

# Example of Hierarchical Modeling: Estimation of Population and Cosmological Parameters of Binary Neutron Stars

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The posterior of  $\vec{\Omega}_P$  and  $\vec{\Omega}_C$  for  $i$ th NS-NS event [Thrane & Talbot],

$$P(\vec{\Omega}_P, \vec{\Omega}_C | D_i) \propto P(\vec{\Omega}_P, \vec{\Omega}_C) L_i(D_i | \vec{\Omega}_P, \vec{\Omega}_C)$$

Quasi Likelihood  $L_i$  is given by,

$$L_i(D_i | \vec{\Omega}_P, \vec{\Omega}_C) = \int_{M_{cz}, D_L} P(D_i | M_{cz}, D_L) P(M_{cz}, D_L | \vec{\Omega}_P, \vec{\Omega}_C) dM_{cz} dD_L$$

which is,

$$L_i(D_i | \vec{\Omega}_P, \vec{\Omega}_C) = \int_{M_{cz}, D_L} P(D_i | M_{cz}, D_L) P(D_L) P(M_{cz} | D_L, \vec{\Omega}_P, \vec{\Omega}_C) dM_{cz} dD_L$$

which involves **Assumption 1**:  $D_L$  and  $(\vec{\Omega}_P, \vec{\Omega}_C)$  are independent. Our model will come through  $z = z(D_L, \vec{\Omega}_C)$ , not by  $D_L = D_L(z, \vec{\Omega}_C)$ .

Now **Assumption 2** is,

$$P(M_{cz} | D_L, \vec{\Omega}_P, \vec{\Omega}_C) = \int_{M_c, z} P(M_{cz} | M_c, z) P(M_c | \vec{\Omega}_P) P(z | D_L, \vec{\Omega}_C) dM_c dz$$

Our models are,

$$P(M_{cz} | z, M_c) = \delta(M_{cz} - M_c(1 + z)) \quad (1)$$

$$P(M_c | \vec{\Omega}_P) = N_{M_c}(\vec{\Omega}_P) \quad (2)$$

$$P(z | D_L, \vec{\Omega}_C) = \delta(z - z(D_L, \vec{\Omega}_C)) \quad (3)$$

with those prior distributions,  $P(M_{cz} | D_L, \vec{\Omega}_P, \vec{\Omega}_C)$  becomes,

$$P(M_{cz} | D_L, \vec{\Omega}_P, \vec{\Omega}_C) = \frac{1}{1 + z(D_L, \vec{\Omega}_C)} N\left(\frac{M_{cz}}{1 + z(D_L, \vec{\Omega}_C)}\right) (\vec{\Omega}_P)$$

Hence our Quasi-Likelihood becomes,

$$L_i(D_i | \vec{\Omega}_P, \vec{\Omega}_C) = \int_{M_{cz}, D_L} P(D_i | M_{cz}, D_L) P(D_L) \frac{1}{(1 + z(D_L, \vec{\Omega}_C))} N\left(\frac{M_{cz}}{1 + z(D_L, \vec{\Omega}_C)}\right) (\vec{\Omega}_P) dM_{cz} dD_L$$

The posterior distribution of  $(\vec{\Omega}_P, \vec{\Omega}_C)$  is given by,

$$P(\vec{\Omega}_P, \vec{\Omega}_C | \bigcap_i D_i) = P(\vec{\Omega}_P, \vec{\Omega}_C) L_i(\bigcap_i D_i | \vec{\Omega}_P, \vec{\Omega}_C) = P(\vec{\Omega}_P, \vec{\Omega}_C) \prod_i L_i(D_i | \vec{\Omega}_P, \vec{\Omega}_C)$$

Hence for  $N$  events the posterior distribution of  $(\vec{\Omega}_P, \vec{\Omega}_C)$  is given by,

$$P(\vec{\Omega}_P, \vec{\Omega}_C | \bigcap_i D_i) = P(\vec{\Omega}_P, \vec{\Omega}_C) \prod_i \int_{M_{cz}, D_L} P(D_i | M_{cz}, D_L) P(D_L) \frac{1}{(1 + z(D_L, \vec{\Omega}_C))} N\left(\frac{M_{cz}}{1 + z(D_L, \vec{\Omega}_C)}\right) (\vec{\Omega}_P) dM_{cz} dD_L \quad (4)$$