



be used visually, such an instrument turns out to be a most powerful astronomical camera. This is not only because of its wide field but also because of its low focal ratio. As all photographers know, a lens opened up to aperture $f/2.5$ is faster than one at $f/3.5$; in fact, the shorter the exposure required the lower the focal ratio must be. Yet even the latest modern reflectors can never operate at lower than $f/3.5$, but Schmidts of $f/2.5$ are not uncommon. Another catadioptric type of reflector with a spherical primary and

using a correcting lens at the front of a very short tube, is that designed by Dmitry Maksutov in the 1940s. First developed as a photographic telescope, it can also be used visually at $f/7$ and $f/8$ and is now favoured by some amateurs because of its extreme portability. Short tube lengths and fast focal ratios are only one aspect in the design of optical telescopes. Light-grasp necessitates a large aperture, and large aperture is also bound up with the vital question of resolution, since the power of a telescope to

The Southern Cross. A star shines because its internal energy source is due to the annihilation of nuclear particles. The energy is immense because it follows the Einstein formula $E = mc^2$.