicular to the equatorial plane.

Mathematically, the **geodetic surface** is a smooth ellipsoid modeling the earth's surface. Clearly, the topography (actual surface height) deviates from this model whenever land is above or below sea level. Less evident is that the actual sea level also deviates from the geodetic model due to local changes in the earth's gravity. Specifically,

mass variations caused by changes in earth density and topography, such as mountains or trenches, change local gravity vectors and therefore sea level relative to the ellipsoid.

Reference to **Mean Sea Level** (MSL) served as the traditional way to express topographic or bathymetric height. Geodesists once considered the sea in balance with the earth's gravity and formed a perfectly regular figure. MSL is usually described as a tidal datum that is the arithmetic mean of hourly water elevations observed over a 19-year (Metonic) cycle. This definition averages out tidal highs and lows caused by the changing effects of the gravitational forces from the moon and sun. MSL defines the zero elevation (vertical datum) for a local area. Because the sea surface conforms to the earth's gravitational field, MSL also has slight hills and valleys similar to the land surface but much smoother. Zero elevation as defined by one nation is often not the same zero elevation defined by another, thus locally defined vertical datums differ from each other.

The **Geoid** is the equipotential surface in the earth's gravity field that coincides most closely with the mean sea level extended continuously under the continents. In other words, it approximates the level of any non-flowing water connected (actually or theoretically) to the seas by waterway or via trenches or tunnels. The geoid surface undulates relative to the geodetic ellipsoid and is perpendicular to the local gravity vector – as seen with a plumb line. Similarly, a spirit level defines the local surface parallel to the geoid, which is tangent to the local horizon. Because the geoid is an equipotential surface, it is the best datum for measuring potential energy and is the true zero surface for measuring elevations. Previously, there was no way to accurately measure the geoid, so heights were measured relative to the similar MSL. EGM96 (Earth Gravity Model 1996) represents the best geoid model currently available and shows smoothly changing surface undulations ranging from +85 to -107 meters relative to the WGS84 ellipsoid.

The geoid surface cannot be directly observed, thus heights above or below it can't be directly measured. Instead the geoid surface is modeled mathematically using gravitational measurements. Although for practical purposes, at the coastline the geoid and MSL surfaces are assumed to be essentially the same, at some spots the geoid can actually differ from MSL by several meters.

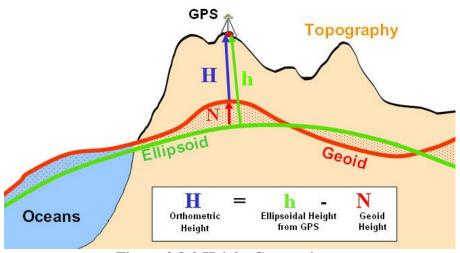


Figure 3.8-2 Height Comparisons

Ellipsoidal height (h) is the same as geodetic height and is the geometric distance between an object and the earth ellipsoid (Figure 3.8-2). This may be a GPS output.

Geoid height (N) is the height of the geoid above or below the ellipsoid. Some GPS devices output this undulation value in the data stream.

Orthometric height (H) is the geometric distance along a gravitational force line from a given point P to the geoid. This is essentially the conventional height measurement because the geoid approximates MSL- the traditional method for determining height.

Modern GPS units typically include a geoid model (e.g. EGM-96) that provides N (geoid height over the WGS ellipsoid) at the current position. Such a unit can provide the height above geoid. If GPS height output is only available relative to the ellipsoid (h), then traditional Orthometric height (H) above the geoid can be obtained by subtracting the geoid height above the ellipsoid. [http://www.esri.com/news/arcuser/0703/geoid1of3.html]

ECEF Transformations

For the purpose of performance, navigation, or noise analysis, flight testers may require distances between two points (the shortest being along the great circle arc) and the average heading of that arc. Calculating these from typical Geodetic System Lat/Long inputs requires conversion to the Earth Centered Earth Fixed (ECEF) coordinate system as shown in Figure 3.8-3.

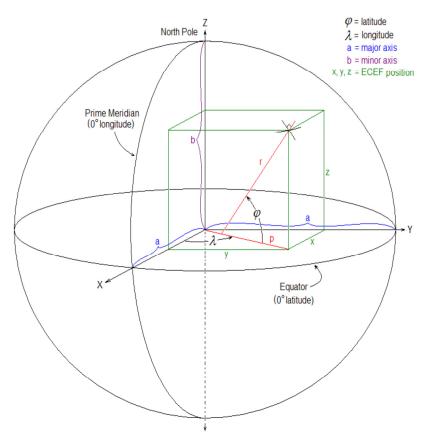


Figure 3.8-3 Earth Centered Earth Fixed Coordinate System

The ECEF coordinate system is a Cartesian system with the origin at the earth's center. In this system, the X-axis is defined by the intersection of the Prime Meridian and equatorial planes. The Z-axis goes through the North Pole. The Y-axis completes a right-handed orthogonal system by a plane 90 degrees east of the X-axis and its intersection with the equator.

Geodetic System (lat/long/height) data converts to ECEF as follows:

$$x = (N + h) \cdot \cos(\phi) \cdot \cos(\lambda)$$

$$y = (N + h) \cdot \cos(\phi) \cdot \sin(\lambda)$$

$$z = (N \cdot (1 - e^2) + h) \cdot \sin(\phi)$$

where,

x = ECEF coordinate parallel to the X-axis

y = ECEF coordinate parallel to the Y-axis

z = ECEF coordinate parallel to the Z-axis

 ϕ = geodetic latitude

 λ = geodetic longitude

h = height above geodetic (ellipsoid) surface

N = Normal radius of curvature; distance from earth axis to any point on the geodetic surface at that latitude (extension of r to axis shown in Figure 3.8-3).

$$N = \frac{a}{\sqrt{1 - e^2 \cdot \sin^2(\phi)}}$$

where,

a = semi-major axis radius (6,378,137 m; 20,925,647 ft)

 e^2 = eccentricity squared; $e^2 = 1 - \left(\frac{a}{b}\right)^2 = 2 \cdot f - f^2 = 0.00669438002290$ (Earth, per WGS84).

Also useful is

M = Meridian radius of curvature; distance from earth axis to any point on the geodetic surface at that longitude.

$$M = \frac{a(1 - e^2)}{[1 - e^2 \cdot \sin^2(\phi)]^{1.5}}$$

Great Circle Calculations

Any plane passing through the center of a spheroid traces a **Great Circle** around the perimeter of that spheroid. The shortest distance between two points on the surface is that portion of the great circle arc encompassing both points (Figure 3.8-3).

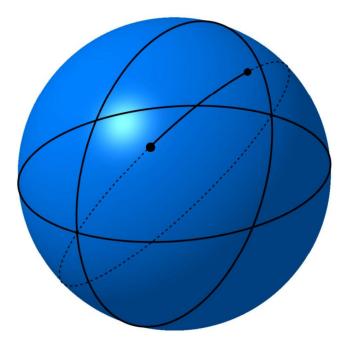


Figure 3.8-3 Great Circle Arc

Except when moving around the equator, navigating along a great circle route has the characteristic of intercepting longitude lines at different angles. In other words, the heading (or bearing) changes along the route. Analysis shows bearing change along a great circle route:

- Is never greater than the longitude difference between the end points.
- Approaches the value of the longitude change as the final latitude approaches a pole (regardless of initial latitude).
- Is smallest when the final latitude is at the equator (for this case, bearing change \approx longitude change x initial latitude/100).

Calculate the **great circle distance** (**D**) between points (subscripts 1 and 2) as

$$\begin{split} P_1 &= \sqrt{x_1^2 + y_1^2 + z_1^2}, \quad P_2 &= \sqrt{x_2^2 + y_2^2 + z_2^2} \\ & \xrightarrow{P_1} \cdot \xrightarrow{P_2} = P_1 \cdot P_2 \cos \varphi = x_1 \cdot x_2 + y_1 \cdot y_2 + z_1 \cdot z_2 \\ & \varphi = \arccos \left(\frac{x_1 \cdot x_2 + y_1 \cdot y_2 + z_1 \cdot Z z_2}{P_1 \cdot P_2} \right) \\ & D &= P_{ava} \cdot \varphi \end{split}$$

where

P = distance from earth center to any point (including height above the spheroid surface).

 $\underset{P}{\rightarrow}$ = Vector from the Earth's center to point P.

 φ = Angle between the two $\underset{p}{\rightarrow}$ vectors

http://www.movable-type.co.uk/scripts/latlong.html provides equations an online tool for calculating great circle initial & final bearings (headings).

For shorter distances typical of local flight testing, the Great Circle model matches the following **Two-dimensional approximations**.

Distance North-South (Northing): $dy = N \cdot \sin(\Delta \phi)$

Earth's radius East-West: $r = N \cdot \cos(\phi)$

Distance East-West (Easting): $dx = r \cdot \sin(\Delta \lambda)$

2-D distance between two points: $D = \sqrt{dx^2 + dy^2}$

Heading between two points (relative to true north) $\psi = \arctan(dy/dx)$

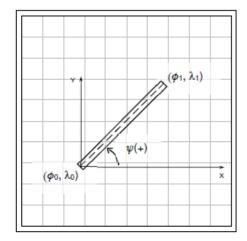
An error analysis of the above 2-D heading approximation shows it consistently lies between the initial and final headings transpiring during great circle navigation and is closest to the final heading. This occurs when considering up to 10 degrees longitude change and is therefore sufficiently accurate for lesser changes that arise in flight testing (e.g. radial from a navigation transmitter).

An error analysis of the above 2-D distance approximation shows accuracy within 0.6% of the great circle distance when changing latitude <u>and</u> longitude 1 degree, and within 3% when changing latitude and longitude 10 degrees. It is accurate to within 0.4% when changing only latitude or longitude 10 degrees.

Local Distance Transformation

Latitude ϕ , longitude λ , and height (typical GPS output data) can be transformed into rectangular (X,Y,Z) coordinates. The following presents a method for applying this to two different coordinate systems, both with the X-Y axes defining the horizontal plane. This is useful when working with local distances associated with typical flight testing such as noise measurement, local navigation, or field performance.

Figure 3.8-4a shows a case where the X-Y coordinate system aligns with the latitude & longitude grid. Figure 3.8-4b shows a case where the X-Y coordinate system aligns with a runway, with Y=0 defining the centerline. In both cases, a designated primary reference datum $[\phi_0, \lambda_0]$, such as the runway centerline threshold, coincides with X=0, Y=0. With the example shown in Figure 3.8-4b, the opposite end of the runway centerline $[\phi_1, \lambda_1]$, coincides with X= *runway length* and Y=0.



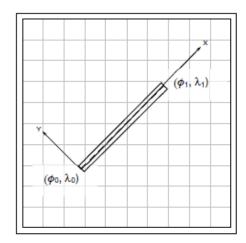


Figure 3.8-4: Local X-Y Coordinate System Aligned with (a)φ,λ Grid; (b) Runway

Because each degree of latitude change is not always exactly 60 nm and because the spacing between degrees of longitude changes markedly with latitude, converting from degrees latitude & longitude change to distance requires scaling factors. First select equatorial distances for each degree change

$$\phi_{\text{scale}}\{\phi=0\} = 110,574 \text{ m} = 362,776.6 \text{ ft} = 59.70518 \text{ nm}$$

$$\lambda_{\text{scale}}\{\phi=0\} = 111,319.5 \text{ m} = 365,221.4 \text{ ft} = 60.10772 \text{ nm}$$

Use average latitude to calculate latitude and longitude scaling factors

$$\emptyset_scale = \frac{\emptyset_scale\{\emptyset=0\}}{\left[1-e^2\cdot sin^2(\phi)\right]^{1.5}} \qquad \lambda_scale = \frac{[\lambda_scale\{\emptyset=0\}]\cdot cos\emptyset}{\sqrt{1-e^2\cdot sin^2(\phi)}}$$

Calculate X and Y components of distances aligned with the latitude & longitude grid (Figure 3.8-4a)

$$Y = \phi _scale \cdot (\phi_1 - \phi_0)$$

$$X = \lambda \ scale \cdot (\lambda_1 - \lambda_0)$$

The 2-D *local* (vice great circle) distance between any two points on the X-Y plane is

$$D = (X^2 + Y^2)^{1/2}$$

For field performance work, it is preferable to align X & Y with the runway as shown in Figure 3.8-4b. To convert from grid-aligned to runway-aligned coordinates, apply a rotation matrix that ensures X remains positive when going from point $[\phi_0, \lambda_0]$ towards $[\phi_1, \lambda_1]$. The rotation angle, ψ , is positive counter-clockwise from true East to the direction of the runway. [Note: do not confuse ψ with runway heading measured relative to magnetic north]. The function

$$\psi = atan2(X, Y)$$

returns rotation angles from - π to π , so that rotating to headings between 90° and 270° results in negative rotation angles.

For any point $[\phi, \lambda]$, calculate runway-aligned X & Y distances using

$$X = \phi _scale \cdot (\phi - \phi_0) \cdot sin(\psi) + \lambda _scale \cdot (\lambda - \lambda_0) \cdot cos(\psi)$$

$$Y = \phi _scale \cdot (\phi - \phi_0) \cdot cos(\psi) - \lambda _scale \cdot (\lambda - \lambda_0) \cdot sin(\psi)$$

This provides a right-handed rectangular coordinate system where X is positive from $[\phi_0, \lambda_0]$ towards $[\phi_1, \lambda_1]$ and Y is positive left of the runway centerline (Figure 3.8-4b).

For multiple tests from a given runway, it is convenient to define the following constants from the above equations

 $K_1 = \phi _scale \cdot sin(\psi)$

 $K_2 = \lambda_scale \cdot cos(\psi)$

 $K_3 = \phi _scale \cdot cos(\psi)$

 $K_4 = \lambda_scale \cdot sin(\psi)$

The overall conversion from $[\phi, \lambda]$ to [X, Y] then reduces to

$$\begin{split} X &= K_1 \cdot (\ \phi - \phi_0) + K_2 \cdot (\lambda - \lambda_0) \\ Y &= K_3 \cdot (\ \phi - \phi_0) - K_4 \cdot (\lambda - \lambda_0) \end{split}$$

Calculating aircraft height Z above the X-Y plane requires defining where the X-Y plane lies. An analyst may define Z=0 at some arbitrary height (i.e. GPS altitude at the beginning of a maneuver) and consider only changes from that reference.

For field performance, it is typical to use the runway altitude as the reference. Because runway altitudes vary however, height should be surveyed and modeled or tabulated as a function runway centerline position, $H\{X\}$. For best accuracy, the runway survey accounts for GPS antenna height above the surface. Airplane height above the X-Y plane (Z) is then

$$Z = \zeta - H\{X\}$$

where ζ is the test GPS antenna's altitude.

Note: Analysts can determine Z using either Orthometric height above the geoid (H) or above the geodetic surface (h) – as long as the runway surface model $H\{X\}$ uses the same reference

Section 3.9 References

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NOTES