Computational Methods in Stochastics

Lecture II

Sampling from Different Distributions aka Stochastic Simulation Logarithmic binning

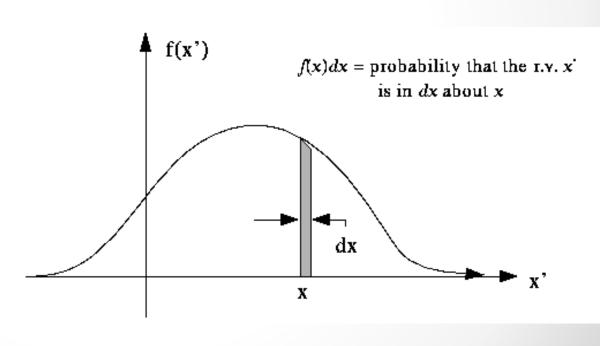
We want to understand the dynamics of a stochastic system.

- 1. We may want to simulate generate states from a stationary distribution
- 2. We may want to determine how an initial distribution of some (continuous) quantity evolves in time. In this case, we determine states from evolving distributions. If we want estimates of expected values, we must sample from many runs of the dynamics and average this is inevitably cruder than computing expected values from a stationary distribution.
- Both 1 and 2 are regarded as stochastic dynamics.
- All models **in this context** are just distributions. All model parameters $\theta = \{\theta_1, \theta_2, ..., \theta_N\}$ are parameters characterising these distributions, for example, mean and variance.

We have a random variable X with probability density function (PDF) f(x) (or $f_X(x)$). We wish to evaluate E(g(X)) for some function $g(\cdot)$: $E(g(X)) = \int_X g(x) f(x) dx$.

Here, think of f(x)dx as the **measure** you are working with: you measure everything with respect to this probability measure.

Think of g(x) as a curve in the dimension orthogonal to x and f(x). Then g(x)f(x)dx is a volume in these coordinates = "probability density element" dp.

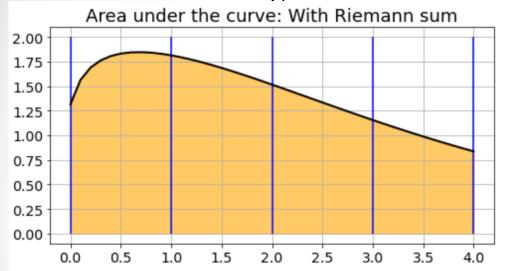


We resort to **Monte Carlo integration**:

We *simulate* realisations $x_1, ..., x_n$ of X and form realisations of the random variable g(X) as $g(x_1), ..., g(x_n)$. Then, provided that the variance of g(X) is finite, the law of large numbers assures that we can approximate the integral by

$$E(g(X)) \simeq \frac{1}{n} \sum_{i=1}^{n} g(x_i)$$
. (x_i are drawn from f)

Basic numerical integration:

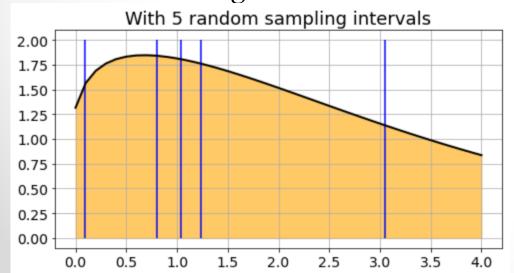


one sums up, for instance, values at midpoints of intervals.

numerical integration

In (non-random)

Monte Carlo integration:



In Monte Carlo integration one takes randomly samples of the area/volume under/inside the curve. MC integration is more effective in dimension 3 or greater.

Monte Carlo integration:

$$E(g(X)) \simeq \frac{1}{n} \sum_{i=1}^{n} g(x_i)$$
. (x_i are drawn from f)

Even if we cannot simulate realisations of X, but can simulate realisations of Y that has PDF $h(\cdot)$, then

$$E(g(X)) = \int_{X} g(x)f(x)dx = \int_{X} \frac{g(x)f(x)}{h(x)}h(x)dx$$

$$\Rightarrow E(g(X)) \simeq \frac{1}{n} \sum_{i=1}^{n} \frac{g(y_{i})f(y_{i})}{h(y_{i})}.$$
 This is importance sampling.

 $(y_i drawn from h)$

Once more to make this clear: We have PDF $f(\cdot)$ that is hard to simulate. So, we simulate $h(\cdot)$ that is easier and use the realisations of Y (samples drawn from that distribution) and compute the values of f(y) and h(y) in the process of computing the expectation - we change the measure.

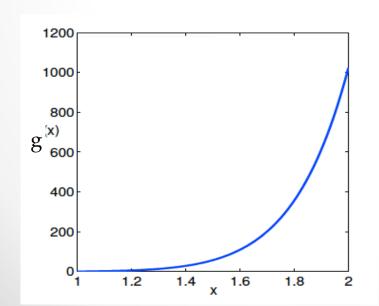
$$E(g(X)) = \int_{X} g(x)f(x)dx = \int_{X} \frac{g(x)f(x)}{h(x)}h(x)dx$$

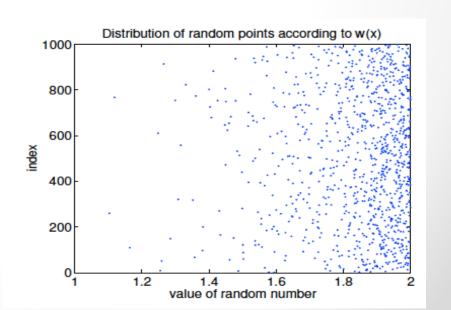
$$\Rightarrow E(g(X)) \simeq \frac{1}{n} \sum_{i=1}^{n} \frac{g(y_{i})f(y_{i})}{h(y_{i})}. \quad (y_{i} \text{ drawn from } h)$$

This may also be written as $E(g(X)) = \frac{1}{n} \sum_{i=1}^{n} g(y_i) w(y_i)$,

where $w(y_i) = \frac{f(y_i)}{h(y_i)}$ is a weight function if $\int_X w(x) dx = 1$.

h(x) should be as similar to f(x) as possible to make the integration computationally effective: this way sampling is done primarily in the regions that contribute most to $E(g(X)) \rightarrow$ weighted integration.





A prerequisite: Change of variable

The mathematical foundation for transformations of (random) variates reads as follows:

The PDF of an arbitrary differentiable invertible transformation Y = g(X) is

$$f_Y(y) = f_X(g^{-1}(y)) \left| \frac{d}{dy} g^{-1}(y) \right|.$$

The last term is the Jacobian of the transformation (in dimension 2 or greater). (For an example in 2D, see the Box-Müller method, page 16.)

"Equal measures": $f_Y(y)dy = f_X(g^{-1}(y))d[g^{-1}(y)].$

"Proof". The event $\{Y \le y\}$ is the same as the event $\{X \le g^{-1}(y)\}$

$$\Rightarrow F_Y(y) = \Pr\{Y \le y\} = \Pr\{X \le g^{-1}(y)\} = F_X(g^{-1}(y))$$

$$y = g(x) \Leftrightarrow x = g^{-1}(y)$$

$$\frac{dg^{-1}}{dy} = \frac{dx}{dy} = \frac{1}{g'(x)}$$

$$f_{Y}(y) = \frac{dF_{Y}(y)}{dy} = \frac{dF_{X}(g^{-1}(y))}{dy} = \frac{dF_{X}(x)}{dx} \frac{dx}{dy} = \frac{1}{g'(x)} f_{X}(x)$$
$$= \frac{dg^{-1}(y)}{dy} f_{X}(g^{-1}(y)) \blacksquare$$

The differentiation $\frac{dg^{-1}(y)}{dy}$ generalises to a Jacobian in multiple dimensions.

The inverse distribution method

When the inverse transformation can be determined, the *inverse* distribution method can – should, for efficiency, - be used:

- 1. Sample y from a uniform distribution $y \in [0,1]$ (notation $Y \sim U_{[0,1]}$, or $Y \sim U(0,1)$).
- 2. Compute $x = F^{-1}(y)$, where $F(x) = \Pr\{X \le x\}$. (If $x \in [a, b]$, then $F(x) = \int_a^x p(x')dx'$, or the sum $F(x) = \sum_{a \le x_i \le x} p(x_i)$ in the discrete case.) X follows the distribution F(x) as desired.

Proposition. If $Y \sim U(0,1)$ and $F(\cdot)$ is a valid invertible cumulative distribution function (CDF), then $X = F^{-1}(Y)$ has CDF $F(\cdot)$. *Proof.*

$$\Pr(X \le x) = \Pr(F^{-1}(Y) \le x) = \Pr(Y \le F(x)) = F_Y(F(x)) = F(x).$$

Example. *Uniform random variates*. Given $U \sim U(0, 1)$, simulate $V \sim U(a, b)$, where a < b.

Now,
$$F^{-1}(u) = v = a + (b - a)u$$
. $\left(F(v) = \frac{v - a}{b - a}, \quad a \le v \le b \right)$

- 1. Sample $u \in [0,1]$.
- 2. Compute v from v = a + (b a)u.

Example. Exponential random variates. Given $U \sim U(0, 1)$, simulate $X \sim Exp(\lambda)$.

PDF
$$f(x) = \lambda e^{-\lambda x}, x \ge 0. \Rightarrow \text{CDF } F(x) = 1 - e^{-\lambda x}, x \ge 0.$$

$$\Rightarrow F^{-1}(u) = x = -\frac{1}{\lambda} \ln(1 - F(x)) = -\frac{1}{\lambda} \ln(1 - u).$$

- 1. Sample $u \in [0,1]$.
- 2. Compute x from $x = -\frac{1}{\lambda} \ln(1 u)$.

Example. Variates from the Lorentzian distribution.

PDF
$$p(x) = \frac{1}{\pi} \frac{1}{1 + x^2}$$
 $(-\infty < x < \infty)$

- 1. Sample $y \in [0,1]$.
- 2. Compute *x* from

$$y = F(x) = \int_{-\infty}^{x} \frac{1}{\pi} \frac{1}{1 + x'^2} dx' = \frac{1}{2} + \frac{1}{\pi} \arctan(x)$$

$$\Rightarrow x = F^{-1}(y) = \tan\left[\pi(y - \frac{1}{2})\right]$$

Learn to do this "without thinking", that is, learn the "notation" and derive F^{-1} using it. It's easy to get mixed up, if you think too much! Assignments \rightarrow

When playing around with distributions it is often useful to identify factors contributing to *scaling* and *location*. In the prototypical case, if Y has PDF f(y) and CDF F(y), and X = aY + b (affine relation), then X has CDF

$$F_X(x) = \Pr(X \le x) = \Pr(aY + b \le x) = \Pr\left(Y \le \frac{x - b}{a}\right) = F\left(\frac{x - b}{a}\right)$$

and PDF

$$f_X(x) = \frac{1}{a} f\left(\frac{x-b}{a}\right).$$

Here, a is the scale parameter and b is the location parameter.

Example. *Gamma random variates.* Simulate $X \sim Ga(n, \lambda)$ random variates for integer n.

Exact inverse $F^{-1}(\cdot)$ does not exist, but use the property that if $Y_i \sim Exp(\lambda)$, and the Y_i are independent, then

$$X = \sum_{i=1}^{n} Y_i \backsim Ga(n, \lambda).$$

In practise, to simulate gamma random variates, simulate exponential random variates and sum them up. $(Ga(1, \lambda) = \text{Exp}(\lambda))$, see Lecture 1) Note: sum of memoryless processes \rightarrow Gamma.

n is the *shape parameter* and λ is the *scale parameter*.

Often library algorithms generate Ga(n, 1), after which one needs to rescale: if $Y \sim Ga(n, 1)$, then $X = Y/\lambda \sim Ga(n, \lambda)$.

Normal random variates

Obviously, efficient simulation of Gaussian random quantities is of great importance. However, the inverse transformation cannot be applied.

We need a technique for simulating $Z \sim N(0, 1)$ random variables, because $X = \mu + \sigma Z \sim N(\mu, \sigma^2)$.

CLT-based method

This approximate method makes use of the central limit theorem.

For example,
$$Z = \sum_{i=1}^{12} U_i - 6$$
, where $U_i \backsim U(0,1), i = 1, 2, ..., 12$

are independent, is <u>approximately</u> normal with E(Z) = 0 and Var(Z) = 1.

Normal random variates

This example case for the CLT-based method has support on [-6, 6] and is poorly behaved in the extreme tails. For $Z \sim N(0,1)$, $P(|Z| > 6) \approx 2 \times 10^{-9}$, so the truncation is not a major problem in many applications. The method is, however, slow: 12 uniform random numbers to get one Gaussian.

Remember: In CLT $Z_n = \frac{\bar{X}_n - \mu}{\sigma/\sqrt{n}}$; then as $n \to \infty$ $Z_n \sim N(0,1)$.

Normal random variates

The Box-Muller method - this is the standard method

- fast and efficient way to generate normal random variates
- 1. Simulate $\Theta \sim U(0,2\pi)$ and $R^2 \sim Exp(1/2)$ independently.
- 2. Compute two independent standard normal random variables $X = R \cos \Theta$ and $Y = R \sin \Theta$.

Proof.
$$f_{X,Y}(x,y) = \frac{1}{2\pi} \exp(-(x^2 + y^2)/2)$$

If $X = R \cos \Theta$ and $Y = R \sin \Theta$,

Box-Müller is an inverse distribution transformation:

$$F(x) = \iint_0^{2\pi} \frac{1}{2\pi} r' e^{-r'^2} dr' d\theta = 1 - e^{-r^2/2} = u_1$$

$$\Rightarrow r = \sqrt{-2 \ln u_1}$$
 and $\theta = 2\pi u_2$

For N(0, 1) distribution.

The Box-Muller algorithm:

- 1. Generate $u_1, u_2 \sim U(0, 1)$ independently.
- 2. Compute $r = \sqrt{-2 \ln u_1}$ and $\theta = 2\pi u_2$
- 3. Compute $x = r \cos(\theta)$ and $y = r \sin(\theta)$
- 4. Multiply x and y by σ_X and σ_Y to get the desired variances σ_X^2 and σ_Y^2 . Add the desired μ .
- 5. Go to 1.

You can use this stuff as an aid when doing Assignment 2.1.

Lookup Methods

The lookup method is **the discrete version of the inverse transformation method**. It is used in simulating a discrete random quantity X with outcome space $S = \{0, 1, 2, ...\}$.

The probability mass function (PMF) of X: $p_k = \Pr\{X = k\}$, k = 0, 1, 2, ...

Define
$$q_k = \Pr\{X \le k\} = \sum_{i=0}^k p_i$$
.

To generate realisations of X, first simulate $u \sim U(0, 1)$, then compute $X = \min\{k | q_k \ge u\}$.

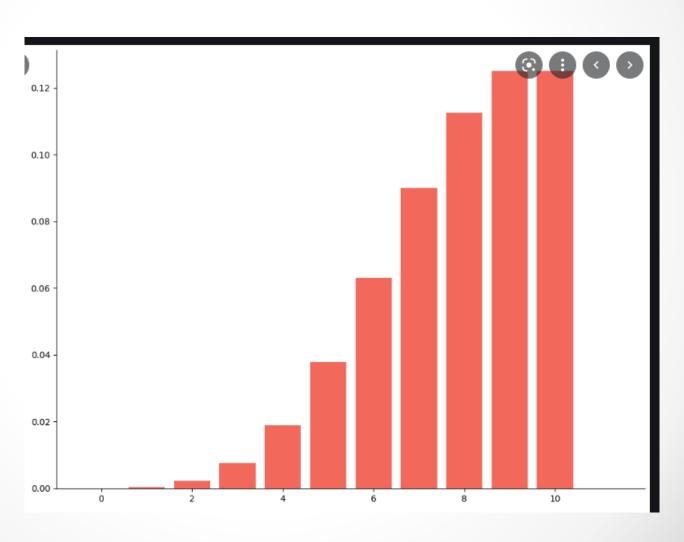
$$\Rightarrow q_{k-1} < u \le q_k$$
, so

$$\Pr\{X = k\} = \Pr\{u \in (q_{k-1}, q_k]) = q_k - q_{k-1} = p_k.$$

(This method is used e.g. for simulating discrete Markov models.)

Lookup Methods

$$q_k = \Pr\{X \le k\}$$



Rejection samplers are used in generating random numbers that follow an arbitrary distribution p(x) that is not analytically invertible.

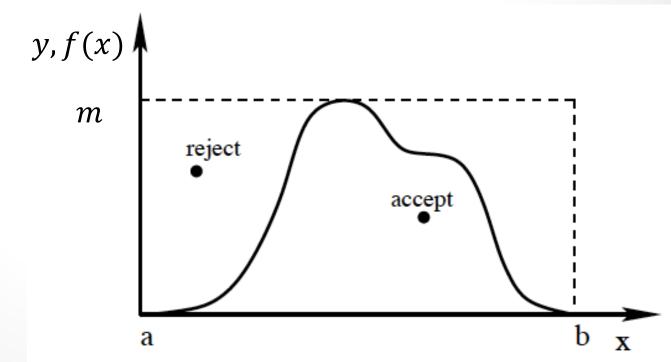
Basic method

We want to simulate from f(x) with finite support on [a, b]. If we can determine m such that $f(x) \le m$, $\forall x \in [a, b]$, we can simulate $X \sim U(a, b)$ and $Y \sim U(0, m)$ and accept X, if Y < f(X), otherwise reject and try again. The accepted X values have PDF f(x).

(For a proof, see e.g. Proposition 4.4. in Stochastic Modelling for Systems Biology.)

The algorithm:

- 1. Generate x in the support of X: $x \in [a, b]$.
- 2. Generate $y \sim U(0, 1)$: if $y \leq f(x)/m$, accept x. (This means: "accept x with probability f(x)/m".)
- 3. Go to 1.



The acceptance probability for this method is

$$\Pr\{\text{Accept}\} = \Pr\{(X, Y) \in A\}$$

$$= \int_{a}^{b} \Pr\{(X, Y) \in A | X = x\} \times \frac{1}{b-a} dx = \int_{a}^{b} \frac{f(x)}{m} \times \frac{1}{b-a} dx$$

$$= \frac{1}{m(b-a)} \int_{a}^{b} f(x) dx = \frac{1}{m(b-a)}.$$

If Pr{Accept} is very low, one can use the *envelope method*.

The envelope method

This method also extends application of the rejection method to distributions with **infinite support**. Decent computational efficiency can be obtained choosing the enveloping region carefully.

To simulate X with PDF $f(\cdot)$ we choose a PDF $h(\cdot)$ such that

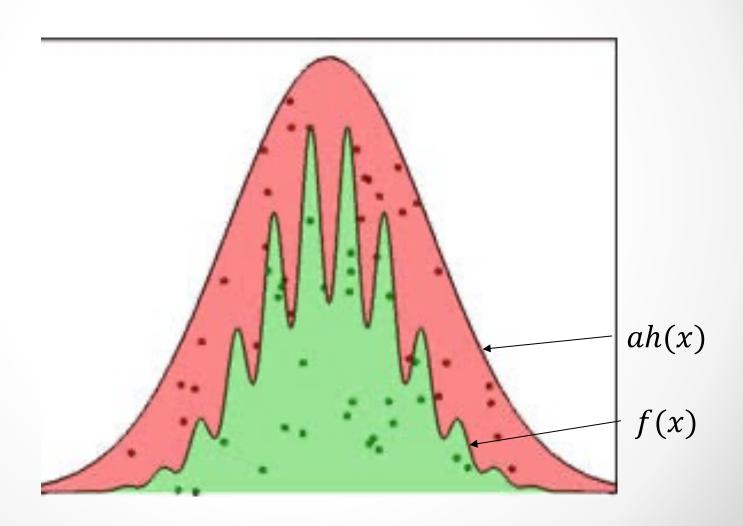
$$f(x) \le ah(x), \forall x$$

and for which we can simulate values of *Y* with the same support as *X*.

a is an upper bound for f(x)/h(x); $a \ge 1$, since both f(x) and h(x) integrate to 1.

Rejection Samplers

The envelope method.



The envelope algorithm.

- 1. Draw Y = y from $h(\cdot)$ and then $U = u \sim U(0, ah(y))$.
- 2. Accept y as a simulated value of X if u < f(y) (with probability f(y)/[ah(y)]*, otherwise reject.
- 3. Go to 1.
 - (* "with probability f(y)/[ah(y)]" means: accept y if z < f(y)/[ah(y)], where $z \sim U[0,1]$)

This procedure distributes points uniformly over a region covering f(x) and keeps ones that lie under f(x).

The overall acceptance probability:
$$\Pr\{U < f(Y)\} = \int_{-\infty}^{\infty} \Pr\{U < f(Y)\} | Y = y\} h(y) dy = \int_{-\infty}^{\infty} \frac{f(y)}{a} dy$$
$$= \frac{1}{a}. \quad \to a \text{ should be as close to 1 as possible.}$$
 (In practise, a should be smaller than 10.)

This stochastic simulation method sort of spawns from importance sampling and the envelope rejection method.

The advantage over the envelope rejection method is that one does not have to "guess" a good envelope and bounding constant.

Instead, any proposal distribution $h(\cdot)$ having the same support as $f(\cdot)$ can be used. In practise, the method works the better the more similar $h(\cdot)$ is to $f(\cdot)$.

Unlike importance sampling, the importance resampling is approximate: the samples are only approximately from $f(\cdot)$. The approximation improves with increasing number of generated samples.

Here, we rewrite the expectation of an arbitrary function in the importance sampling

$$E(g(X)) \simeq \frac{1}{n} \sum_{i=1}^{n} \frac{g(y_i)f(y_i)}{h(y_i)}$$

as

$$E(g(X)) \simeq \frac{1}{n} \sum_{i=1}^{n} w_i g(y_i),$$

where $w_i = f(y_i)/h(y_i)$.

The procedure:

- 1. generate samples from the proposal $h(\cdot)$
- 2. resample from the sample using the weights w_i .
- \rightarrow the new sample is distributed *approximately* according to $f(\cdot)$.

The algorithm:

- 1. Sample $y_1, y_2, ..., y_n \sim h(\cdot)$.
- 2. Compute the weights $w_k = f(y_k)/h(y_k)$, k = 1, 2, ..., n.
- 3. Compute the sum of the weights $w_0 = \sum_{j=1}^n w_j$.
- 4. Compute the normalised weights $w'_k = w_k/w_0$, k = 1, 2, ..., n.
- 5. Sample n times with replacement from the set $\{y_1, y_2, ..., y_n\}$ using the probabilities $\{w'_1, w'_2, ..., w'_n\}$ (using e.g. the lookup method) to generate a new sample $\{x_1, x_2, ..., x_n\}$.
- 6. Return the new sample $\{x_1, x_2, ..., x_n\}$ as an approximate sample from $f(\cdot)$.

Note: "Sample with replacement" means that if you pick for example y_3 , you "mentally" put it back to the set, so you may pick it again.

So, samples from $h(\cdot)$ are used as if they were samples from $f(\cdot)$, only they are re-weighted by w_i . So, generate samples from the proposal $h(\cdot)$ and resample from the sample using the weights w_i . One more effort to explain the resample part: Once you have the list $\{y_1, y_2, ..., y_n\}$, you resample m times from this list using the weights $\{w_1, w_2, ..., w_m\}$ as probabilities to get a new list $\{x_1, x_2, ..., x_n\}$ and you use this as an *approximate* sample from the original $f(\cdot)$. Because there is replacement, you might have more than one instances of the same y_i . What's the point? To generate several approximate samples from the original – that is, different instances. To see the "sense" in this method, one might think of using the limited sample as if there were a number of them.

One should try to find $h(\cdot)$ as closely reminiscent to $f(\cdot)$ as possible.

After all this fancy talk: Importance resampling just uses the already generated ensemble of samples and takes "new" batches of samples from it. This is done when the number of available samples is limited. This is all you need to remember of it in addition to the fact that such a trick exists; look into it and use it when you need it.

The formal description of this method is a typical example of how some people in computer science like to present things: Be so formal that no one will understand how simple - and of how limited use - the trick you are presenting is.

In order to analyse/identify distributions and make them more "presentable" one often needs to **bin the data**. The data is assorted to intervals along the abscissa (the first axis). There are two ways to bin the data:

1. Linear binning: The range [a, b] in the abscissa is divided into L intervals of equal width $\Delta = (b - a)/L$. The data element x_j belongs to the ith bin, if $x_j \in [a + (i - 1)\Delta, a + i\Delta)$. After assorting the data there will be N_i elements in the ith bin. One point represents the data in each bin. The binned data:

$$\check{x}_i = a + \frac{\Delta}{2} + (i - 1)\Delta \quad ;$$

$$y_i = \frac{N_i}{\Delta}$$
, where N_i is the number of $x_j \in [a + (i-1)\Delta, a + i\Delta)$

2. Logarithmic binning: The range R in the abscissa is divided into intervals of width $\Delta_i = \exp(i/r) - \exp[(i-1)/r]$, where $i \in \{-R, -(R-1), ..., R-1, R\}$. R and r define the range and resolution, respectively. The data element x_j belongs to the ith bin, if $x_j \in [\exp\left[\frac{(i-1)}{r}\right]$, $\exp(i/r)$). After assorting the data there will be N_i elements in the ith bin. One point represents the data in each bin. The values of these points are calculated as

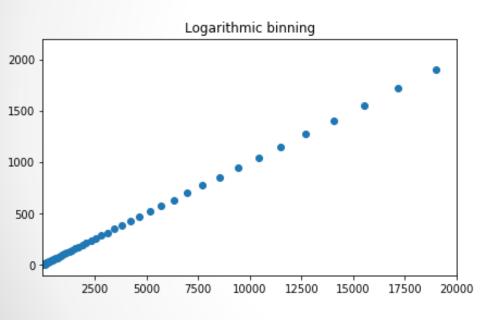
$$\check{x}_i = \frac{\exp[(i-1)/r] + \exp(i/r)}{2};$$

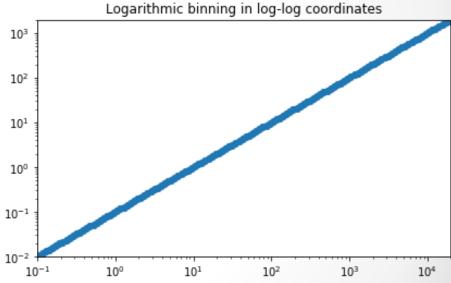
$$y_i = \frac{N_i}{N_i}, \text{ where } N_i \text{ is the number of } x_j \in \left[\exp\left[\frac{(i-1)}{r}\right], \exp(i/r)\right).$$

(You generate and plot points (\check{x}_i, y_i) to represent the data.)

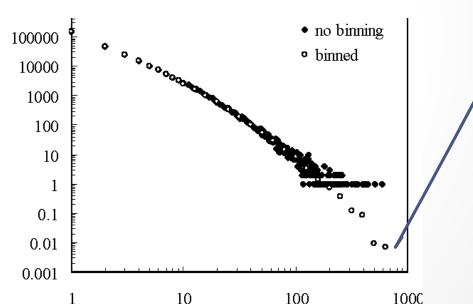
For the layman: The bins are really small for small *i* and grow exponentially with *i*. Plot in log-log coordinates.

Here are Δ_i vs interval midpoints in linear (left) and logarithmic (right) coordinates.





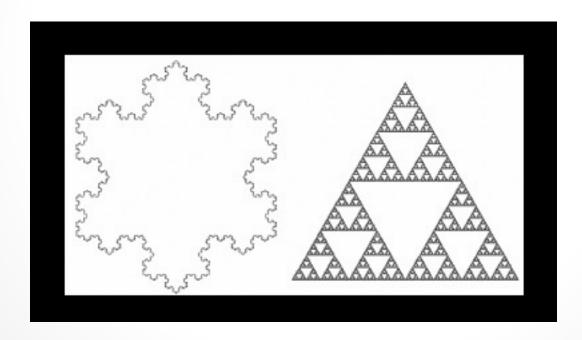
Log-binning allows you to see the functional dependence of a fat-tailed distribution (density) more precisely at small x. It also extends the visibility of the functional dependence to larger x values (for real data).



To avoid hassle and have faster algorithms we use python histograms and tune the plotting a bit \rightarrow

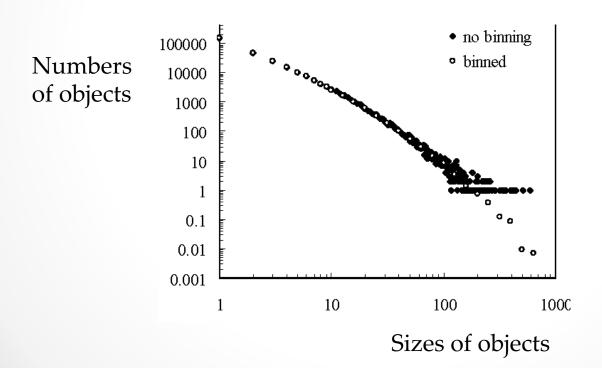
Power-Law Distributions

Log-binning is the way to determine exponents α of **power-law distributions**, whose PDFs are of form $f(x) \propto x^{-\alpha}$. Such distributions are the hall-mark of scale-invariance: things look the same at different (length) scales (e.g. fractals).



Power-Law Distributions

In the real world power-law distributions are usually accompanied by an exponential cut-off: $f(x) \propto x^{-\alpha} \exp(-x/A)$ – because the magnitude (the size of objects, communities, etc.) in the abscissa does not g to infinity.



About axes and plotting

When trying to understand a given data, plot in linear, semilog, and log-log axes to see what works.

In order to plot values at centres of the bins instead of histogram bars do the following in python first generate the bins, then compute the centre points, and plot.

For linear binning:

```
bins = np.linspace(start=START_VALUE, stop=STOP_VALUE, num=NUMBER_OF_BINS)
```

For logarithmic binning:

```
\label{eq:bins} \verb| bins = np.logspace(start=np.log10(START_VALUE), stop=np.log10(STOP_VALUE), num=NUMBER_OF_BINS)| \\
```

Edges \rightarrow **Centres** \rightarrow **Plot**:

```
y,edges = np.histogram(DATA, bins, density=True)
centers = 0.5*(edges[1:]+ edges[:-1])
plt.plot(centers,y,'*')
```