

International Space Weather Summer Camp 2022: How to turn measurements into an analytical model – at the example of VLF data

David Wenzel

German Aerospace Center (DLR)
Institute of Solar-Terrestrial Physics
Space Weather Observations

E-Mail: david.wenzel@dlr.de

A large, curved image of the Earth from space occupies the right half of the slide. It shows a portion of the Arctic region with white ice and surrounding green landmasses and blue oceans. The curvature of the planet is clearly visible.

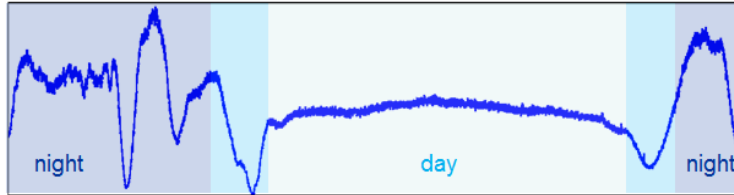
Knowledge for Tomorrow

Daily variation



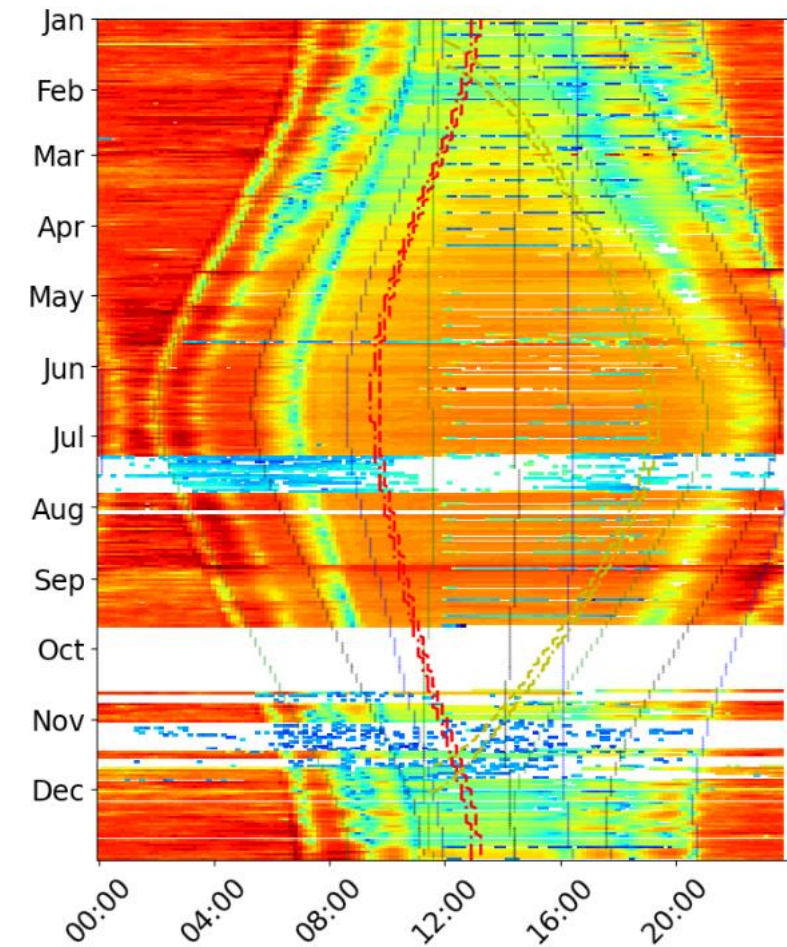
Measurements

- !!! Screenshots from previous day (color plot and a 24hr curve)
- quiet vs. oscillating segments suggest diurnal dependence



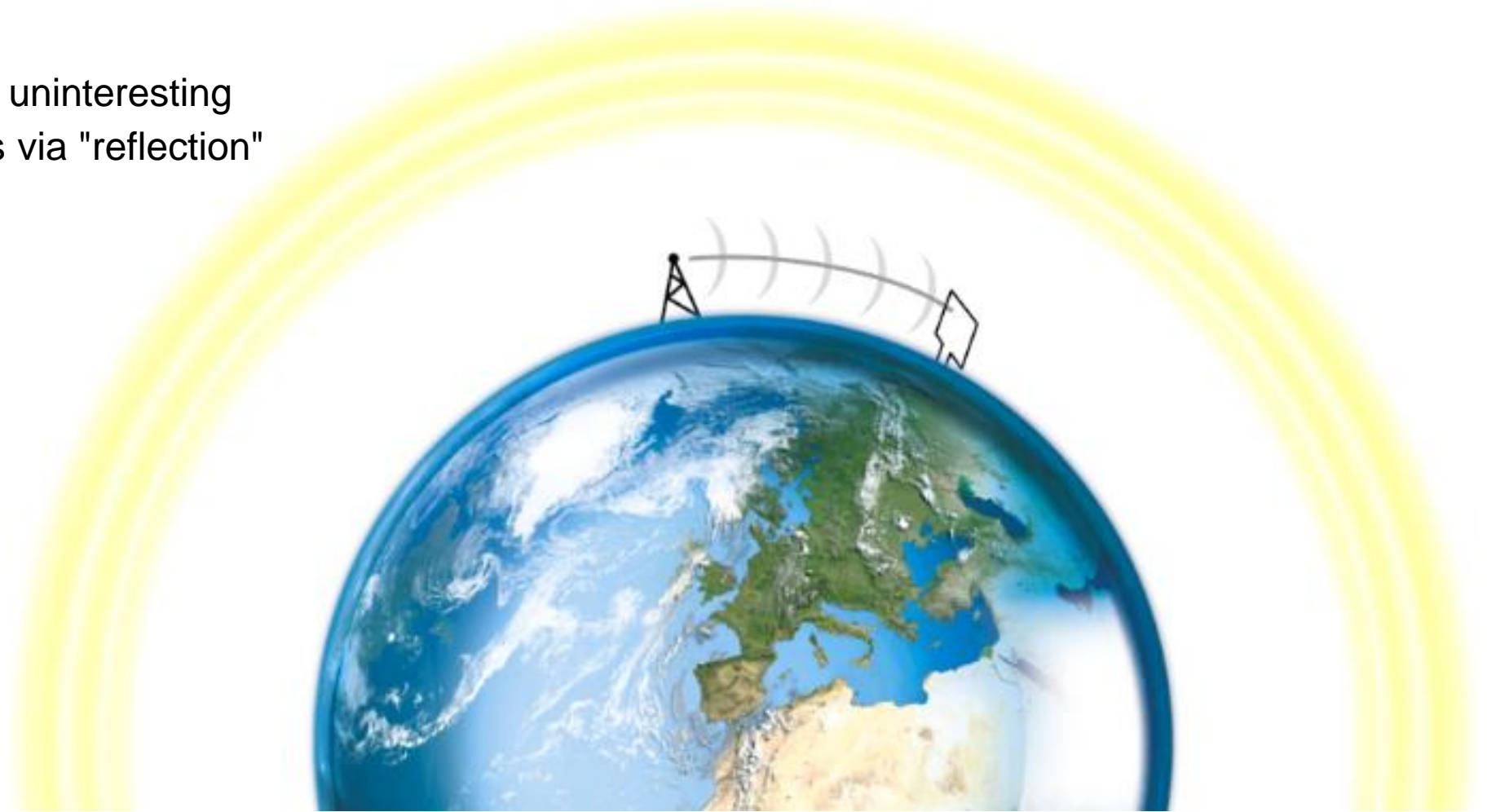
2018

GIFDS: NAA - NTZ



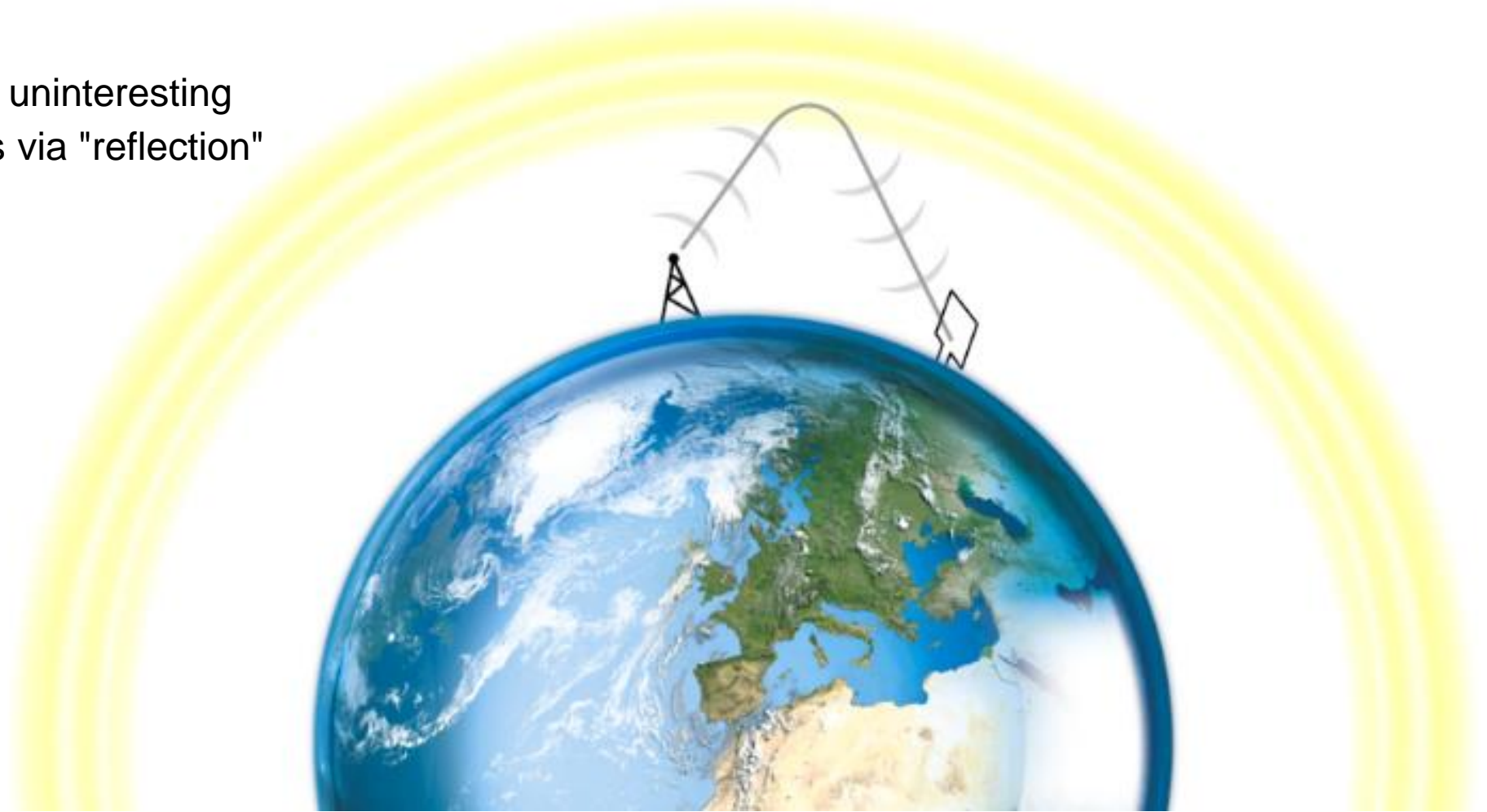
The way of the signal

- direct (or ground) waves uninteresting
- higher-order (sky) waves via "reflection"
- rays gradually refracted
- paths sum up



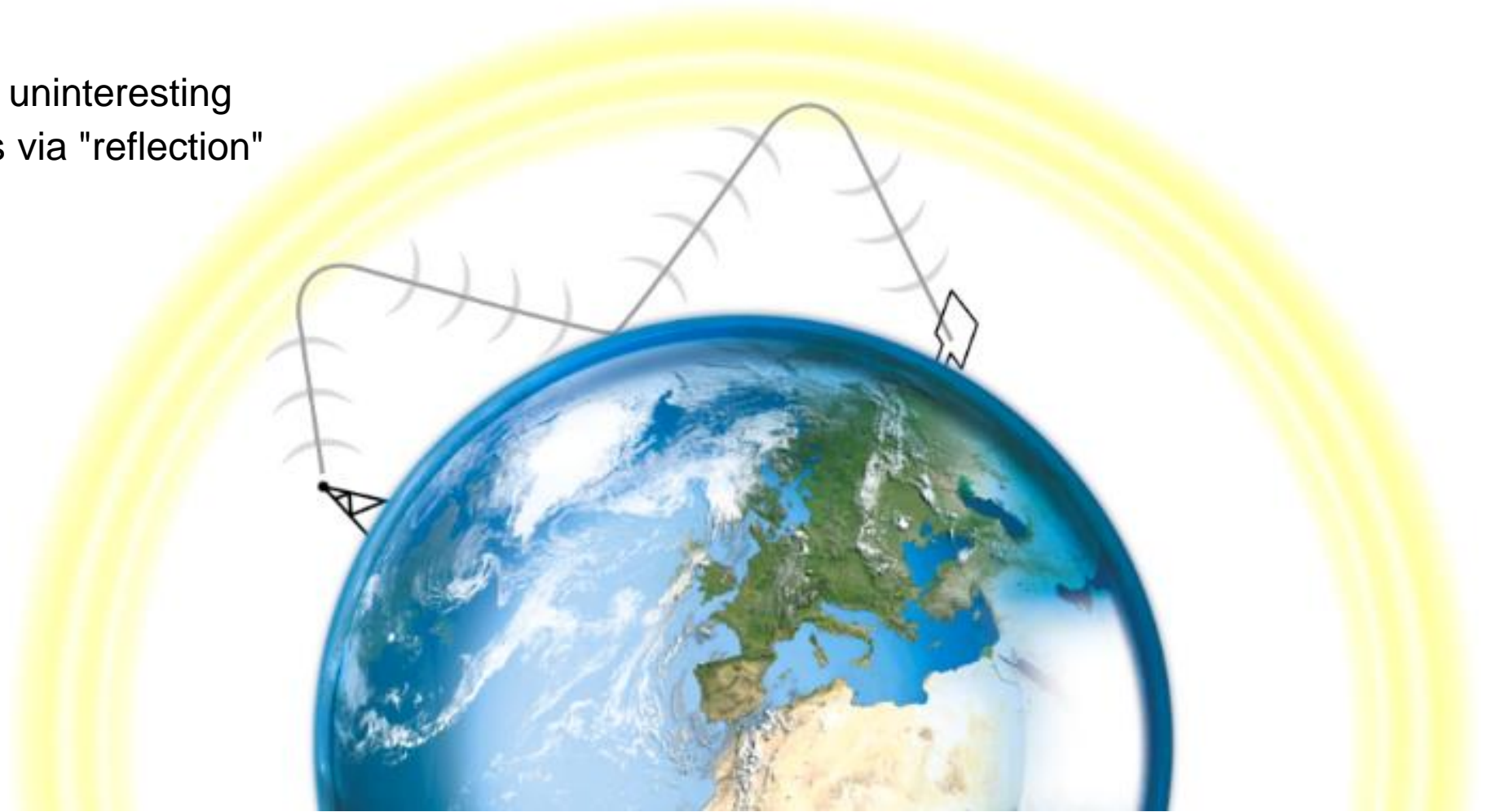
The way of the signal

- direct (or ground) waves uninteresting
- higher-order (sky) waves via "reflection"
- rays gradually refracted
- paths sum up



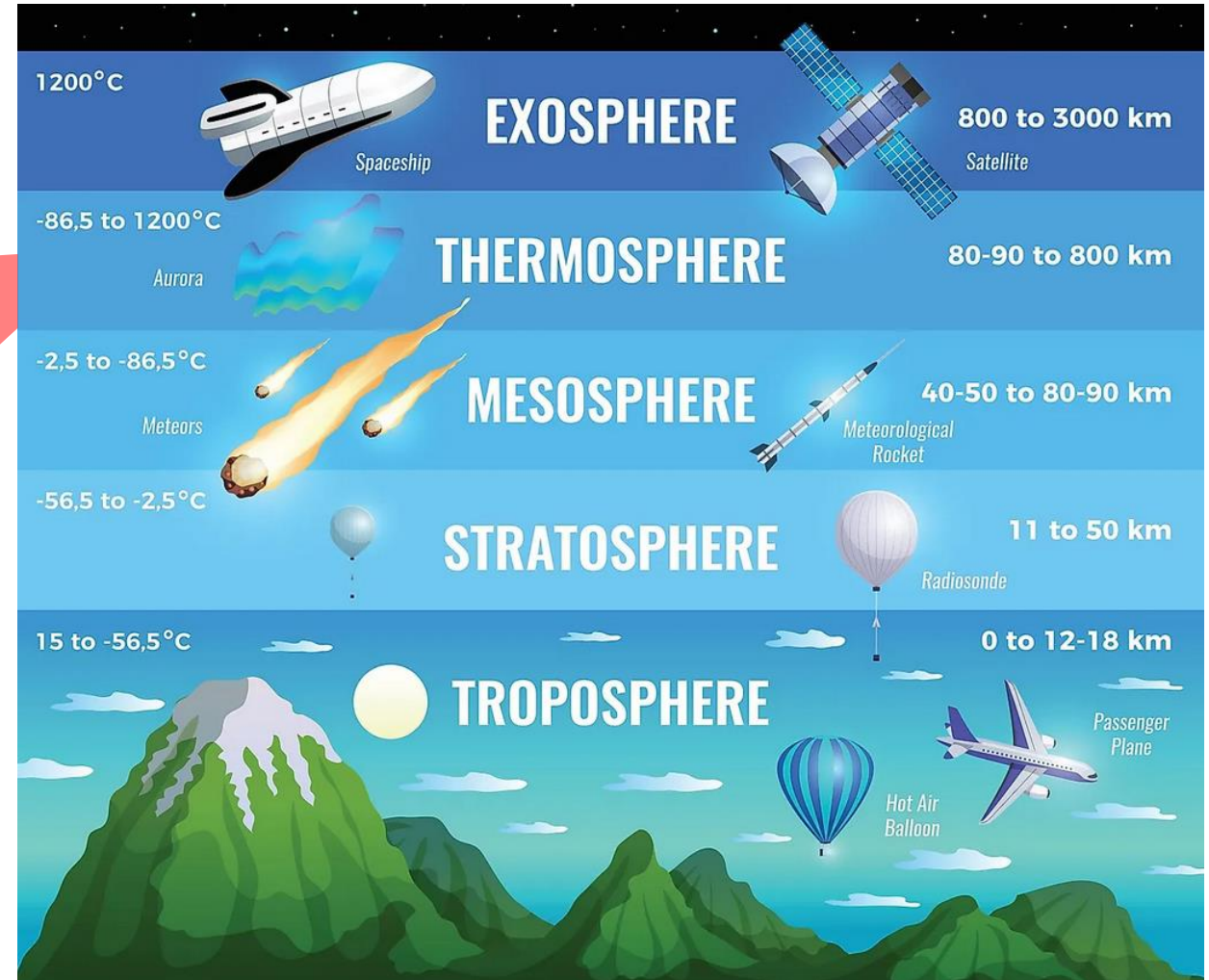
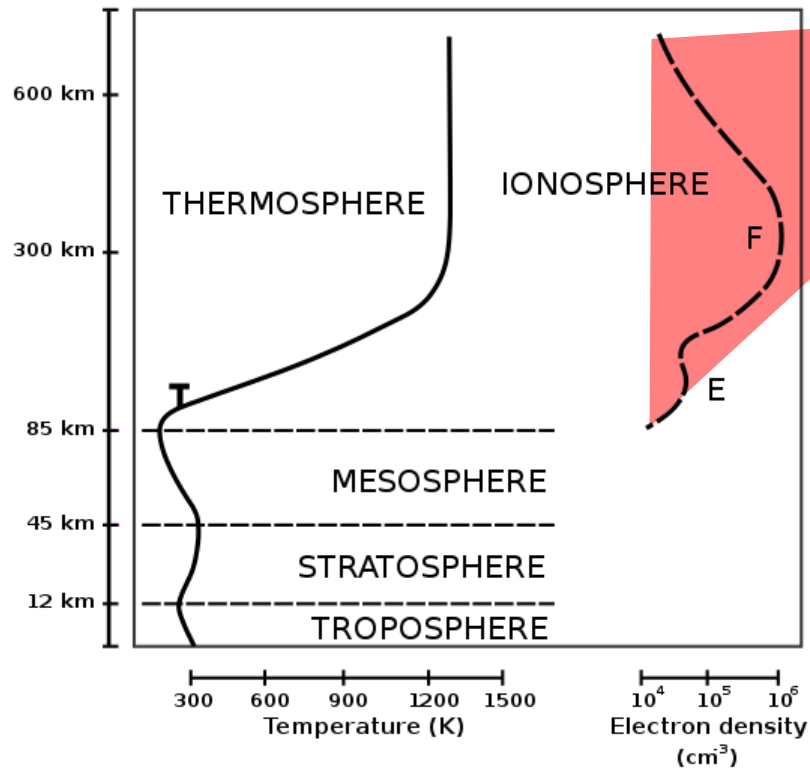
The way of the signal

- direct (or ground) waves uninteresting
- higher-order (sky) waves via "reflection"
- rays gradually refracted
- paths sum up



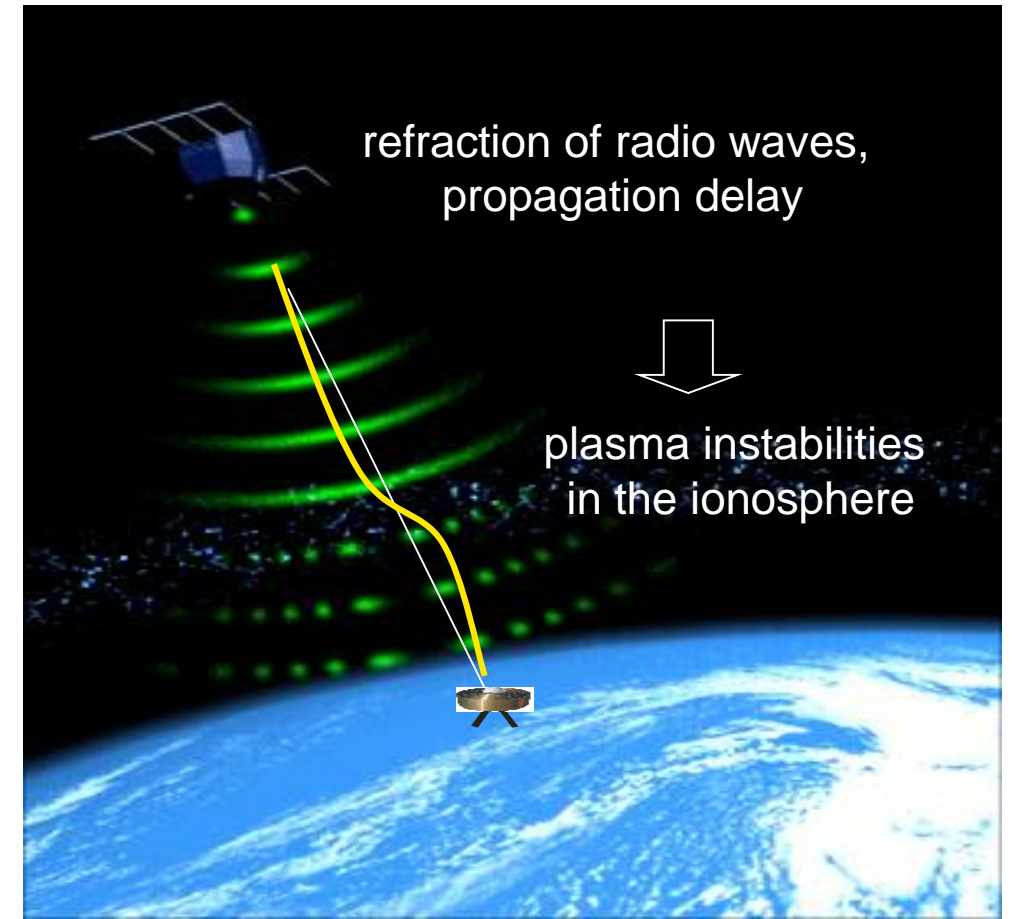
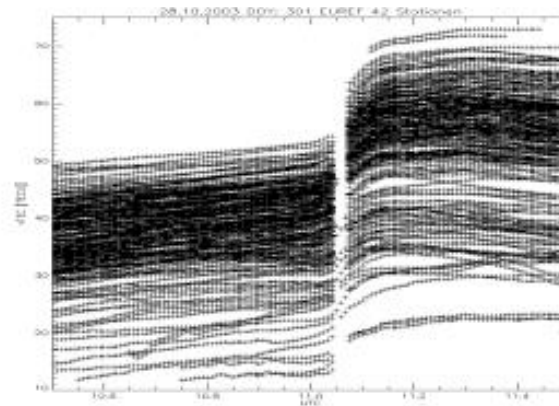
Atmospheric layers

lower thermosphere = ionosphere



A kind of distorting filter

- plasma (ionized gas) causes a delay of radio signals
→ pretending excess in distance between satellite and site
 - plasma instability causes signal strength fluctuations and defocussing of the signal
→ possible loss of the signal
- challenge for navigation, communication, Earth observation
- interplay between solar irradiation and ionosphere



Composition of the ionosphere

- via photoionisation neutral gas is split into electrons and ions (plasma)
- monitoring via GNSS → VTEC

F region (150 - 500 km)

- ionisation by EUV (20 to 90 nm)

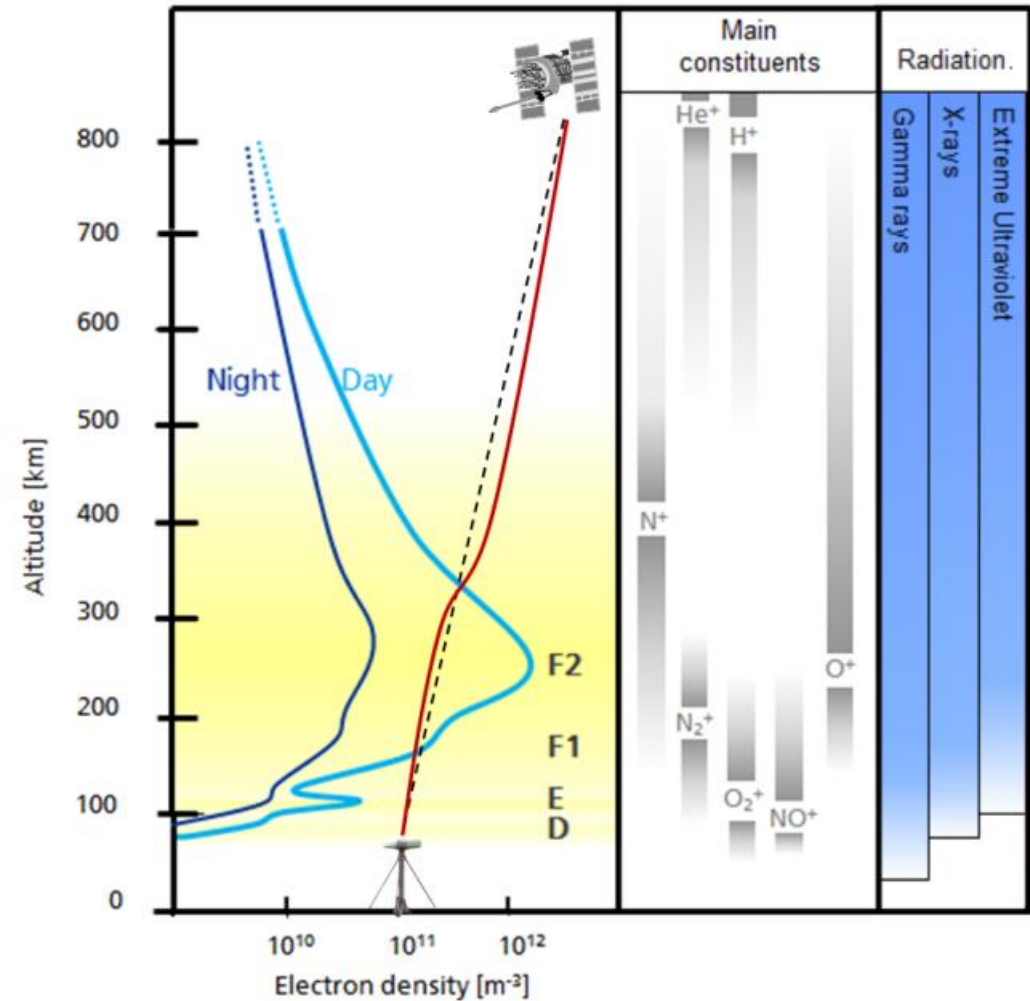
E region (90 - 150 km)

- ionisation by EUV (80 to 102.7 nm)

D region = "VLF mirror"

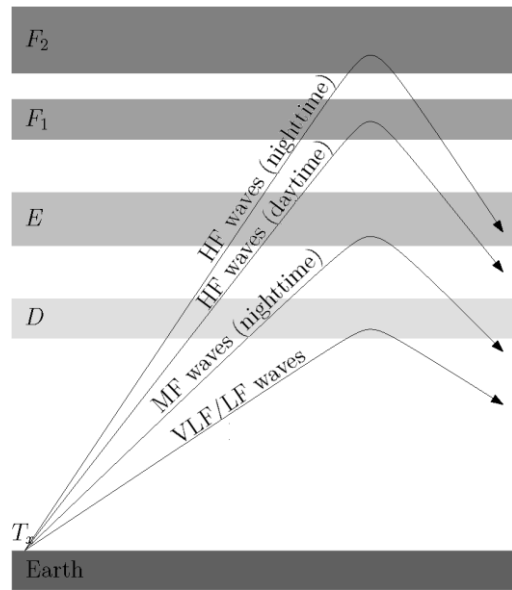
D region (60 - 90 km)

- ionisation by Lyman- α (121.6 nm) and soft X-rays (below 1 nm)

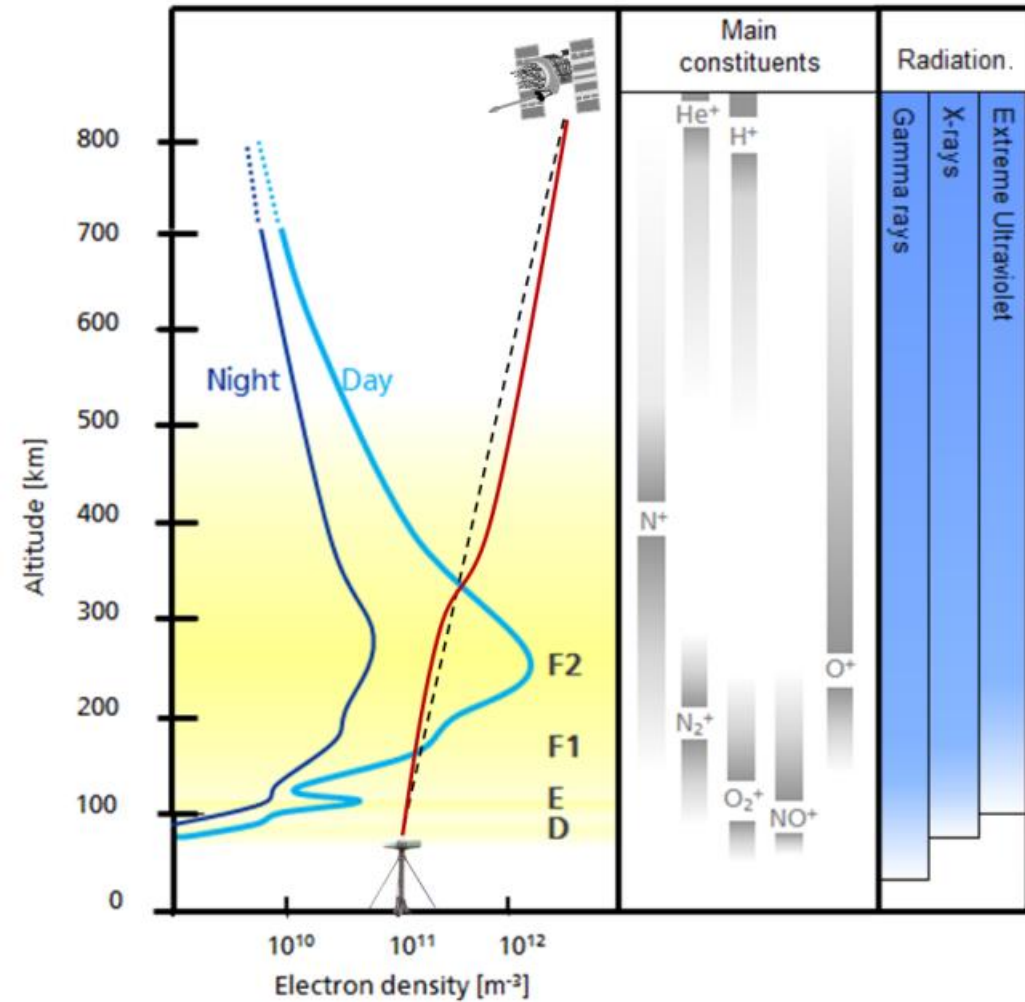


Composition of the ionosphere

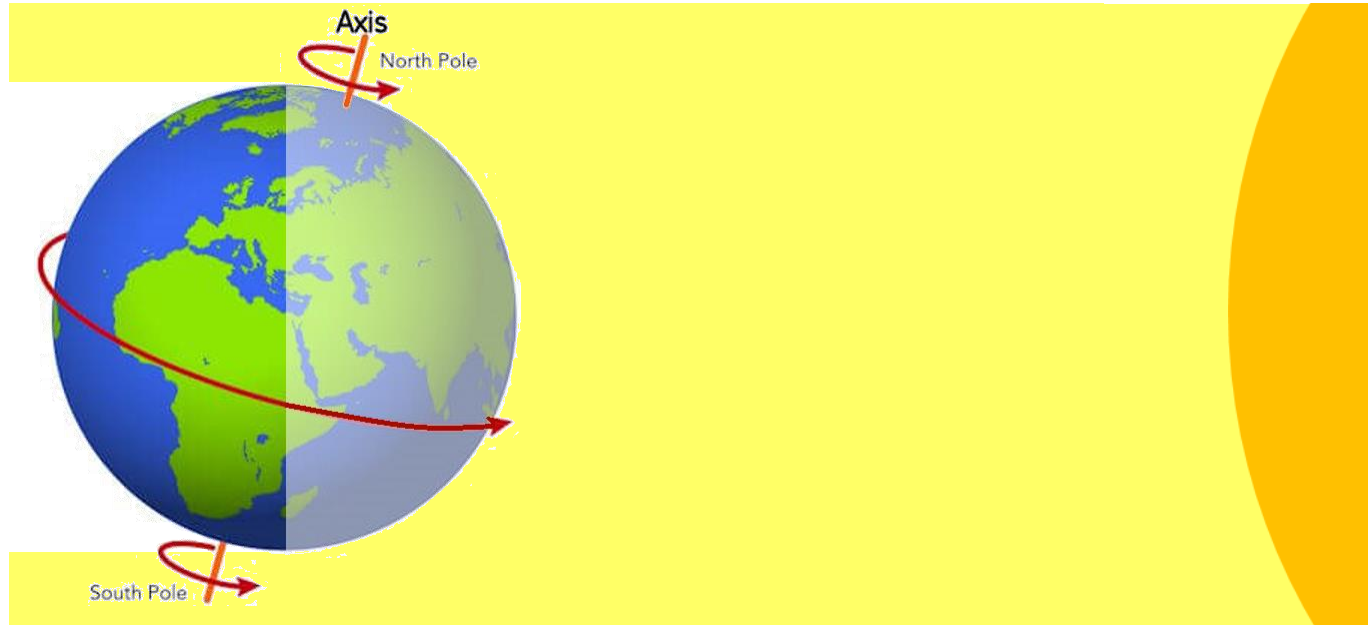
- via photoionisation neutral gas is split into electrons and ions (plasma)
- monitoring via GNSS → VTEC



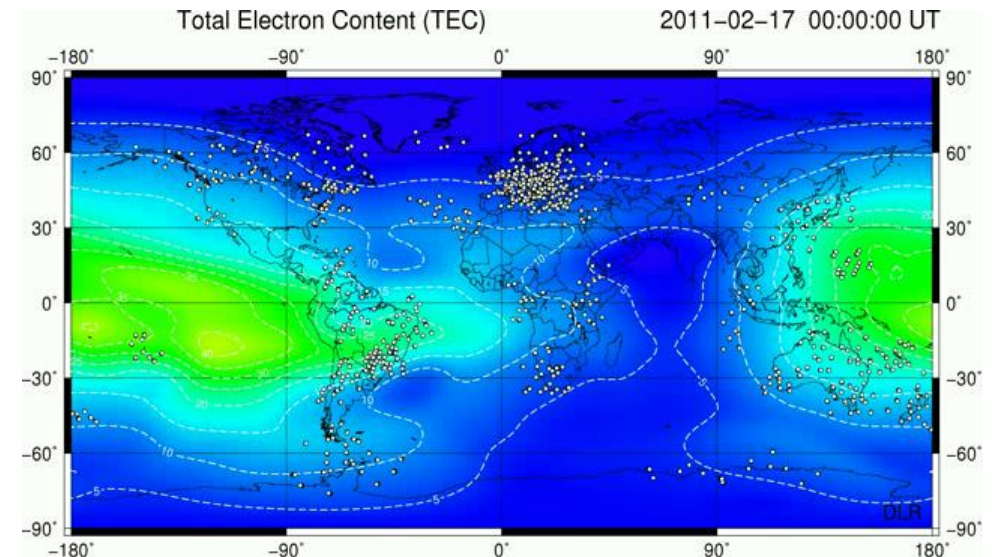
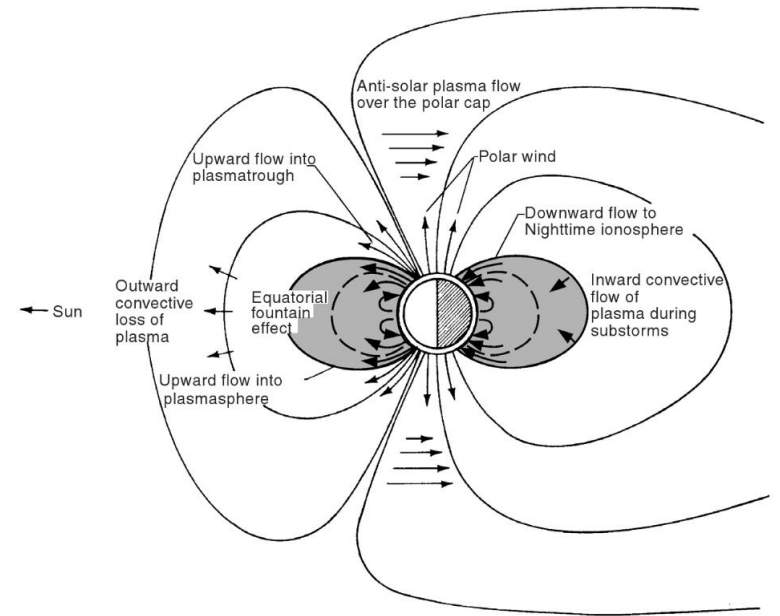
D region = "VLF mirror"



Earth's rotation



- spinning with a 24 hr period → diurnal variation
- only spotlighted parts → dayside affected
- varying striking angle → non-constant trend
- tilted axis → different day lengths

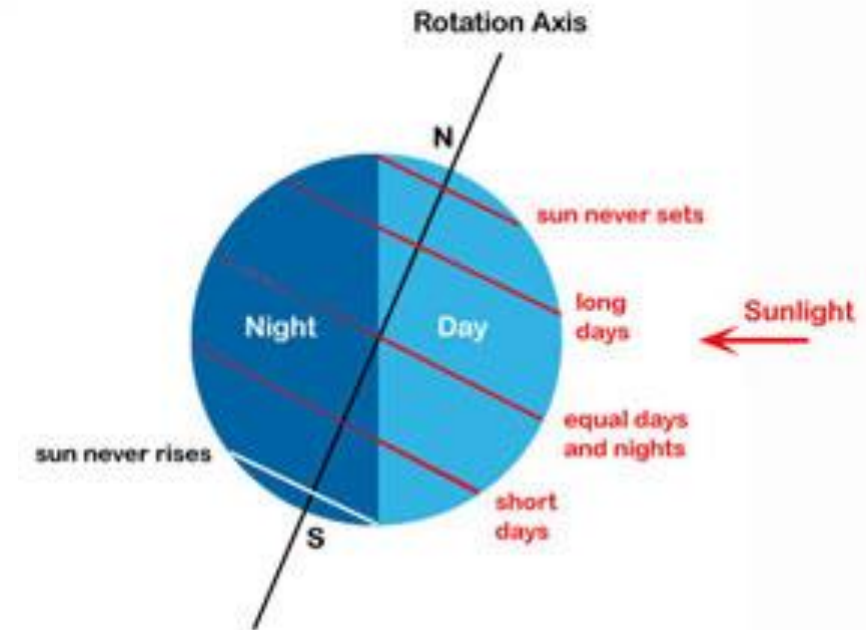
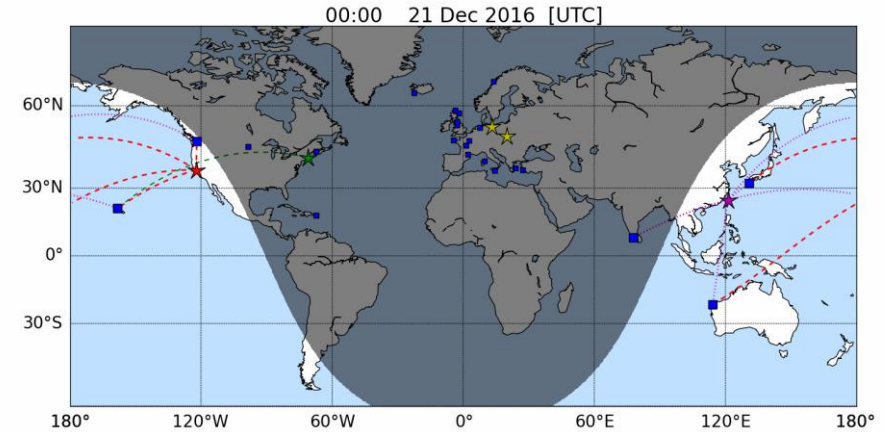


Tasks for processing measurements

- Where are your Tx and Rx?
- When is your path on the dayside?
- How to throw away the night?
→ Find a strategy.
- Where is the zenith?
→ Plot the "noon" curve.

Useful packages in Python:

- geographiclib, geopy, haversine
- ephem, astral, skyline



Disturbances



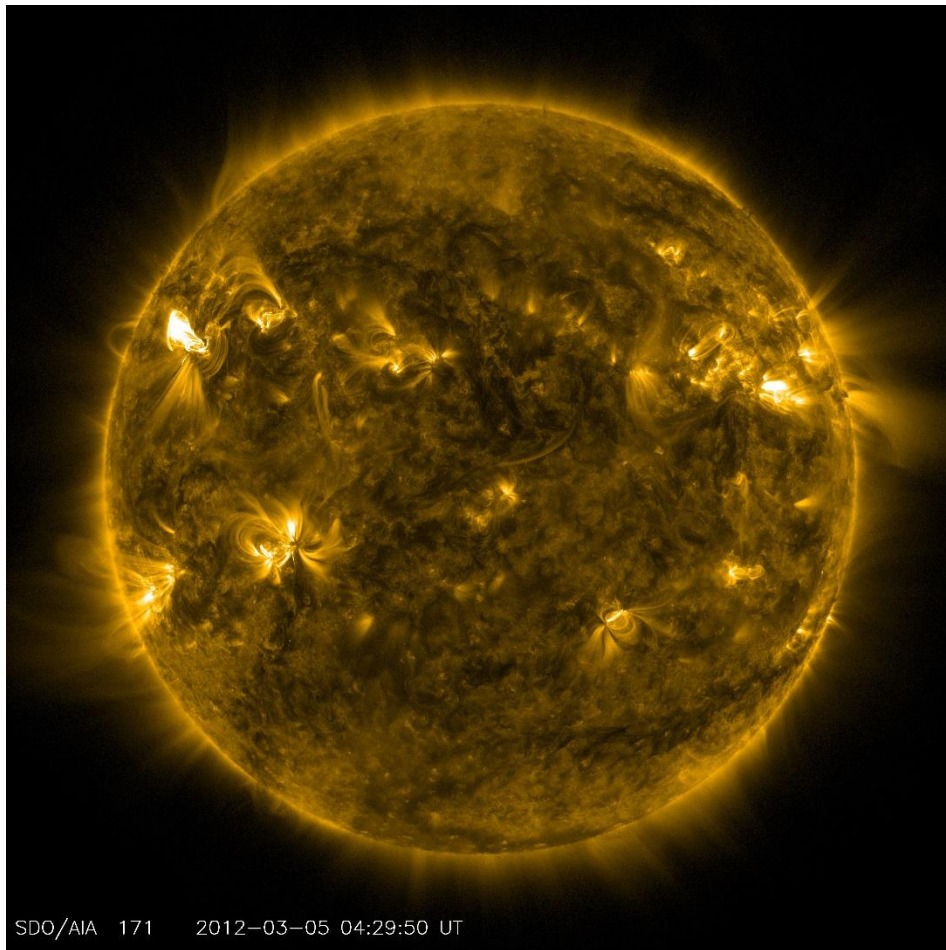
Expectation vs. reality

Daylight ionosphere is quite stable (small fluctuations observed as noise)

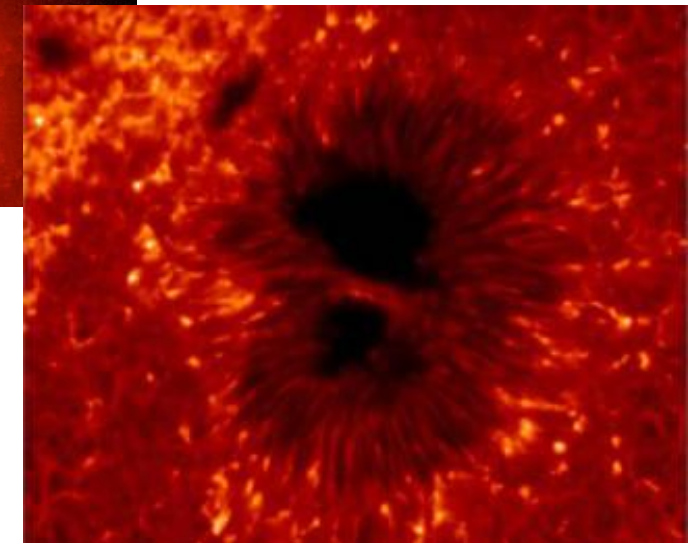
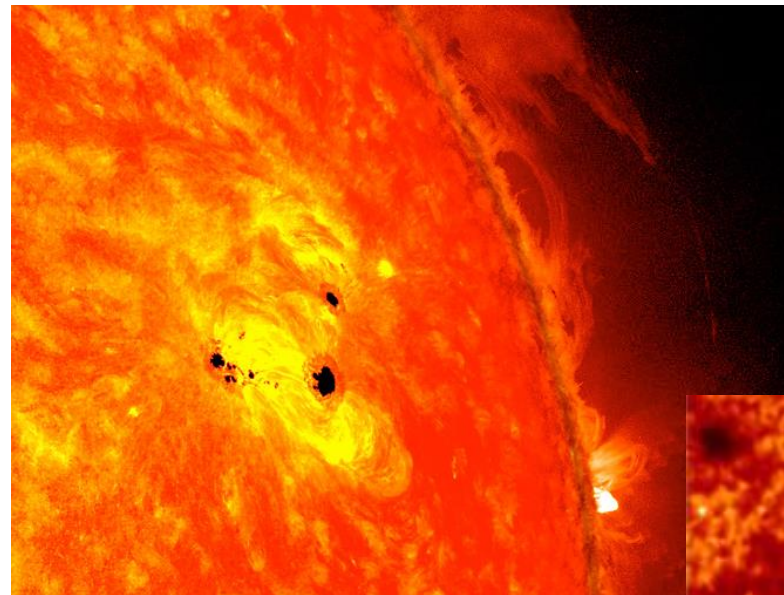
→ Impacting radiation from the Sun may vary!



The Sun



Source: NASA/SDO/AIA



Sources: NASA/SDO/AIA/HMI/Goddard Space Flight Center and NAOJ/NASA/Hinode



From the Sun to the Earth

Energetic particles
/ CMEs

Arrival: 2- 4 days

Duration: several days

High energetic particles

Arrival: minutes to hours

Duration: several days

Electromagnetic radiation

X-rays up to radio waves

Arrival: 8 min

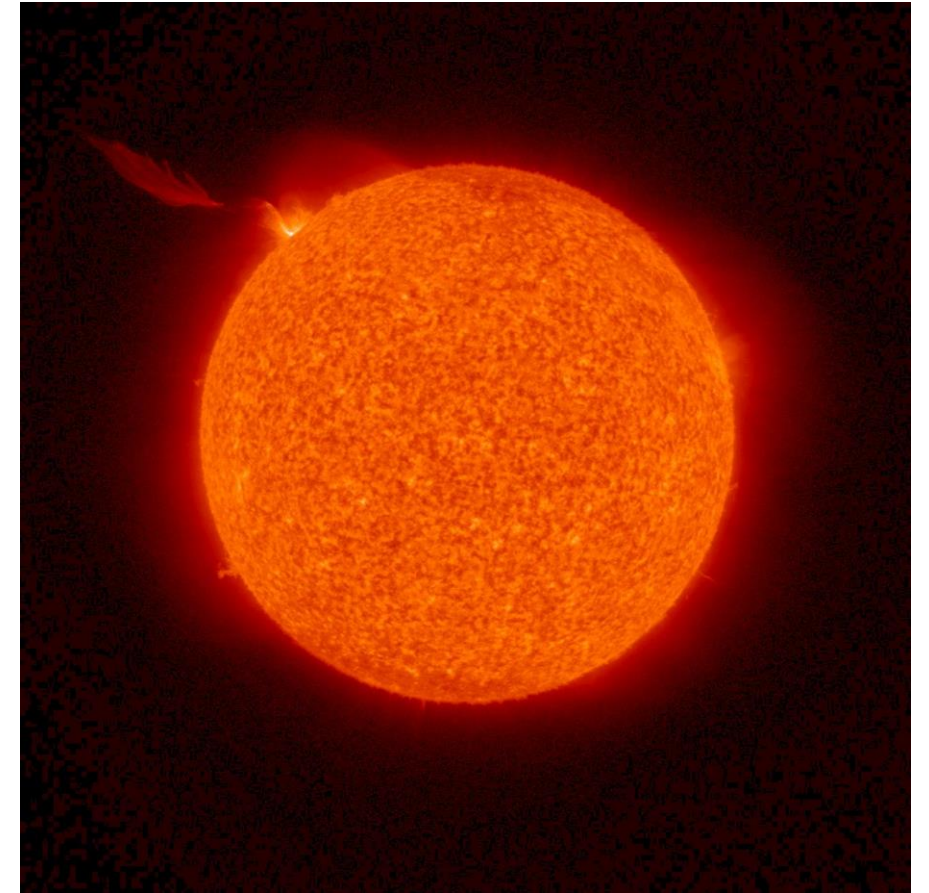
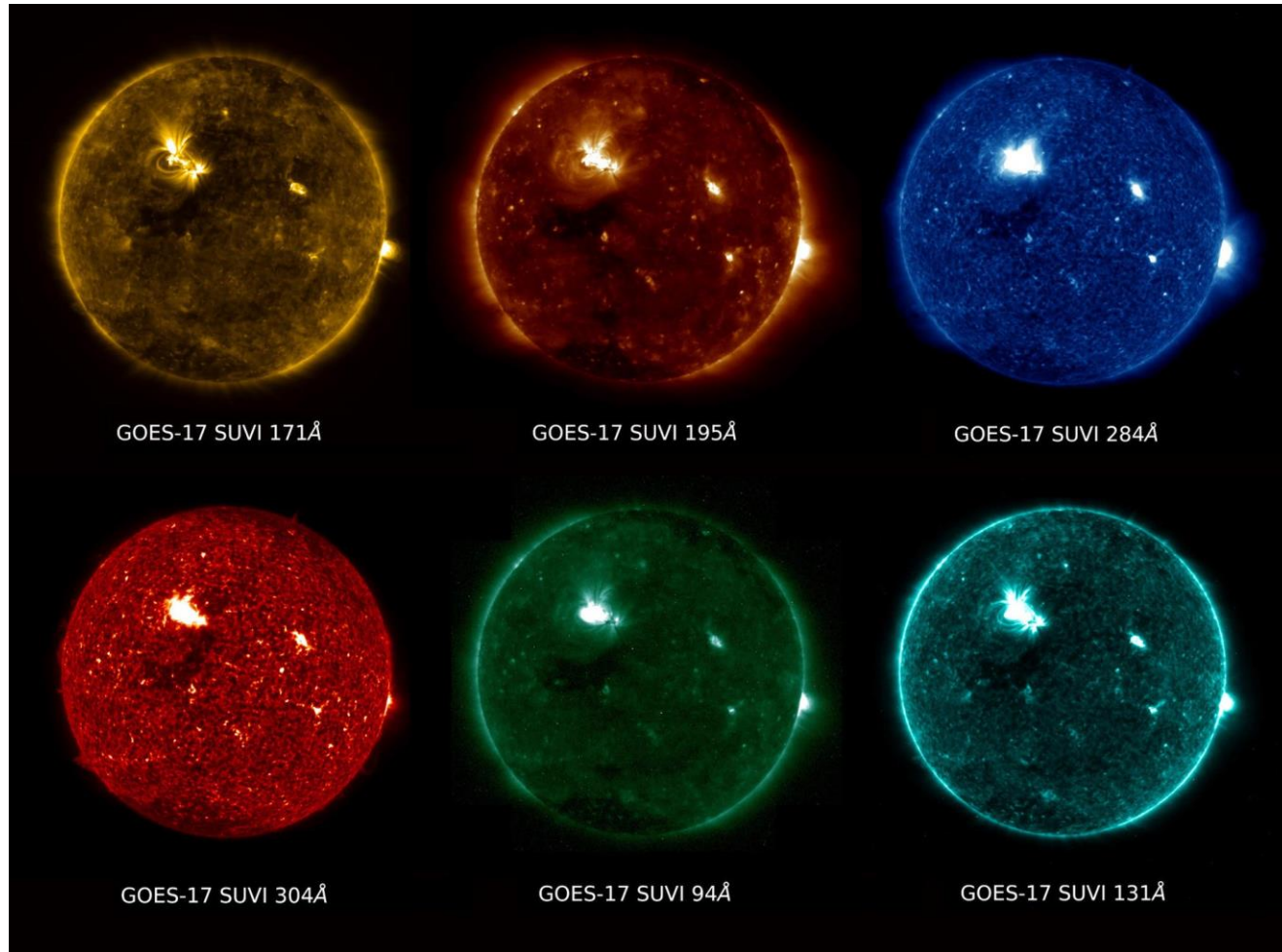
Duration: 1- 2 hours

ACE / DSCOVR

Early warning

GNSS Satellite

Solar flares

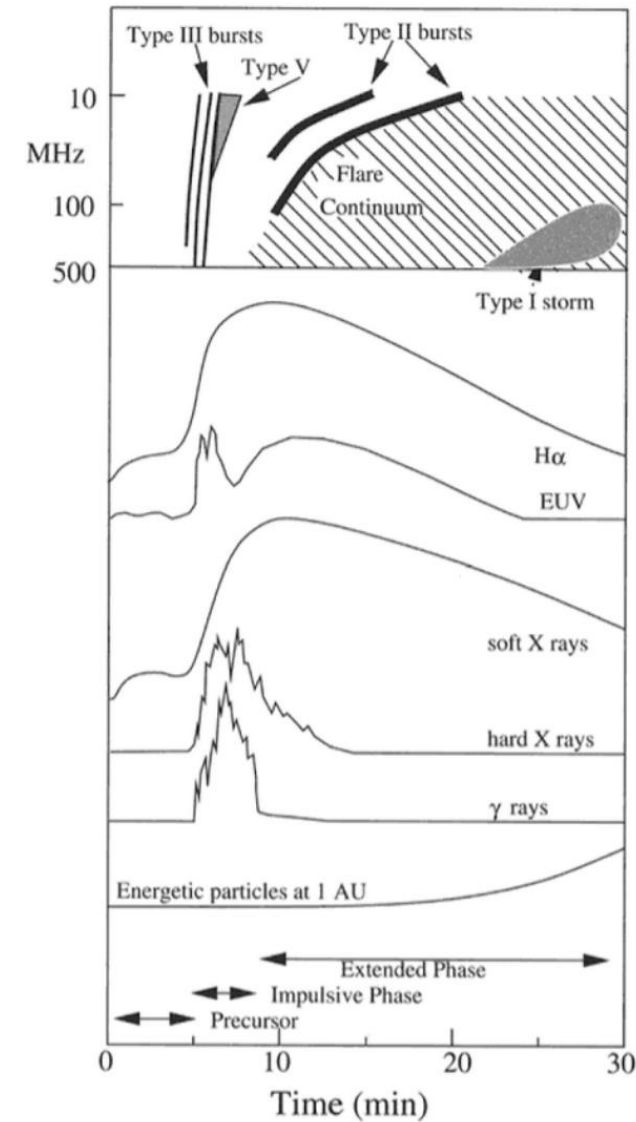
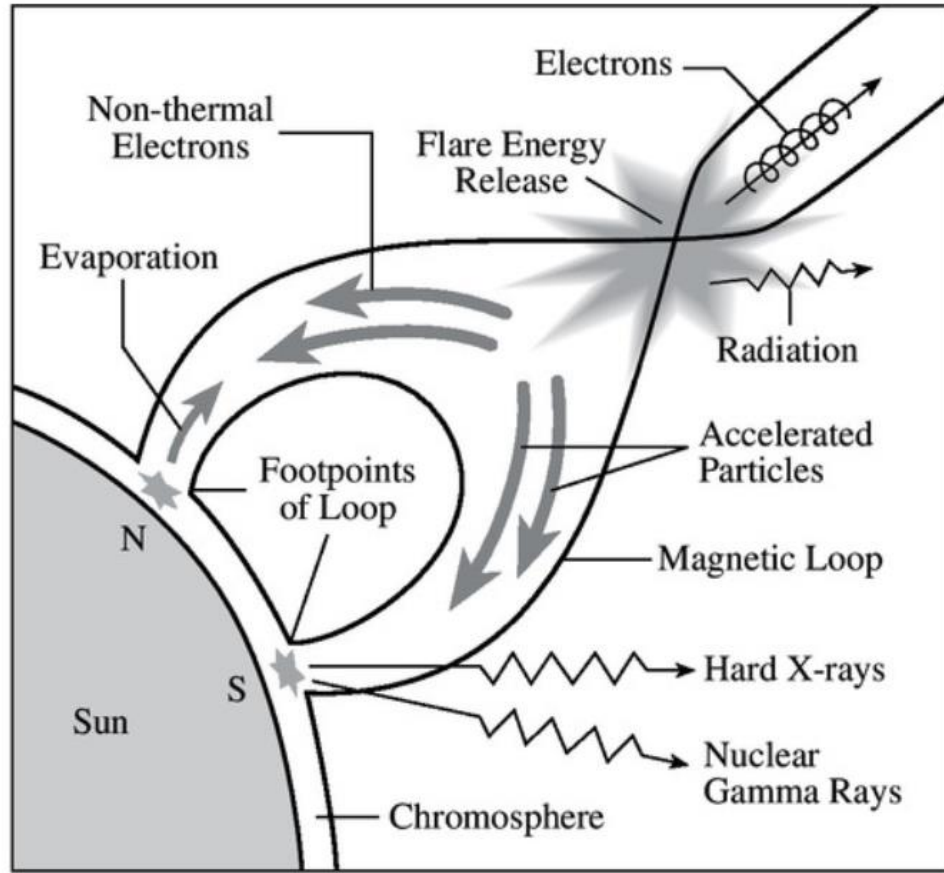


Snapshots of solar flares

Sources: GOES-16 and GOES-17



Solar flares

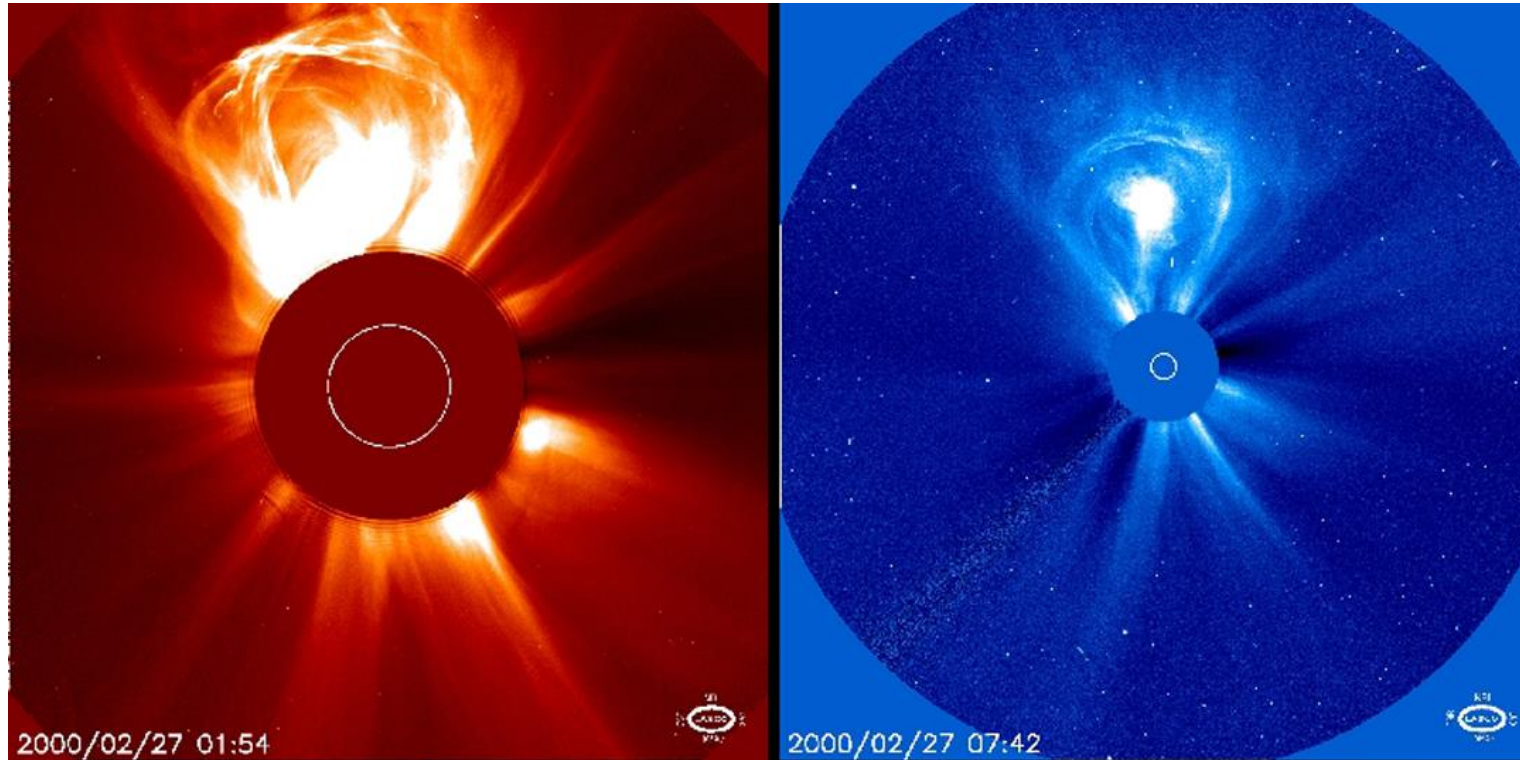


Left: Standard flare model. Right: A flare in different wavelength.

➔ X-ray flares ↔ ionosphere D layer ↔ VLF signal



Coronal Mass Ejections



CME on 200-02-27 taken by SOHO LASCO (ESA&NASA/SOHO)

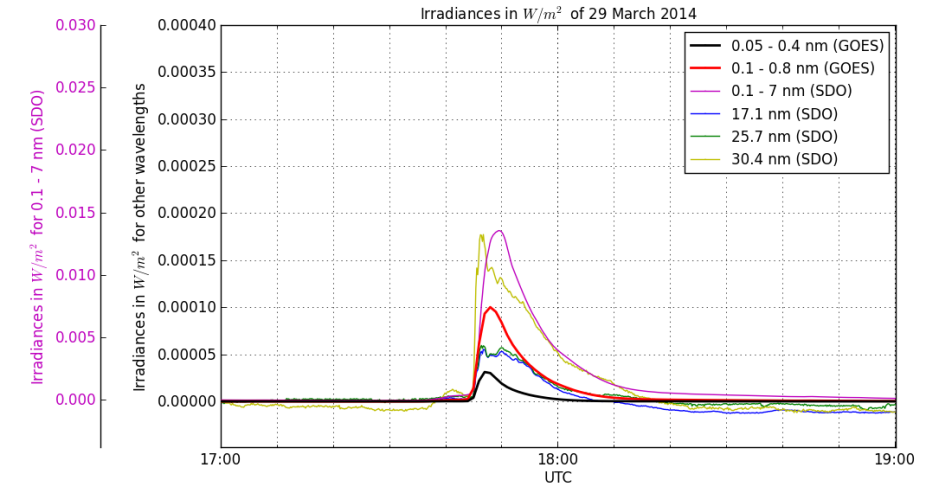
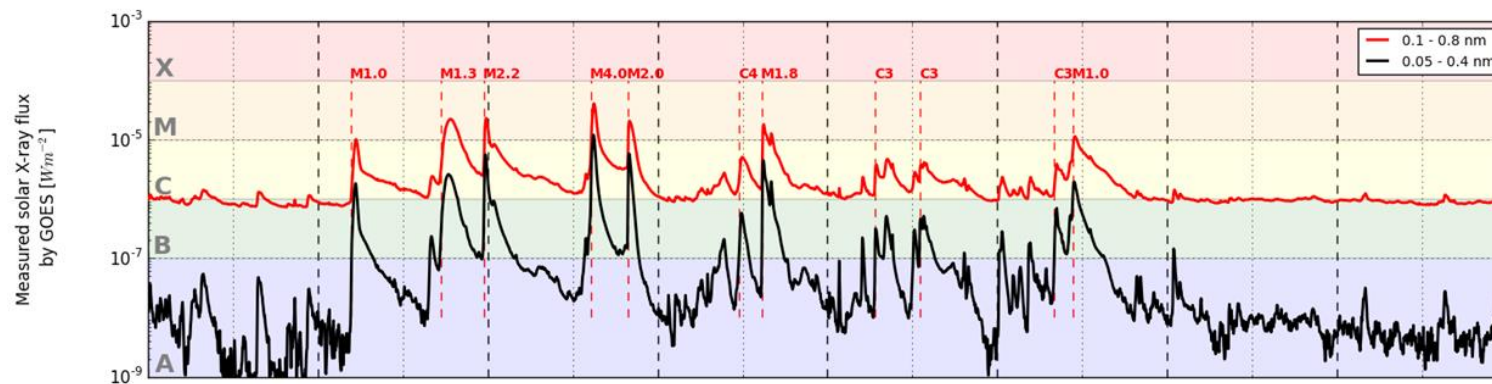


Source ...

... can be measured directly only in space

GOES – Geostationary Operational Environmental Satellite

- network for monitoring Earth from above
- also X-ray detectors before ionospheric influence
- stable basic level coming from Sun
- events like flares distributed over the spectrum



... and effect

... can be observed ground-based:

Radiation bursts at EUV and X-ray considerably increase plasma density in ionosphere
→ reaction visible in VLF measurements

What about the "normal" impact?

- no D layer in nighttime!
- non-constant progression in daytime

