# Basic description of physical data

deterministic: described by explicit mathematical relation



$$x(t) = X \cos(\sqrt{\frac{k}{t}}t)$$

non deterministic: no way to predict an exact value at a future instant of time



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## Spectral Analysis and Time Series

Andreas Lagg



#### on time series Part I: fundamentals

- classification
- prob. density func.
- auto-correlation
- power spectral density
- cross-correlation
- applications
- pre-processing
- sampling
- trend removal

### Part II: Fourier series

- method

definition

- properties
- convolution
- correlations
- leakage / windowing
- irregular grid
- noise removal

### Part III: Wavelets

- why wavelet transforms?
- fundamentals: FT, STFT and
- resolution problems
- multiresolution analysis: CWT
- DWT

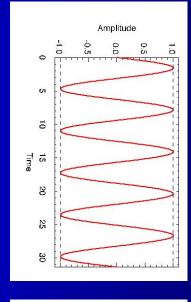
Exercises

## Sinusoidal data

$$x(t) = X \sin(2\pi f_0 t + \Theta)$$
$$T = 1/f_0$$

#### time history

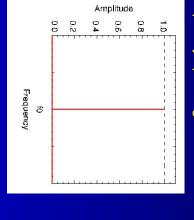


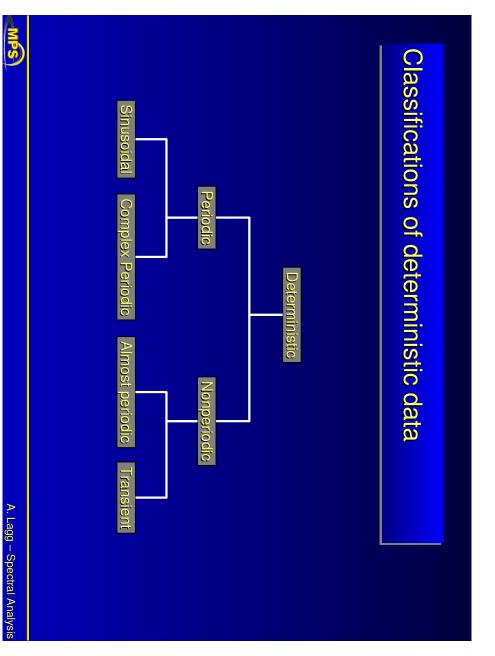


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### frequency spectrogram





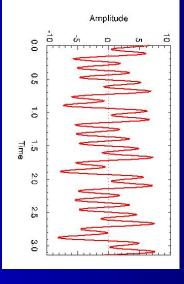
## Almost periodic data

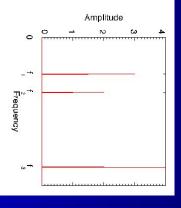
$$x(t) = X_1 \sin(2t + \Theta_1) + X_2 \sin(3t + \Theta_2) + X_3 \sin(\sqrt{50}t + \Theta_3)$$

no highest common divisor -> infinitely long period T

time history

frequency spectrogram





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## Complex periodic data

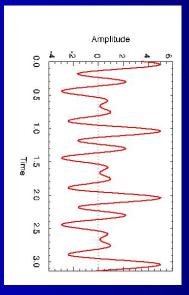
$$x(t) = x(t \pm nT)$$
  $n = 1, 2, 3, ...$ 

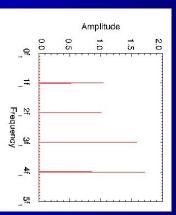
$$x(t) = \frac{a_0}{2} + \sum (a_n \cos 2\pi n f_1 t + b_n \sin 2\pi n f_1 t)$$

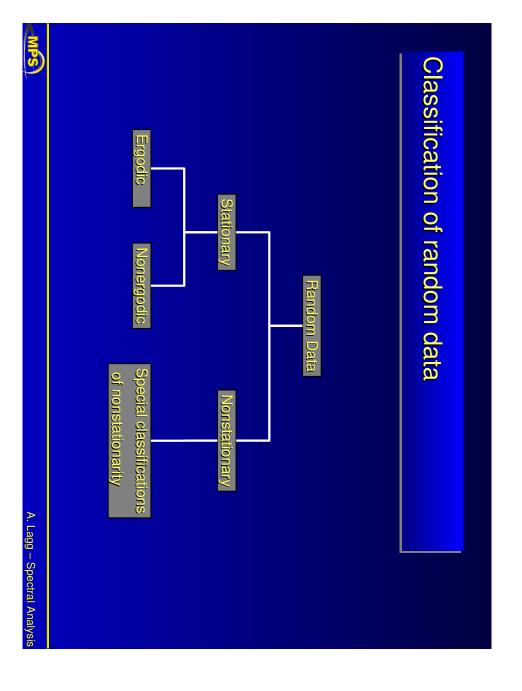
(T = fundamental period)

time history

### frequency spectrogram







## Transient non-periodic data

all non-periodic data other than almost periodic data

$$x(t) = \begin{cases} Ae^{-at} & t \ge 0 \\ 0 & t < 0 \end{cases}$$

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$$x(t) = \begin{cases} Ae^{-at} \cos bt & t <$$

0

 $c \ge t \ge 0$ c < t < 0

Amplitude

1.6 0.8 0.4 0.2

> Amplitude C.C.2

0.00 0.05 0.10 0.15 0.20 0.25 0.30 Frequency

## ergodic / non ergodic

Ergodic random process:

properties of a stationary random process described by computing averages over only **one single sample function** in the ensemble

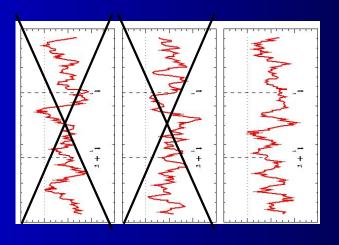
mean value of k-th sample function:

$$\mu_{x}(k) = \lim_{T \to \infty} \frac{1}{T} \int_{0}^{T} x_{k}(t) dt$$

autocorrelation function (joint moment):

$$R_{x}(\tau,k) = \lim_{T \to \infty} \frac{1}{T} \int_{0}^{t} x_{k}(t) x_{k}(t+\tau) dt$$

ergodic:  $\mu_x(k) = \mu_x$ ,  $R_x(\tau, k) = R_x(\tau)$ 



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## stationary / non stationary

collection of sample functions = ensemble

data can be (hypothetically) described by computing ensemble averages (averaging over multiple measurements / sample functions)

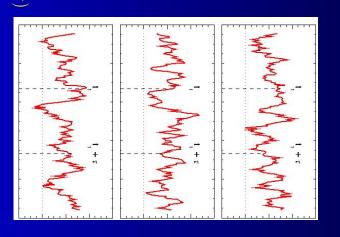
mean value (first moment):

$$\mu_{x}(t_{1}) = \lim_{N \to \infty} \frac{1}{N} \sum_{k=1}^{N} x_{k}(t_{1})$$

autocorrelation function (joint moment):

$$R_x(t_1, t_1 + \tau) = \lim_{N \to \infty} \frac{1}{N} \sum_{k=1}^{N} x_k(t_1) x_k(t_1 + \tau)$$

stationary:  $\mu_x(t_1) = \mu_x$ ,  $R_x(t_1, t_1 + \tau) = R_x$ weakly stationary:  $\mu_x(t_1) = \mu_x$ ,  $R_x(t_1, t_1 + \tau) = R_x(\tau)$ 



## Mean square values

(mean values and variances)

describes general intensity of random data:

$$\Psi_x^2 = \lim_{T \to \infty} \frac{1}{T} \int_0^T x^2(t) dt$$

rout mean square value:  $|\Psi_x^{ms}| = \sqrt{\Psi_x^2}$ 

$$\sqrt{\Psi_{
m v}^2}$$

often convenient:

static component described by mean value:  $\mu_x = \lim_{T o \infty} \frac{1}{T} \int_0^T x(t) dt$ 

standard deviation:  $\sigma_x = \sqrt{\sigma_x^2}$ 

■ dynamic component described by variance: 
$$\sigma_x^2 = \lim_{T \to \infty} \frac{1}{T} \int_0^T [x(t) - \mu_x]^2 dt$$
 standard deviation:  $\sigma_x = \sqrt{\sigma_x^2} = \sqrt{\sigma_x^2}$  =  $Y_x^2 - \mu_x^2$ 



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# Basic descriptive properties of random data

- mean square values
- propability density function
- autocorrelation functions
- power spectral density functions

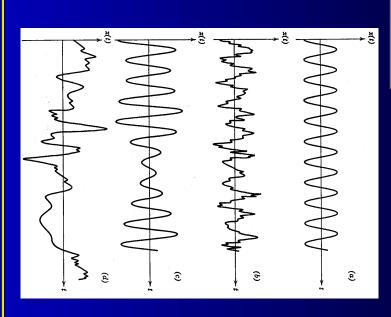
(from now on: assume random data to be stationary and ergodic)

# Illustration: probability density function

sample time histories:

- sine wave (a)
- sine wave + random noise
- narrow-band random noise
- wide-band random noise

all 4 cases: mean value  $\mu_x = 0$ 





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## Probability density functions

range at any instant of time describes the probability that the data will assume a value within some defined

$$\operatorname{Prob}[x < x(t) \le x + \Delta x] = \lim_{T \to \infty} \frac{T_x}{T}, \quad T_x = \sum_{i=1}^k \Delta t_i$$

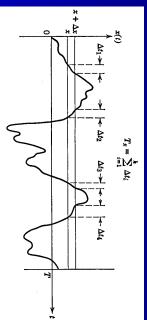
for small  $\Delta x$ : Prob $\left[x < x(t) \le x + \Delta x\right] \approx p(x)\Delta x$ 

probability density function
n(x) = 1

$$p(x) = \lim_{\Delta x \to 0} \frac{\text{Prob}[x < x(t) \le x + \Delta x]}{\Delta x} = \lim_{\Delta x \to 0} \frac{1}{\Delta x} \left[ \lim_{T \to \infty} \frac{T_x}{T} \right]$$

probability distribution function

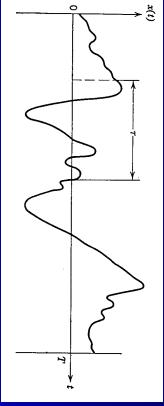
$$p(x) = \text{Prob}[x(t) \le x]$$
  
=  $\int_{-\infty}^{x} p(\xi) d\xi$ 



## Autocorrelation functions

describes the general dependence of the data values at one time on the values at another time.

$$R_{x}(\tau) = \lim_{T \to \infty} \frac{1}{T} \int_{0}^{T} x(t) x(t+\tau) dt$$



$$\mu_x = \sqrt{R_x(\infty)}$$

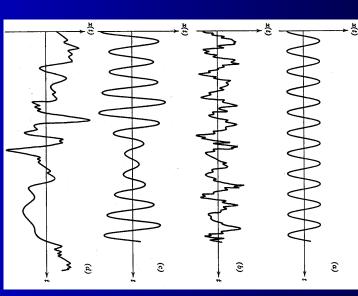
$$\Psi_x^2 = R_x(0)$$
 (not for special cases like sine waves)

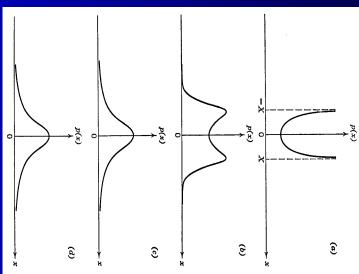
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# Illustration: probability density function

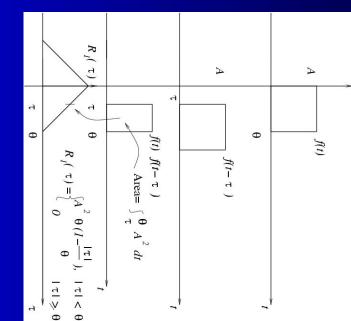
probability density function

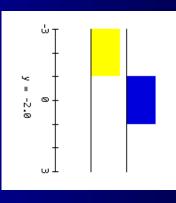




### Illustrations

## autocorrelation function of a rectangular pulse





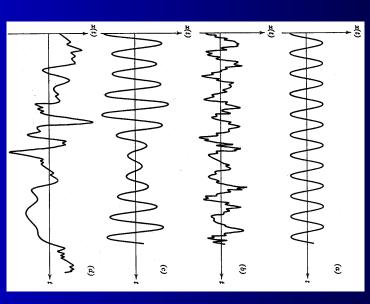
 $x(t)x(t-\tau)$ 

autocorrelation function

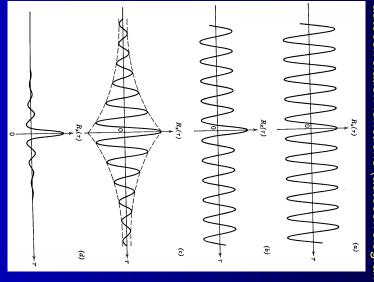


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### Illustration: ACF



## autocorrelation functions (autocorrelogram)



### Illustration: PSD

Dirac delta function at f=f<sub>0</sub>

sine wave

power spectral density functions

ê

6

+ random noise sine wave

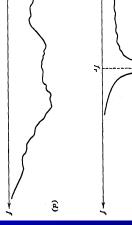
narrowband noise

broadband noise

spectrum is uniform over

"white" noise:

all frequencies





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# Power spectral density functions (also called autospectral density functions)

describe the general frequency composition of the data in terms of the spectral density of its mean square value

mean square value in frequency range  $(f,\ f+\Delta f)$  :

$$\Psi_x^2(f, \Delta f) = \lim_{T \to \infty} \frac{1}{T} \int_0^T \frac{x(t, f, \Delta f)^2 dt}{t}$$

portion of x(t) in  $(f,f+\Delta f)$ 

definition of power spectral density function: 
$$\begin{aligned} G_x(f) &= \lim_{\Delta f \to 0} \frac{\Psi_x^2(f,\Delta f)}{\Delta f} = \lim_{\Delta f \to 0} \frac{1}{\Delta f} \left[ \lim_{T \to \infty} \frac{1}{T} \int\limits_0^T x(t,f,\Delta f)^2 dt \right] \end{aligned}$$

autocorrelation function by a Fourier transform: important property: spectral density function is related to the

$$G_x(f) = 2\int_{-\infty}^{\infty} R_x(\tau)e^{-i2\pi f \tau}d\tau = 4\int_{0}^{\infty} R_x(\tau)\cos 2\pi f \tau d\tau$$

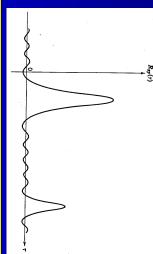
## Cross-correlation function

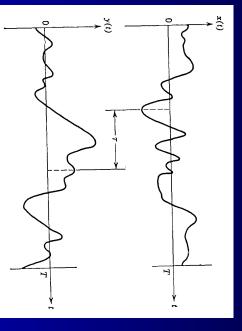
describes the general dependence of one data set to another

$$R_{xy}(\tau) = \lim_{T \to \infty} \frac{1}{T} \int_{0}^{T} x(t) y(t+\tau) dt$$

similar to autocorrelation function

 $R_{_{\mathrm{Jy}}}( au) {=} 0$  functions are uncorrelated





cross-correlation measurement

typical cross-correlation plot (cross-correlogram): sharp peaks indicate the existence of a correlation between x(t) and y(t) for specific time displacements



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# Joint properties of random data

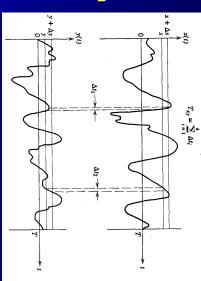
until now: described properties of an individual random process

## Joint probability density functions

joint properties in the amplitude domain

## **Cross-correlation functions**

joint properties in the time domain

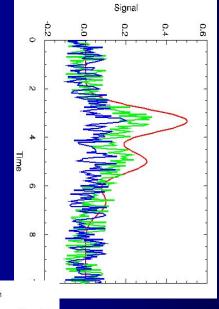


joint probability measurement

## Cross-spectral density functions

joint properties in the frequency domain

### **Applications**

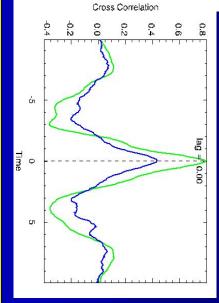


determine if theoretical signal is cross correlation can be used to present in data

### signals in noise Detection and recovery from

3 signals:

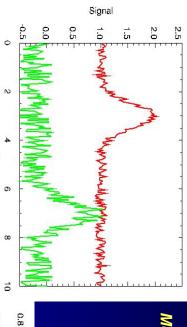
- noise free replica of the signal (e.g. model)
- 2 noisy signals





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### **Applications**



## Measurement of time delays

2 signals:

- different offset
- different S/N
- time delay 4s



-0.4 0.2

#### 0.0 0.6 0.2 0.4

'discrete' cross correlation coefficient lag =1, for l >=0:

Cross Correlation

often used:

Time

Time

### Trend removal

spectral analysis often desirable before performing a

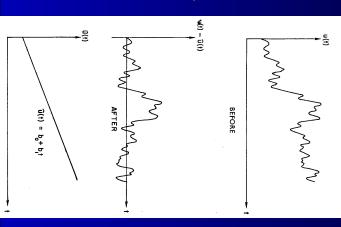
## Least-square method:

time series: u(t)

desired fit esired fit  $\hat{u} = \sum_{k=0}^{K} bk (nh)^k$  n=1,2,...,N

Lsq-Fit: minimize  $Q(b) = \sum_{n=1}^{\infty} (u_n - \hat{u}_n)^2$ 

- set partial derivatives to 0:  $\frac{\partial \mathcal{Q}}{\partial b_l} = \sum_{n=1}^{N} 2(u_n - \hat{u}_n)[-(nh)^l]$ - K+1 equations:  $\sum_{k=0}^{K} b_k \sum_{n=1}^{N} (nh)^{k+l} = \sum_{n=1}^{N} u_n (nh)^l$ 



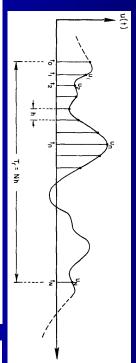
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## Pre-processing Operations

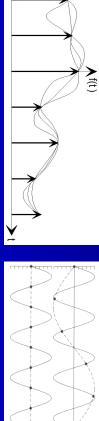
- sampling considerations
- trend removal
- filtering methods

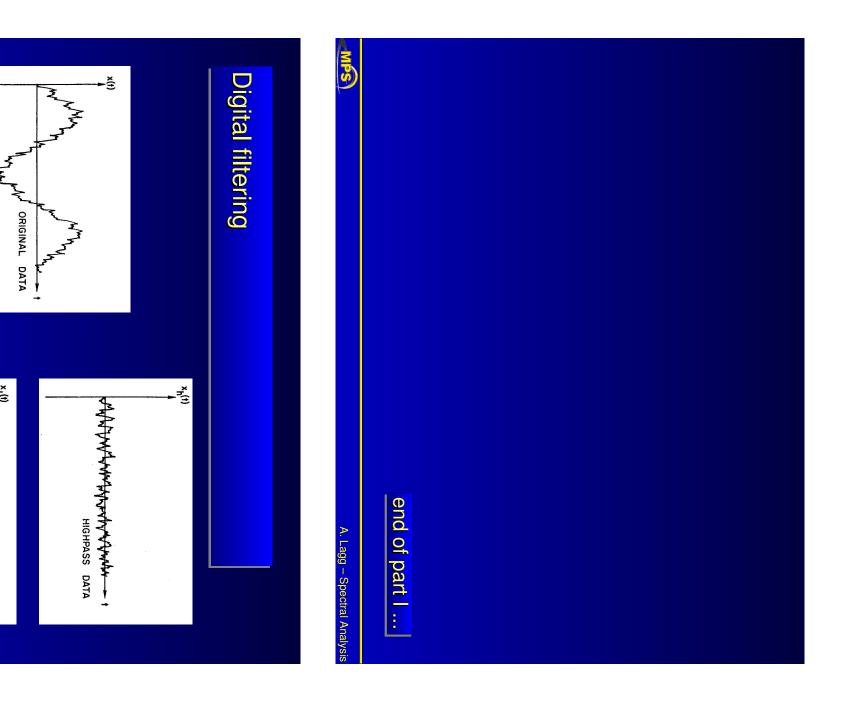
#### sampling



frequency or folding frequency) cutoff frequency (=Nyquist

$$f_c = \frac{1}{2h}$$





LOWPASS DATA

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# Fourier Series and Fast Fourier Transforms

Standard Fourier series procedure:

frequency  $f_1=1/T_p$ ), then x(t) can be represented by the Fourier series: if a transformed sample record x(t) is periodic with a period  $T_{_{p}}$  (fundamental

$$x(t) = \frac{a_0}{2} + \sum_{q=1}^{\infty} (a_q \cos 2\pi \, q \, f_1 t \, + \, b_q \sin 2\pi \, q \, f_1 t)$$
 here 
$$a_q = \frac{2}{T} \int_0^T x(t) \cos 2\pi \, q \, f_1 t \, dt \quad q = 0, 1, 2, \dots$$
 
$$b_q = \frac{2}{T} \int_0^T x(t) \sin 2\pi \, q \, f_1 t \, dt \quad q = 1, 2, 3, \dots$$



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## **Spectral Analysis and Time Series**

Andreas Lagg



on time series Part I: fundamentals

- classification
- prob. density func.
- auto-correlation
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### Part II: Fourier series

- method

definition

- properties
- convolution
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- irregular grid leakage / windowing
- noise removal

### Part III: Wavelets

- why wavelet transforms?
- fundamentals: resolution problems FT, STFT and
- multiresolution analysis: CWT
- DWT

Exercises

# Fourier Transforms - Properties

Linearity 
$$\{x_n\} \qquad \stackrel{DFT}{\Leftrightarrow} \qquad \{X_k\}$$
 
$$\{y_n\} \qquad \stackrel{DFT}{\Leftrightarrow} \qquad \{Y_k\}$$
 
$$a\{x_n\} + b\{y_n\} \qquad \Leftrightarrow \qquad a\{X_k\} + b\{Y_k\}$$

$$\begin{array}{ccc} \boxed{ \textbf{Symmetry} } & \{X_k\} & = & \{X_{-k}^*\} \\ \Re\{X_k\} & \text{is even} & \Im\{X_k\} & \text{is odd} \end{array}$$

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## Fourier series procedure - method

sample record of finite length, equally spaced sampled:

$$x_n = x(nh)$$
  $n=1,2,\ldots$ 

Fourier series passing through these N data values:

$$x(t) = A_0 + \sum_{q=1}^{N/2} A_q \cos\left(\frac{2\pi qt}{T_p}\right) + \sum_{q=1}^{N/2-1} B_q \sin\left(\frac{2\pi qt}{T_p}\right)$$
Fill in particular points:  $t = nh$ ,  $n = 1, 2, ..., N$ ,  $T_p = Nh$ ,  $x_n = x(nh) = ...$ 

• coefficients 
$$A_{q}$$
 and  $B_{q}$ :  $A_{0} = \frac{1}{N} \sum_{n=1}^{N} x_{n} = \overline{x}$   $A_{N/2} = \frac{1}{N} \sum_{n=1}^{N} x_{n} \cos n \pi$ 

$$A_{q} = \frac{2}{N} \sum_{n=1}^{N} x_{n} \cos \frac{2\pi q n}{N} \qquad q=1,2,..., \frac{N}{2}-1$$

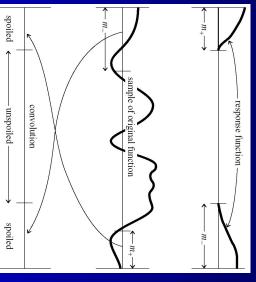
$$B_{q} = \frac{2}{N} \sum_{n=1}^{N} x_{n} \sin \frac{2\pi q n}{N} \qquad q=1,2,..., \frac{N}{2}-1$$

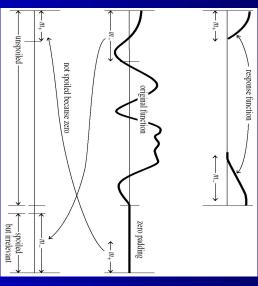
inefficient & slow => Fast Fourier Trafos developed

# Treatment of end effects by zero padding

constraint 1: simply expand response function to length N by padding it with zeros

constraint 2: extend data at one end with a number of zeros equal to the max. positive / negative duration of r (whichever is larger)







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## **Using FFT for Convolution**

$$r * s \equiv \int_{-\infty}^{\infty} r(\tau) s(t-\tau) d\tau$$

Convolution Theorem:

$$r *_S \stackrel{FT}{\Leftrightarrow} R(f)S(f)$$

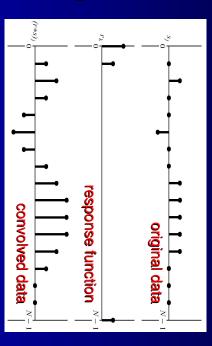
Fourier transform of the convolution is product of the individual Fourier transforms

### discrete case:

$$(r*s)_j \equiv \sum_{k=-N/2+1}^{\infty} s_{j-k} r_k$$

Convolution Theorem:

$$\sum_{k=-N/2+1}^{N/2} s_{j-k} r_k \stackrel{FT}{\Leftrightarrow} R_n S_n$$



(note how the response function for negative times is wrapped around and stored at the extreme right end of the array)

#### constraints:

- duration of r and s are not the same
- signal is not periodic

# Correlation / Autocorrelation with FFT

definition of correlation / autocorrelation see first lecture

$$Corr(g,h) = g*h = \int_{-\infty}^{\infty} g(t+\tau)h(\tau)d\tau$$

2.5

Correlation Theorem:

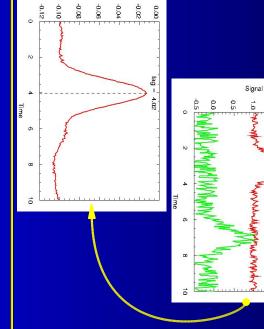
$$Corr(g,h) \stackrel{FT}{\Leftrightarrow} G(f)H^*(f)$$

**Auto-Correlation:** 

$$Corr(g,g) \stackrel{FT}{\Leftrightarrow} |G(f)|^2$$

discrete correlation theorem:

Cross Correlation



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## **FFT for Convolution**

- zero-pad data
- 5 zero-pad response function
- (-> data and response function have N elements)
- မှ calculate FFT of data and response function
- multiply FFT of data with FFT of response function
- 5 calculate inverse FFT for this product

### Deconvolution

-> undo smearing caused by a response function

use steps (1-3), and then:

- divide FFT of convolved data with FFT of response function
- ပ္ပာ calculate inverse FFT for this product

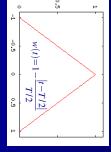
## reducing leakage by windowing (1)

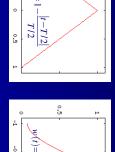
Applying windowing (apodizing) function to data record:

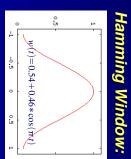
$$\overline{x}(t) = x(t)w(t)$$
 (original data record x windowing function)  $\overline{x}_n = x_n w_n$ 

### No Window:

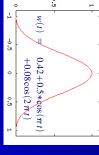
### Bartlett Window:



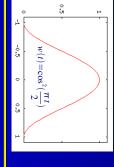




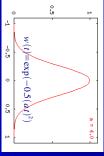
### Blackman Window:



### Hann Window:



Gaussian Window:

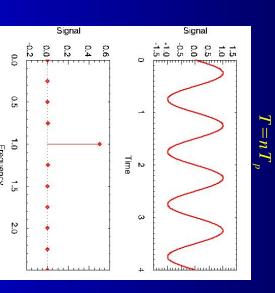


#### MPS

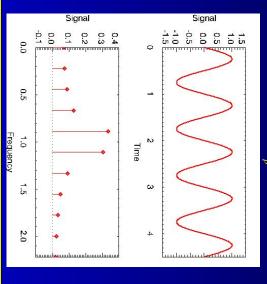
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## Fourier Transform - problems

### spectral leakage



#### $T \neq nT_p$



## No constant sampling frequency

equal distances) Fourier transformation requires constant sampling (data points at

-> not the case for most physical data

## Solution: Interpolation

- linear:
- inear interpolation between  $\boldsymbol{y}_k$  and  $\boldsymbol{y}_{k+1}$
- IDL> idata=interpol(data,t,t\_reg)
- quadratic:
- quadratic interpolation using  $y_{k-1}$ ,  $y_k$  and  $y_{k+1}$
- IDL> idata=interpol(data,t,t\_reg,/quadratic)
- least-square quadratic
- least-square quadratic fit using  $y_{k+1}$ ,  $y_k$ ,  $y_{k+1}$  and  $y_{k+2}$
- IDL> idata=interpol(data,t,t\_reg,/lsq)
- spline
- IDL> idata=interpol(data,t,t\_reg,/spline)
- IDL> idata=spline(t,data,t\_reg[,tension])

#### important:

interpolation changes sampling rate!

-> careful choice of new (regular) time grid

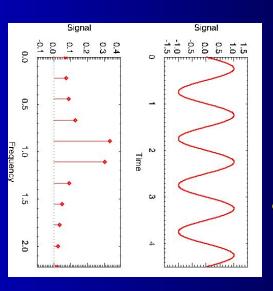
necessary



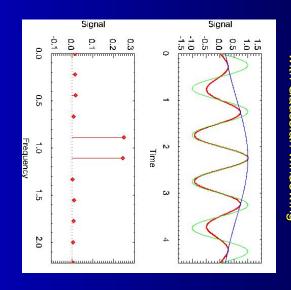
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## reducing leakage by windowing (2)

### without windowing

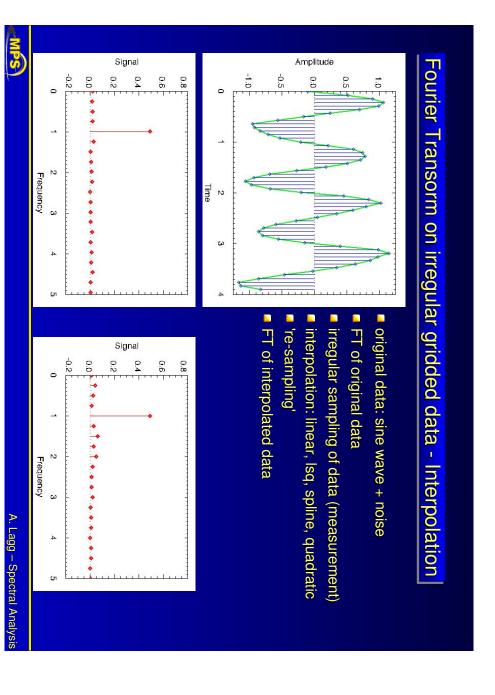


## with Gaussian windowing





#### set high frequencies to 0 transorm back to time make FT of data Frequency threshold (lowpass) Noise removal domain -0.10 0.30 0.50 0.00 0.10 0.40 Frequency σ $\infty$ Amplitude -1.0 0.0 0 4 Time A. Lagg - Spectral Analysis



## Optimal Filtering with FFT

normal situation with measured data:

underlying, uncorrupted signal u(t)

 $r(t-\tau)u(\tau)d\tau$ 

- + response function of measurement r(t)
- = smeared signal s(t)
- + noise n(t)
- = smeared, noisy signal c(t)

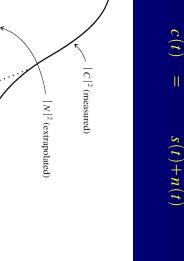
estimate true signal u(t) with:

$$\tilde{U}(f) = \frac{C(f)\Phi(f)}{R(f)}$$

log scale

 $\Phi(f), \varphi(t) = \text{optimal filter}$ (Wiener filter)

 $|S|^2$  (deduced)





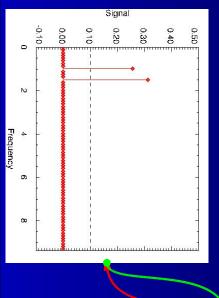
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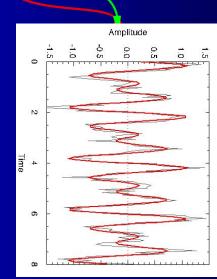
### Noise removal

## signal threshold for

weak frequencies (dB-threshold)

- make FT of data
- set frequencies with amplitudes below a given threshold to 0
- transorm back to time domain





# Using FFT for Power Spectrum Estimation

discrete Fourier transform of c(t)

-> Fourier coefficients:

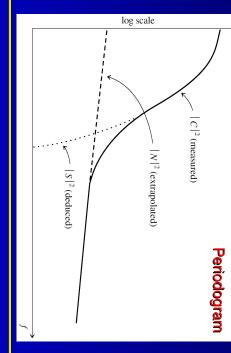
$$C_k = \sum_{j=0}^{N-1} c_j e^{2\pi i jklN} \quad k = 0, ..., N-1$$

-> periodogram estimate of power spectrum:

$$P(0) = P(f_0) = \frac{1}{N^2} |C_0|^2$$

$$P(f_k) = \frac{1}{N^2} [|C_k|^2 + |C_{N-k}|^2]$$

$$P(f_c) = P(f_{N/2}) = \frac{1}{N^2} |C_{N/2}|^2$$





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## Calculation of optimal filter

reconstructed signal and uncorrupted signal should be close in least-square sense: -> minimize

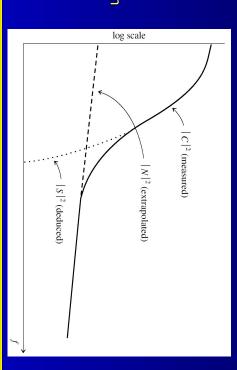
 $\int_{-\infty}^{\infty} |\tilde{u}(t) - u(t)|^2 dt = \int_{-\infty}^{\infty} |\tilde{U}(f) - U(f)|^2 df$ 

$$\Rightarrow \frac{\partial}{\partial \Phi(f)} \left| \frac{\left[ S(f) + N(f) \right] \Phi(f)}{R(f)} - \frac{S(f)}{R(f)} \right|^2 = 0$$

$$\Rightarrow \frac{\Phi(f)}{\partial \Phi(f)} - \frac{\left[ S(f) \right]^2}{\left[ S(f) \right]^2} = 0$$

$$\Rightarrow \Phi(f) = \frac{|S(f)|^2}{|S(f)|^2 + |N(f)|^2}$$

additional information:
power spectral density can often
be used to disentangle noise
function N(f) from smeared
signal S(f)



## Spectral Analysis and Time Series

Andreas Lagg



#### on time series Part I: fundamentals

- classification
- prob. density func.
- auto-correlation
- power spectral density
- cross-correlation
- applications
- pre-processing
- sampling
- trend removal

## Part II: Fourier series

- definition
- method
- properties
- convolution
- correlations
- leakage / windowing
- irregular grid
- noise removal

### Part III: Wavelets

- why wavelet transforms?
- fundamentals: resolution problems FT, STFT and
- multiresolution analysis: CWT
- DWT

### Exercises

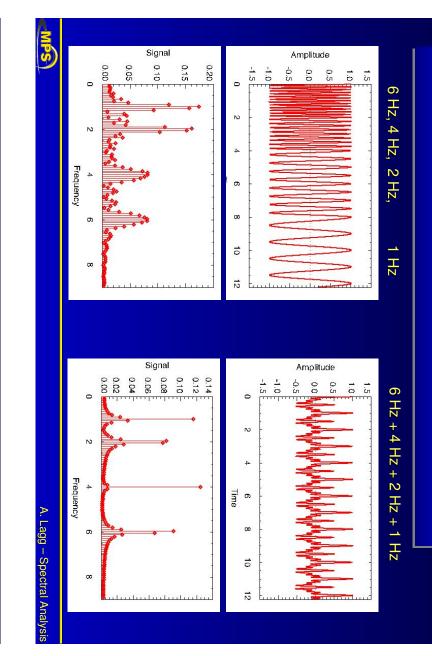


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end of FT



## Fourier: lost time information



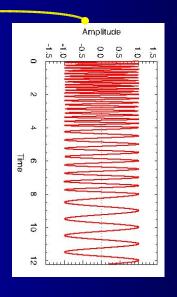
## Introduction to Wavelets

- why wavelet transforms?
- fundamentals: FT, short term FT and resolution problems
- multiresolution analysis: continous wavelet transform
- multiresolution analysis: discrete wavelet transform

#### frequency information! STFT-spectrogram shows both time and **Short Time Fourier Transform** Amplitude 0.4 0.6 AMPLITUE 30, STFT FREQUENCY A. Lagg - Spectral Analysis TIME



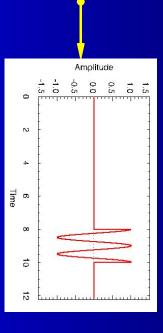
## Solution: Short Time Fourier Transform (STFT)



perform FT on 'windowed' function:

- example: rectangular window
- move window in small steps over data
- perform FT for every time step

$$STFT(f,t') = \int_{t} [x(t)\omega(t-t')]e^{-i2\pi ft}dt$$



Amplitude

0.0

1.5 0.5

N

4

6 Time

00

10

0.5

1.5

# Solution: Wavelet Transformation

approach: analyze the signal at different frequencies with different resolutions time vs. frequency resolution is intrinsic problem (Heisenberg Uncertainty Principle)

## -> multiresolution analysis (MRA)

## Continuous Wavelet Transform

### similar to STFT:

- signal is multiplied with a function (the wavelet)
- transform is calculated separately for different segments of the time domain

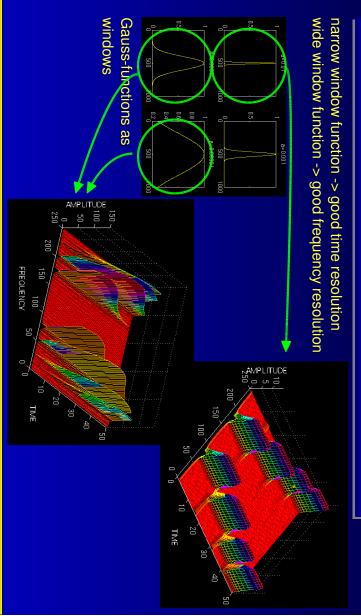
#### but:

- the FT of the windowed signals are not taken
- (no negative frequencies)
- The width of the window is changed as the transform is computed for every single spectral component



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# Short Time Fourier Transform: Problem



MPS

### The Scale

similar to scales used in maps:

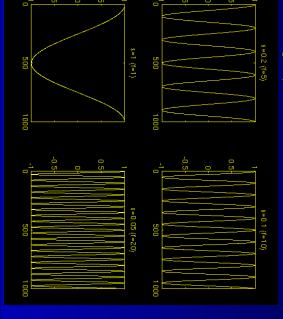
- high scale = non detailed global view (of the signal)
- low scale = detailed view

in practical applications:

- low scales (= high frequencies) appear usually as short bursts or spikes
- high scales (= low frequencies) last for entire signal

compresses a signal: scaling dilates (stretches out) or

s > 1 -> dilation s < 1 -> compression





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# Continuous Wavelet Transform

$$CWT_x^{\psi} = \Psi_x^{\psi}(\tau, s) = \frac{1}{\sqrt{|s|}} \int_t x(t) \psi^* \left(\frac{t - \tau}{s}\right) dt$$

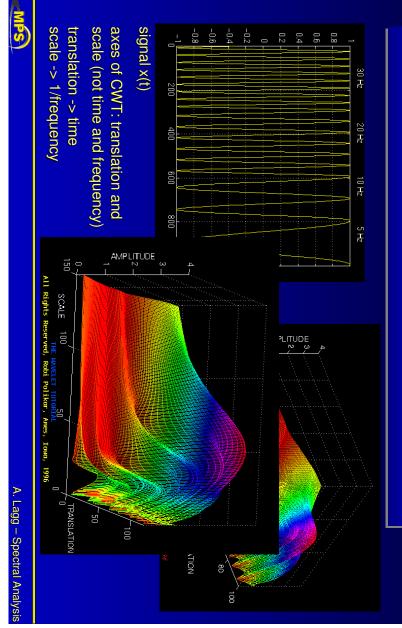
 $\psi(t)$  ... mother wavelet (= small wave) τ ... translation parameter, s ... scale parameter

mother wavelet:

- finite length (compactly supported) ->'let'
- oscillatory ->'wave'
- functions for different regions are derived from this function -> 'mother'

scale parameter's replaces frequency in STFT

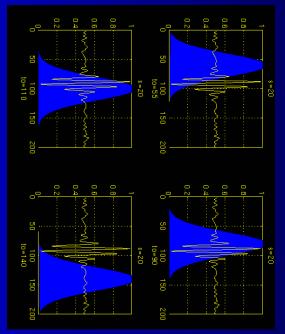
## CWT - Example



## Computation of the CWT

signal to be analyzed: x(t), mother wavelet: Morlet or Mexican Hat

- start with scale s=1 (lowest scale, highest frequency)
- -> most compressed wavelet
- shift wavelet in time from t<sub>0</sub> to t<sub>1</sub>
- increase s by small value
- shift dilated wavelet from t<sub>0</sub> to t<sub>1</sub>
- repeat steps for all scales



# Wavelets: Mathematical Approach

WL-transform:

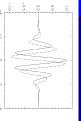
$$CWT_x^{\psi} = \Psi_x^{\psi}(\tau, s) = \frac{1}{\sqrt{|s|}} \int_t x(t) \psi^* \left(\frac{t-\tau}{s}\right) dt$$

Mexican Hat wavelet:

$$\nu(t) = \frac{1}{\sqrt{s\pi\sigma^3}} e^{\frac{-t^2}{2\sigma^2}} \left(\frac{t^2}{\sigma^2} - 1\right)$$

Morlet wavelet:

$$\psi(t) = e^{iat} e^{-\frac{t^2}{2\sigma}}$$



inverse WL-transform:

$$x(t) = \frac{1}{c_{\psi}^2} \int_{s} \int_{\tau} Y_x^{\psi}(\tau, s) \frac{1}{s^2} \psi\left(\frac{t - \tau}{s}\right) d\tau ds$$

admissibility condition:

$$c_{\psi} = \left\{2\pi \int_{-\infty}^{\infty} \frac{|\hat{\psi}(\xi)|^2}{|\xi|} d\xi\right\} < \infty \quad \text{with} \quad \hat{\psi}(\xi) \iff \psi(t)$$

#### MIPS

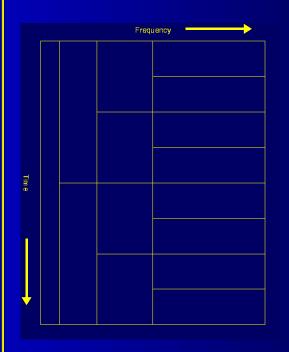
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# Time and Frequency Resolution

every box corresponds to a value of the wavelet transform in the time frequency plane

- all boxes have constant area
   Δf Δt = const.
- low frequencies: high resolution in f, low time resolution
- high frequencies: good time resolution

STFT: time and frequency resolution is constant (all boxes are the same)



## Discrete Wavelet Transform

discretized continuous wavelet transform is only a sampled version of the CWT implementation in computers. The discrete wavelet transform (DWT) has significant advantages for

excellent tutorial:

http://users.rowan.edu/~polikar/WAVELETS/WTtutorial.html

**IDL-Wavelet Tools:** 

IDL> wv\_applet

Wavelet expert at MPS:

Rajat Thomas

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# Discretization of CWT: Wavelet Series

-> sampling the time - frequency (or scale) plane

advantage:

- sampling high for high frequencies (low scales) scale s, and rate N,
- sampling rate can be decreased for low scale s<sub>2</sub> and rate N<sub>2</sub> frequencies (high scales)

$$N_{2} = \frac{s_{1}}{s_{2}} N_{1}$$

$$N_{2} = \frac{f_{2}}{f_{1}} N_{1}$$

continuous wavelet

### Exercises

## Part I: Fourier Analysis (Andreas Lagg)

Instructions:

http://www.linmpi.mpg.de/~lagg

### Part II: Wavelets (Rajat Thomas)

Seminar room Time: 15:00



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end of Wavelets

### Literature

Random Data: Analysis and Measurement Procedures Bendat and Piersol, Wiley Interscience, 1971

The Analysis of Time Series: An Introduction Chris Chatfield, Chapman and Hall / CRC, 2004

Time Series Analysis and Its Applications Shumway and Stoffer, Springer, 2000

Numerical Recipies in C Cambridge University Press, 1988-1992 http://www.nr.com/

The Wavelet Tutorial Robi Polikar, 2001 http://users.rowan.edu/~polikar/WAVELETS/WTtutorial.html



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# Exercise: Galileo magnetic field

data set from Galileo magnetometer (synthesized)

file: gll\_data.sav, contains:

- total magnetic field
- radial distance
- time in seconds

### your tasks:

- Which ions are present?
- Is the time resolution of the magnetometer sufficient to detect electrons or protons?

#### Tips:

- restore,'gll\_data.sav'
- use IDL-FFT
- remember basic plasma physics formula for the ion cyclotron wave:

$$\omega_{gyro} = \frac{q B}{m}, \quad f_{gyro} = \frac{\omega_{gyro}}{2 \pi}$$

#### Background:

If the density of ions is high enough they will excite ion cyclotron waves during gyration around the magnetic field lines. This gyration frequency only depends on mass per charge and on the magnitude of the magnetic field

In a low-beta plasma the magnetic field dominates over plasma effects. The magnetic field shows only very little influence from the plasma and can be considered as a magnetic dipole.

p://www.sciencemag.org/cgi/content/full/274/5286/396