ν_d

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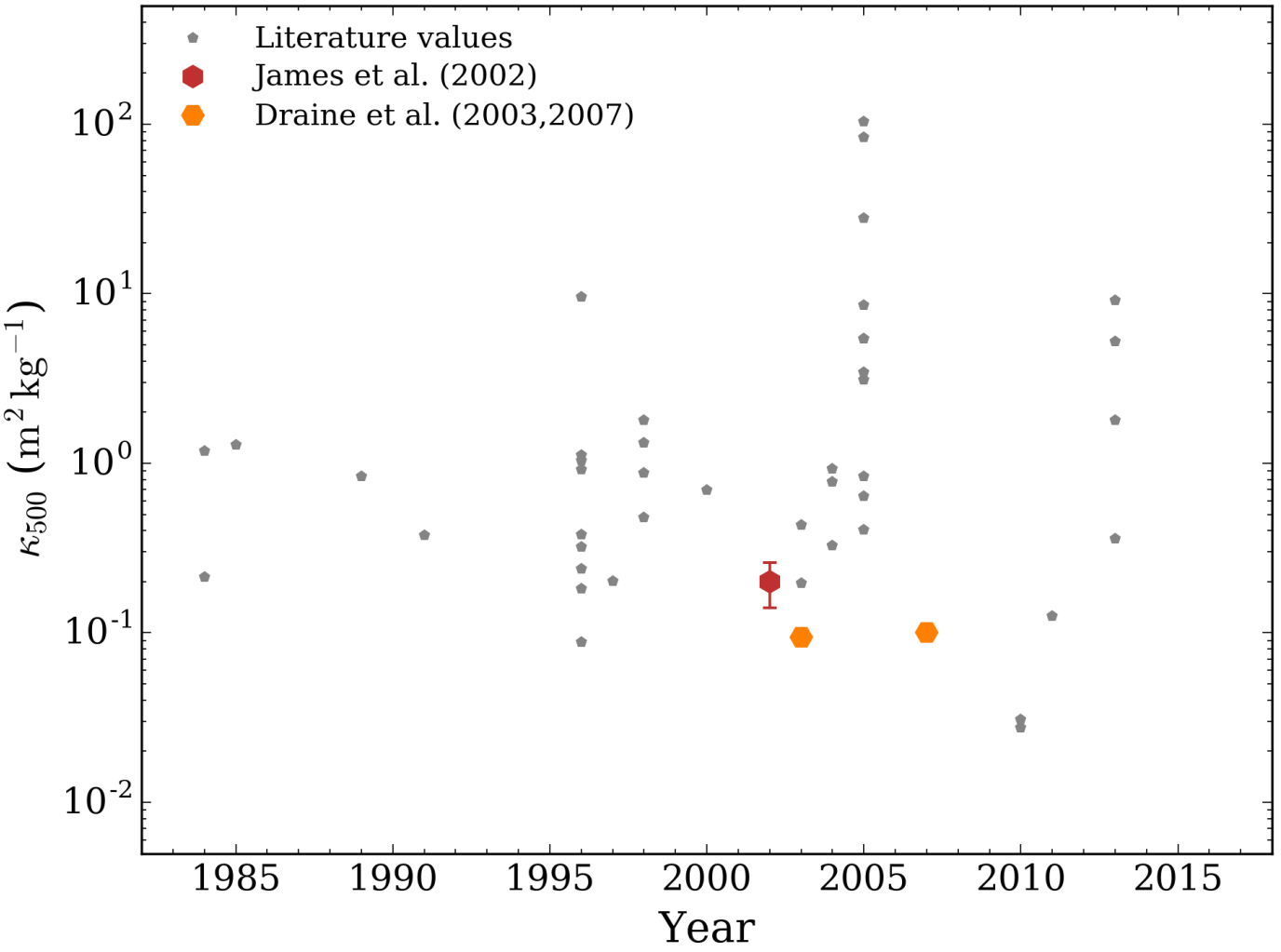
The Elusive $\varkappa_d$



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$$M_d = M_g \, \varepsilon_d \, f_Z$$

M_d = Dust mass

M_g = Gas mass

ε_d = ISM dust-to-metals ratio = 0.5 +/- 0.1

f_Z = ISM metal mass fraction (ie, ISM metallicty)

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Improvements Over James et al (2002)

Using data from the HRS

- We now have *Herschel* photometry; James et al only had IRAS & SCUBA, and so could not constrain dust temperature.
- We have integrated (and normalised) ISM metallicites from drift-scan spectroscopy
- We have integrated CO maps; they only had central pointings.
- James et al did not consider the effect of oxygen depletion on their metallicites – we do.
- James et al did not account for Helium or gas-phase metals in their ISM masses – we do.

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$$\kappa_\lambda = \frac{D^2}{\xi \left(M_{HI} + M_{H_2}\right) \varepsilon_d \, f_{Z_\odot} \, Z} \left(\frac{S_{\lambda_w}}{B_\lambda(T_w)} + \frac{S_{\lambda_c}}{B_\lambda(T_c)} \right)$$

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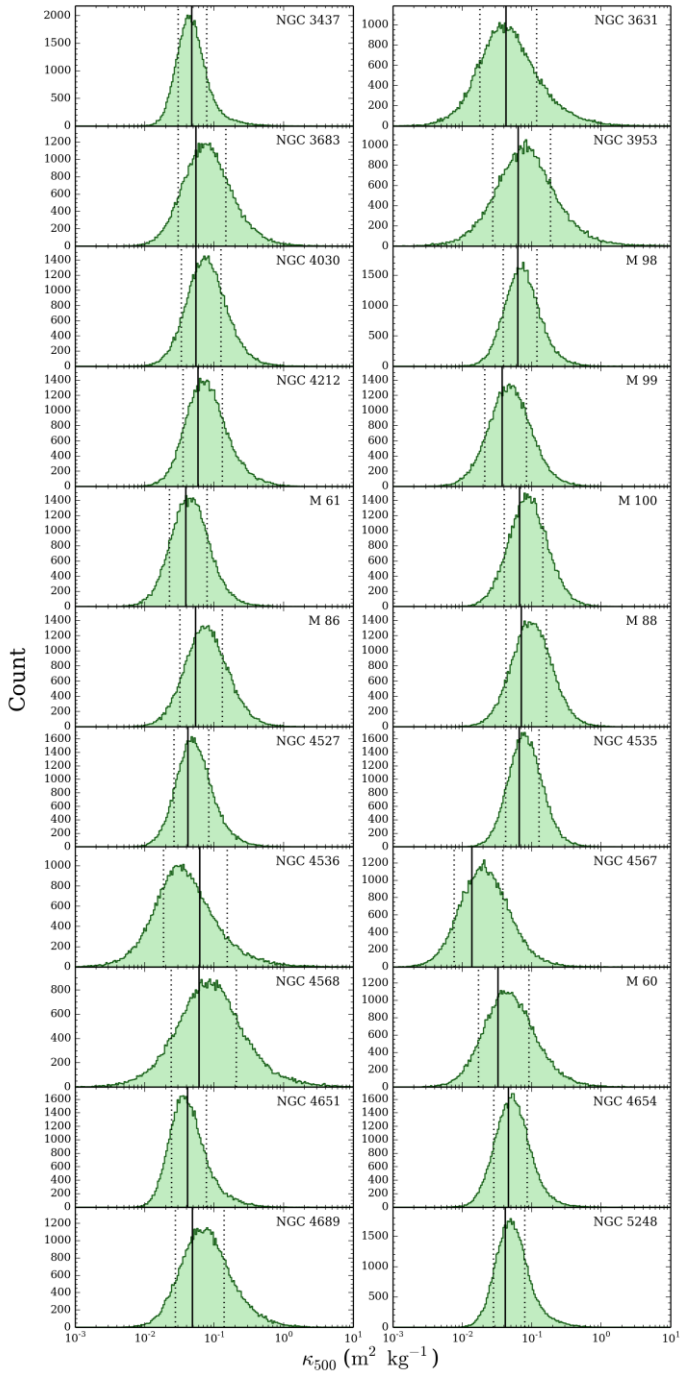
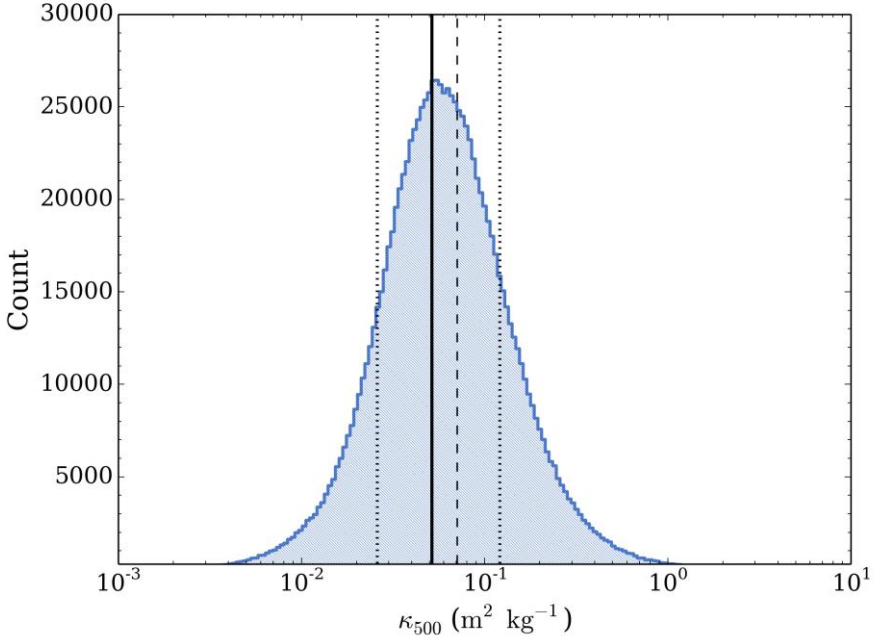

NGC 3437
NGC 3631
NGC 3683
NGC 3953
NGC 4030
M 98
NGC 4212
M 99
M 61
M 100
M 86
M 88
NGC 4527
NGC 4535
NGC 4536
NGC 4567
NGC 4568
M 60
NGC 4651
NGC 4654
NGC 4689
NGC 5248
Count
κ_{500} (m^2 kg^{-1})


Count
κ_{500} (m^2 kg^{-1})

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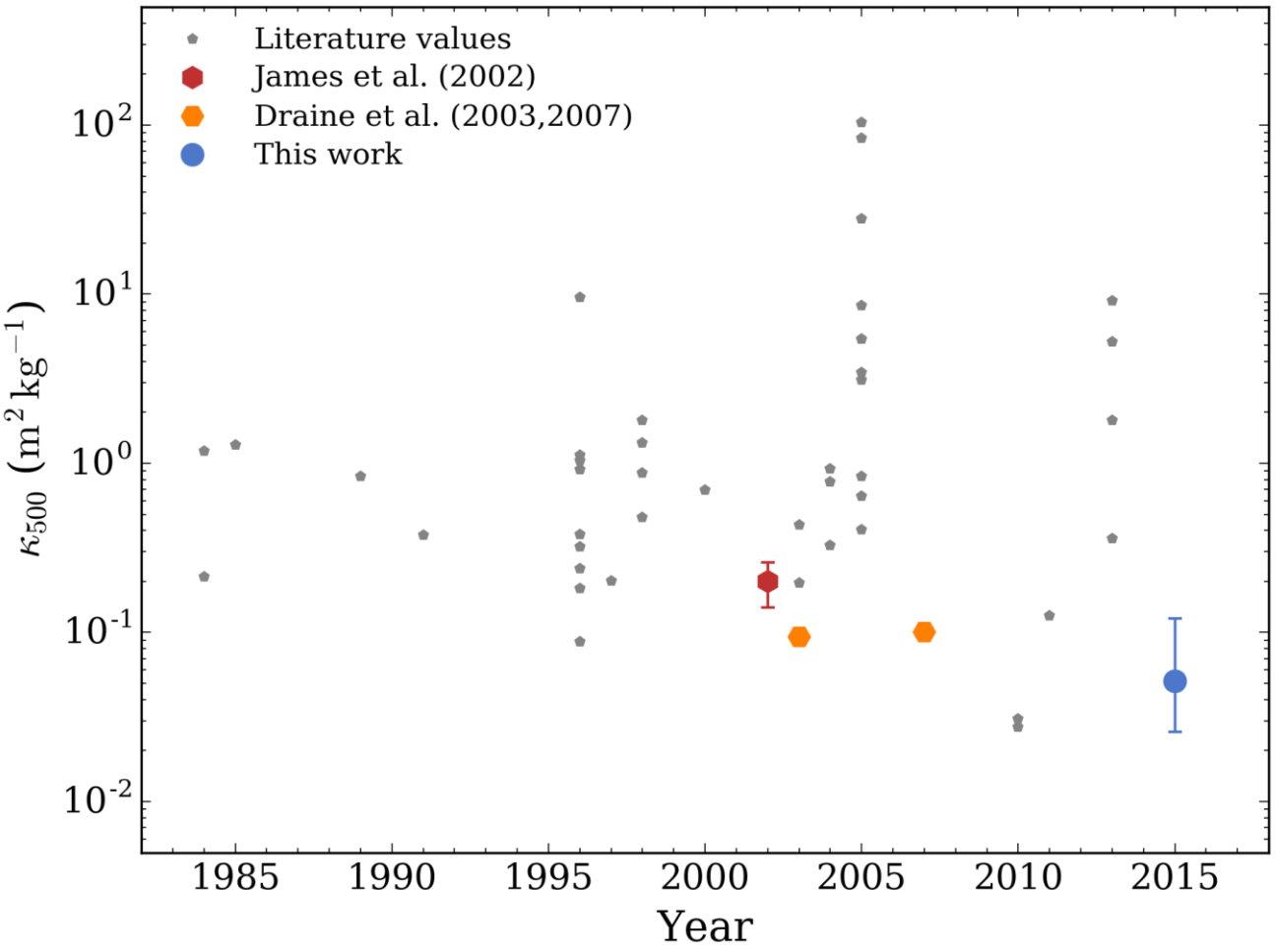
An Empirical Value For κ_d

K_{500} = 0.051 m^2 kg^{-1} (+/- ~0.24 dex)



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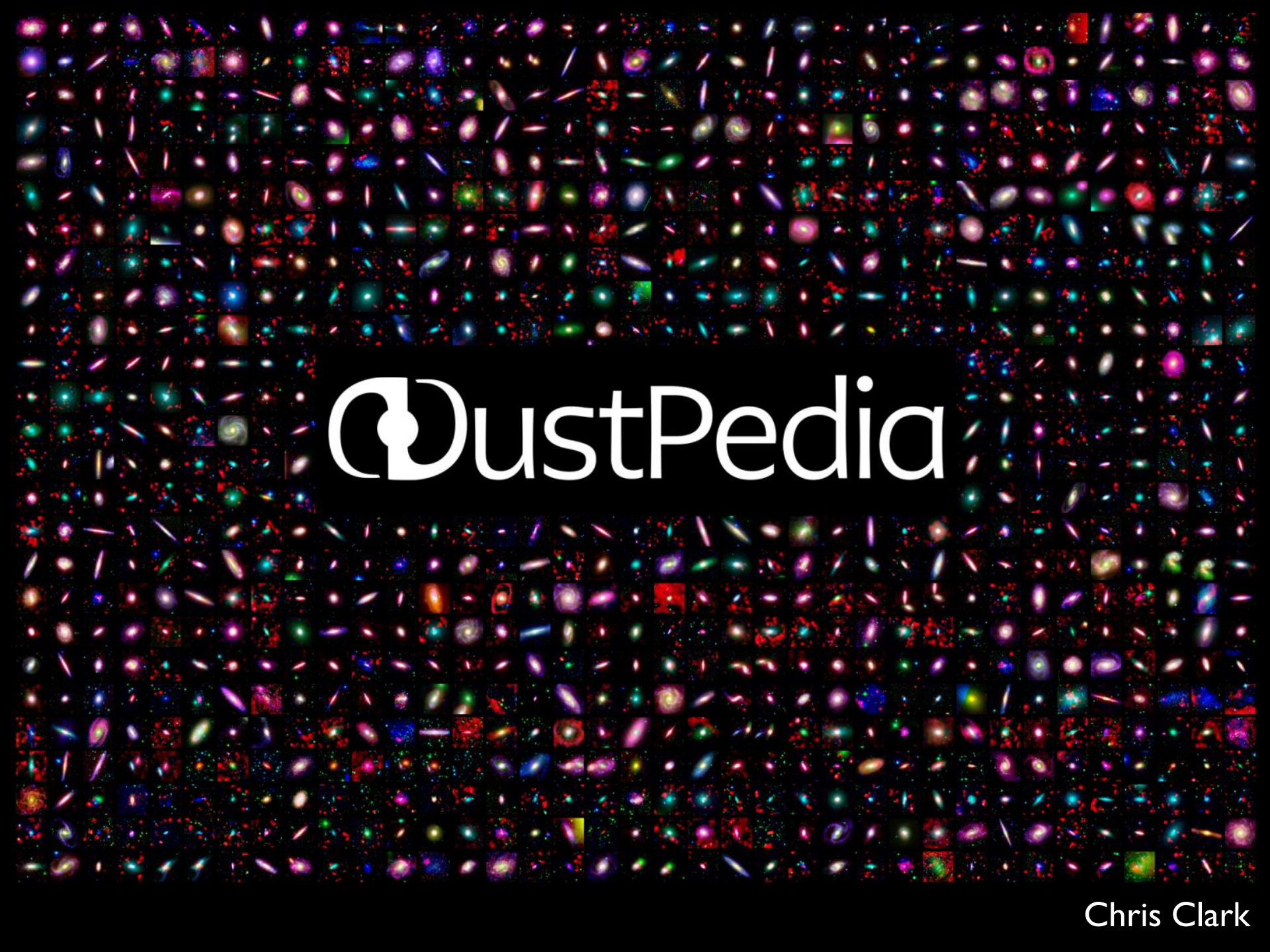

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DustPedia Overview

- THEMIS – A physically-motivated dust model based upon laboratory observations of astrophysical minerals (see papers by A Jones, N Ysard).
- SKIRT – Radiative transfer modelling suite that can fit a dust model to galaxy observations in 3D (see papers by M Baes, P De Camps, S Verstoken).
- A hierarchical Bayesian SED fitting toolkit that can be used to model both galaxies both globally and pixel-by-pixel (see papers by F Galliano).
- A database of multiwavelength imagery and photometry database for all 876 nearby galaxies observed by *Herschel* for which V<3000 km/s and D25<1' (see upcoming paper by me).

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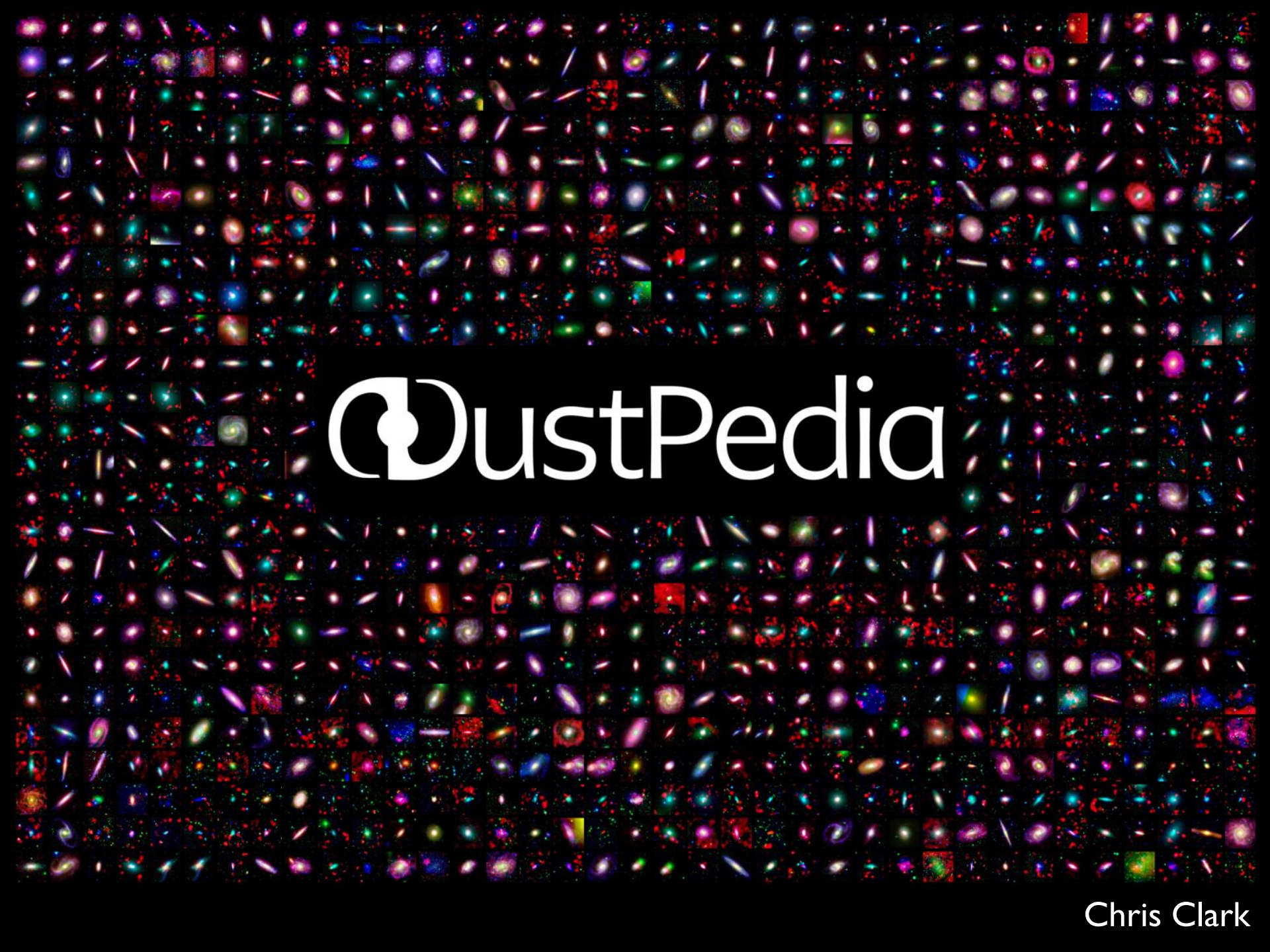

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