

Comparison of results from different codes: URGENT, US, SRW and pySRU for zero emittance calculations

This document compares the results of undulator emission by different codes (US, URGENT, SRW and pySRU).

The calculations are:

- Flux emission or radiance at an aperture (x,y) placed at a given distance for a monochromatic photon energy E_0
- Flux spectrum integrated over the aperture
- Power density distribution (versus x,y) at the aperture.

Zero emittance has been used for all calculations in this document.

The purpose of the document is a validation of our pySRU code. This code calculates the emission at an aperture $I(x,y)$ for a photon energy E_0 , thus $I(E_0,x,y)$. For different energies one can calculate the stack array $I(E,x,y)$. Integration over the zero axis gives the power density $P(x,y)$. Integration over axes 1 and 2 gives the energy spectrum $I(E)$. Certainly, this is not the most efficient way for computing the spectrum and the power density, but in this document it is done in this way for testing the pySRU results. Other more efficient formulas will be added to pySRU for quick spectral and power density calculations, as used in the other codes.

The calculated cases are 1) the figure 2.5 of the X-ray Data Booklet http://xdb.lbl.gov/Section2/Sec_2-1.html, and 2) the ESRF electron beam with undulators at various K values: $K=0.25$ (SHADOW_DEFAULTS), $K=1.68$ (ESRF_NEW_OB), and $K=4.0$ (ID16_NA).

Parameters of the cases studied

XRAY_BOOKLET

ElectronBeamDivergenceH = $1e-30$

ElectronBeamDivergenceV = $1e-30$

ElectronBeamSizeH = $1e-30$

ElectronBeamSizeV = $1e-30$

ElectronEnergySpread = $1e-30$

ElectronCurrent = 1.0

ElectronEnergy = 1.3

Kv = 1.87

NPeriods = 14

PeriodID = 0.035

distance = 100.0

gapH = 0.2

gapV = 0.2

ID16_NA

ElectronBeamDivergenceH = $1e-30$

ElectronBeamDivergenceV = $1e-30$

ElectronBeamSizeH = $1e-30$

ElectronBeamSizeV = $1e-30$

ElectronEnergySpread = $1e-30$

ElectronCurrent = 0.2

ElectronEnergy = 6.04

Kv = 4.0

NPeriods = 77

PeriodID = 0.026

distance = 20.0

gapH = 0.015

gapV = 0.015

ESRF_NEW_OB

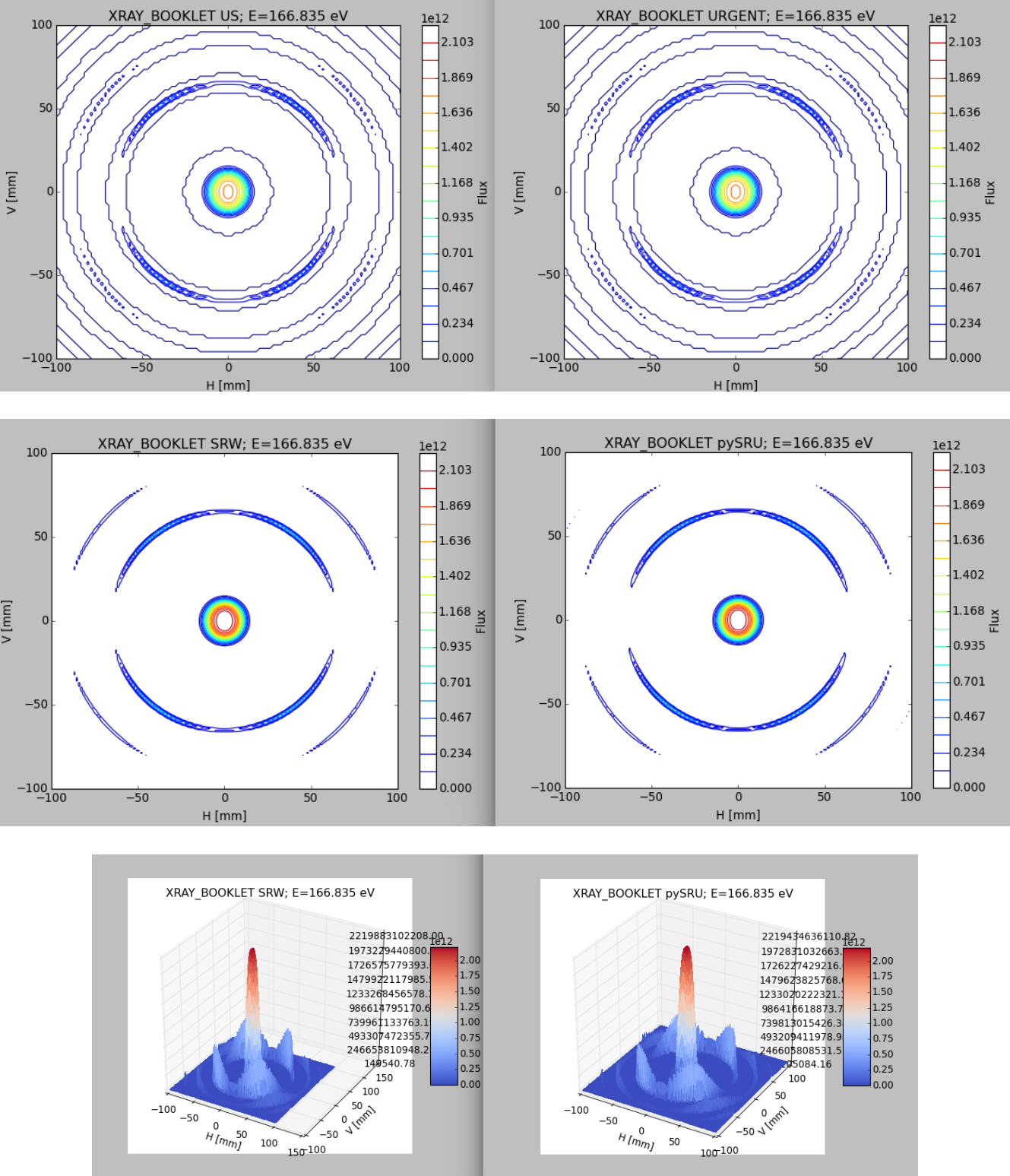
ElectronBeamDivergenceH = 1e-30
ElectronBeamDivergenceV = 1e-30
ElectronBeamSizeH = 1e-30
ElectronBeamSizeV = 1e-30
ElectronEnergySpread = 1e-30
ElectronCurrent = 0.2
ElectronEnergy = 6.0
Kv = 1.68
NPeriods = 222
PeriodID = 0.018
distance = 30.0
gapH = 0.001
gapV = 0.001

SHADOW_DEFAULTS

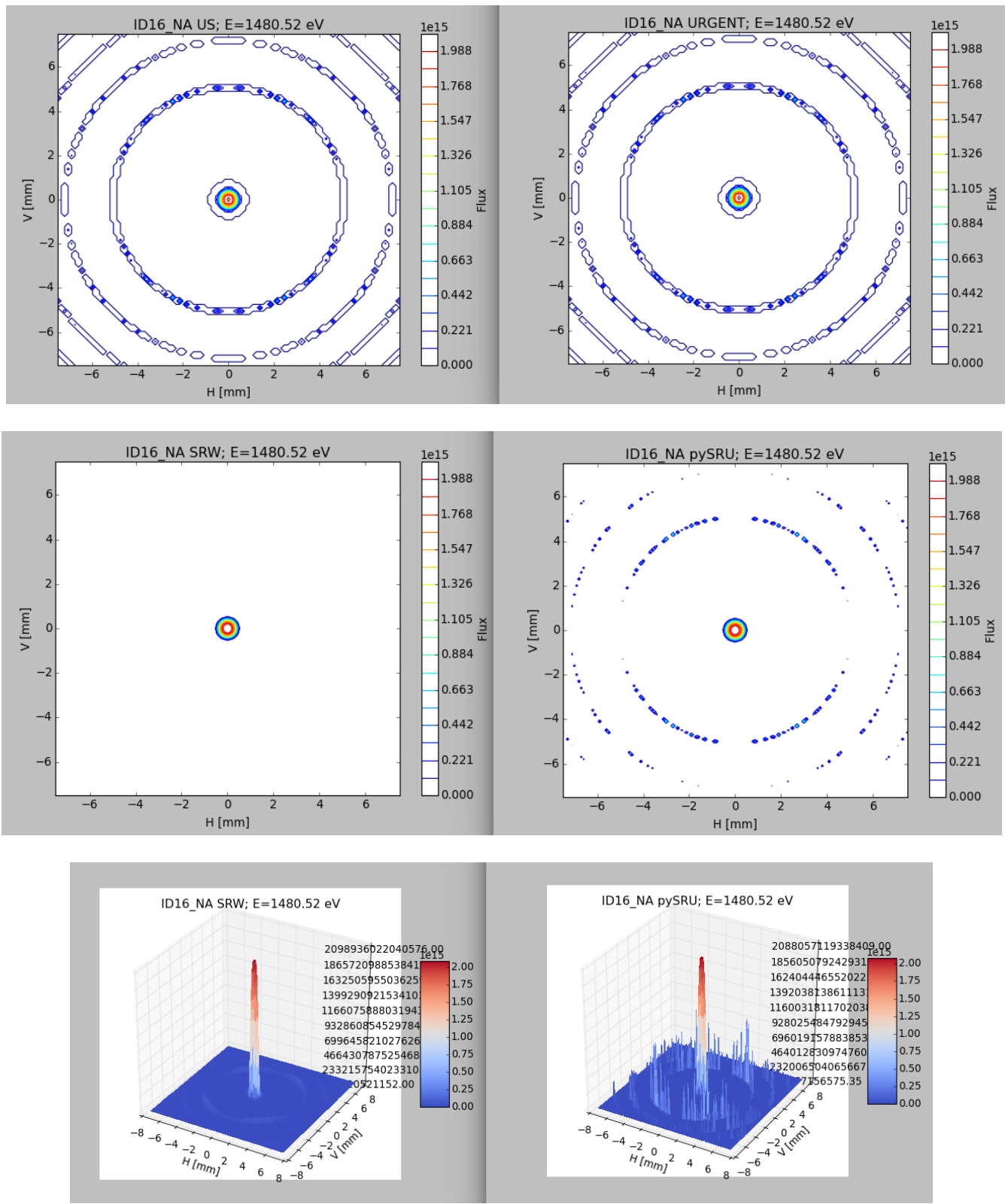
ElectronBeamSizeH = 1e-30
ElectronBeamSizeV = 1e-30
ElectronBeamDivergenceH = 1e-30
ElectronBeamDivergenceV = 1e-30
ElectronEnergySpread = 1e-30
ElectronCurrent = 0.2
ElectronEnergy = 6.04
Kv = 0.25
NPeriods = 50.0
PeriodID = 0.032
distance = 10.0
gapH = 0.002
gapV = 0.002

Radiance at first harmonic

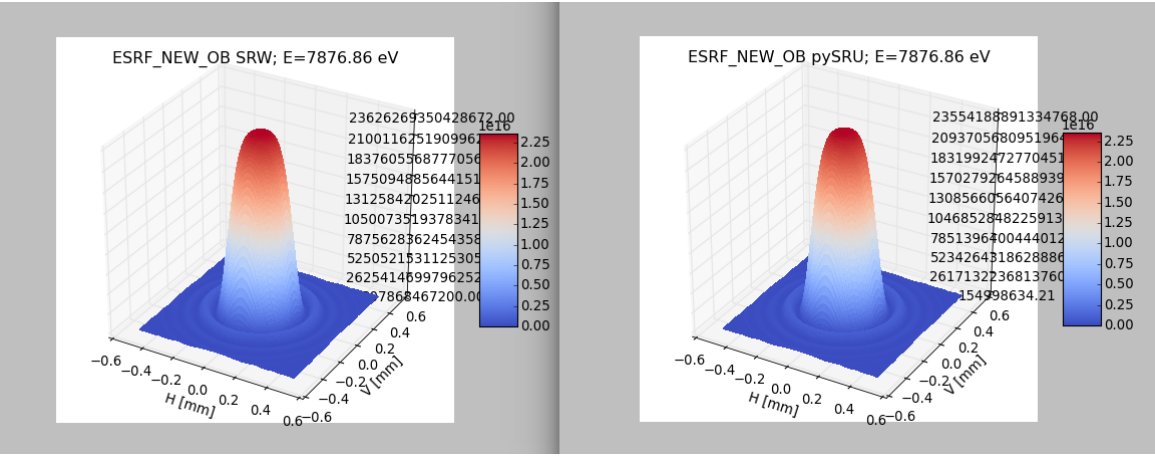
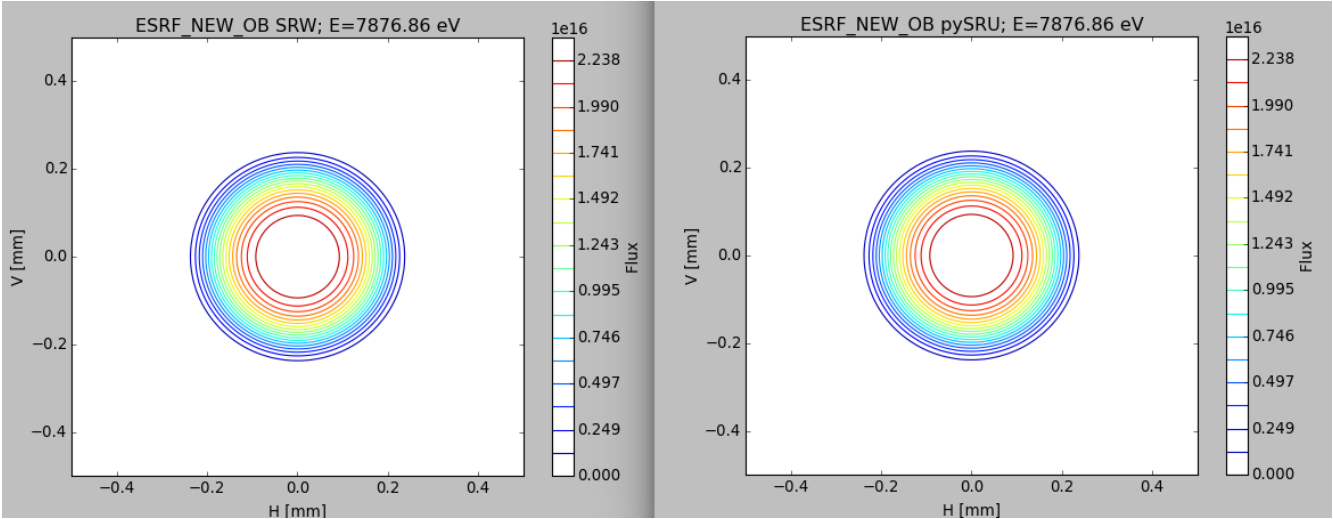
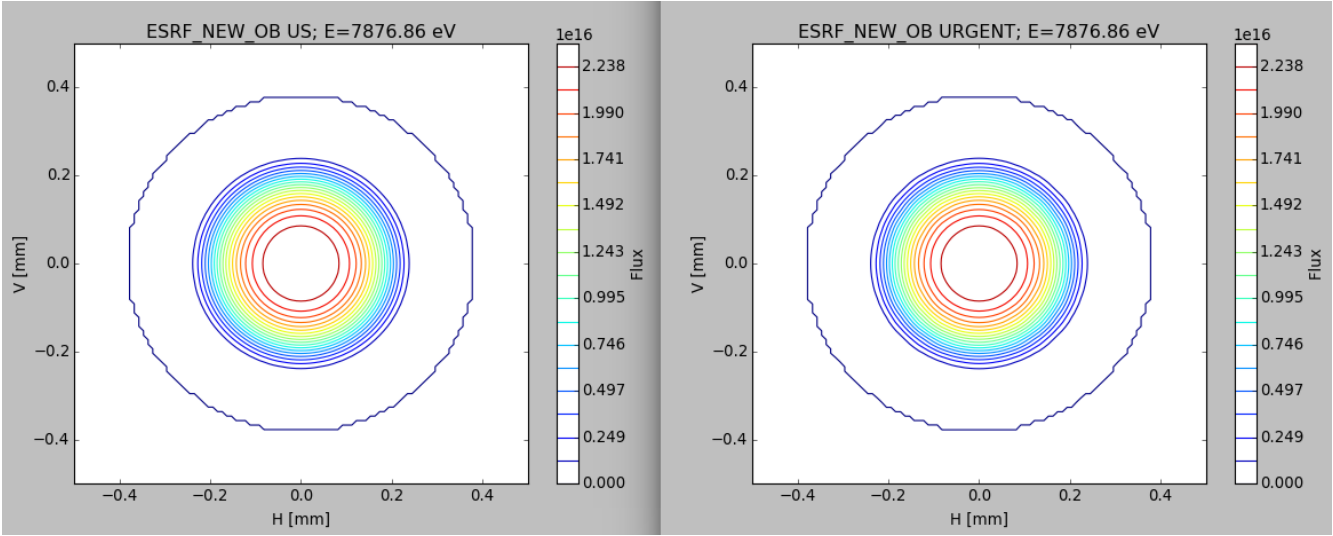
XRAY_BOOKLET



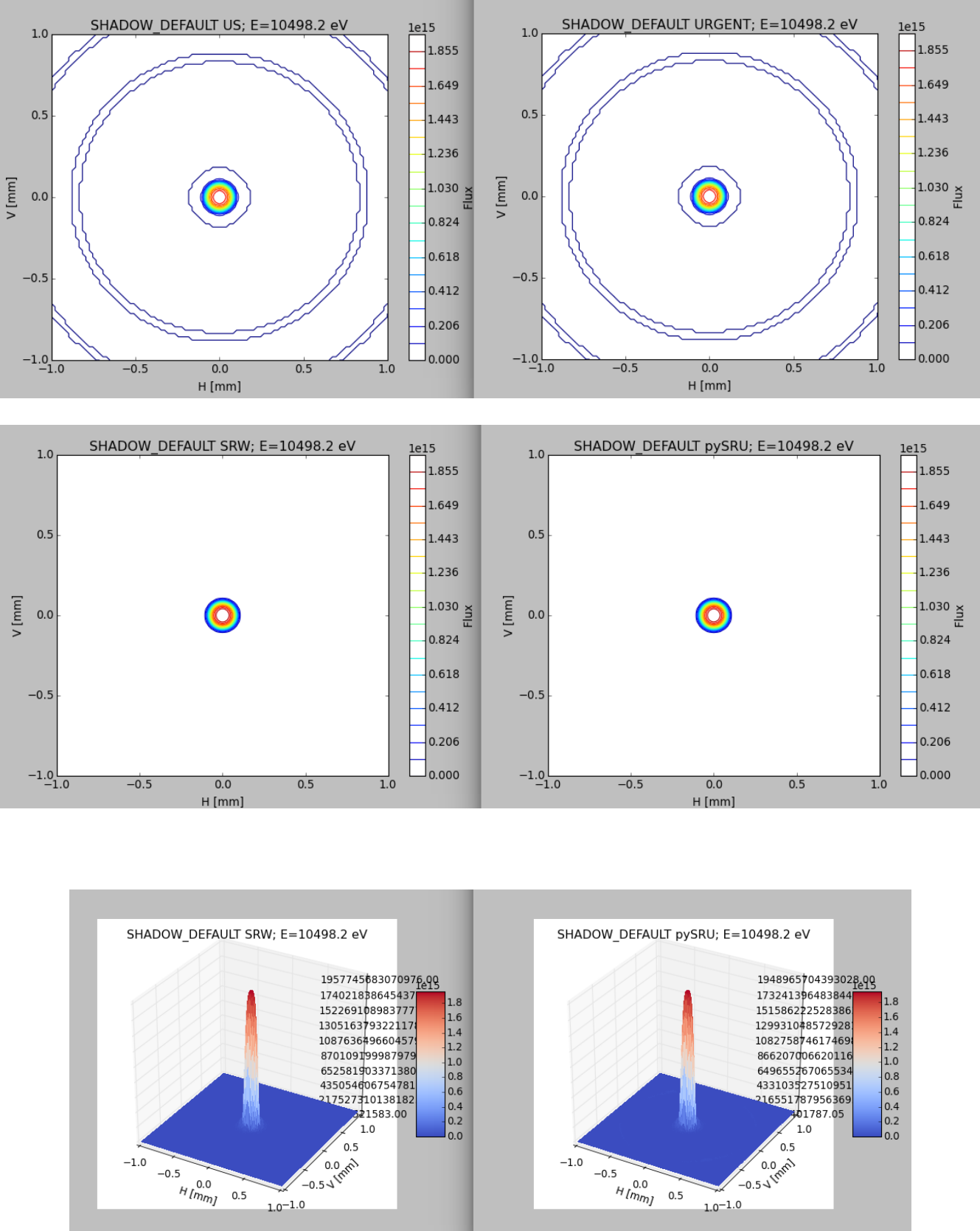
ID16_NA



ESRF_NEW_OB



SHADOW_DEFAULT

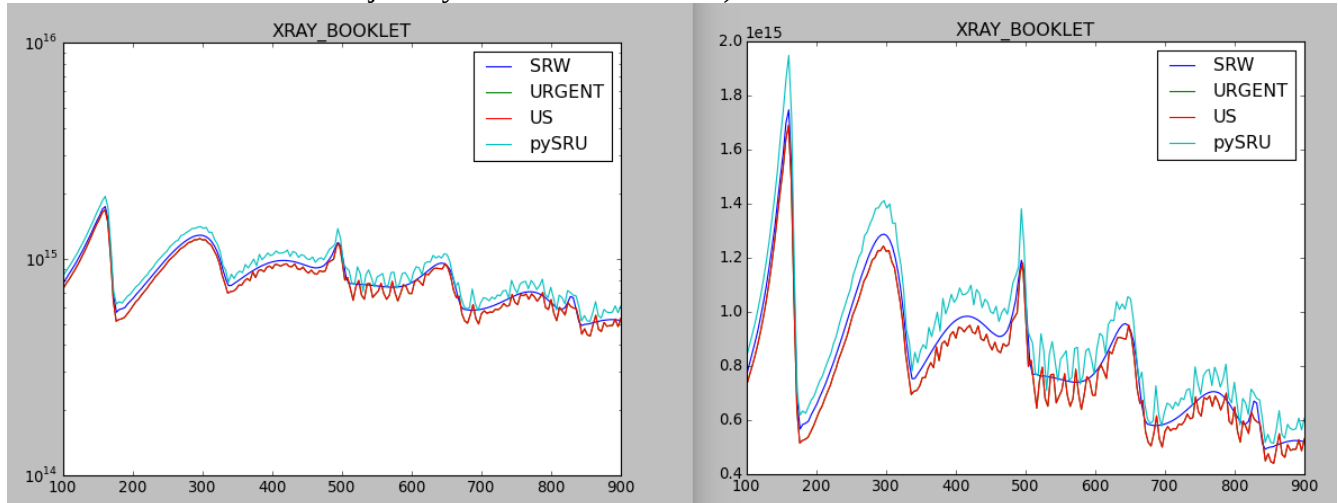


Energy spectrum

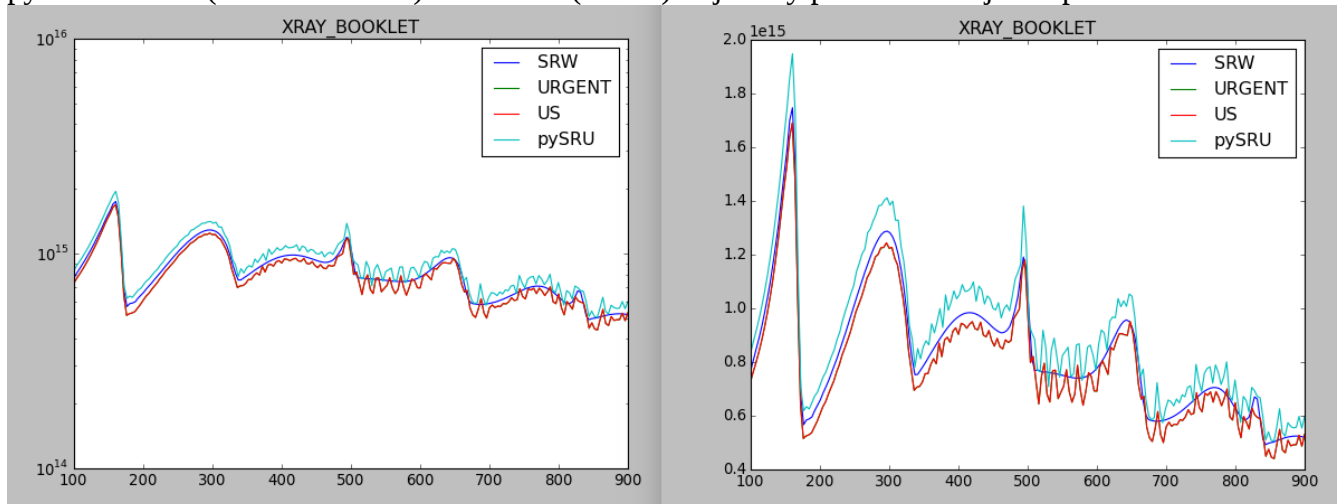
XRAY_BOOKLET

200 energy points. pySRU default (ODE+NEAR F)

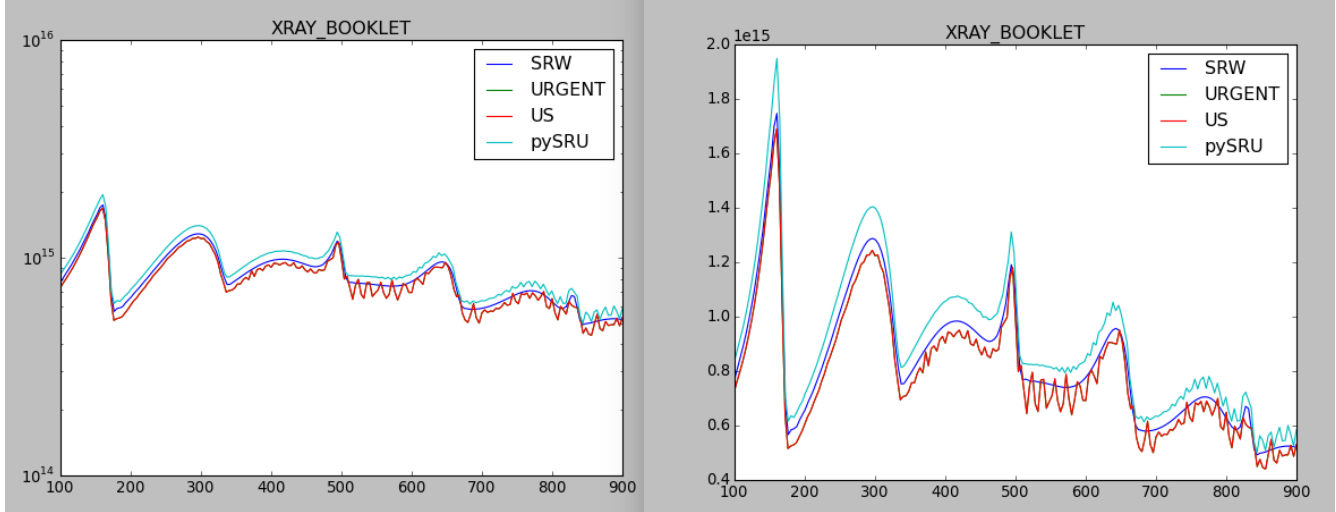
(the pySRU flux is a bit higher because it used the full magnetic field with edges added -ODE- and SRW uses the theoretical trajectory for flux calculations)



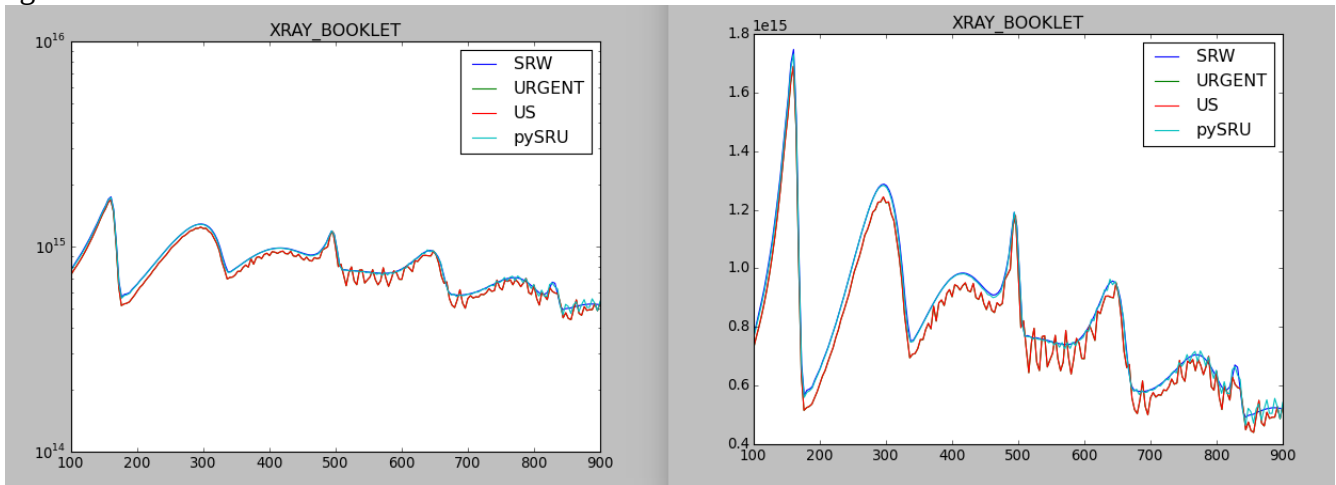
pySRU default (ODE+NEAR F) with more (20000) trajectory points: no major improvement



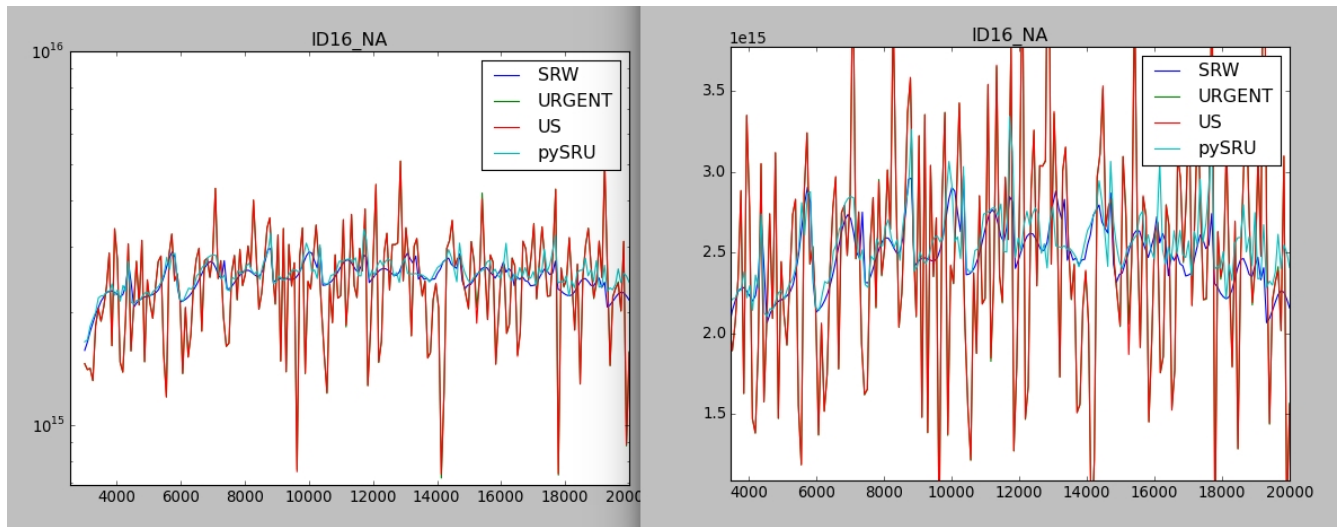
with more mesh points (101x101) we reduce the pySRU noise:



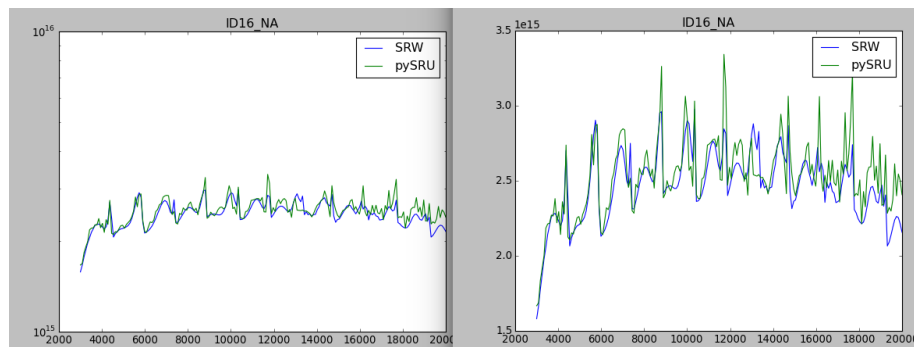
Using in pySRU analytical trajectory (N=14) and (101x101) we get a very good pySRU/SRW agreement:



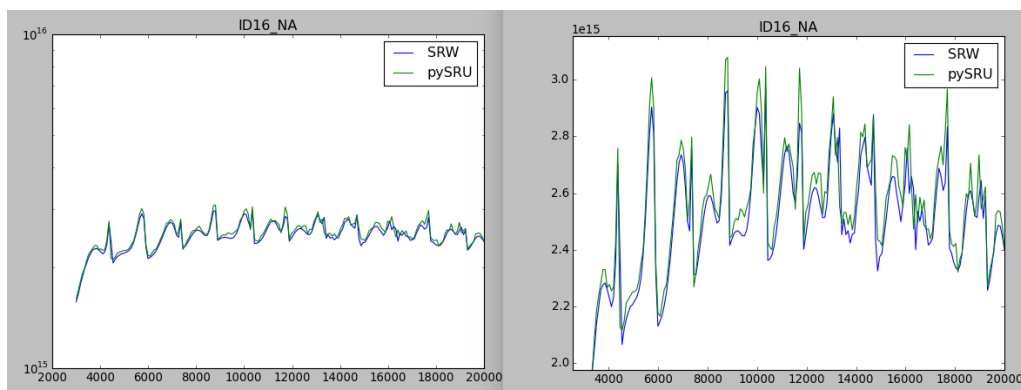
ID16_NA



US and URGENT fail to get something reasonable, at least for the inputs I used. This is the comparison pySRU and SRW only. Certainly pySRU needs more grid points, and SRW needs to increase a bit more the max harmonic number (used 31 here, before we used 21).

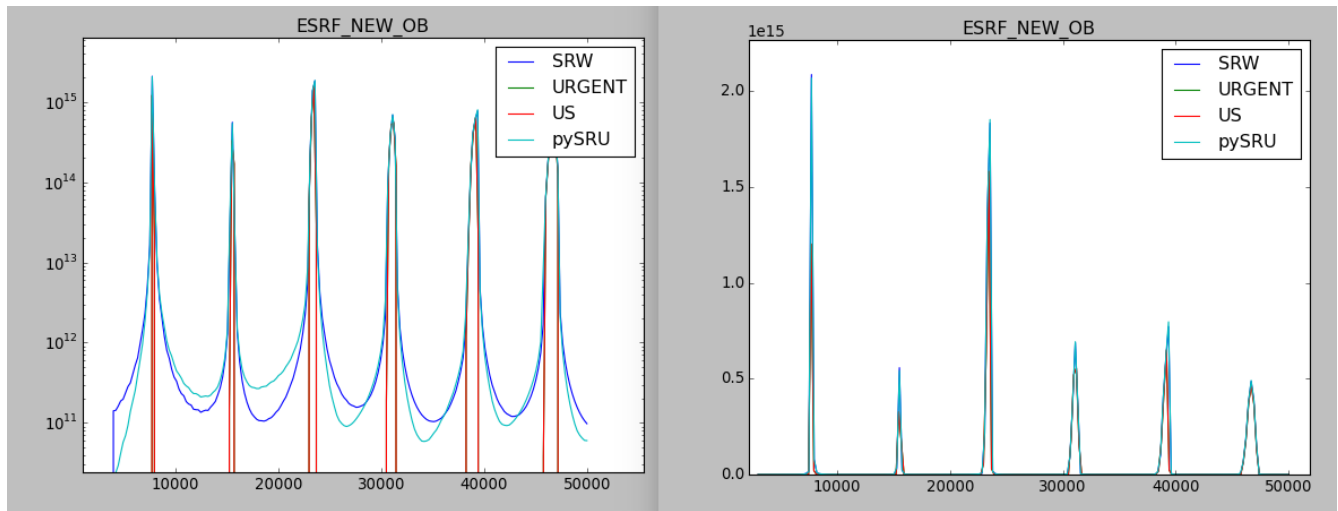


using up to 41 harmonics in SRW and 101x101 grid in pySRU we get:

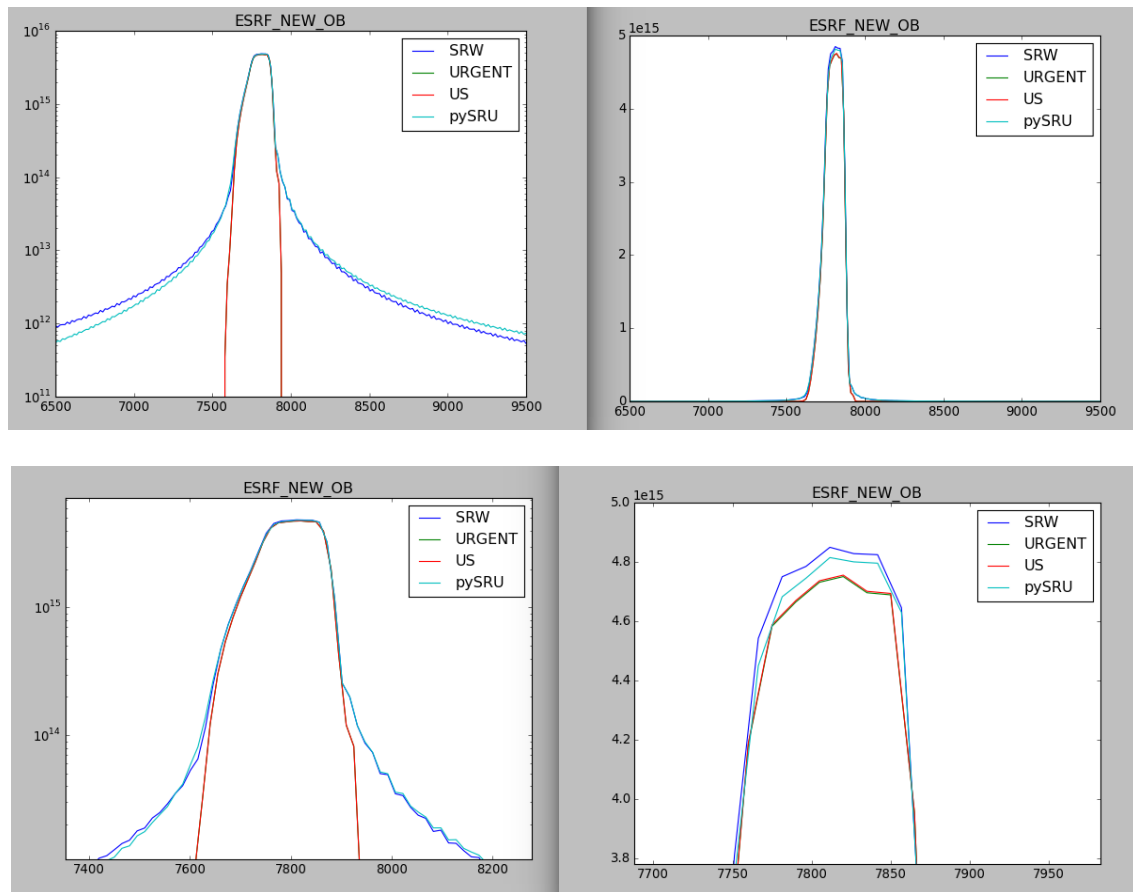


ESRF_NEW_OB

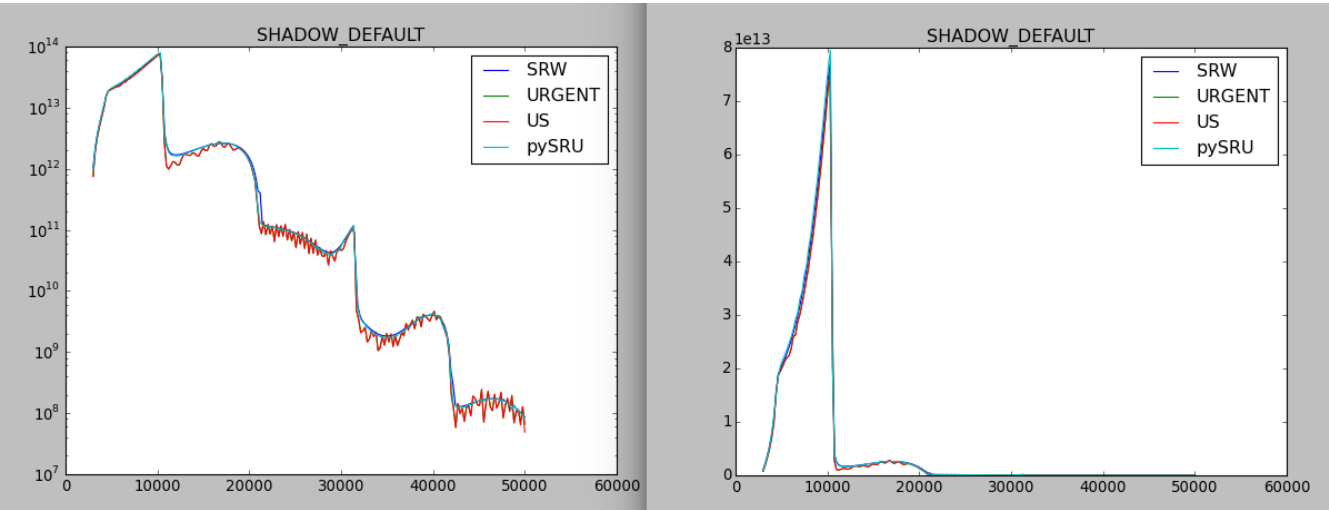
SRW Nmax=61; 101x101 mesh in pySRU:



First peak:



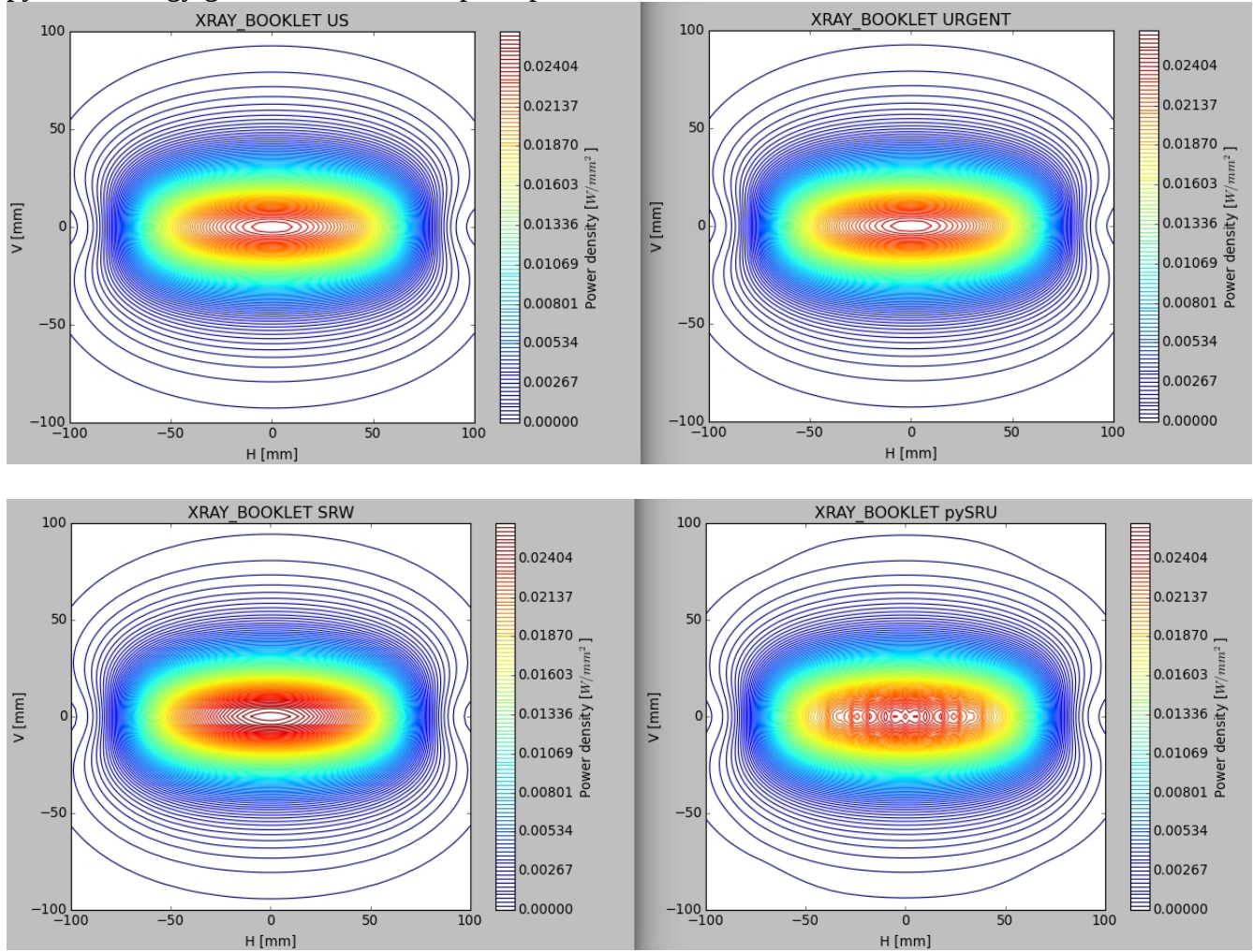
SHADOW_DEFAULT



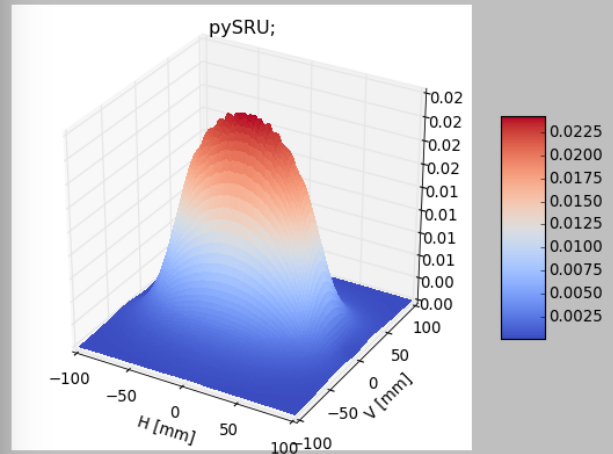
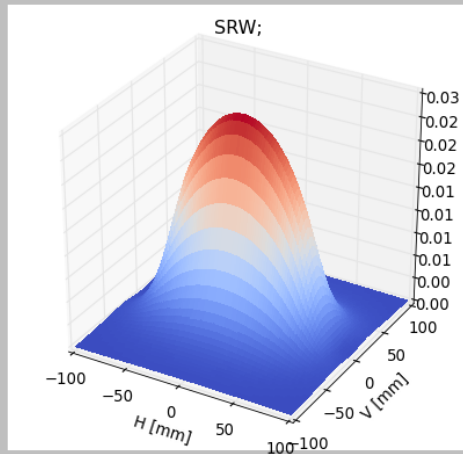
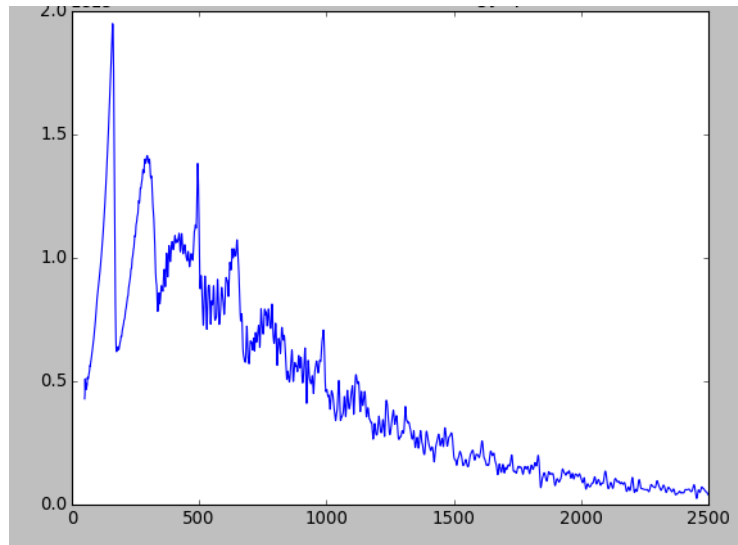
Power density

XRAY_BOOKLET

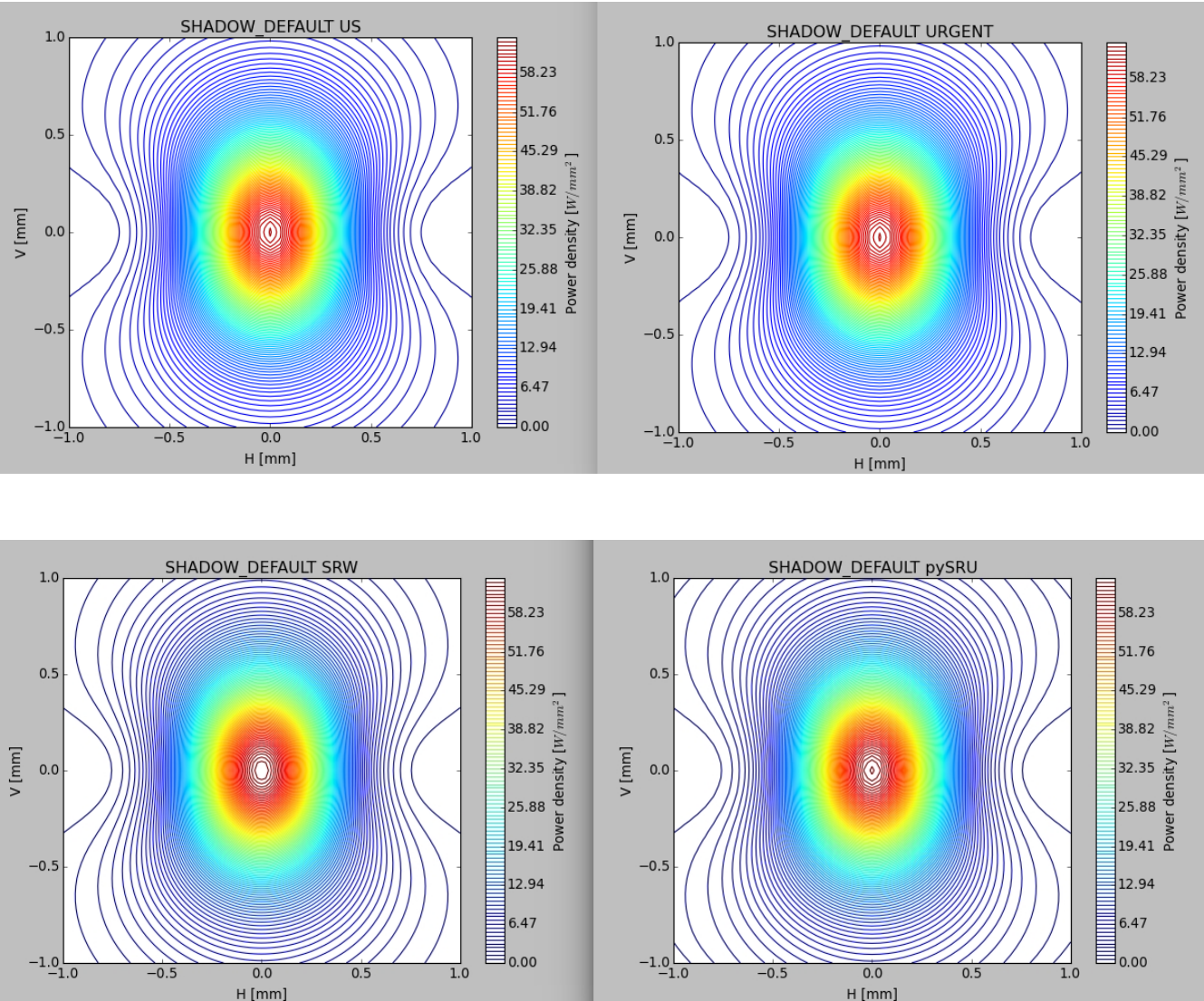
pySRU: Energy grid 90-2500eV one point per eV:



Spectrum by pySRU



SHADOW_DEFAULT



Spectrum by pySRU

