

Fig. 9-1 right:

A single lens brings rays passing close to its edge to focus at a different point from those passing near the centre (a). This is because the surfaces of the lens are spherically curved, and this fault is known as **spherical aberration**. A single lens also brings rays of different colours to focus at different points (b). This is because the lens disperses white light into its separate colours. This is known as **chromatic aberration**. A two component lens, with one component having convex surfaces and the other a plane surface and a concave one (c) may be designed to minimize spherical aberration and also to overcome chromatic aberration for two specific colours.

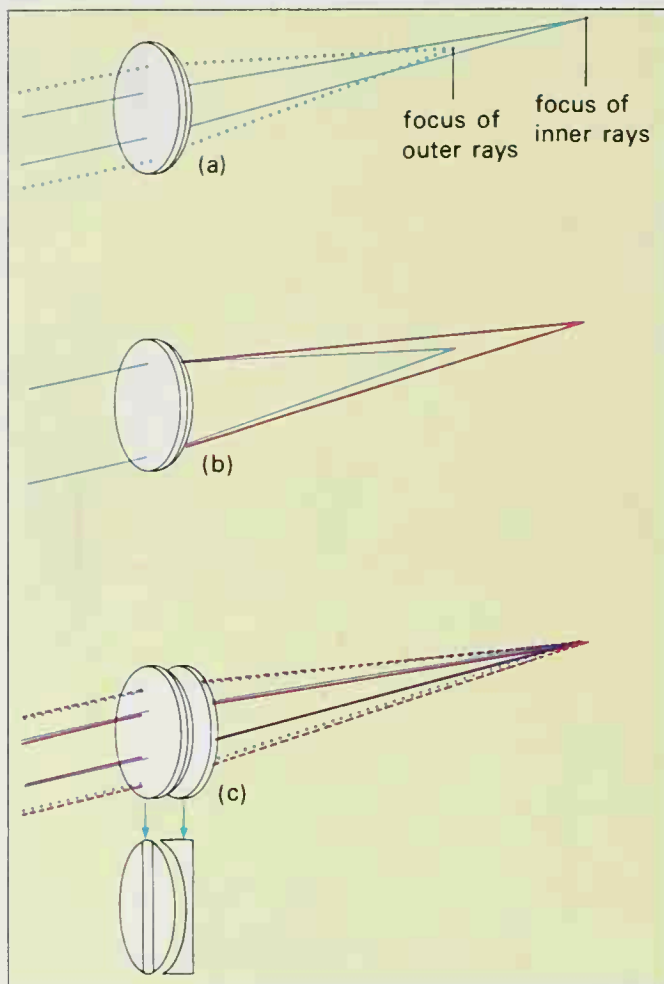


Fig. 9-2 top right:

Use of a concave mirror to bring incoming light to a focus.

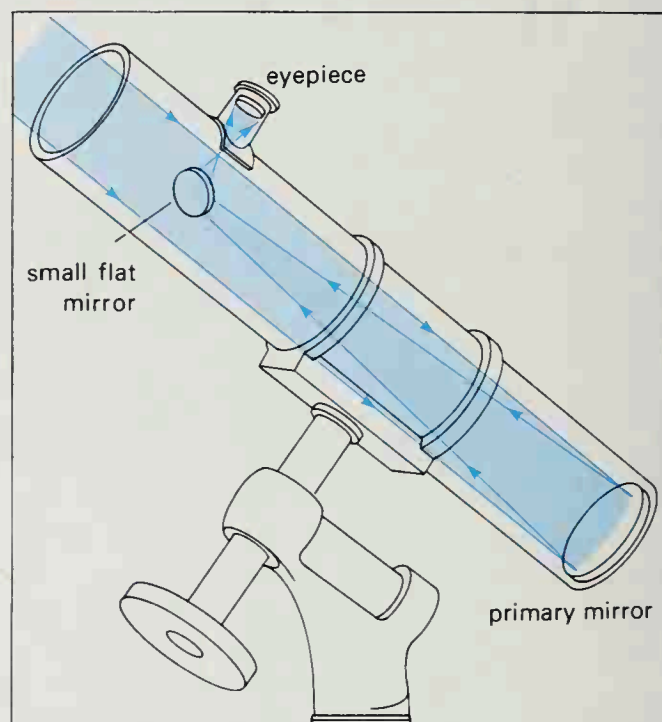
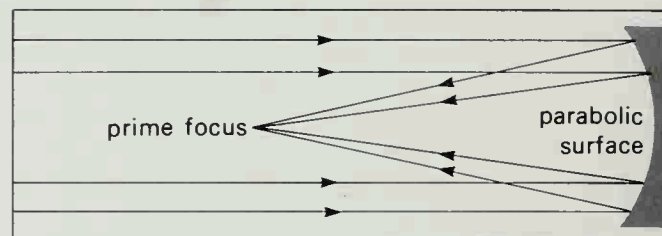


Fig. 9-4 far right, centre: The Newtonian focus as used in many amateur reflecting telescopes.

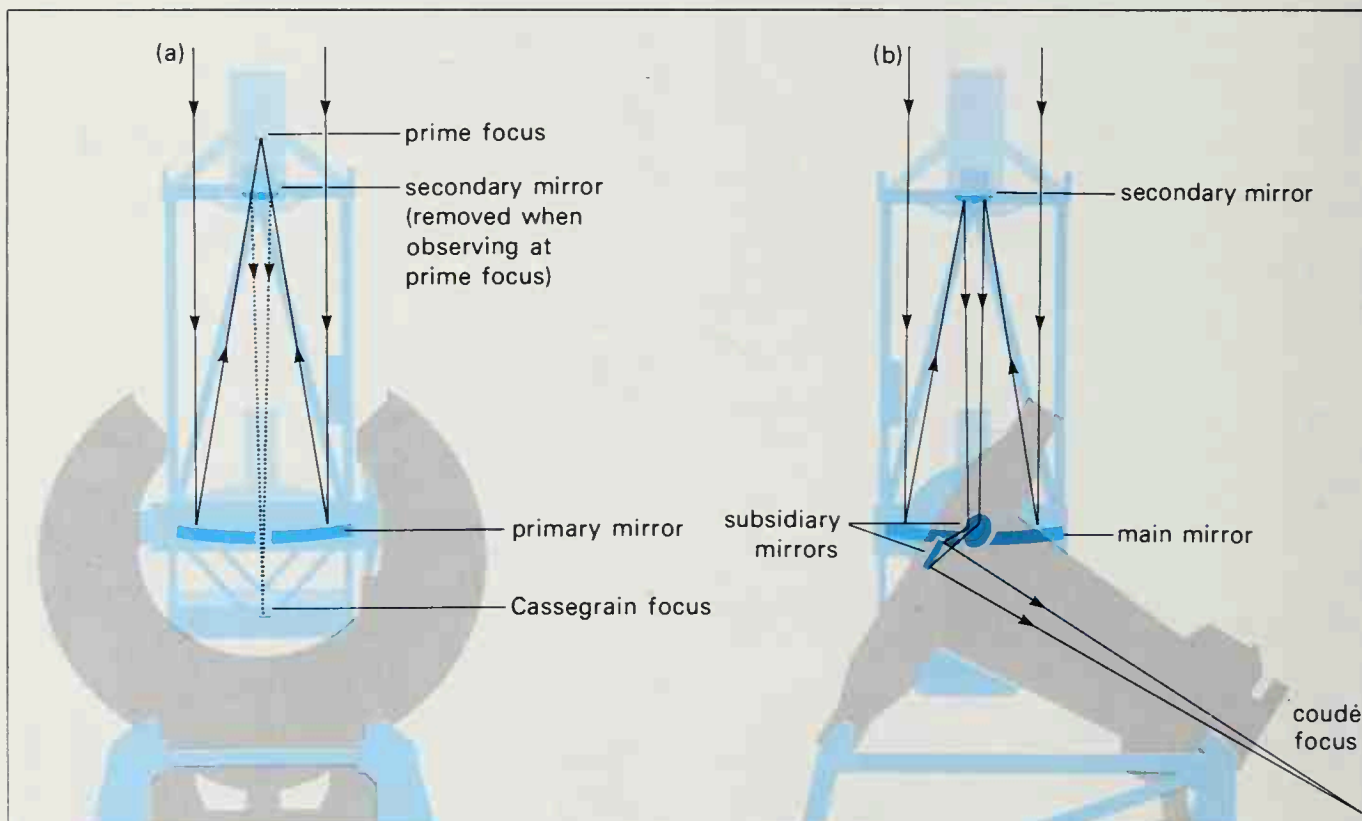


Fig. 9-3

Two drawings of a modern telescope showing the prime focus and Cassegrain focus (which lies below the primary mirror) in (a), and the fixed coudé focus (b).

## The Schmidt and other telescopes

Even with all modern refinements, the field of view of a large reflecting telescope is small. The area of really sharp definition on a photograph is about  $1^\circ$  square and although this is perfectly adequate for the detailed examination of specific objects, it is too small an area for any large scale surveys or statistical work. Something with a wider field of view is required for these. In the 1930s the Estonian optician Bernhard

Schmidt developed a new wide-angle telescope especially for photographic work, and since then the Schmidt telescope has become one of the optical astronomer's most powerful tools. The instrument uses a spherical primary mirror and at the front of the tube there is a correcting plate. This has a complex shape (Fig. 9-5, page 232) and can be said to parabolize the light as it enters the tube. The spherical primary gives a wide field coverage, typically some  $40^\circ$  square and although the telescope cannot