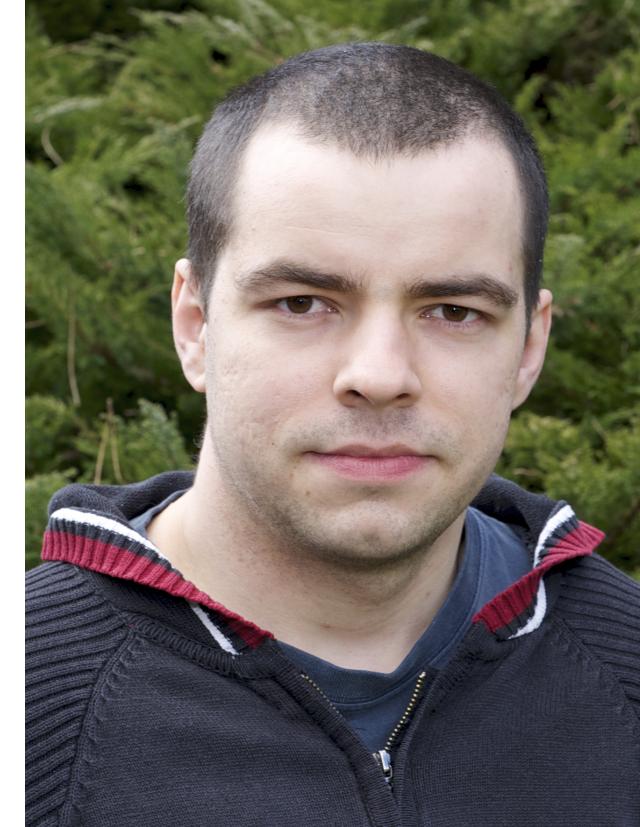


# 2D PSF characterization

Neven Caplar, Joshua Mayers (Princeton University)



PSF of the Prime Focus Spectrograph will be caused by 3 different components: telescope pupil illumination, focal ratio degradation in the fibers and the camera. Goal of the project is to understand and to characterize the contribution of the camera imperfections to the PSF. This will also enable us to model separately different contributions to the PSF. We have started analysing defocused images taken by the red arm camera which has been mounted at LAM. At the moment, we are using package Galsim and forward modelling of the wavefront to estimate Zernike coefficients describing the wavefront aberrations. In this early stage, we are also investigating different methods and procedures on how to best model the system.

To get the data, 10 fibers are illuminated with HgAr lamp (see Figures on the right and below). Slit is moved in order to produce defocused images.

Figure below shows 100 spots across the detector (data taken on November 2)

- Numbers above each image denote the id of each fiber and the wavelength of the light in nm, respectively
- Note the movement of the detector and three supporting struts as a function of the position on the detector
- Also visible is the images below the slit (not shown on the right) which intersect the beam

