

# A First-Principles Origin of the Inverse-Square Law in Swirl-String Theory: Three Derivations from Local Field Mediation and Momentum-Flux Conservation

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## Abstract

A recurrent objection to flat-background, emergent-gravity programs is that the inverse-square distance law is often imported rather than derived. We close this gap in the weak-field, static monopole sector by exhibiting three independent derivations of the  $1/r$  potential and  $1/r^2$  flux. (I) A Gauss-law scalar effective field theory (EFT) for the far-field mediator yields the Poisson equation whose Green's function in  $\mathbb{R}^3$  is  $1/r$ . (II) Identifying the specific SST far-field degree of freedom as a foliation/clock scalar, we compute its quadratic EFT stress tensor  $T_{ij}$  and show the associated conserved radial flux density scales as  $1/r^2$ , with charge  $Q \propto \int \rho_m d^3x$ . (III) We replace the Newtonian potential  $\chi$  by a foliation ("khronon-like") scalar and show that, in the static weak-field monopole limit, the EFT reduces to the same Gauss-law scalar form as in (I), so the  $1/r^2$  law follows automatically. These results demonstrate that once SST commits to a local mediator in three spatial dimensions, inverse-square behavior is not an assumption but a consequence of locality, symmetry, and the  $\mathbb{R}^3$  Green's function.

## 1 Setup and scope

We work in a flat operational background with Minkowski causal structure and consider the weak-field, static, spherically symmetric (monopole) sector. Let  $\rho_m(\mathbf{x})$  be the rest-mass density of a compact source with total mass

$$M = \int_{\mathbb{R}^3} \rho_m(\mathbf{x}) d^3x. \quad (1)$$

The goal is to derive, from first principles, that the far-field gravitational influence (whatever SST field represents it) must exhibit

$$\Phi(\mathbf{x}) \sim \frac{1}{r}, \quad \nabla \Phi \sim \frac{1}{r^2}, \quad r = \|\mathbf{x}\|. \quad (2)$$

We present three derivations that differ in emphasis but agree in content.

## 2 Derivation I: Gauss-law scalar EFT $\Rightarrow 1/r$ Green's function

### 2.1 Specify the mediator and write the quadratic EFT

In the static weak-field monopole sector, the minimal local mediator is a scalar field  $\phi(\mathbf{x})$  coupled linearly to the source density. The most general rotationally invariant quadratic functional (Euclidean

static limit of a Lorentzian EFT) is

$$S_{\text{stat}}[\phi] = \int_{\mathbb{R}^3} d^3x \left[ \frac{\kappa}{2} (\nabla\phi)^2 - \lambda \phi \rho_m(\mathbf{x}) \right], \quad (3)$$

with constants  $\kappa > 0$  and coupling  $\lambda$ .

## 2.2 Euler–Lagrange equation: Poisson form

Varying (3):

$$\delta S_{\text{stat}} = \int d^3x [\kappa \nabla\phi \cdot \nabla(\delta\phi) - \lambda \rho_m \delta\phi] \quad (4)$$

$$= \int d^3x [-\kappa (\nabla^2\phi) \delta\phi - \lambda \rho_m \delta\phi] \quad (\text{integrate by parts, drop boundary term}) \quad (5)$$

so stationarity for arbitrary  $\delta\phi$  gives

$$\kappa \nabla^2\phi(\mathbf{x}) = -\lambda \rho_m(\mathbf{x}). \quad (6)$$

Define the “Gauss-law charge density”  $\rho_Q := \rho_m$  and total charge

$$Q := \int \rho_Q d^3x = \int \rho_m d^3x = M, \quad (7)$$

so the source is monopolar with charge  $Q$ .

## 2.3 Solve using the Green’s function of $\nabla^2$ on $\mathbb{R}^3$

The Green’s function  $G(\mathbf{x})$  satisfying

$$\nabla^2 G(\mathbf{x}) = -4\pi \delta^{(3)}(\mathbf{x}) \quad (8)$$

is

$$G(\mathbf{x}) = \frac{1}{\|\mathbf{x}\|}. \quad (9)$$

This is a standard result:  $G(r) = 1/r$  is the unique (up to addition of harmonic functions) spherically symmetric fundamental solution on  $\mathbb{R}^3$  [1, 2].

Convoluting (6) with  $G$  gives

$$\phi(\mathbf{x}) = \frac{\lambda}{4\pi\kappa} \int d^3x' \frac{\rho_m(\mathbf{x}')}{\|\mathbf{x} - \mathbf{x}'\|}. \quad (10)$$

In the far field  $r \gg$  source size,  $\|\mathbf{x} - \mathbf{x}'\| \approx r$  and

$$\phi(r) \simeq \frac{\lambda}{4\pi\kappa} \frac{1}{r} \int \rho_m(\mathbf{x}') d^3x' = \frac{\lambda}{4\pi\kappa} \frac{Q}{r}. \quad (11)$$

Thus

$$\nabla\phi(r) \simeq -\frac{\lambda Q}{4\pi\kappa} \frac{\hat{\mathbf{r}}}{r^2}, \quad (12)$$

which is the inverse-square flux law.

## 2.4 Gauss law from divergence theorem

Define the flux density

$$\mathbf{J} := -\kappa \nabla \phi. \quad (13)$$

Then (6) is precisely

$$\nabla \cdot \mathbf{J} = \lambda \rho_m. \quad (14)$$

Integrating over a ball  $B_R$  and applying divergence theorem:

$$\oint_{S_R} \mathbf{J} \cdot d\mathbf{A} = \int_{B_R} \nabla \cdot \mathbf{J} d^3x = \lambda \int_{B_R} \rho_m d^3x \xrightarrow{R \rightarrow \infty} \lambda Q. \quad (15)$$

Spherical symmetry implies  $\mathbf{J} = J_r(r)\hat{\mathbf{r}}$ , hence

$$4\pi r^2 J_r(r) = \lambda Q \quad \Rightarrow \quad J_r(r) = \frac{\lambda Q}{4\pi} \frac{1}{r^2}. \quad (16)$$

Since  $\mathbf{J} \propto \nabla \phi$ , this is equivalent to (12).

**Interpretation (10-year-old analogy).** Imagine “influence” leaving a source and spreading equally in all directions. The surface area of a sphere grows like  $r^2$ , so the influence per square meter must drop like  $1/r^2$ .

## 3 Derivation II: Identify the SST far-field carrier, compute $T_{ij}$ , and extract the $1/r^2$ flux

### 3.1 Which SST field carries far-field momentum flux?

In SST language, the long-range static field is taken to be carried by a *clock/foliation* mode: a scalar that labels preferred-time hypersurfaces (“swirl-clock”). Denote this field by  $T(x)$  and consider small perturbations about an inertial foliation:

$$T(x) = t + \tau(x), \quad (17)$$

where  $t$  is the operational background time coordinate and  $\tau$  is a weak perturbation sourced by matter.

At quadratic order, the most general Lorentz-invariant action for  $\tau$  (ignoring higher derivatives) is

$$S[\tau] = \int d^4x \left[ \frac{\kappa}{2} \partial_\mu \tau \partial^\mu \tau - \lambda \tau \rho_m(\mathbf{x}) \right], \quad (18)$$

with  $\rho_m$  treated as static,  $\partial_t \rho_m = 0$ . The static sector of (18) reduces to (3) with  $\phi \equiv \tau$ .

### 3.2 Compute the stress-energy tensor

From (18), the symmetric stress-energy tensor (metric variation, or canonical symmetrized) for the free part is

$$T_{\mu\nu}^{(\tau)} = \kappa \left( \partial_\mu \tau \partial_\nu \tau - \frac{1}{2} \eta_{\mu\nu} \partial_\alpha \tau \partial^\alpha \tau \right). \quad (19)$$

In the static regime,  $\partial_0 \tau = 0$ , so  $\partial_\alpha \tau \partial^\alpha \tau = -(\nabla \tau)^2$  and

$$T_{ij}^{(\tau)} = \kappa \left( \partial_i \tau \partial_j \tau - \frac{1}{2} \delta_{ij} (\nabla \tau)^2 \right), \quad T_{00}^{(\tau)} = \frac{\kappa}{2} (\nabla \tau)^2. \quad (20)$$

### 3.3 Monopole solution and the *conserved* radial flux density

From Derivation I, for  $r$  outside the source,

$$\tau(r) = \frac{\lambda Q}{4\pi\kappa} \frac{1}{r}, \quad \partial_r \tau(r) = -\frac{\lambda Q}{4\pi\kappa} \frac{1}{r^2}. \quad (21)$$

Now define the *Gauss-law flux* associated to the foliation scalar (this is the conserved radial “momentum-like” flux density in the static sector):

$$\mathcal{F}_r(r) := -\kappa \partial_r \tau(r). \quad (22)$$

Using (21),

$$\mathcal{F}_r(r) = \frac{\lambda Q}{4\pi} \frac{1}{r^2}. \quad (23)$$

This is the requested monopole scaling:

$$\boxed{\mathcal{F}_r(r) \propto \frac{1}{r^2}, \quad Q \propto \int \rho_m d^3x.}$$

The proportionality to the enclosed mass is explicit via  $Q = M$  from (7). The crucial point is that this  $1/r^2$  law is enforced by  $\nabla \cdot (\kappa \nabla \tau) = -\lambda \rho_m$  and the divergence theorem, i.e. by local field mediation in  $\mathbb{R}^3$ .

### 3.4 How $T_{ij}$ encodes momentum transport

Although the conserved Gauss-law flux (22) is *linear* in  $\nabla \tau$  and therefore scales as  $1/r^2$ , the mechanical stress carried by the field is quadratic and scales as  $(\nabla \tau)^2 \sim 1/r^4$ . For completeness, the radial traction (pressure/tension) on a sphere is

$$t_r(r) := \hat{r}_i T_{ij}^{(\tau)} \hat{r}_j = \kappa \left[ (\partial_r \tau)^2 - \frac{1}{2} (\partial_r \tau)^2 \right] = \frac{\kappa}{2} (\partial_r \tau)^2 \propto \frac{1}{r^4}. \quad (24)$$

However, the *integrated* traction force over the sphere involves the area factor:

$$F_{\text{field}}(r) = \int_{S_r} t_r dA \sim 4\pi r^2 \times \frac{1}{r^4} \sim \frac{1}{r^2}. \quad (25)$$

Thus the field’s stress tensor is consistent with the inverse-square scaling of net force, while the conserved Gauss-law flux density (23) provides the most direct “flux” statement.

**Interpretation (10-year-old analogy).** The field is like a stretched rubber sheet. The *slope* (gradient) gets smaller like  $1/r^2$  because the “amount of slope” has to spread over bigger spheres. The *stretching energy* depends on slope squared, so it falls off faster.

## 4 Derivation III: Replace $\chi$ by a foliation scalar; static weak-field ⇒ Gauss-law scalar form

### 4.1 Foliation scalar as the gravitational potential variable

Many covariant “preferred foliation” theories introduce a scalar  $T(x)$  whose gradient defines a timelike direction. Define the unit timelike vector

$$u_\mu := \frac{\partial_\mu T}{\sqrt{\partial_\alpha T \partial^\alpha T}}, \quad (26)$$

and consider a general quadratic action in derivatives of  $u_\mu$  (the usual Einstein–*aether*/khronon class) [3, 4]. In the weak-field, static, monopole limit, only the scalar (spin-0) sector contributes at leading order, and one can parameterize the effective action for  $\tau$  by the leading operator  $(\nabla\tau)^2$ .

Concretely, expand about  $T = t$ :

$$T = t + \tau, \quad \partial_\mu T = (1, \nabla\tau), \quad \sqrt{\partial_\alpha T \partial^\alpha T} = \sqrt{1 - (\nabla\tau)^2} \approx 1 - \frac{1}{2}(\nabla\tau)^2. \quad (27)$$

To quadratic order, the effective static Lagrangian necessarily contains

$$\mathcal{L}_{\text{eff,stat}} = \frac{\kappa}{2} (\nabla\tau)^2 - \lambda \tau \rho_m + (\text{higher derivatives / higher powers}). \quad (28)$$

This is exactly the Gauss-law scalar functional (3). Therefore, independent of the microphysical foliation completion, the static weak-field monopole sector reduces to the same Poisson equation (6).

## 4.2 Identify $\chi$ as a reparameterization of $\tau$

Define the Newtonian potential variable  $\chi$  as a linear rescaling of  $\tau$ :

$$\chi := \alpha \tau, \quad (29)$$

for some constant  $\alpha$  chosen so that test-particle equations match the conventional Newtonian limit. Then (28) becomes

$$\mathcal{L}_{\text{eff,stat}} = \frac{\kappa}{2\alpha^2} (\nabla\chi)^2 - \frac{\lambda}{\alpha} \chi \rho_m. \quad (30)$$

Variation yields

$$\nabla^2 \chi = - \left( \frac{\alpha \lambda}{\kappa} \right) \rho_m. \quad (31)$$

Thus, with the identification

$$\frac{\alpha \lambda}{\kappa} = 4\pi G, \quad (32)$$

the foliation scalar reproduces the Gauss-law scalar Poisson equation of Newtonian gravity. Its monopole solution is therefore

$$\chi(r) = -\frac{GM}{r}, \quad \nabla\chi(r) = +GM \frac{\hat{\mathbf{r}}}{r^2}, \quad (33)$$

which is the inverse-square law.

**Interpretation (10-year-old analogy).** Calling the field  $\chi$  or calling it a “clock field” does not change the math. If it obeys the same Poisson equation, it must have the same  $1/r$  shape.

## 5 Discussion and outlook (SST identification of $G$ kept here)

The three derivations show that once SST specifies a *local* far-field mediator (identified here with a foliation/clock scalar in the monopole sector), the  $1/r^2$  law is forced by:

1. the form of the quadratic EFT and its Euler–Lagrange equation;
2. the unique  $\mathbb{R}^3$  Green’s function  $G(r) = 1/r$  for  $\nabla^2$ ;
3. flux conservation (divergence theorem) implying  $4\pi r^2 \mathcal{F}_r = \text{const.}$

In SST, one additionally seeks to *derive the normalization*  $G$  (not merely the  $1/r^2$  shape) from microphysical constants. In the SST program, this corresponds to matching the EFT coefficient combination in (32) to a derived expression  $G_{\text{swirl}}$  constructed from SST's canonical constants. This matching fixes  $\lambda/\kappa$  (and the rescaling  $\alpha$ ) in terms of SST parameters and thereby provides a first-principles value for  $G$  in the low-energy monopole limit. The present paper establishes the distance-law part independently of that normalization step.

## Appendix A: How this connects to other SST manuscripts (non-cited in main text)

This appendix is informational and does not enter the main derivations. The “Kelvin-mode suppression” result supports treating the far-field source as effectively rigid in ordinary conditions (so the monopole approximation is stable). Thermodynamic SST results provide a microscopic route to compute the EFT stiffness  $\kappa$  and coupling  $\lambda$  by coarse-graining the condensate’s equation of state. Variational particle-structure results motivate that unique stable configurations select unique effective couplings, suggesting that  $G$  can be obtained by matching SST microphysics to the EFT coefficients in (32).

## References

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