

QFT Notes - Chapter 2

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Expressions used: Stress Energy Tensor

$$T_{\nu}^{\mu} = \frac{\partial \mathcal{L}}{\partial(\partial_{\mu}\phi)} \partial_{\nu}\phi - \mathcal{L}\delta_{\nu}^{\mu}$$

Hamiltonian Density

$$H = \int T^{00} d^3x = \int \mathcal{H} d^3x$$

1 2.1 The Klein Gordon Field as Harmonic Oscillations

The simplest type of field is the **Real Klein-Gordon Field**. We will start with a classical field theory and then quantize it.

Second Quantization

Reinterpret the Dynamic Variables as operators that obey canonical commutation relations

We will then solve the theory by **finding the eigenvalues and eigenstates of the Hamiltonian using the Harmonic Oscillator as an analogy**.

To move from classical to quantum field theory, we do what we do with all dynamical systems: **We promoted ϕ and π to operators** and we impose suitable commutation relations.

$$[q_i, p_j] = i\delta_{ij} \tag{1}$$

$$[p_i, p_j] = [q_i, q_j] = 0 \tag{2}$$

For a continuous system, the generalization is quite natural. Because $\pi(x)$ is a the momentum **density**, we get a dirac delta rather than a Kronecker:

$$[\phi(x), \pi(y)] = i\delta^{(3)}(x-y) \quad (3)$$

$$[\phi(x), \phi(y)] = [\pi(x), \pi(y)] = 0 \quad (4)$$

((For now ϕ and π do not depend on time).

Now it is time for the **Hamiltonian**. First of all let's write the Klein-Gordon field in Fourier Space:

$$\phi(x, t) = \int \frac{d^3p}{(2\pi)^3} e^{ip \cdot x} \phi(p, t) \quad (5)$$

(Here we have $\phi^*(p) = -\phi(p)$ so that $\phi(x)$ is real.)

The Klein Gordon Equation becomes

$$\left[\frac{\partial^2}{\partial t^2} + (|p|^2 + m^2) \right] \phi(p, t) = 0 \quad (6)$$

This is the same as the equation of motion for a simple harmonic oscillator with frequency $\omega = \sqrt{|p|^2 + m^2}$

SHO is something we know pretty well.

$$H_{SHO} = \frac{1}{2}p^2 + \frac{1}{2}\omega^2\phi^2 \quad (7)$$

Now to find the eigenvalues of H_{SHO} , we write p and ϕ in terms of ladder operators.

$$\phi = \frac{1}{\sqrt{2\omega}}(a + a^\dagger) \quad (8)$$

$$p = -i\sqrt{\frac{\omega}{2}}(a - a^\dagger) \quad (9)$$

The Hamiltonian can now be written as

$$H_{SHO} = \omega(a^\dagger a + \frac{1}{2}) \quad (10)$$

The state $|0\rangle$ is the ground-state with eigenvalue $\frac{1}{2}\omega$.

Furthermore, the Commutators $[H_{SHO}, a^\dagger] = \omega a^\dagger$ and $[H_{SHO}, a] = -\omega a$
From Quantum Mechanics we know that

$$|n\rangle = (a^\dagger)^n |0\rangle \quad \text{with eigenstate } (n + \frac{1}{2})\omega \quad (11)$$

We can find the spectrum of the Klein Gordon Hamiltonian with the same trick, but now, each Fourier mode of the field is treated as an independent oscillator with it's own creating and annihilation operators.

$$\phi(x) = \int \frac{d^3p}{(2\pi)^3} \frac{1}{\sqrt{2\omega_p}} (a_p e^{ip \cdot x} + a_p^\dagger e^{-ip \cdot x}) \quad (12)$$

$$\pi(x) = \int \frac{d^3p}{(2\pi)^3} -i \sqrt{\frac{\omega_p}{2}} (a_p e^{ip \cdot x} - a_p^\dagger e^{-ip \cdot x}) \quad (13)$$

We will rearrange the above equation as they will be more useful to us in the future.

$$\phi(x) = \int \frac{d^3p}{(2\pi)^3} \frac{1}{\sqrt{2\omega_p}} (a_p + a_{-p}^\dagger) e^{ip \cdot x} \quad (14)$$

$$\pi(x) = \int \frac{d^3p}{(2\pi)^3} -i \sqrt{\frac{\omega_p}{2}} (a_p - a_{-p}^\dagger) e^{ip \cdot x} \quad (15)$$

The commutation relation now becomes

$$[a_p, a_{p'}^\dagger] = i\delta^{(3)}(p - p') \quad (16)$$

Using these we can verify that

$$[\phi(x), \pi(x')] = \int \frac{d^3p d^3p'}{(2\pi)^6} \frac{-i}{2} \sqrt{\frac{\omega_{p'}}{\omega_p}} ([a_{-p'}^\dagger, a_{p'}] - [a_p, a_{-p}^\dagger]) e^{i(p \cdot x + p' \cdot x')} = i\delta^{(3)}(x - x') \quad (17)$$

Now, we can express the Hamiltonian in terms of ladder operators.

$$H = \int d^3x \int \frac{d^3p d^3p'}{(2\pi)^6} e^{i(p+p') \cdot x} \left\{ -\frac{\sqrt{\omega_p \omega_{p'}}}{4} (a_p - a_{-p}^\dagger)(a_{p'} - a_{-p'}^\dagger) + \frac{-p \cdot p' + m^2}{4\sqrt{\omega_p \omega_{p'}}} (a_p + a_{-p}^\dagger)(a_{p'} + a_{-p'}^\dagger) \right\} \quad (18)$$

$$= \int \frac{d^3p}{(2\pi)^3} \omega_p (a_p^\dagger a_p + \frac{1}{2} [a_p, a_p^\dagger]) \quad (19)$$

To find this, all we do is plug the equations (1,20 to the Hamiltonian Expression:

$$H = \int d^3x \left[\frac{1}{2} \pi^2 + \frac{1}{2} (\nabla \phi)^2 + \frac{1}{2} m^2 \phi^2 \right] \quad (20)$$

$$\begin{aligned} &= \frac{1}{2} \int \frac{d^3x d^3p d^3q}{(2\pi)^6} \left[-\frac{\sqrt{\omega_{\vec{p}} \omega_{\vec{q}}}}{2} (a_{\vec{p}} e^{i\vec{p} \cdot \vec{x}} - a_{\vec{p}}^\dagger e^{-i\vec{p} \cdot \vec{x}}) (a_{\vec{q}} e^{i\vec{q} \cdot \vec{x}} - a_{\vec{q}}^\dagger e^{-i\vec{q} \cdot \vec{x}}) \right. \\ &\quad + \frac{1}{2\sqrt{\omega_{\vec{p}} \omega_{\vec{q}}}} (i\vec{p} a_{\vec{p}} e^{i\vec{p} \cdot \vec{x}} - i\vec{p} a_{\vec{p}}^\dagger e^{-i\vec{p} \cdot \vec{x}}) \cdot (i\vec{q} a_{\vec{q}} e^{i\vec{q} \cdot \vec{x}} - i\vec{q} a_{\vec{q}}^\dagger e^{-i\vec{q} \cdot \vec{x}}) \\ &\quad \left. + \frac{m^2}{2\sqrt{\omega_{\vec{p}} \omega_{\vec{q}}}} (a_{\vec{p}} e^{i\vec{p} \cdot \vec{x}} + a_{\vec{p}}^\dagger e^{-i\vec{p} \cdot \vec{x}}) (a_{\vec{q}} e^{i\vec{q} \cdot \vec{x}} + a_{\vec{q}}^\dagger e^{-i\vec{q} \cdot \vec{x}}) \right] \quad (21) \end{aligned}$$

$$\begin{aligned}
&= \frac{1}{4} \int \frac{d^3 p}{(2\pi)^3} \frac{d^3 q}{(2\pi)^3} \left[-\sqrt{\omega_{\vec{p}} \omega_{\vec{q}}} \left(a_{\vec{p}} a_{\vec{q}} \delta(\vec{p} + \vec{q}) - a_{\vec{p}}^\dagger a_{\vec{q}} \delta(-\vec{p} + \vec{q}) - a_{\vec{p}} a_{\vec{q}}^\dagger \delta(\vec{p} - \vec{q}) + a_{\vec{p}}^\dagger a_{\vec{q}}^\dagger \delta(-\vec{p} - \vec{q}) \right) \right. \\
&+ \frac{1}{\sqrt{\omega_{\vec{p}} \omega_{\vec{q}}}} \left(-\vec{p} \cdot \vec{q} a_{\vec{p}} a_{\vec{q}} \delta(\vec{p} + \vec{q}) + \vec{p} \cdot \vec{q} a_{\vec{p}}^\dagger a_{\vec{q}} \delta(-\vec{p} + \vec{q}) + \vec{p} \cdot \vec{q} a_{\vec{p}} a_{\vec{q}}^\dagger \delta(\vec{p} - \vec{q}) - \vec{p} \cdot \vec{q} a_{\vec{p}}^\dagger a_{\vec{q}}^\dagger \delta(-\vec{p} - \vec{q}) \right) \\
&\left. + \frac{m^2}{\sqrt{\omega_{\vec{p}} \omega_{\vec{q}}}} \left(a_{\vec{p}} a_{\vec{q}} \delta(\vec{p} + \vec{q}) + a_{\vec{p}}^\dagger a_{\vec{q}} \delta(-\vec{p} + \vec{q}) + a_{\vec{p}} a_{\vec{q}}^\dagger \delta(\vec{p} - \vec{q}) + a_{\vec{p}}^\dagger a_{\vec{q}}^\dagger \delta(-\vec{p} - \vec{q}) \right) \right] \quad (22)
\end{aligned}$$

$$\begin{aligned}
&= \frac{1}{4} \int \frac{d^3 p}{(2\pi)^3} \left[-\omega_{\vec{p}} a_{\vec{p}} a_{-\vec{p}} + \omega_{\vec{p}} a_{\vec{p}}^\dagger a_{-\vec{p}} + \omega_{\vec{p}} a_{\vec{p}} a_{\vec{p}}^\dagger - \omega_{\vec{p}} a_{\vec{p}}^\dagger a_{-\vec{p}}^\dagger \right. \\
&+ \frac{1}{\omega_{\vec{p}}} \vec{p}^2 a_{\vec{p}} a_{-\vec{p}} + \frac{1}{\omega_{\vec{p}}} \vec{p}^2 a_{\vec{p}}^\dagger a_{\vec{p}} + \frac{1}{\omega_{\vec{p}}} \vec{p}^2 a_{\vec{p}} a_{\vec{p}}^\dagger + \frac{1}{\omega_{\vec{p}}} \vec{p}^2 a_{\vec{p}}^\dagger a_{-\vec{p}}^\dagger \\
&\left. + \frac{m^2}{\omega_{\vec{p}}} a_{\vec{p}} a_{-\vec{p}} + \frac{m^2}{\omega_{\vec{p}}} a_{\vec{p}}^\dagger a_{\vec{p}} + \frac{m^2}{\omega_{\vec{p}}} a_{\vec{p}} a_{\vec{p}}^\dagger + \frac{m^2}{\omega_{\vec{p}}} a_{\vec{p}}^\dagger a_{-\vec{p}}^\dagger \right] \quad (23)
\end{aligned}$$

$$\begin{aligned}
&= \frac{1}{4} \int \frac{d^3 p}{(2\pi)^3} \frac{1}{\omega_{\vec{p}}} \left[(-\omega_{\vec{p}}^2 + \vec{p}^2 + m^2) a_{\vec{p}} a_{-\vec{p}} + (-\omega_{\vec{p}}^2 + \vec{p}^2 + m^2) a_{\vec{p}}^\dagger a_{-\vec{p}}^\dagger + (\omega_{\vec{p}}^2 + \vec{p}^2 + m^2) a_{\vec{p}} a_{\vec{p}}^\dagger + (\omega_{\vec{p}}^2 + \vec{p}^2 + m^2) a_{\vec{p}}^\dagger a_{\vec{p}} \right] \quad (24)
\end{aligned}$$

$$\begin{aligned}
&= \frac{1}{4} \int \frac{d^3 p}{(2\pi)^3} \frac{1}{\omega_{\vec{p}}} \left[(-\omega_{\vec{p}}^2 + \vec{p}^2 + m^2) (a_{\vec{p}} a_{-\vec{p}} + a_{\vec{p}}^\dagger a_{-\vec{p}}^\dagger) + (\omega_{\vec{p}}^2 + \vec{p}^2 + m^2) (a_{\vec{p}} a_{\vec{p}}^\dagger + a_{\vec{p}}^\dagger a_{\vec{p}}) \right] \quad (25)
\end{aligned}$$

$$= \frac{1}{2} \int \frac{d^3 p}{(2\pi)^3} \omega_{\vec{p}} [a_{\vec{p}} a_{\vec{p}}^\dagger + a_{\vec{p}}^\dagger a_{\vec{p}}] \quad (26)$$

$$= \int \frac{d^3 p}{(2\pi)^3} \omega_{\vec{p}} [a_{\vec{p}}^\dagger a_{\vec{p}} + \frac{1}{2} [a_{\vec{p}}, a_{\vec{p}}^\dagger]] \quad (27)$$

The term on the far right is proportional to $\delta(0)$, an infinite c-number. It is the sum over all modes of the zero-point energies $\frac{\omega_{\vec{p}}}{2}$, so its presence is expected. The infinite energy shift cannot be detected experimentally, since experiments measure only energy differences from the ground state of H. Therefore, we will simply ignore this energy shift for now. (*look at epilogue*)

$$= \int \frac{d^3 p}{(2\pi)^3} \omega_{\vec{p}} [a_{\vec{p}}^\dagger a_{\vec{p}} + \frac{1}{2} (2\pi)^3 \delta^{(3)}(0)] \quad (28)$$

Using the expression for the Hamiltonian in terms of $a_{\vec{p}}$ and $a_{\vec{p}}^\dagger$ it is easy to evaluate the commutators:

$$[H, a_{\vec{p}}^\dagger] = \omega_{\vec{p}} a_{\vec{p}}^\dagger \quad [H, a_{\vec{p}}] = -\omega_{\vec{p}} a_{\vec{p}} \quad (29)$$

We can now write the spectrum of theory, just as for the harmonic oscillator. The ground state $|0\rangle$ is defined by the relation $a_p|0\rangle = |0\rangle \forall p$. This is also called the **vacuum state**.

All other energy eigenstates can be built by the creation operators acting on the vacuum.

In general, the state $a_{p_1}^\dagger a_{p_2}^\dagger a_{p_3}^\dagger \dots |0\rangle$ is an eigenstate of H with energy $\sum_j \omega_{p_j}$.

Having found the spectrum of the Hamiltonian we can now interpret its eigenstates.

We can write the total momentum operator with the use of the following equation.

$$P^i = \int T^{0i} d^3x = - \int \pi \partial_i \phi d^3x \quad (30)$$

Carrying out the operators we get:

$$\mathbf{P} = - \int d^3x \pi(\mathbf{x}) \nabla \phi(x) = \int \frac{d^3p}{(2\pi)^3} \mathbf{p} a_{\mathbf{p}}^\dagger a_{\mathbf{p}} \quad (31)$$

That mean that the operator a_p^\dagger creates momentum p and energy $\omega_p = \sqrt{|p|^2 + m^2}$. It is natural to call these excitation **particles**, since they are discrete entities that have the proper relativistic energy-momentum relation. **From now on we will refer to ω_p as E_p or simply E , as it really is the energy of the particle.**

We choose to normalize the vacuum state so that $\langle 0|0\rangle = 1$. The one-particles states $|p\rangle \propto a_p^\dagger |0\rangle$ will also appear quite often, and it is worthwhile to adopt a convention for their normalization. The simplest one is $\langle p|q\rangle = (2\pi)^3 \delta^{(3)}(p-q)$. This is not Lorentz invariant, which we can demonstrate by considering the effect of a boost in the 3-direction.¹

¹Peskin and Schroeder pages 22-24

2 2.2 The Klein Gordon Field in Space-Time

In this section, we will switch from the Schroedinger picture to the Heisenberg. Here, we need to introduce time-dependency in our operators.

$$\phi(x) = \phi(x, t) = e^{iHt} \phi(x) e^{-iHt} \quad (32)$$

and similarly for $\pi(x) = \pi(x, t)$.

The Heisenberg equation of motion allows to compute the time dependence of ϕ and π .

$$i \frac{\partial}{\partial t} \mathcal{O} = [\mathcal{O}, H] \quad (33)$$

$$\begin{aligned} i \frac{\partial}{\partial t} \phi(x, t) &= [\phi(x, t), \int d^3 x' \{ \frac{1}{2} \pi^2(x', t) + \frac{1}{2} (\nabla \phi(x, t))^2 + \frac{1}{2} m^2 \phi^2(x', t) \}] \quad (34) \\ &= \int d^3 x' (i \delta^{(3)}(x - x') \pi(x', t)) = i \pi(x, t) \end{aligned}$$

$$\begin{aligned} i \frac{\partial}{\partial t} \pi(x, t) &= [\pi(x, t), \int d^3 x' \{ \frac{1}{2} \pi^2(x', t) + \frac{1}{2} (\nabla \phi(x, t))^2 + \frac{1}{2} m^2 \phi^2(x', t) \}] \quad (35) \\ &= \int d^3 x' (-i \delta^{(3)}(x - x') (-\nabla^2 + m^2) \phi(x', t)) = i (-\nabla^2 + m^2) \phi(x, t) \end{aligned}$$

Combining the two results we get the **Klein Gordon Equation**

$$\frac{\partial^2}{\partial t^2} \phi = (\nabla^2 - m^2) \phi \quad (36)$$

It is easier to understand the time dependence of ϕ and π by writing them both in terms of creation and annihilation operators.

Note that for any n,

$$H a_p = a_p (H - E_p) \quad \text{and} \quad H^n a_p = a_p (H - E_p)^n$$

(similarly for a_p^\dagger , but instead of - we get +)

This means we have now derived the identities:

$$e^{iHt} a_{\mathbf{p}} e^{-iHt} = a_{\mathbf{p}} e^{-iE_{\mathbf{p}}t}$$

$$e^{iHt} a_{\mathbf{p}}^\dagger e^{-iHt} = a_{\mathbf{p}}^\dagger e^{iE_{\mathbf{p}}t}$$

We can now use this to find the Heisenberg Operator $\phi(x, t)$, which is defined in the equation 32.

$$\phi(x, t) = \int \frac{d^3p}{(2\pi)^3} \frac{1}{\sqrt{E_{\mathbf{p}}}} [a_{\mathbf{p}} e^{-ipx} + a_{\mathbf{p}}^\dagger e^{ipx}] \Big|_{p^0=E_{\mathbf{p}}} \quad (37)$$

and obviously

$$\pi(x, t) = \partial_t \phi(x, t) \quad (38)$$

We can use the same exact manipulations for \mathbf{P} instead of H to relate $\phi(x)$ to $\phi(0)$

$$\phi(x) = e^{iP \cdot x} \phi(0) e^{-iP \cdot x} \text{ (Unitarity in QM)}$$

where $P^\mu = (H, \mathbf{P})$

\mathbf{P} is the momentum operator, whose eigenvalue is the total momentum of the system. On the other hand, \mathbf{p} is the momentum of a single Fourier mode of the field, which we interpret as the momentum of a particle in that mode.

Equation 37 shows us the wave-particle duality of the quantum field. We see it as an operator in a Hilbert Space, and then we interpret it as a linear combination of the solutions $e^{\pm ip \cdot x}$, to the Klein-Gordon Equation. Both signs of the time dependence in the exponential appear: We find both $e^{-ip_0 t}$ and $e^{+ip_0 t}$, although p_0 is always positive. This is equivalent to positive and negative energies, that we interpret as particles and antiparticles. If these were single-particle wavefunctions, they would correspond to states of positive and negative energy; let us refer to them more generally as positive- and negative-frequency modes. The connection between the particle creation operators and the waveforms displayed here is always valid for free quantum fields: A positive-frequency solution of the field equation has as its coefficient the operator that destroys a particle in that single-particle wavefunction. A negative-frequency solution of the field equation, being the Hermitian conjugate of a positive-frequency solution, has as its coefficient the operator that creates a particle in that positive-energy single-particle wavefunction. In this way, the fact that relativistic wave equations have both positive- and negative-frequency solutions is reconciled with the requirement that a sensible quantum theory contain only positive excitation energies.

Causality

Let's go back to the important question that we raised earlier:

We are still working in the Heisenbergian framework, and the probability of a particle propagating from y to x , in vacuum states is given by:

$$\langle 0 | \phi(x) \phi(y) | 0 \rangle$$

We shall call this quantity $D(x-y)$

$$D(x-y) = \langle 0 | \phi(x) \phi(y) | 0 \rangle$$

Each operator $\phi(x)$, as seen in equation 37, are sums of ladder operators, so we see that, using equation 37, $D(x-y)$ evaluates to

$$\langle 0|\phi(x)\phi(y)|0\rangle = \int \frac{d^3p}{(2\pi)^3} \frac{1}{2E_p} e^{ip \cdot (x-y)}$$

We have seen (Equation. 2.40 Peskin and Schroeder) that integrals of this type are Lorentz invariant, which intuitively makes a lot of sense.

Let's try to evaluate the Integral for some values of $x-y$. Consider an example where this difference is purely in the 0th index, meaning time:

$$x^0 - y^0 = t \quad \mathbf{x} - \mathbf{y} = 0$$