

# HW9 of Plasma

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1. standard Ampere law in GS unit is:

$$\nabla \times B = \frac{4\pi}{c} J + \frac{1}{c} \partial_t E$$

where the displacement part comes from  $E = -\frac{1}{c} u \times B$  given MHD framework, so approximately the extra  $(\nabla \times B)_\Delta$  has order of:

$$(\nabla \times B)_\Delta \sim \frac{1}{c^2} \partial_t u \times B$$

When  $|u| \ll c$ ,  $\partial_t u$  is always very small compared to  $c$ , the approximation is very good. On the other hand, even  $|u| \sim c$  sometimes  $\partial_t u$  is still very small, so in relativistic MHD, the approximation is not bad sometimes.

2. (a) Invariant flux has form of:

$$\Phi_m = B \cdot 4\pi R^2$$

since  $\Phi_m = \Phi_m'$ , new magnetic field  $B'$  is:

$$B' = B_* \cdot \frac{R_*^2}{R_{ns}^2}, \text{ calculated result is } B' = 10^{12} \text{ Gauss}$$

(b) cyclotron frequency of new  $B'$  is

$$\omega = \frac{eB'}{m} = 1.759 \times 10^{19} \text{ rad/s}$$

with  $f = \omega/2\pi = 2.8 \times 10^{18} \text{ Hz}$ , it correspond to X-rays (Hard or soft)

3. first of all, the magnetic field  $B$  in cylindrical coordinate is

$$\vec{B} = \frac{\mu_0 I}{2\pi r} \hat{e}_\varphi$$

the magnetic pressure force  $\vec{f}_p = -\nabla \left( \frac{B^2}{2\mu_0} \right)$

$$\text{so } \vec{f}_p = \frac{\mu_0 I^2}{4\pi^2} \cdot \frac{1}{R^3} \hat{e}_r$$

and magnetic tension force  $\vec{f}_T = \frac{1}{\mu_0} (B \cdot \nabla) B$

$$\vec{f}_T = \frac{B^2}{\mu_0} (\hat{e}_B \cdot \nabla) \hat{e}_B = -\frac{\mu_0 I^2}{4\pi^2 R^2} \cdot \frac{1}{R} \hat{e}_r ; (\hat{e}_\varphi \cdot \nabla) \hat{e}_\varphi = -\frac{\hat{e}_r}{R}$$

two forces have same strength, and the net force is zero.

4. outside the surface of earth, magnetic force has form of

$$B = \frac{B_0 R_m^3}{r^3}, \text{ and its pressure is:}$$

$$P_B = \frac{1}{2\mu_0} B^2 = \frac{B_0^2 R_m^6}{2\mu_0} \cdot \frac{1}{r^6}$$

on the other hand, pressure force from solar wind is

$$P_s = \rho u^2, \text{ when } P_B \geq P_s, \text{ the magnetic field could}$$

stop solar wind, so

$$\frac{B_0^2 R_m^6}{2\mu_0} \cdot \frac{1}{r^6} = \rho \cdot u^2$$

$$r = 4.578 \times 10^9 \text{ cm} \quad \sim 7.15 R_m$$