

# Homework 2: Motion

24-760 Robot Dynamics & Analysis  
Fall 2023

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**Note:** For homework submission, please submit the PDF of the written portion to “Homework 2” and a zipped folder of your Matlab code to “Homework 2 Programming” in Gradescope.

## Problem 1) Lunar Motion

Consider a simplified model of the motion of the earth and the moon. Attach a stationary coordinate frame to the center of the earth ( $s$ ), a frame with the same origin that rotates with the earth ( $e$ ), and a frame to the moon ( $m$ ). The axes of rotation are all aligned with each other and pointing in the  $+z$  direction of each frame. Assume the moon’s orbit around the earth is circular with radius  $l_m$ . The earth’s radius is  $r_e$  and the moon’s radius is  $r_m$ . The moon rotates about the earth at a rate of 1 revolution per 28 days, and about its own axis at a rate of 1 revolution per 28 days. The earth rotates about its own axis at a rate of 1 revolution per day. *Hint: Draw a figure to keep track of the different frames.*

**1.1)** Just consider the earth’s rotation to start. At time  $t$ , assume the earth is rotated so that the earth’s  $+x$  axis is aligned with the stationary  $-y$ . What is  $R_{se}$ ? What is  $g_{se}$ ? Use this configuration for the other parts of this question.

**1.2)** For a point  $q$  on the surface of the earth,  $q_e = [0, r_e, 0]^T$ , calculate the location of this point in the stationary frame using a rigid body transformation.

**1.3)** What is the body velocity of the earth’s rotation,  $V_{se}^b$ ? What is the spatial velocity  $V_{se}^s$ ?

**1.4)** Using that body and spatial velocity, what is the instantaneous velocity of the point  $q_e$  in the earth’s frame,  $v_{qe}$ ? What is the velocity in the stationary frame,  $v_{qs}$ ?

**1.5)** Now consider the position of the moon relative to the earth. Assume at time  $t$  that the moon is located at  $[l_m, 0, 0]^T$  in the stationary frame, with the moon’s  $x$  axis pointing to the earth. Calculate  $g_{sm}$ , then calculate  $g_{em}$  based on  $g_{sm}$  and  $g_{se}$ .

**1.6)** What is the body velocity of the moon’s motion,  $V_{sm}^b$ ?

**1.7)** Calculate the spatial velocity,  $V_{sm}^s$ , using an adjoint operation. What is special about  $v_{sm}^s$ ?

**1.8)** Based on the rotation of the earth and the orbit of the moon, how long is a lunar day on earth? That is, from the earth’s perspective, how long does the moon take to come back over the same spot on the surface of the earth?

# I. LUNAR MOTION

I.1 Given that;

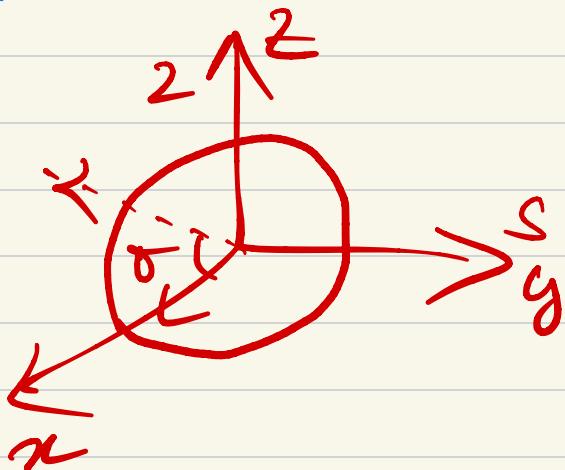
e = Rotates with the earth.

S = Stationary.

m = Frame of moon.

→ In this configuration; Frame e is aligned with frame S, but it is rotated by an angle  $\theta$  around the z-axis of frame S, since the earth's rotation axis is the z-axis of frame S.

→ The rotation angle ' $\theta$ ' can be calculated as  $\theta = \omega t$ , where  $\omega$  is the angular velocity of the earth's rotation and 't' is the time.



Now Rotation matrix R<sub>SE</sub> is;

$$R_{Z(\theta)} = \begin{bmatrix} \cos \theta & -\sin \theta & 0 \\ \sin \theta & \cos \theta & 0 \\ 0 & 0 & 1 \end{bmatrix}$$

However, the frame is rotated by  
 $\theta = 90^\circ$ ,

$$\therefore R_{SE} = \begin{bmatrix} 0 & +1 & 0 \\ -1 & 0 & 0 \\ 0 & 0 & 1 \end{bmatrix}$$

$3 \times 3$

Now, for  $\mathbf{g}_{SE}$ , which represents the acceleration due to gravity in Frame 'e'. Gravity points towards the center of the Earth and is represented as  $(0, 0, -g)$ .

However, the origins are same for both frames.

$$\therefore \mathbf{g} = \begin{bmatrix} R_{SE} : 0 \\ 0 \ 0 \ 0 : 1 \end{bmatrix}$$

$$\therefore \bar{\mathbf{g}}_{SE} = \begin{bmatrix} 0 & 1 & 0 & 0 \\ -1 & 0 & 0 & 0 \\ 0 & 0 & 1 & 0 \\ 0 & 0 & 0 & 1 \end{bmatrix}$$

$4 \times 4$

1.2 Given pt;  $\vec{q}_{ce} = \begin{bmatrix} 0 \\ \vec{r}_c \\ 0 \end{bmatrix}$

We can use a rigid body transformation.

Now, for  $\vec{q}_{cs} \hookrightarrow \vec{q}_{se} \vec{q}_{re}$

Now,

$$\vec{q}_{cs} = \begin{bmatrix} 0 & 1 & 0 & 0 \\ -1 & 0 & 0 & 0 \\ 0 & 0 & 1 & 0 \\ 0 & 0 & 0 & 1 \end{bmatrix} \begin{bmatrix} \vec{r}_e \\ \vec{r}_e \\ 0 \\ 1 \end{bmatrix} = \begin{bmatrix} \vec{r}_e \\ 0 \\ 0 \\ 1 \end{bmatrix}$$

$4 \times 4$        $4 \times 1$        $4 \times 1$

Wrong

$\vec{q}_{cs} = \begin{bmatrix} \vec{r}_e \\ 0 \\ 0 \end{bmatrix}$

In 's'-coordinate frame

1.3 To calculate the body velocity of the earth's rotation ( $\vec{v}_{se}^b$ ) and the spatial velocity ( $\vec{v}_{se}$ ) of a point on the earth's surface due to its rotation, we will consider the rotation of the earth.

$$\vec{v}_{se} = \vec{g}_{se}^{-1} \times \vec{g}_{se}$$

$$\vec{v}_{se}^b = \begin{bmatrix} R^{-1} & P^T \\ 0 & 1 \end{bmatrix} \begin{bmatrix} R & P \\ 0 & 0 \end{bmatrix} = \begin{bmatrix} RP & PP \\ 0 & 0 \end{bmatrix}$$

$$\vec{v}_k = \begin{bmatrix} \cos\theta & -\sin\theta & 0 \\ \sin\theta & \cos\theta & 0 \\ 0 & 0 & 1 \end{bmatrix} \quad 3 \times 3$$

1.4 Instantaneous Velocity of point  $qe$  in the earth's frame ( $\vec{v}_{qe}$ ):

$$\vec{v}_{qe} = \omega \times \vec{r}_{qe}$$

And,  $\vec{v}_{qs} \rightarrow$  Velocity of point  $qe$  in the stationary frame ('q/s')

$$\text{Now, } \vec{v}_{qs} = \vec{v}_{se}^n \times \vec{r}_{qe}$$

$$\vec{V}_{\text{rel}} \Rightarrow \begin{bmatrix} 0 & -\theta & 0 & 0 \\ \theta & 0 & 0 & 0 \\ 0 & 0 & 0 & 0 \\ 0 & 0 & 0 & 0 \end{bmatrix} \begin{bmatrix} \text{re} \\ 0 \\ 0 \\ 1 \end{bmatrix} = \begin{bmatrix} 0 \\ \theta e \\ 0 \\ 0 \end{bmatrix}$$

$$\vec{V}_{\text{rel}} = \begin{bmatrix} 0 \\ 2\pi \theta e \\ 0 \\ 0 \end{bmatrix}$$

Now,  $\vec{V}_{\text{rel}} = \vec{V}_{\text{ce}} + \vec{v}_{\text{ave}} = \begin{bmatrix} 0 & -\theta & 0 & 0 \\ \theta & 0 & 0 & 0 \\ 0 & 0 & 0 & 0 \\ 0 & 0 & 0 & 0 \end{bmatrix} \begin{bmatrix} \text{re} \\ 0 \\ 0 \\ 1 \end{bmatrix}$

$$\vec{V}_{\text{ave}} = \begin{bmatrix} -2\pi \theta e \\ 0 \\ 0 \\ 0 \end{bmatrix}$$

$\underline{\underline{g_{\text{sm}}} = [0, -g_{\text{moon}}, 0]}$  acceleration due to gravity on moon's surface.

Gravitational acceleration of the earth relative to the moon in the stationary frame, based on  $g_{\text{sm}} \approx g_{\text{ce}}$

The origin of the moon is only translated along x-axis only.

$$\therefore \bar{P}_{sm} = \begin{bmatrix} em \\ 0 \\ 0 \\ 1 \end{bmatrix}$$

When  $\theta = 90^\circ$ , (z-axis);  $R_{sm} = \begin{bmatrix} -1 & 0 & 0 \\ 0 & 1 & 0 \\ 0 & 0 & 1 \end{bmatrix}$

Now,  $\begin{bmatrix} R_{sm} \ P_{sm} \\ 0 \ 0 \ 0 \end{bmatrix}, \quad \downarrow \quad \} \Rightarrow g_{sm}$

$$\begin{bmatrix} -1 & 0 & 0 & em \\ 0 & 1 & 0 & 0 \\ 0 & 0 & 1 & 0 \\ 0 & 0 & 0 & 1 \end{bmatrix} = g_{sm}$$

LxH

Now, for  $g_{em} \Rightarrow g_{se} \times g_{sm}$

$$\begin{bmatrix} R^t : P \\ 0 \ 0 \ 0 : 1 \end{bmatrix} \Rightarrow g_{se} \times g_{sm}$$

$$\Rightarrow \begin{bmatrix} 0 & 1 & 0 & 0 \\ 1 & 0 & 0 & 0 \\ 0 & 0 & 1 & 0 \\ 0 & 0 & 0 & 1 \end{bmatrix} \begin{bmatrix} 1 & 0 & 0 & em \\ 0 & 1 & 0 & 0 \\ 0 & 0 & 1 & 0 \\ 0 & 0 & 0 & 1 \end{bmatrix}$$

LxH                                    LxH

$$\therefore g_{sm} = \begin{bmatrix} 0 & 1 & 0 & 0 \\ -1 & 0 & 0 & 0 \\ 0 & 0 & -1 & 0 \\ 0 & 0 & 0 & 1 \end{bmatrix}_{4 \times 4}$$

1.b Finding  $V_{sm}^b$ , we need to use

$$V_{sm}^b = g_{sm}^{-1} \times g_{sm}$$

$$\Rightarrow \begin{bmatrix} R_{sm}^T & p \\ 0 & 1 \end{bmatrix}_{4 \times 4} \begin{bmatrix} R_{sm} & p \\ 0 & 0 \end{bmatrix}_{4 \times 4} = \begin{bmatrix} R_{sm}^T R_{sm} & p.p \\ 0 & 1 \end{bmatrix}_{4 \times 4}$$

$$\text{Now, } \hat{P} = \begin{bmatrix} \hat{x} \\ \hat{y} \\ \hat{z} \end{bmatrix} = \begin{bmatrix} -cm \sin\theta \cdot \dot{\theta} \\ cm \cos\theta \cdot \dot{\theta} \\ 0 \end{bmatrix} = \begin{bmatrix} 0 \\ -cm\pi/14 \\ 0 \end{bmatrix}$$

$$\text{Now, } R_{sm}^T R_{sm} = \begin{bmatrix} 1 & 0 & 0 \\ 0 & 1 & 0 \\ 0 & 0 & 1 \end{bmatrix}_{3 \times 3} \begin{bmatrix} 0 & \dot{\theta} & 0 \\ 0 & 0 & 0 \\ 0 & 0 & 0 \end{bmatrix}_{3 \times 3}$$

$$\Rightarrow \begin{bmatrix} 0 & -\dot{\theta} & 0 \\ 0 & 0 & 0 \\ 0 & 0 & 0 \end{bmatrix}_{3 \times 3}$$

Now,  $-R_{Sm}^{-1} \cdot P$

$$\Rightarrow \begin{bmatrix} -1 & 0 & 0 \\ 0 & -1 & 0 \\ 0 & 0 & 0 \end{bmatrix} \begin{bmatrix} 0 \\ -em\pi/14 \\ 0 \end{bmatrix} = \begin{bmatrix} 0 \\ em\pi/14 \\ 0 \end{bmatrix}$$

Now,

$$V_{Sm}^b = \begin{bmatrix} 0 & 0 & 0 \\ 0 & 0 & em\pi/14 \\ 0 & 0 & 0 \\ 0 & 0 & 0 \end{bmatrix} \quad 4 \times 3$$

$$V_{Sm}^b = \begin{bmatrix} V_{Sm}^b \\ \omega_{Sm} \end{bmatrix} \Rightarrow$$

$$\begin{bmatrix} 0 \\ em\pi/14 \\ 0 \\ 0 \\ 0 \\ -\pi/14 \end{bmatrix} \quad 6 \times 1$$

Q7 Calculate the moon's body velocity  
in its own rotating frame

$$V_{\text{Bm}} = \text{adj}(g_{\text{sm}}) V_{\text{sm}}^b$$

$$\text{adj}(g_{\text{sm}}) = \begin{bmatrix} R_{\text{sm}} & P \cdot R_{\text{sm}} \\ 0 & R_{\text{sm}} \end{bmatrix}$$

where,  $R_{\text{sm}} = \begin{bmatrix} 1 & 0 & 0 \\ 0 & 1 & 0 \\ 0 & 0 & 1 \end{bmatrix}$ ;  $P = \begin{bmatrix} 0 & 0 & 0 \\ 0 & 0 & -lm \\ 0 & lm & 0 \end{bmatrix}$

$$P \cdot R_{\text{sm}} = \begin{bmatrix} 0 & 0 & 0 \\ 0 & 0 & -lm \\ 0 & lm & 0 \end{bmatrix} \begin{bmatrix} 1 & 0 & 0 \\ 0 & 1 & 0 \\ 0 & 0 & 1 \end{bmatrix}$$

$$P \cdot R_{\text{sm}} \Rightarrow \begin{bmatrix} 0 & 0 & 0 \\ 0 & 0 & -lm \\ 0 & -lm & 0 \end{bmatrix}$$

$$\text{adj}(g_{\text{sm}}) = \begin{bmatrix} 1 & 0 & 0 & 0 & 0 & 0 \\ 0 & 1 & 0 & 0 & 0 & -lm \\ 0 & 0 & 1 & 0 & -lm & 0 \\ 0 & 0 & 0 & 1 & 0 & 0 \\ 0 & 0 & 0 & 0 & 1 & 0 \\ 0 & 0 & 0 & 0 & 0 & 1 \end{bmatrix}$$

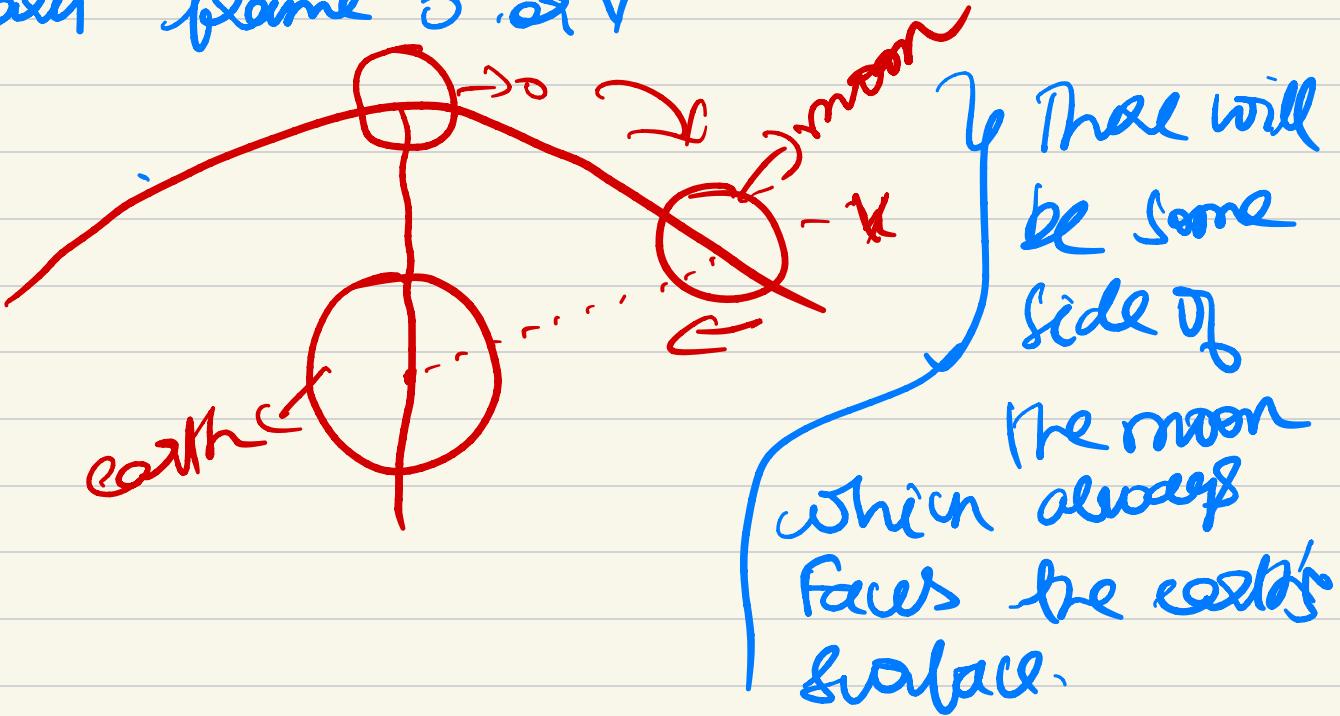
6x6

Now,

$$V_{Sm} = \text{adj}(g_{Sm}) \times$$
  
$$6 \times C$$

$$\begin{bmatrix} 0 \\ \ln \frac{\pi}{14} \\ 0 \\ 0 \\ -\frac{\pi}{14} \\ 6 \times 1 \end{bmatrix} \Rightarrow \begin{bmatrix} 0 \\ 0 \\ 0 \\ 0 \\ -\frac{\pi}{14} \\ 6 \times 1 \end{bmatrix}$$

Now, The speciality of  $V_{Sm}$  is that the velocity of translation of moon will be stationary frame of ref. or  $V$ .



1.8 A Lunar day, is the time it takes for the moon to return to same position in the sky relative to the earth & the sun. This is approx 29.5 days.

→ Length of a lunar day is longer than the moon's orbital period around the earth, which is about 27-3 days. This difference is due to combined motion of the earth and the moon in their orbits around the Sun.

→ So only considering rotation vectors

$$\omega_{\text{re}} - \omega_{\text{Sun}} = \begin{bmatrix} 0 \\ 0 \\ \frac{2\pi}{T} \end{bmatrix}$$

→ Since the angle  $\theta$  earth will be  $2\pi$ .

$$\Rightarrow \text{Time} = \frac{\text{angle}}{\text{angular velocity}} = \frac{2\pi}{\frac{2\pi}{T}} = \frac{T}{2} = 1.036$$

∴ Lunar Day = 1.036 days

## 2.1)

```
% Combined script for angvel2skew, skew2angvel, and unit test

% angvel2skew.m

function W_hat = angvel2skew(w)

    % Create a 3x3 skew-symmetric matrix from a 3-element vector w

    W_hat = [0, -w(3), w(2);
              w(3), 0, -w(1);
              -w(2), w(1), 0];

end

% skew2angvel.m

function w = skew2angvel(W_hat)

    % Extract the angular velocity vector w from a 3x3 skew-symmetric matrix W_hat

    w = [W_hat(3,2); W_hat(1,3); W_hat(2,1)];

end

% Unit test for angvel2skew and skew2angvel functions

% Generate a random angular velocity vector w

w = rand(3, 1);

% Call angvel2skew to convert w to a skew-symmetric matrix W_hat

W_hat = angvel2skew(w);

% Call skew2angvel to convert the skew-symmetric matrix W_hat back to an angular velocity vector w_prime

w_prime = skew2angvel(W_hat);

% Check if w and w_prime are approximately equal (within a tolerance)

tolerance = 1e-6;

is_equal = all(abs(w - w_prime) < tolerance);

if is_equal

    disp('Test is passed as angvel2skew and skew2angvel are inverse to each other.');

else

    disp('Test is failed as angvel2skew and skew2angvel are not inverse to each other.');

end
```

## 2.2)

```
% Combined script for twist2rbvel, rbvel2twist, and unit test

% twist2rbvel.m

function V_hat = twist2rbvel(V)

    % Create a 4x4 rigid body velocity matrix in homogeneous coordinates from a 6-element twist vector V

    V_hat = [skewSymmetricMatrix(V(1:3)), V(4:6);

              zeros(1, 4)];

end

function W_hat = skewSymmetricMatrix(w)

    % Create a 3x3 skew-symmetric matrix from a 3-element vector w

    W_hat = [0, -w(3), w(2);

              w(3), 0, -w(1);

              -w(2), w(1), 0];

end

% rbvel2twist.m

function V = rbvel2twist(V_hat)

    % Extract the 6-element twist vector V from a 4x4 rigid body velocity matrix in homogeneous coordinates V_hat

    V = [V_hat(1:3, 4); unskewSymmetricMatrix(V_hat(1:3, 1:3))];

end

function w = unskewSymmetricMatrix(W_hat)

    % Extract a 3-element vector w from a 3x3 skew-symmetric matrix W_hat

    w = [W_hat(3, 2); W_hat(1, 3); W_hat(2, 1)];

end

% Unit test for twist2rbvel and rbvel2twist functions

% Generate a random twist vector V

V = rand(6, 1);

% Call twist2rbvel to convert V to a 4x4 rigid body velocity matrix in homogeneous coordinates V_hat

V_hat = twist2rbvel(V);
```

```
% Call rbvel2twist to convert the rigid body velocity matrix V_hat back to a twist vector V_prime  
V_prime = rbvel2twist(V_hat);  
  
% Check if V and V_prime are approximately equal (within a tolerance)  
tolerance = 1e-6;  
  
is_equal = all(abs(V - V_prime) < tolerance);  
  
if is_equal  
    disp('Test is passed as twist2rbvel and rbvel2twist are inverse of each other.');//  
else  
    disp('Test is failed as twist2rbvel and rbvel2twist are not inverse of each other.');//  
end
```

**2.3)**

```
% tform2adjoint.m

function Adg = tform2adjoint(g)

    % Extract the rotation matrix R and translation vector p from the homogeneous transformation matrix g

    R = g(1:3, 1:3);

    p = g(1:3, 4);

    % Create the 6x6 adjoint transformation matrix Adg

    Adg = zeros(6, 6);

    % Populate the upper-left 3x3 block with R

    Adg(1:3, 1:3) = R;

    % Populate the lower-right 3x3 block with R

    Adg(4:6, 4:6) = R;

    % Compute the 3x3 skew-symmetric matrix p_hat from the translation vector p

    p_hat = [0, -p(3), p(2);

              p(3), 0, -p(1);

              -p(2), p(1), 0];

    % Populate the upper-right 3x3 block with p_hat

    Adg(1:3, 4:6) = p_hat;

end
```

## 2.4)

```
% compare_twist.m

function compare_twist()

    % Generate a random 4x4 homogeneous transformation matrix g
    g = random_homogeneous_matrix();

    % Compute the spatial velocity Vs_hat and body velocity Vb_hat
    Vs_hat = compute_spatial_velocity(g);

    Vb_hat = compute_body_velocity(g);

    % Convert the spatial velocity Vs_hat to a spatial twist Vs
    Vs = rbvel2twist(Vs_hat);

    % Convert the body velocity Vb_hat to a body twist Vb
    Vb = rbvel2twist(Vb_hat);

    % Display the generated transformation matrix g
    disp('Generated Transformation Matrix g:');
    disp(g);

    % Display the spatial and body twists
    disp('Spatial Twist Vs:');
    disp(Vs);
    disp('Body Twist Vb:');
    disp(Vb);

    % Check if the spatial twist and body twist are approximately equal (within a tolerance)
    tolerance = 1e-6;
    is_equal = all(abs(Vs - Vb) < tolerance);

    if is_equal
        disp('Test is passed as Spatial Twist and Body Twist are equal.');
    else
        disp('Test is Failed as Spatial Twist and Body Twist are not equal.');
    end
end
```

```

function g = random_homogeneous_matrix()

    % Generate a random 4x4 homogeneous transformation matrix

    R = random_rotation_matrix();

    p = rand(3, 1);

    g = eye(4);

    g(1:3, 1:3) = R;

    g(1:3, 4) = p;

end

function R = random_rotation_matrix()

    % Generate a random 3x3 rotation matrix

    theta = rand() * 2 * pi;

    v = rand(3, 1);

    v = v / norm(v);

    K = [0, -v(3), v(2);

          v(3), 0, -v(1);

          -v(2), v(1), 0];

    R = eye(3) + sin(theta) * K + (1 - cos(theta)) * K^2;

end

function Vs_hat = compute_spatial_velocity(g)

    % Compute the spatial velocity Vs_hat from the transformation matrix g

    Vs_hat = eye(4);

    Vs_hat(1:3, 4) = g(1:3, 4);

end

function Vb_hat = compute_body_velocity(g)

    % Compute the body velocity Vb_hat from the transformation matrix g

    Vb_hat = g;

    Vb_hat(1:3, 4) = [0; 0; 0];

end

```

## 2.5)

```
% compare_twist.m

function [g, Vs_Adg, Vs, Vb] = compare_twist()

% Generate a random 4x4 homogeneous transformation matrix g
g = random_homogeneous_matrix();

% Compute the spatial velocity Vs_hat and body velocity Vb_hat
Vs_hat = compute_spatial_velocity(g);

Vb_hat = compute_body_velocity(g);

% Convert the spatial velocity Vs_hat to a spatial twist Vs
Vs = rbvel2twist(Vs_hat);

% Convert the body velocity Vb_hat to a body twist Vb
Vb = rbvel2twist(Vb_hat);

% Compute the adjoint transformation matrix Adg
Adg = tform2adjoint(g);

% Compute Vs_Adg by applying the adjoint transformation
Vs_Adg = Vs * Adg;

% Display the generated transformation matrix g
disp('Generated Transformation Matrix g:');
disp(g);

% Display Vs, Vs_Adg, and Vb
disp('Spatial Twist Vs:');
disp(Vs);

disp('Vs_Adg (Spatial Twist after Adjoint Transformation):');
disp(Vs_Adg);

disp('Body Twist Vb:');
disp(Vb);

% Check if Vs and Vs_Adg are approximately equal (within a tolerance)
tolerance = 1e-6;
is_equal = all(abs(Vs - Vs_Adg) < tolerance);
```

```
if is_equal  
    disp('Test is passed as Vs and Vs_Adg are identical.');
```

```
else  
    disp('Test is failed as Vs and Vs_Adg are not identical.');
```

```
end
```

```
end
```

```
% Rest of the functions (random_homogeneous_matrix, random_rotation_matrix,  
% compute_spatial_velocity, compute_body_velocity, tform2adjoint) remain the same as in the previous answer.
```