



Search For
Delayed
Photons
Using
Timing.

Tambe E.
Norbert

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and Decay

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and
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Summary

Search For Delayed Photons Using Timing.

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Roger Rusack¹

¹University Of Minnesota

**Long-Lived Meeting,
November 29, 2014**



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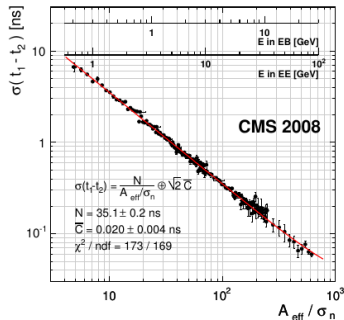
Summary

• Long-Lived Particle Models

- ★ Gauge Mediated Supersymmetry Breaking (GMSB)
 - ▷ Next-to-lightest SUSY (NLSP) is **Neutralino** ($\tilde{\chi}_1^0$)
 - ▷ $eV - keV$ Lightest-SUSY particle (LSP) is **Gravitino** (\tilde{G}).
 - ▷ Gravitino is a Dark Matter Candidate.
- ★ General Gauge Mediation (GGM)
 - ▷ NLSP is a mixture of fermions (Bino, Wino, Higgsino).
 - ▷ Several SUSY particles can be NLSP.

• ECAL Resolution

- † ECAL timing resolution
 $\sigma_t < 500$ ps.
- † Use timing to identify
photons and electrons from
long-lived decay.



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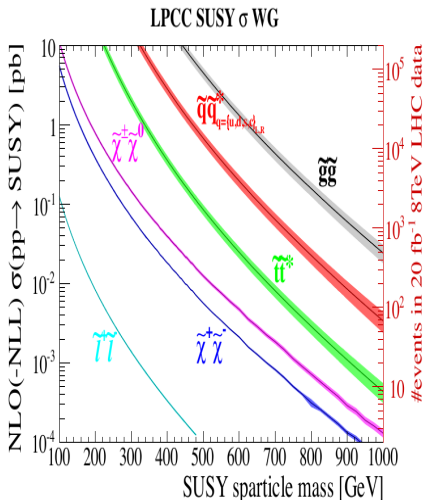
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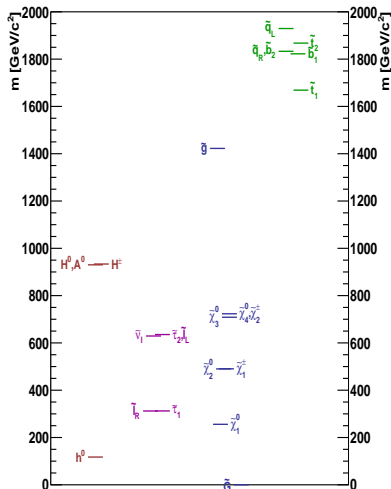
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<https://twiki.cern.ch/twiki/bin/view/LHCPhysics/SUSYCrossSections>

arXiv:1206.2892



SUSY production mostly in strong interactions at LHC.

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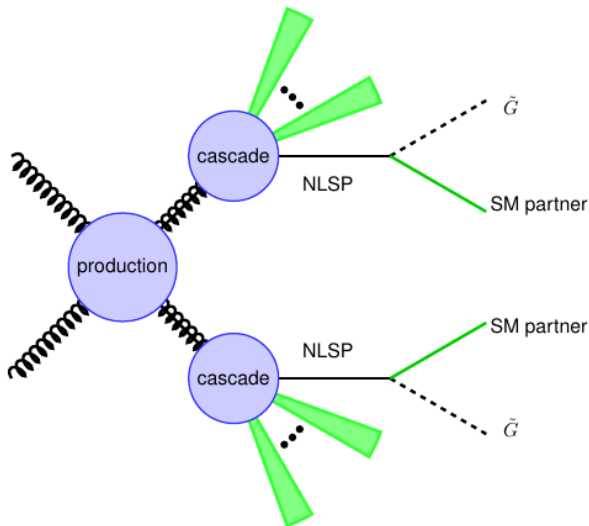
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Y. Kats et al: [arXiv:1110.6444v2](https://arxiv.org/abs/1110.6444v2)

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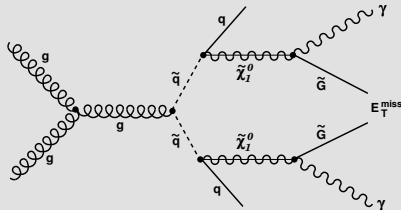
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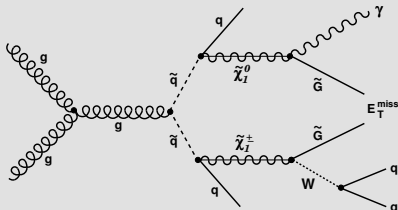
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Double Photon



2 Photons, 2 Jets, Large MET

Single Photon



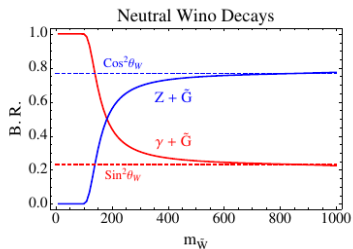
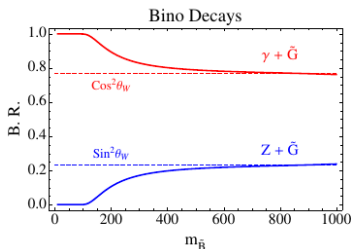
1 Photon, Jets, Large MET

Distance Travelled

$$L_T = c\tau \cdot (\gamma\beta_T) = c\tau \cdot \left(\frac{p_T}{m}\right)$$

Proper Decay Length

$$c\tau_{\text{NLSP}} = C_{\text{grav}}^2 \frac{1}{\kappa} \left(\frac{m_{\text{NLSP}}}{\text{GeV}}\right)^{-5} \left(\frac{\sqrt{F}}{\text{TeV}}\right)^4$$





Datasets



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• Data ($19.1 fb^{-1}$)

Dataset Name	Recorded Luminosity [fb^{-1}]
/Run2012B/SinglePhoton/EX0DisplacedPhoton-PromptSkim-v3	5.1
/Run2012C/SinglePhoton/EX0DisplacedPhoton-PromptSkim-v3	6.9
/Run2012D/SinglePhoton/EX0DisplacedPhoton-PromptSkim-v3	7.1
/Run2012C/Cosmics/Run2012C-22Jan2013-v1/RECO	3130384(events)
/Run2012D/Cosmics/Run2012C-22Jan2013-v1/RECO	52430 (events)
/SingleElectron/Run2012A-22Jan2013-v1/AOD	5.2
/DoubleElectron/Run2012C-22Jan2013-v1/AOD	4.8

• Signal MC [GMSB (SPS8)]

Λ [TeV]	100	120	140	160	180	300
$M_{\tilde{\chi}_1^0}$ [GeV/c^2]	140	169	198	227	256	430
$c\tau$	215	325	130	245	185	
(mm)	425	645	515	490	365	495
	1700	1290	1030	975	730	
	3400	1935	2060	1945	1100	995
	5100	2955	2920	2930	2195	2960
	6000	3870	3985	3910	3950	
	9300	5985	6000	5875	5980	6000
		9825	10450	9815	10450	10450

• $\gamma +$ Jets MC

\hat{p}_T [GeV / c]	σ_{LO} (pb)	Number of events
50 – 80	3322.3	1995062
80 – 120	558.3	1992627
120 – 170	108.0	2000043
170 – 300	30.1	2000069
300 – 470	2.1	2000130
470 – 800	0.212	1975231

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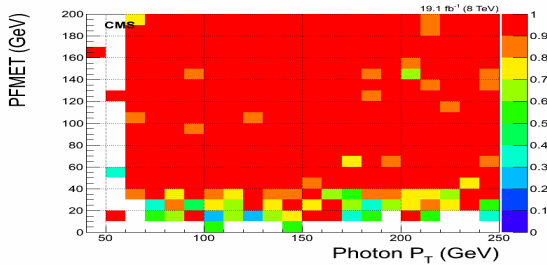
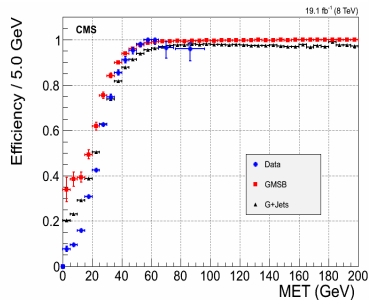
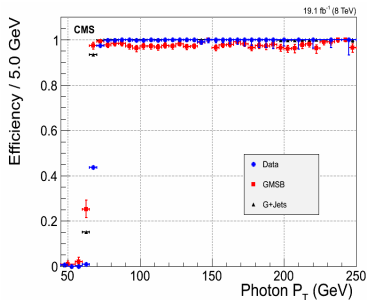
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- **HLT_DisplacedPhoton65_CaloldVL_IsoL_PFMET25**
 - HLT_Photon50_CaloldVL_IsoL (Study Trigger)



• Time Reconstruction

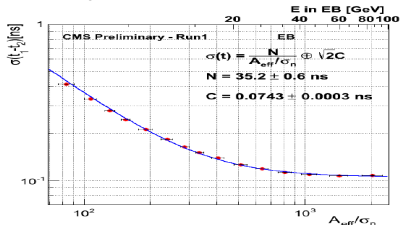
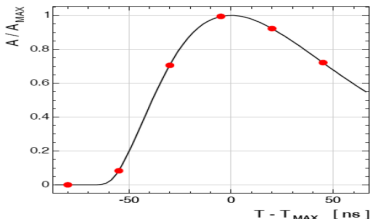
- 10 digitized samples used.
- Fit and Weighted methods used to extract time.

• Time Measurement

$$T_{MAX} = \frac{\sum_i \frac{T_{MAX,i}}{\sigma_i^2}}{\sum_i \frac{1}{\sigma_i^2}}$$

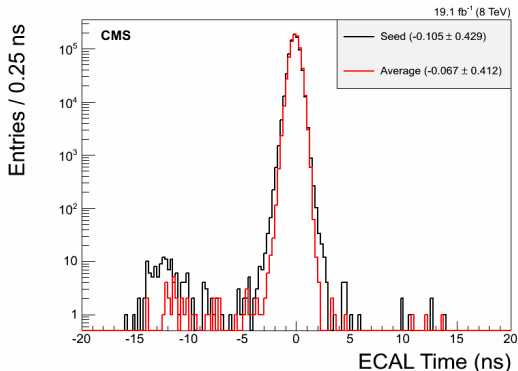
• Time Performance

- Time resolution better than 200 ps for $E > 30$ GeV



• Photon Timing

- T_γ = Average Time of all Crystals.
- T_γ = Seed (most energetic) Crystal Time.



- Similar behavior seen in Seed and Average Time.
- We use seed time as Photon Measured Time in this analysis.

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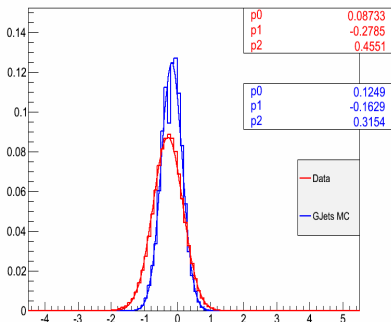
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Ecal Time from Seed Crystal



Ecal Time from Seed Crystal

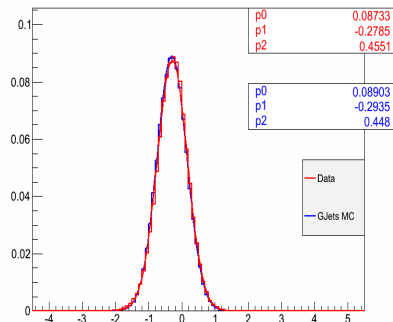
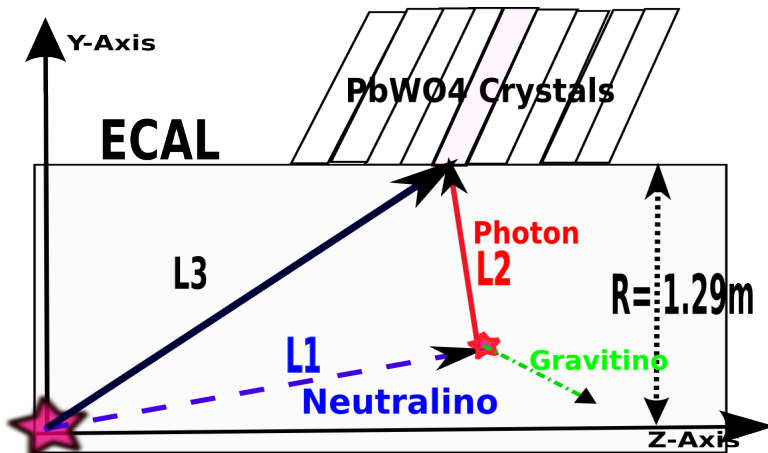


Figure : (LEFT): Before (RIGHT): After

- Timing corrections from data applied to γ + Jets MC.
- γ + Jets MC timing aligns better with data after corrections are applied.

• Source of Delayed Photon?

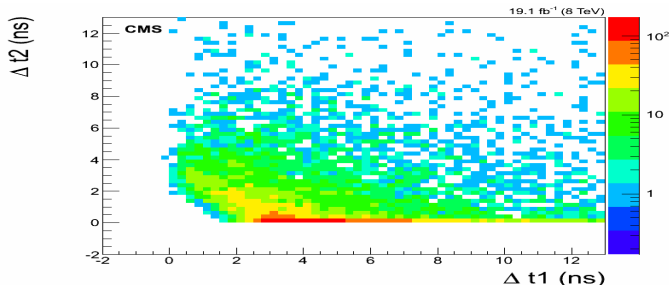
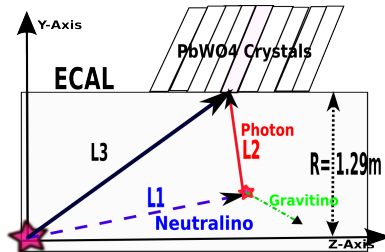
- Slow moving particle; $\beta \ll 1$,
- Non-nominal flight path,
- Stopped in subdetectors,



Photon Arrival Time

$$\Delta t_1 = (L1/c\beta) - (L1/c)$$

$$\Delta t_2 = (L1 + L2 - L3)/c$$



Delayed photons mostly from slow moving neutralino decays.



Event Selection



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Object Selection Criteria

Variable	Selection Cuts
Photon $p_T(\gamma^{1(2)})$	$> 80(45) \text{ GeV}$
$ \eta_\gamma , (\text{EB only}),$	$< 3.0(< 1.5)$
Semi-minor axis(S_{Minor})	$0.12 \leq S_{Minor} \leq 0.38$
H/E	< 0.05
Track Veto, $\Delta R(\gamma, track)$	> 0.6
HCAL, ECAL, Track, Isolation	$< 4.0, < 4.5, < 0.2$
Cone Size(Iso γ) $\Delta R(\gamma, SC)$	< 0.4
Spike Swiss-Cross	$1 - E_4/E_1 < 0.98$
Jets must satisfy	JetID Requirements
Leading Jet p_T	$> 35 \text{ GeV}$
Number Of Constituents	> 1
$\Delta R(\gamma, jet) = \sqrt{(\phi_\gamma - \phi_{jet})^2 + (\eta_\gamma - \eta_{jet})^2}$	> 0.3
E_T^{miss}	$> 25 \text{ GeV}$

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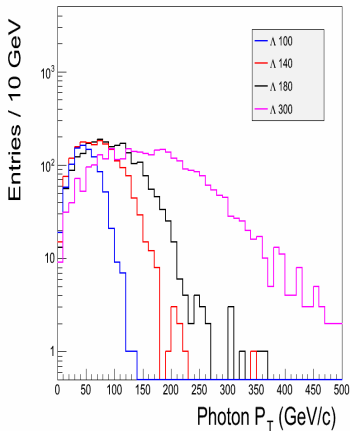


Figure : Photon p_T

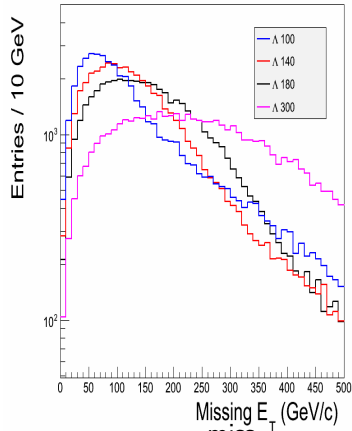


Figure : E_T^{miss}

- Different Λ values with the same $c\tau(10 \text{ m})$. Photon p_T is harder with higher values of Λ .

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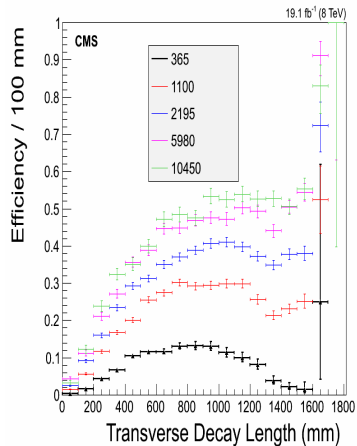
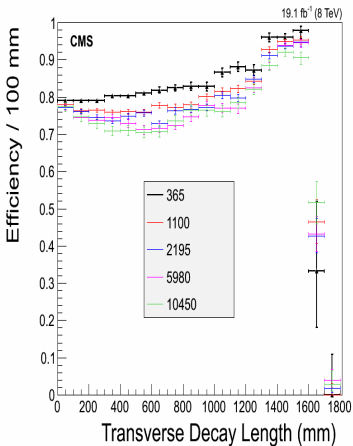
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Sharp drop in efficiency immediately beyond ECAL radius for slow moving neutralino decay as source of delayed photon.

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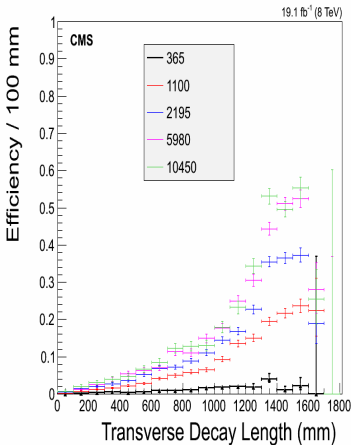


Figure : Slow Moving

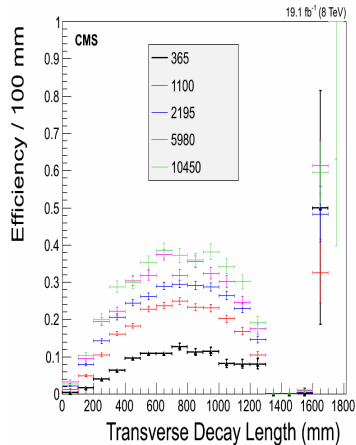


Figure : Off-Pointing

Acceptance peaks at transverse decay length 800 mm with delayed photons from off-pointing neutralino decays.

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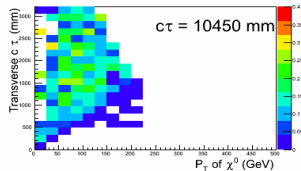
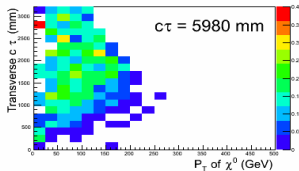
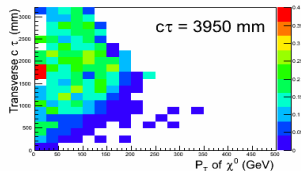
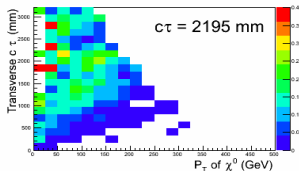
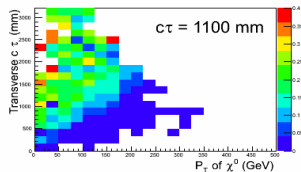
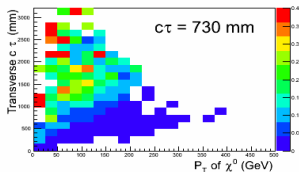


Figure : 2 Dim Efficiency



Background Estimation



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Background estimation is Data driven. Thus, most of a systematics come from signal, including:

Experimental Systematics

- Definition of Absolute or Zero time,
- ECAL time Resolution,
- Unclustered Energy,
- Jet energy scale,
- Jets energy resolution,
- Photon energy scale,
- Luminosity. We use standard CMS luminosity uncertainty.

Theoretical Systematics

- Choice of PDF.
- Re-normalization group equations.

Systematic Uncertainties

Source	Uncertainty(%)
Absolute time(Zero time)	10 ~ 6
Unclustered Energy	10 ~ 4
Photon Energy Scale	4 ~ 2
ECAL Time Resolution	5 ~ 2
Jet Energy Scale	9 ~ 3
Jet Energy Resolution	9 ~ 2
Luminosity	2.6
Choice of PDF	< 1

- Systematics is obtained by studying the effects of varying by a few amount of a particular source of systematic on the total number of objects passing object selection cuts.



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Events Passing Final Selection

Sample	Lifetime($c\tau$)[mm]	Number Of Events
GMSB $\Lambda = 180$ TeV	10500	
GMSB $\Lambda = 180$ TeV	6000	
GMSB $\Lambda = 180$ TeV	4000	
GMSB $\Lambda = 180$ TeV	3000	
GMSB $\Lambda = 180$ TeV	2000	
GMSB $\Lambda = 180$ TeV	1000	
GMSB $\Lambda = 180$ TeV	500	
Data	1.00	
Background Total	0.014	

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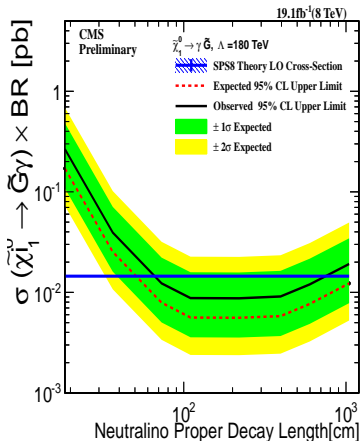


Figure : $c\tau$ Limits

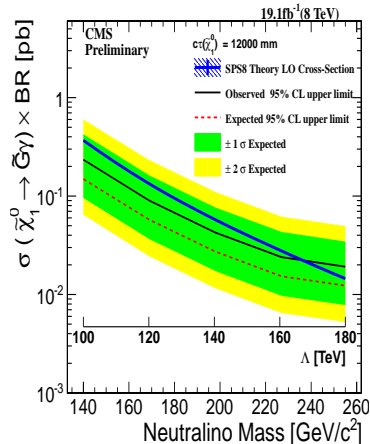


Figure : Mass Limit

sample is $c\tau = 12000$ mm but we measure $c\tau \approx 10500$ mm

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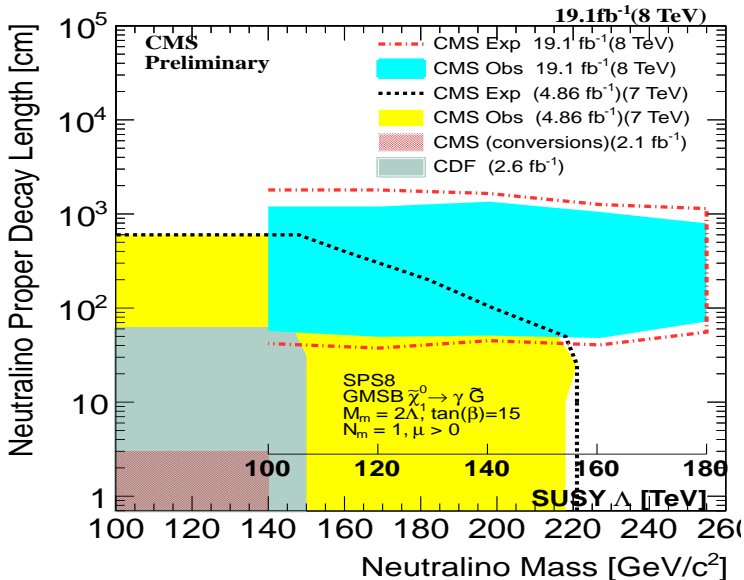
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