

INDEXES

n : grid point index

p : pixel index

l : line index

MODIFIERS

\vec{q} : sequence

\tilde{q} : model approximation

\hat{q} : continuous value

SYMBOLS

ρ : stellar properties

S : PHOENIX spectrum

\vec{G} : PHOENIX grid

$\vec{\phi}$: Flux vector

$\vec{\lambda}$: Wavelength axis

μ : Line center

A : Line amplitude

σ : Gaussian line width

γ : Lorentzian line width

M : Manifold model

Φ : Model flux vector

$$\rho = \left[T_{eff}, \log{(g)}, [\mathrm{Fe}/\mathrm{H}]\right]$$

$$\rho_n \mapsto S_n$$

$$S_n = \begin{bmatrix} \vec{\phi}_n := (\phi_{np})_{p=1}^P \\ \vec{\lambda} := (\lambda_p)_{p=1}^P \end{bmatrix}$$

$$\vec{G} = (S_n)_{n=1}^N$$

$$\vec{\phi}_n \mapsto \left(\begin{bmatrix} \mu_l & A_{nl} & \sigma_{nl} & \gamma_{nl} \end{bmatrix} \right)_{l=1}^L$$

$$M_l:\vec{\rho}\rightarrow \begin{bmatrix} \vec{A}_l & \vec{\sigma}_l & \vec{\gamma}_l \end{bmatrix}:=\left(\begin{bmatrix} A_{nl} & \sigma_{nl} & \gamma_{nl} \end{bmatrix}\right)_{n=1}^N$$

$$\tilde{M}_l \approx M_l$$

$$\tilde{M}_l:\vec{\rho}\rightarrow \begin{bmatrix} \vec{\tilde{A}}_l & \vec{\tilde{\sigma}}_l & \vec{\tilde{\gamma}}_l \end{bmatrix}$$

$$\vec{\tilde{M}} = (\tilde{M}_l)_{l=1}^L$$

$$\vec{\tilde{M}}(\hat{\rho}) = \left(\begin{bmatrix} \hat{A}_l & \hat{\sigma}_l & \hat{\gamma}_l \end{bmatrix}\right)_{l=1}^L$$

$$\left[\left(\mu_l\right)_{l=1}^L\mid \vec{\tilde{M}}(\hat{\rho})\right]\mapsto \vec{\Phi}$$