

*Displaced Photon Background  
Studies Update  
LL Meeting Nov 02 2012*

*Tambe E. Norbert*, *Shin-chuan  
Kao, Yuichi Kubota, Giovanni  
Franzoni, Roger Rusack*

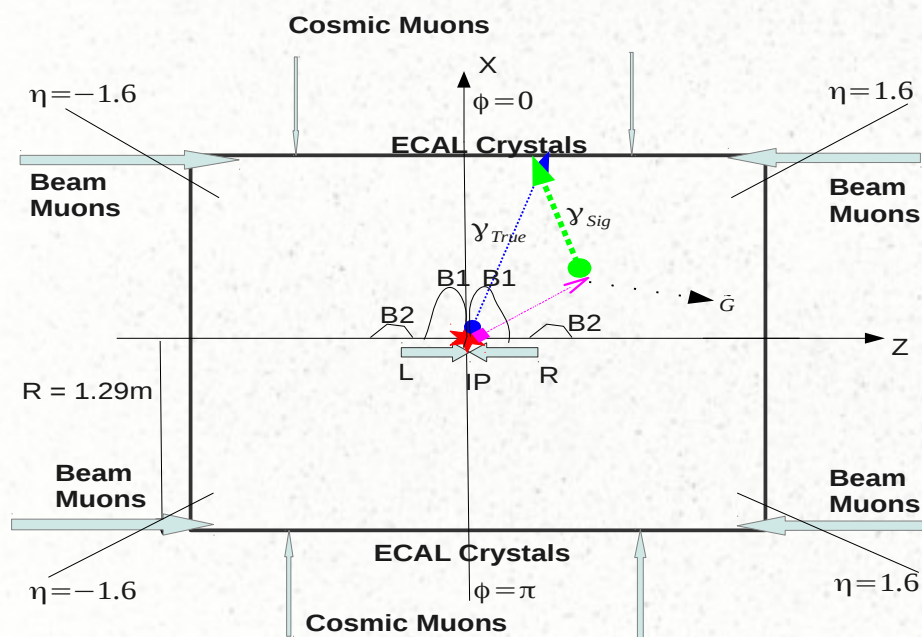
# *DataSet & Triggers*

- **DataSets:**
  - /SinglePhoton/Run2012C/Prompt-Reco V1&2
  - Int Lumi =
- **HLT Trigger:**
  - HLT\_DisplacedPhoton65\_CaloIdVL\_IsoL\_PFMET25



# Signal from Background

- **Ecal Time Calibration:**  
 $\langle t_{\gamma_{true}} \rangle \simeq 0$  but  $t_{\gamma_{Sig}} \neq 0$
- **2 Types of Backgrounds**

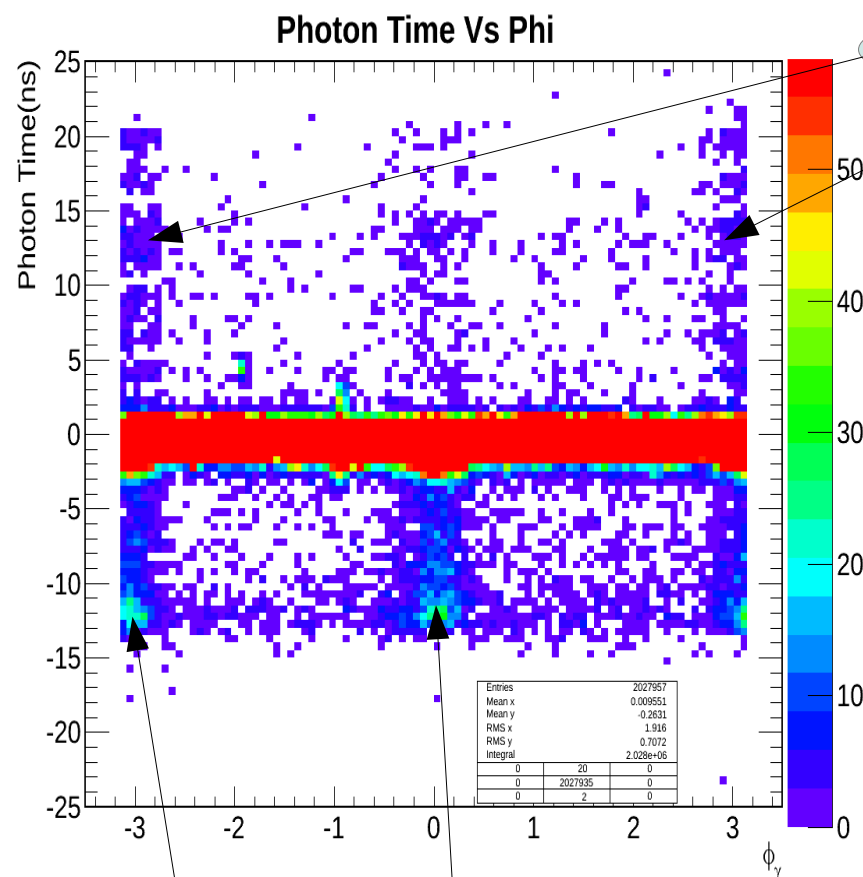


- **Collision :**
  - Left B1 w/ Right B2
  - Left B2 w/ Right B1
  - Left B2 w/ Right B2
- **Non-Collision:**
  - Beam halo (beam dump or  $P+Gas \rightarrow$  muons which Brem/shower in ECAL.
  - Cosmic muons.

# *Event Selection*

- **Selection :**
  - $\text{Gamma\_Pt1(2)} > 60(45)\text{GeV}$
  - $|\text{eta}| < 2.5, \text{Jet\_pt} > 35 \text{ GeV}$
  - Egamma VL Iso cuts,
  - $\text{MET} > 0 \text{ GeV}$
- **Photon Tagging:**
  - CSC Segment  $|\text{eta}| > 1.6$
  - Halo tagged if  $\text{dPhi}(\text{cscSeg}, \text{gamma}) < 0.6$

# Photon time vs $\Phi$



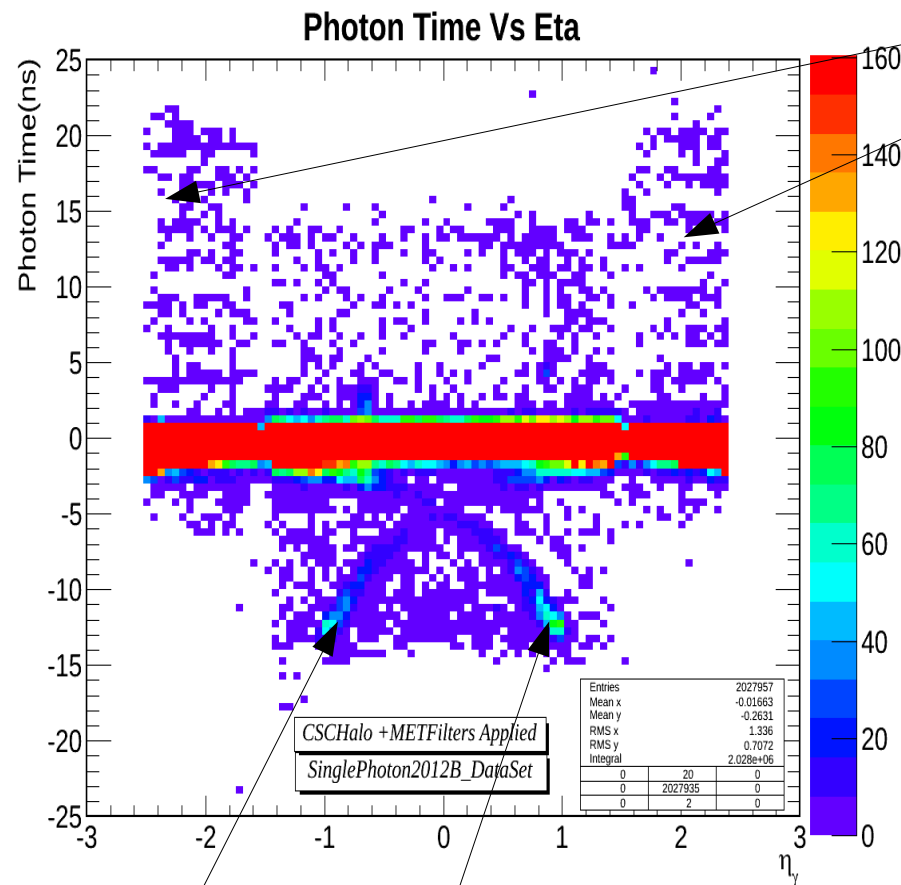
Cosmic muons?

- Observed in **2012B dataset**
- High intensity in  $\Phi = 0 \pm \pi$
- Most arriving early in Ecal time.
- +ve and -ve might not all come from same source.

Beam halo muons?



# Photon time vs Eta

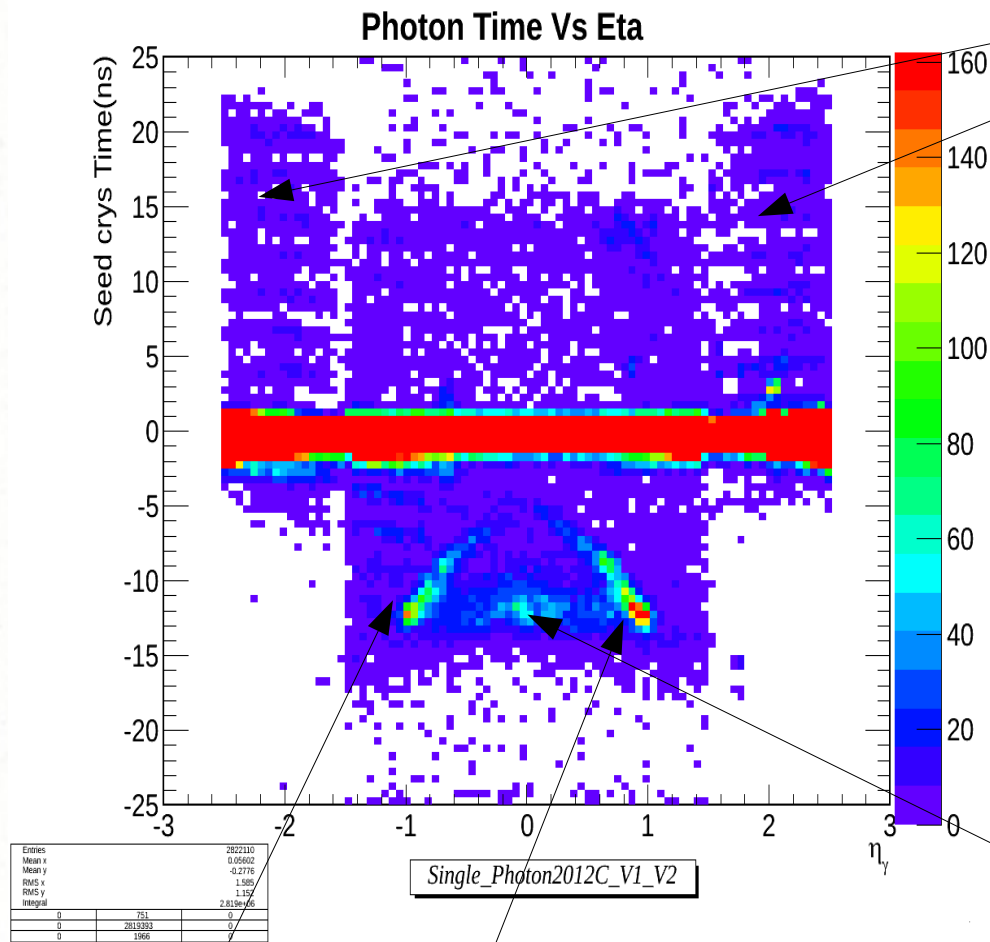


Cosmic muons?

- Intense at impact point in EB then slowly decreases towards IP.
- Eta dependence in early Ecal time.
- Surely +ve and -ve cannot all come from a **unique** source.

Beam halo muons?

# Photon time vs Eta



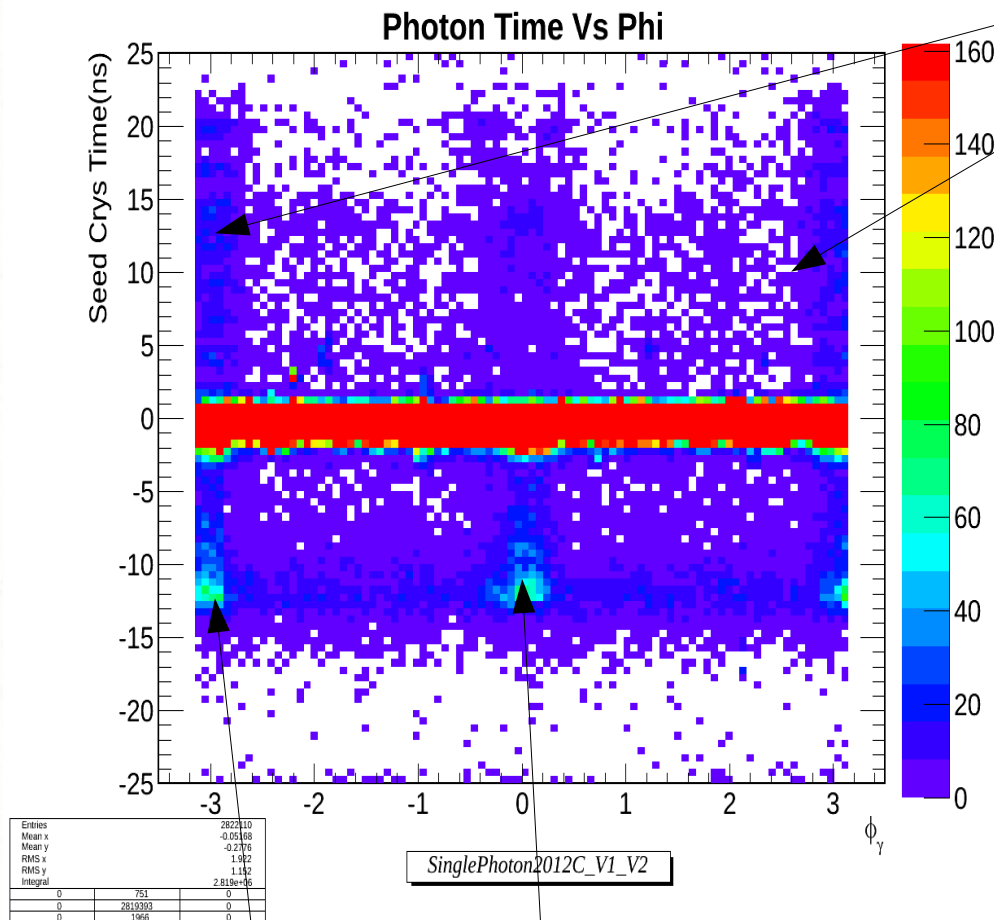
Cosmic muons?

- Similar phenomenon in **2012C dataset**.
- Increased intensity with luminosity
- The fore it must be real.

Beam halo muons?

Must be spikes failing spike cleaning.

# Photon time vs Phi



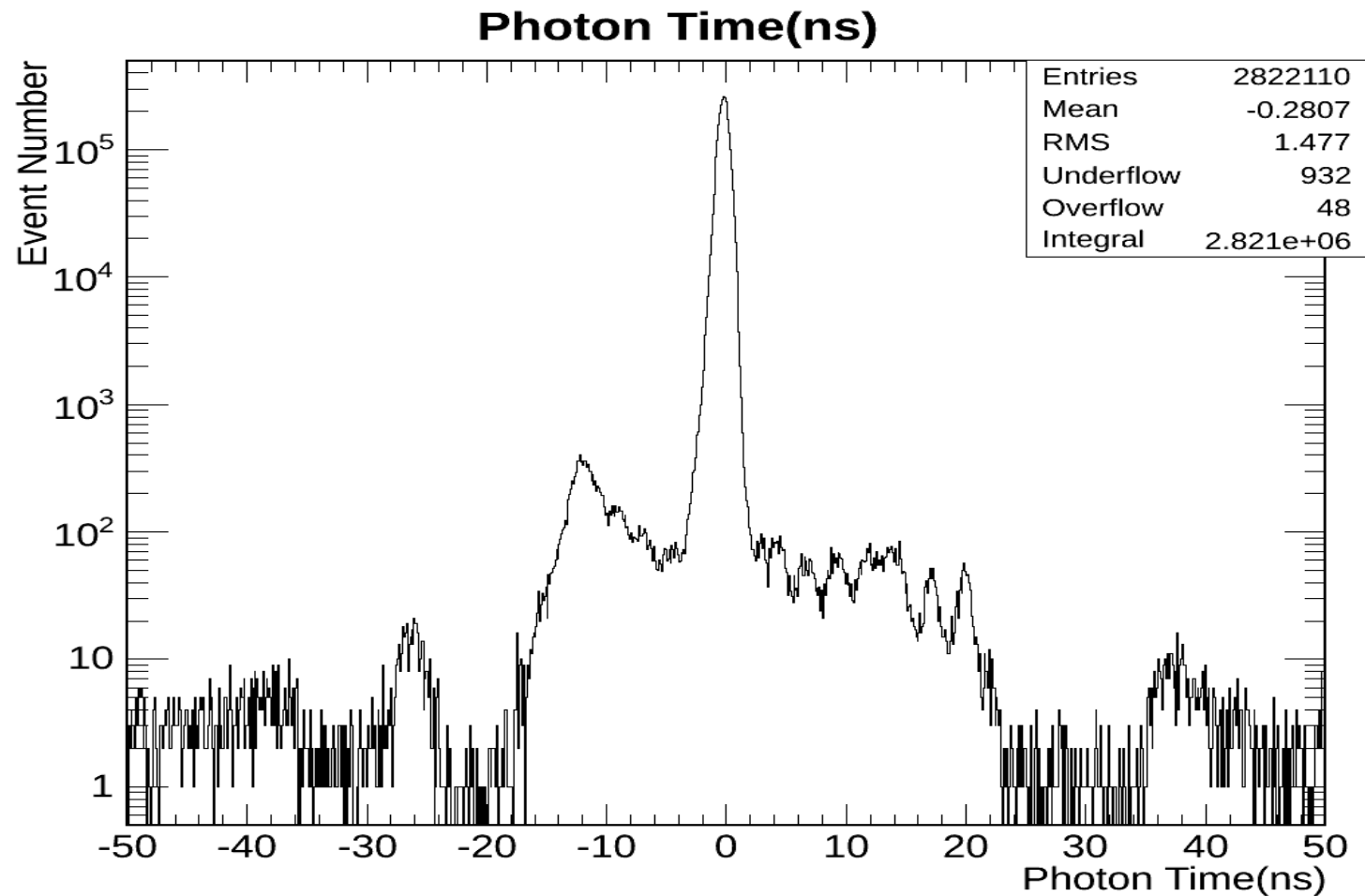
Cosmic muons?

- Similar phenomenon in **2012C dataset**.
- Increased intensity with luminosity.
- Phi dependence in +ve and -ve Ecal time of photons.

Beam halo muons?



# Photon Time



Cosmic, spikes, halo, and what else? All here!

# *Definition*

- **Egamma Photon :**

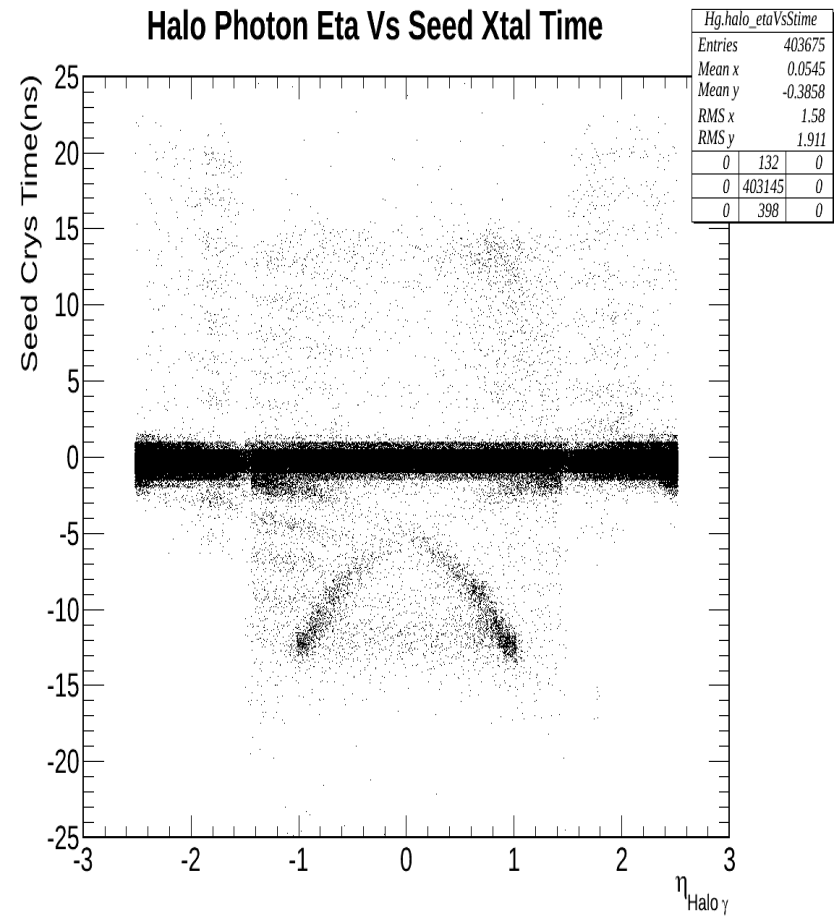
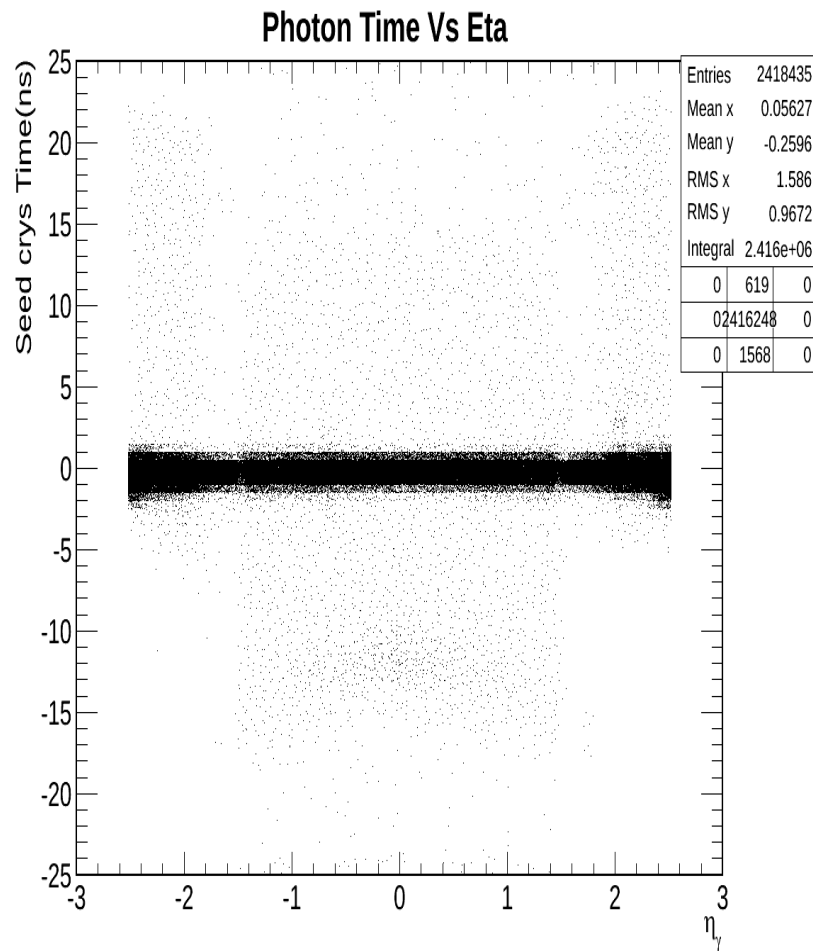
- \* Egamma + VL Iso selection criteria + Sminor.

- **Halo Photon :**

- Tag as Halo photon if:

- \* CSC Segment matching:  $d\phi(\text{cscsegment}, \text{gamma}) < 0.6$

# EB & EE Photons



Egamma Photon

Halo Photon

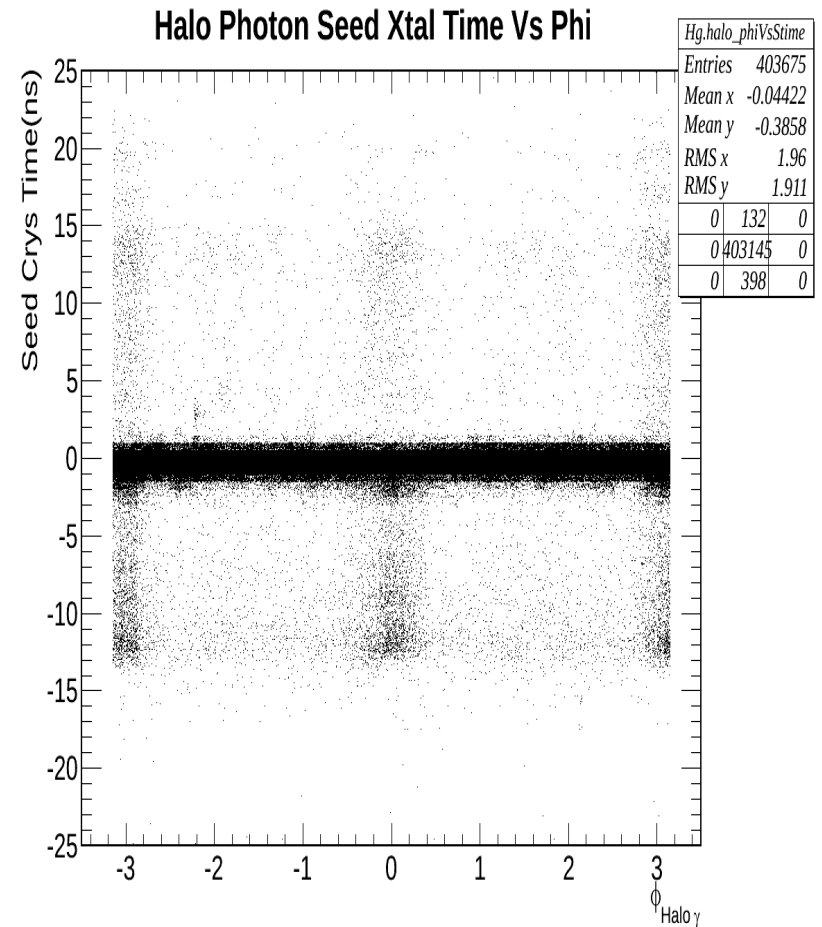
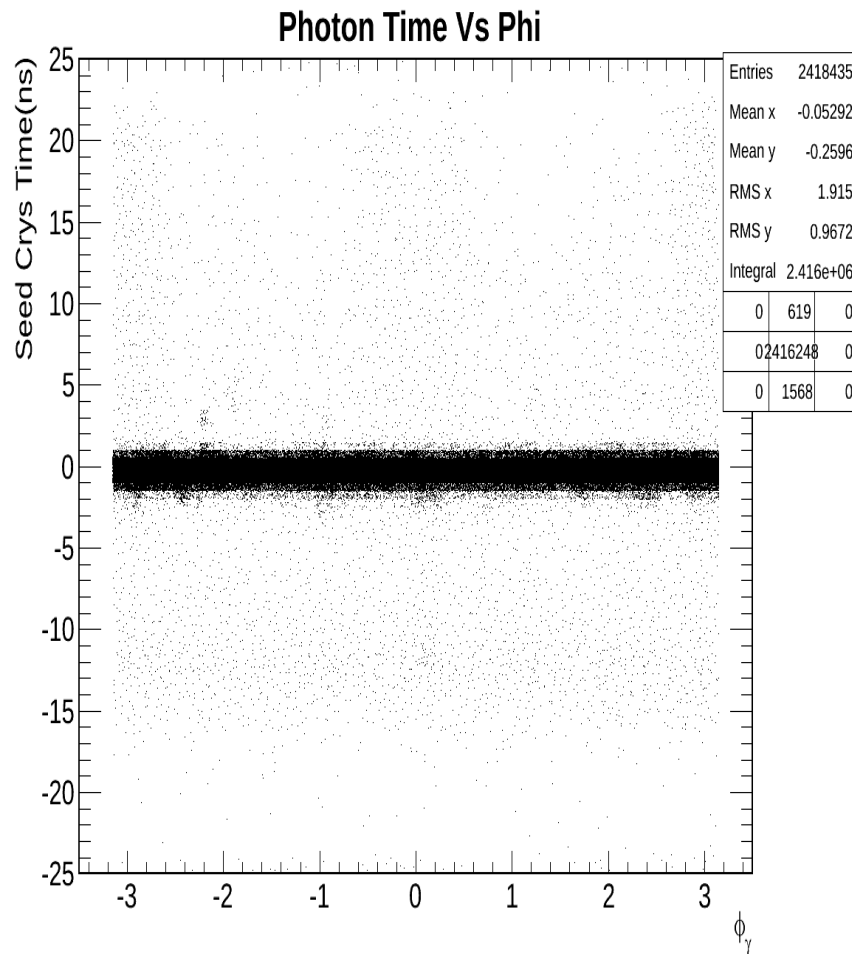
02/11/2012

Tambe E. Norbert(UMN US)

11



# EB & EE Photons



Egamma Photon

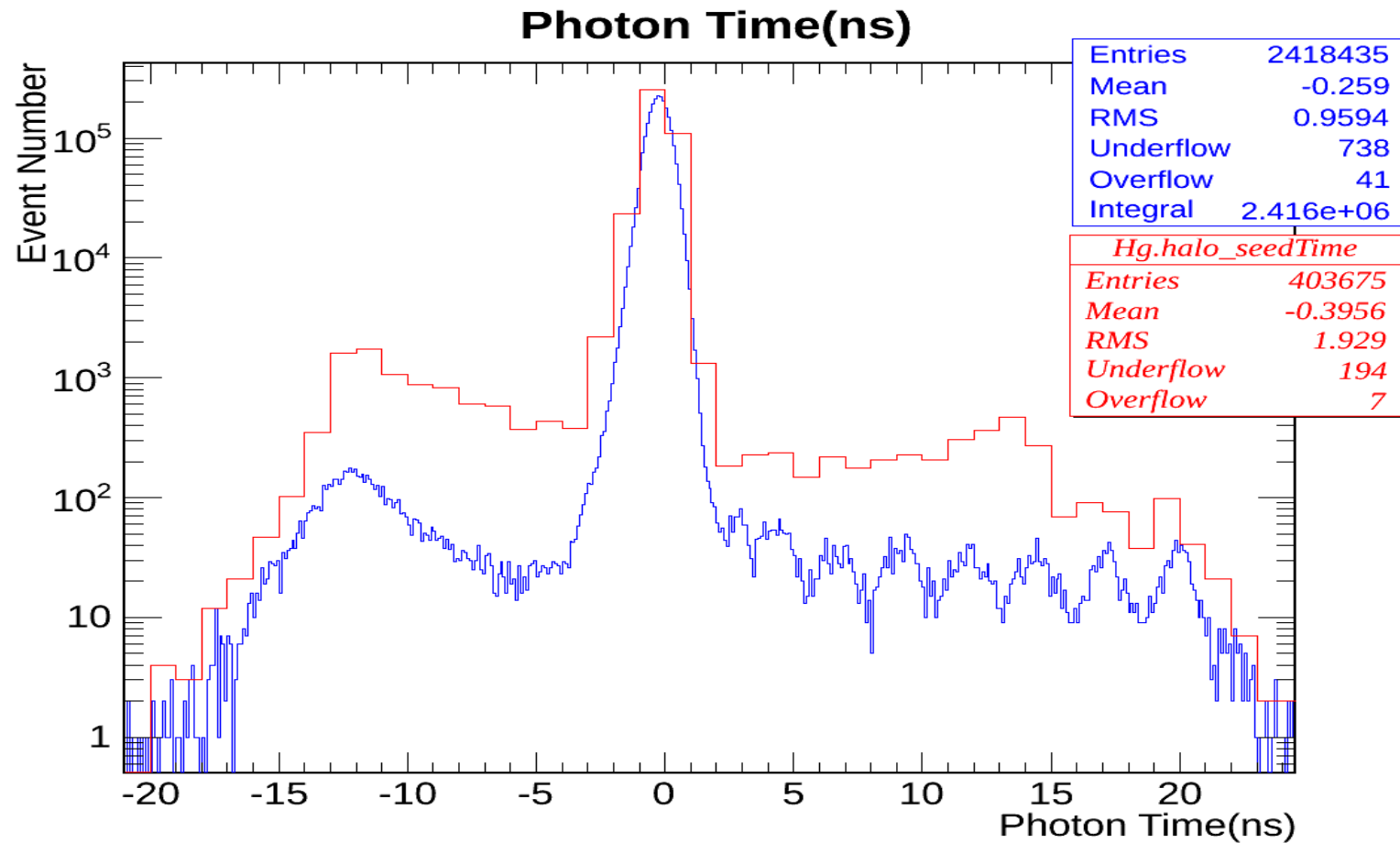
Halo Photon

02/11/2012

Tambe E. Norbert(UMN US)

12

# EB & EE Photon Time

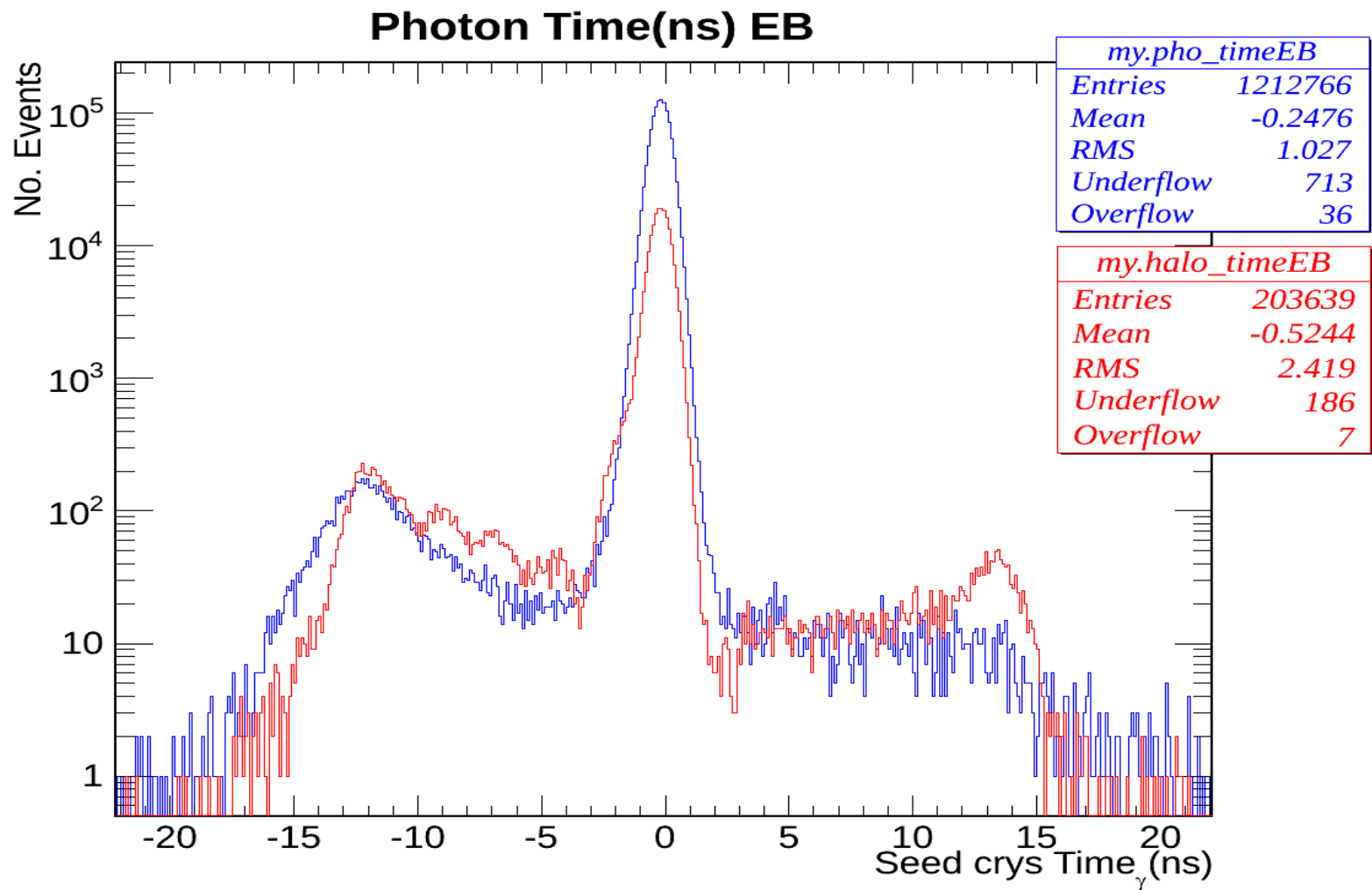


- Strange structure in later time

02/11/2012

Tamara E. Norbert(UMIN-US)

# Photon Time EB



02/11/2012

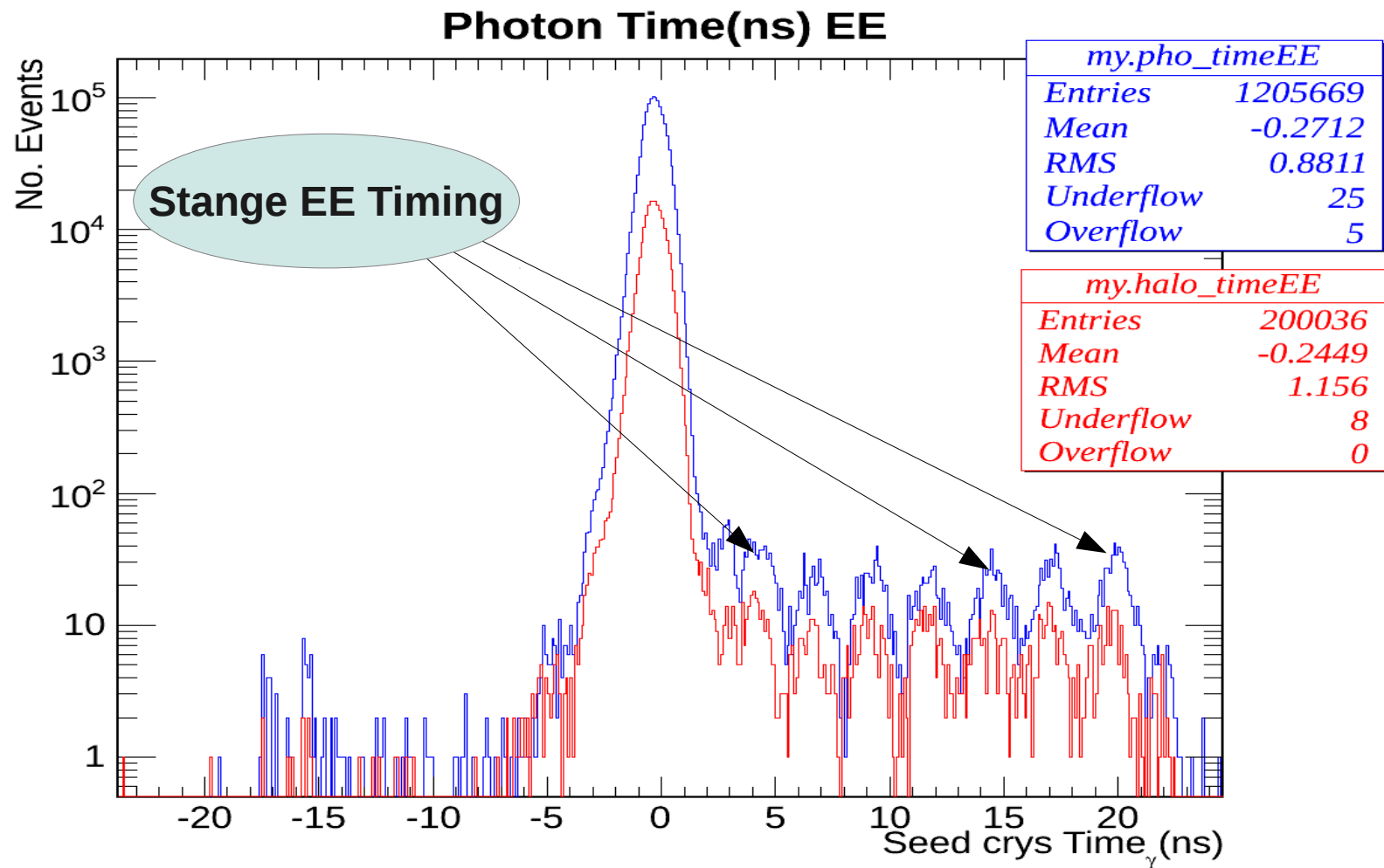
Egamma(Blue) Halo(Red) Photon Ecal time

Tamara E. Norbert (UMN US)

14

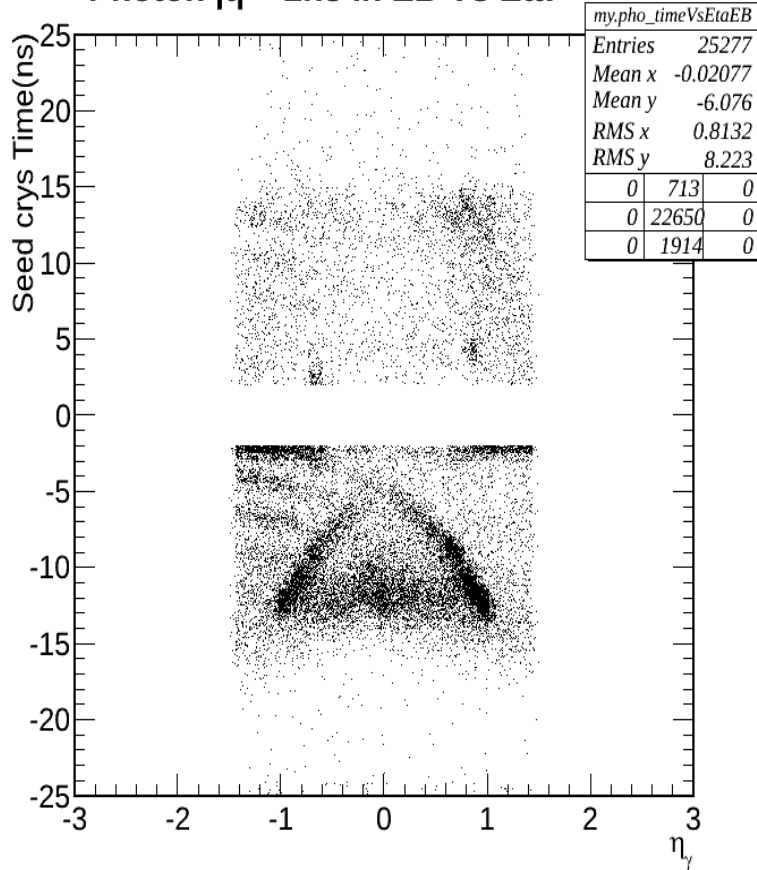


# Photon Time EE

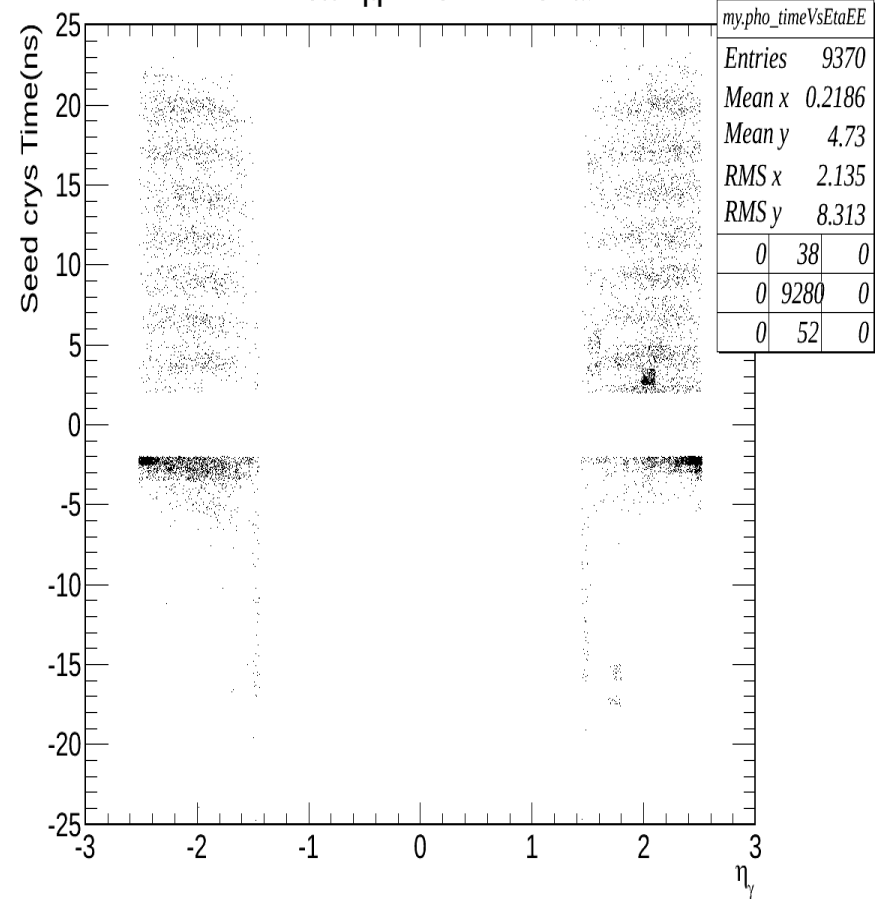


# Region of Interest EB & EE

Photon  $|t| > 2\text{ns}$  in EB Vs Eta



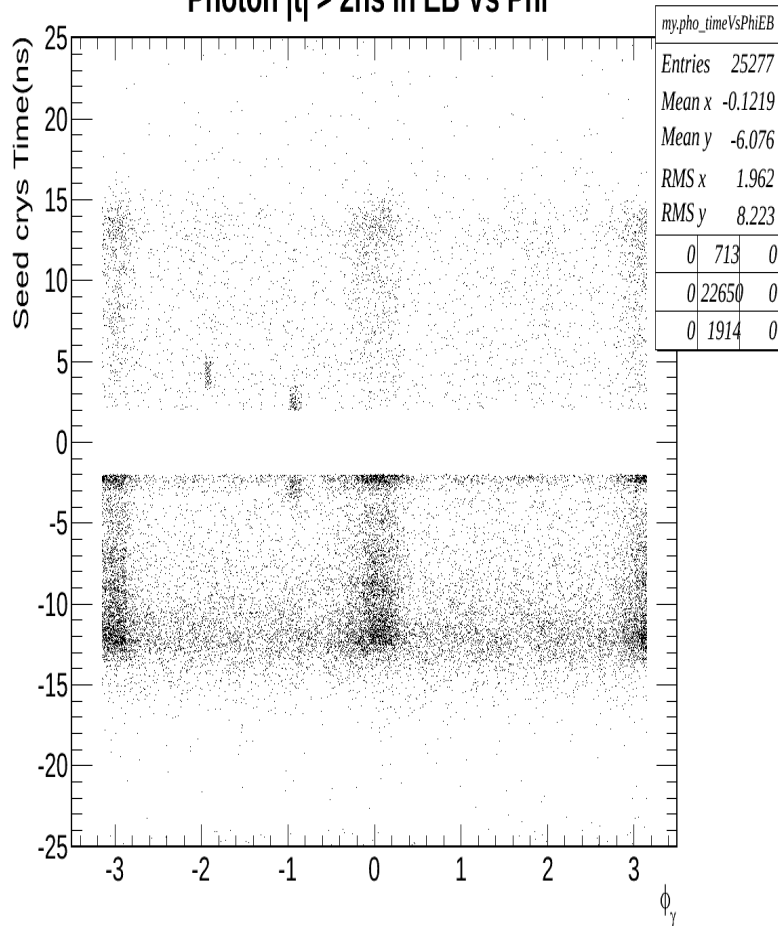
Photon  $|t| > 2\text{ns}$  in EE Vs Eta



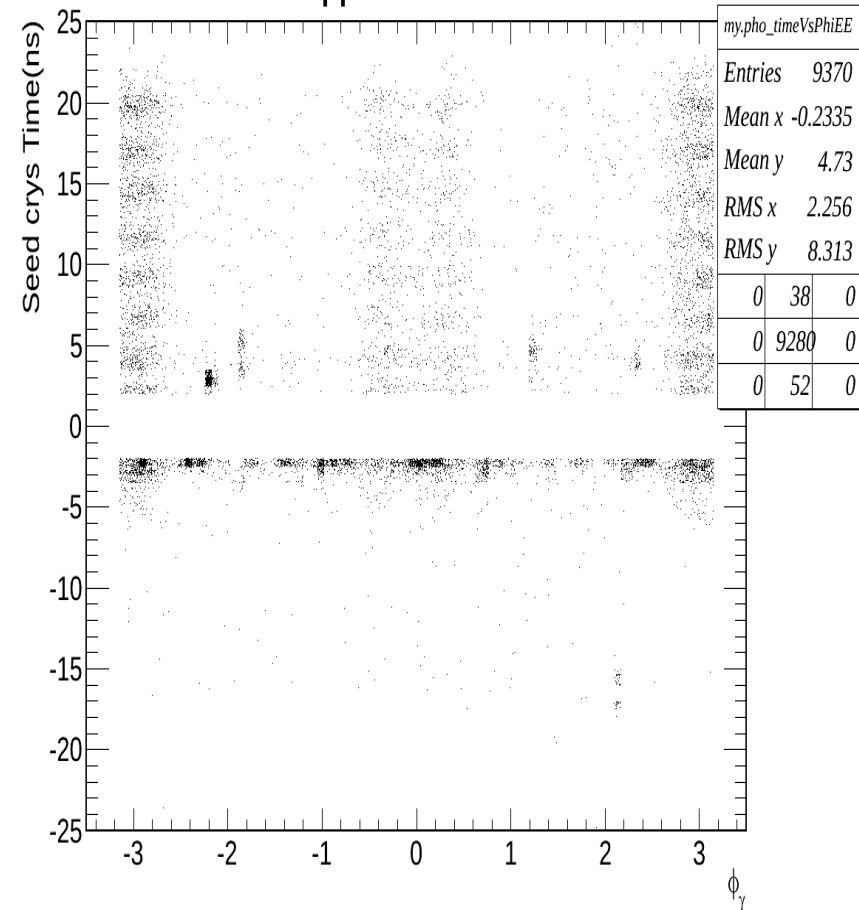
Egamma + Halo Photons time Vs Eta

# Region of Interest EB & EE

Photon  $|t| > 2\text{ns}$  in EB Vs Phi



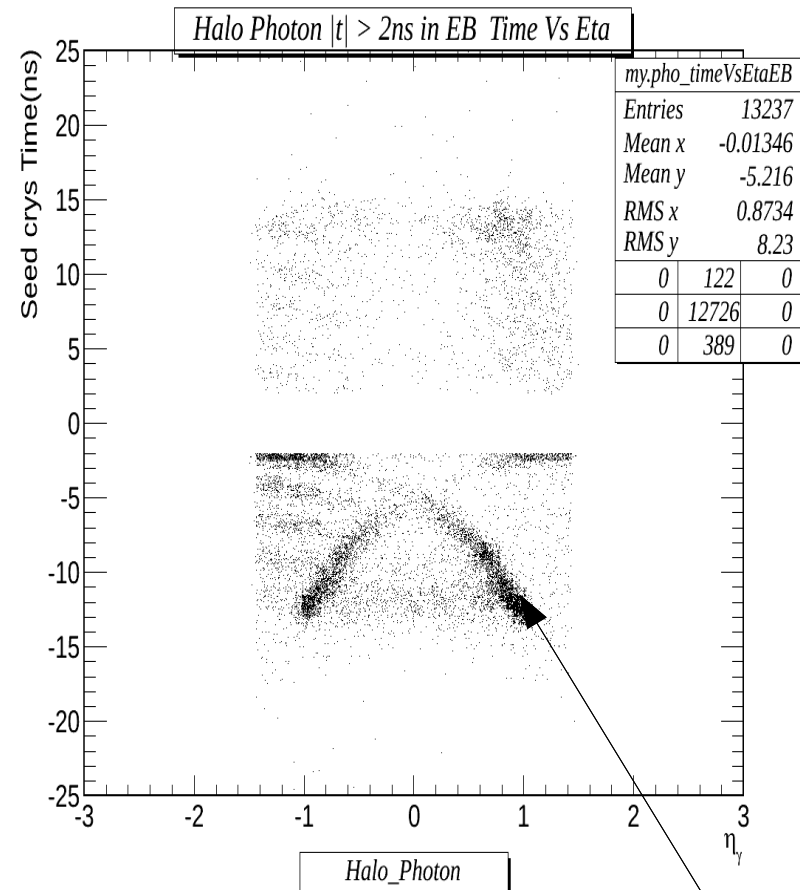
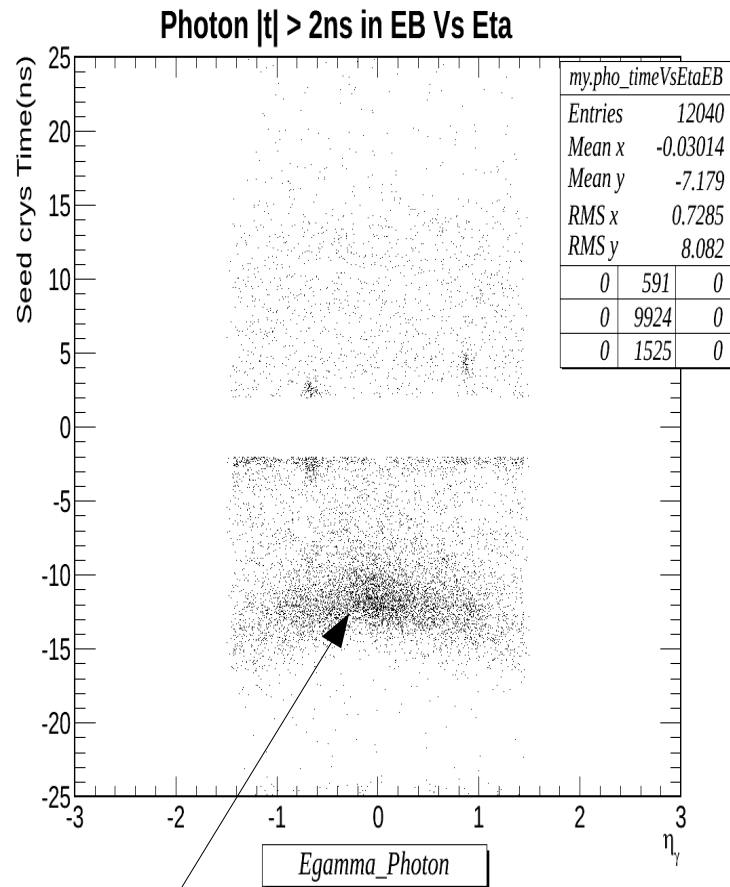
Photon  $|t| > 2\text{ns}$  in EE Vs Phi



02/11/2012 Egamm + Halo Photons time Vs Phi



# Region of Interest: EB



Spikes?

Beam Halos

02/11/2012

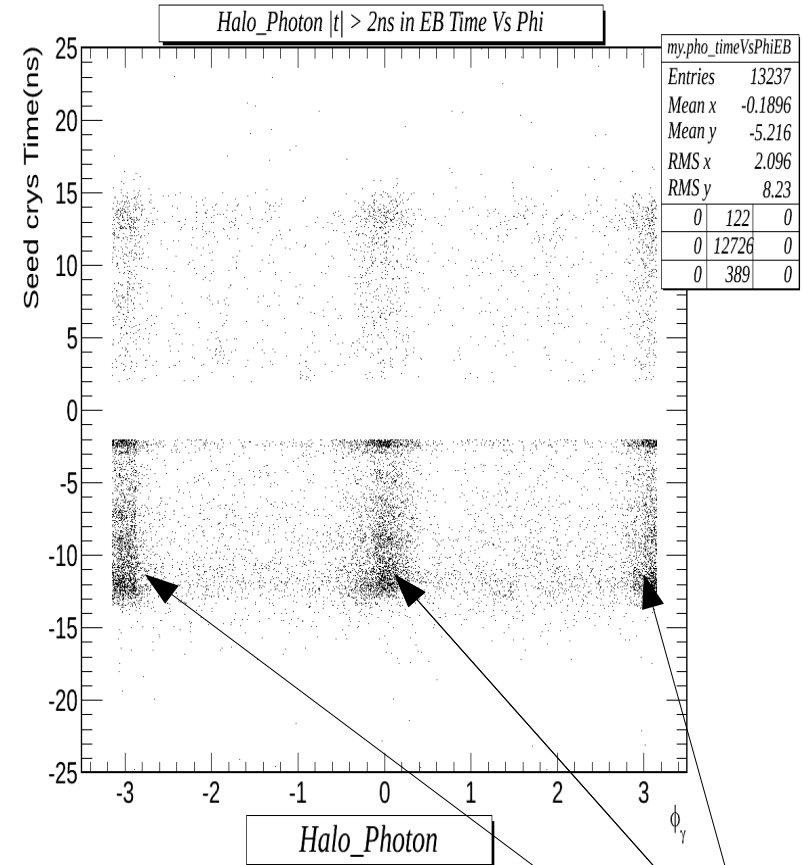
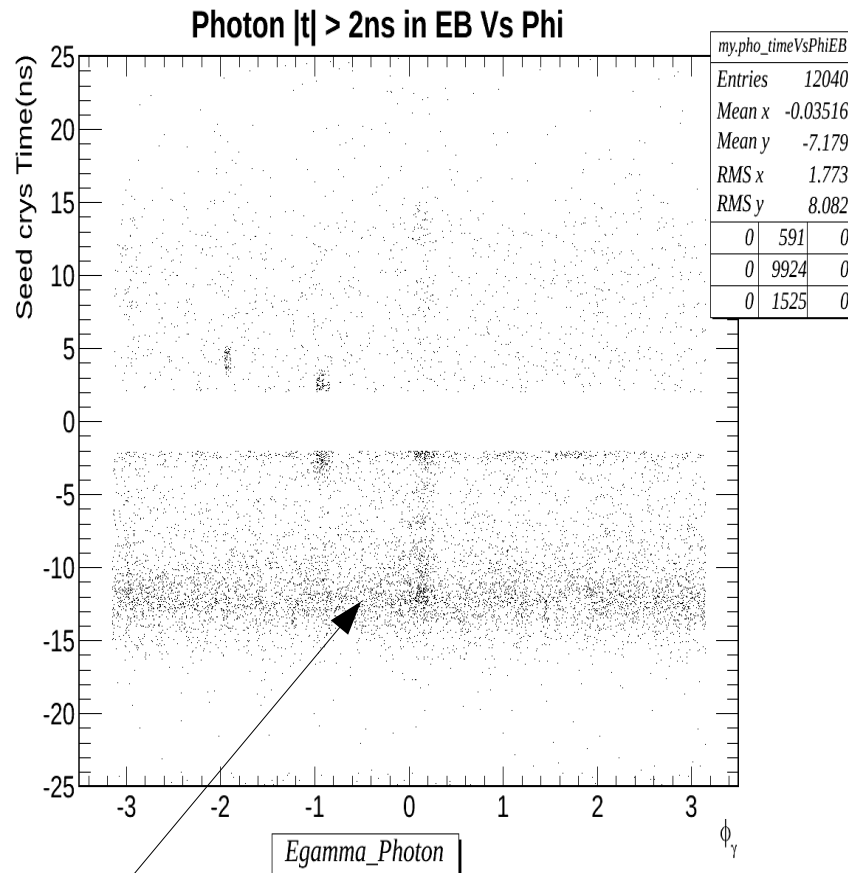
Egamma photons

Franke E. Norbert(UMN US)

Halo photons.

18

# Region of Interest: EB



Spikes?

Beam Halos

02/11/2012

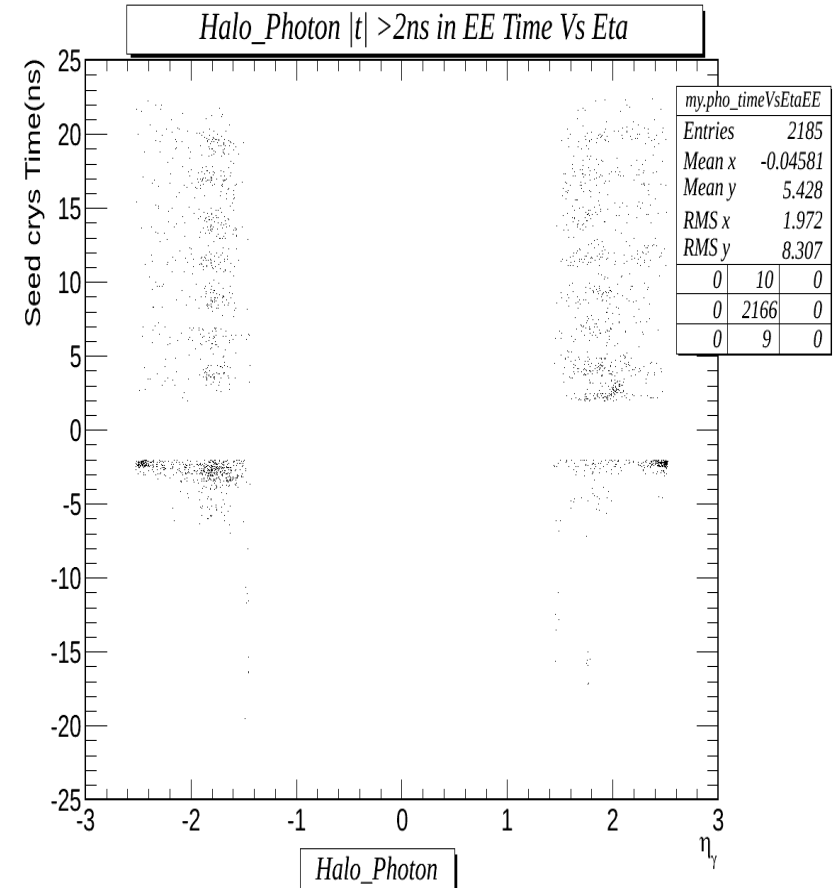
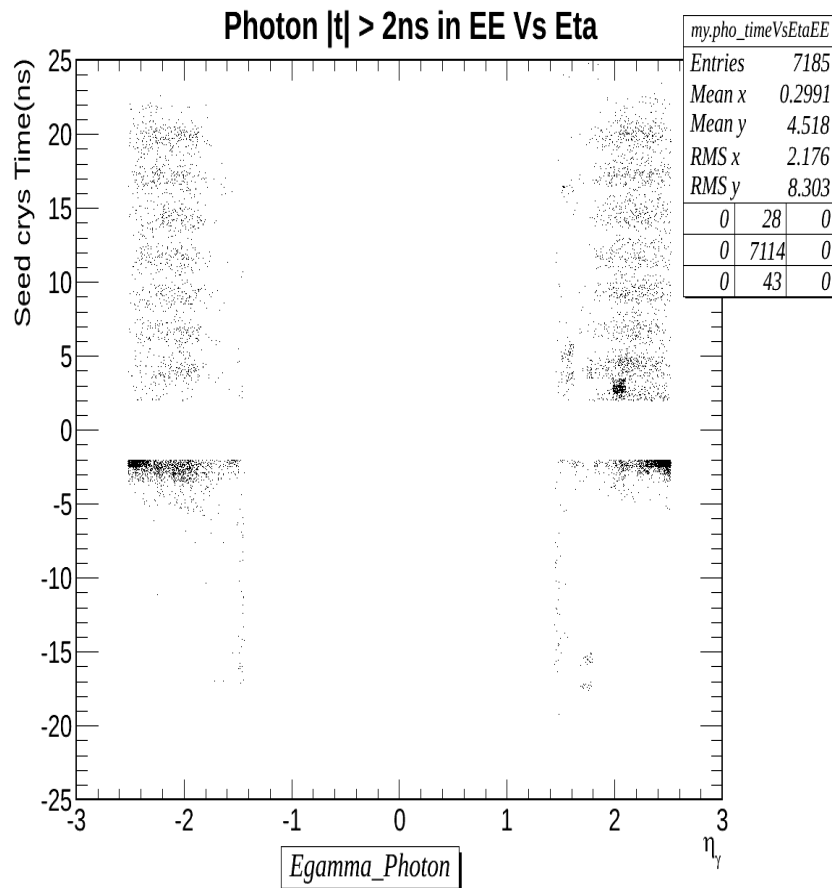
Egamma photons

Frank E. Norbert(UMN US)

Halo photons.

19

# Region of Interest: $EE$



Few early time Halo photon than in EB

02/11/2012

Egamma photons

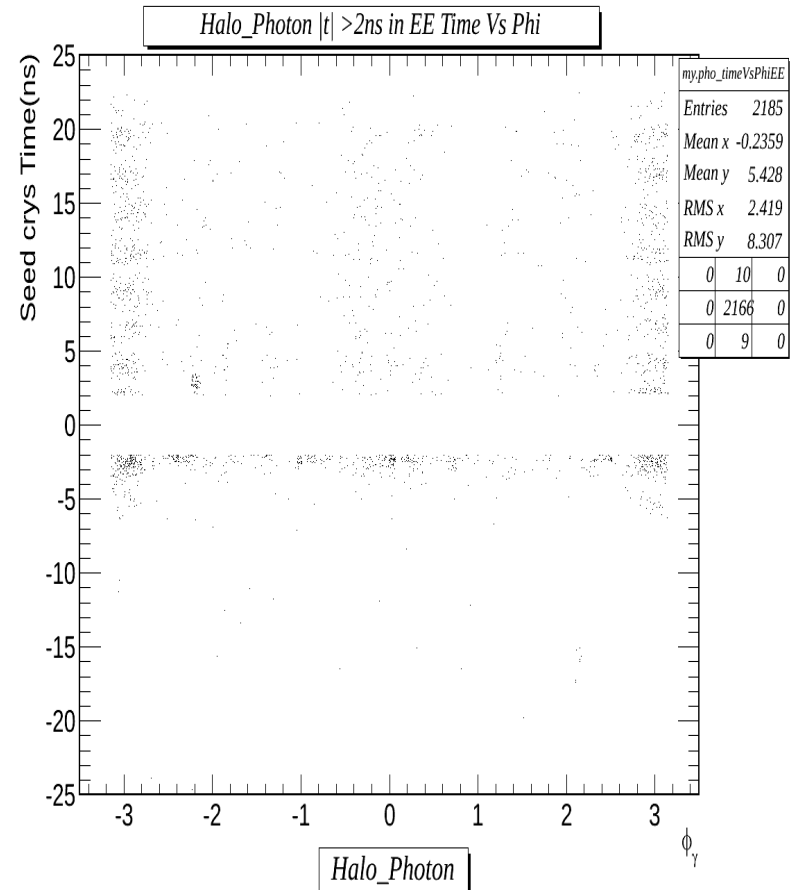
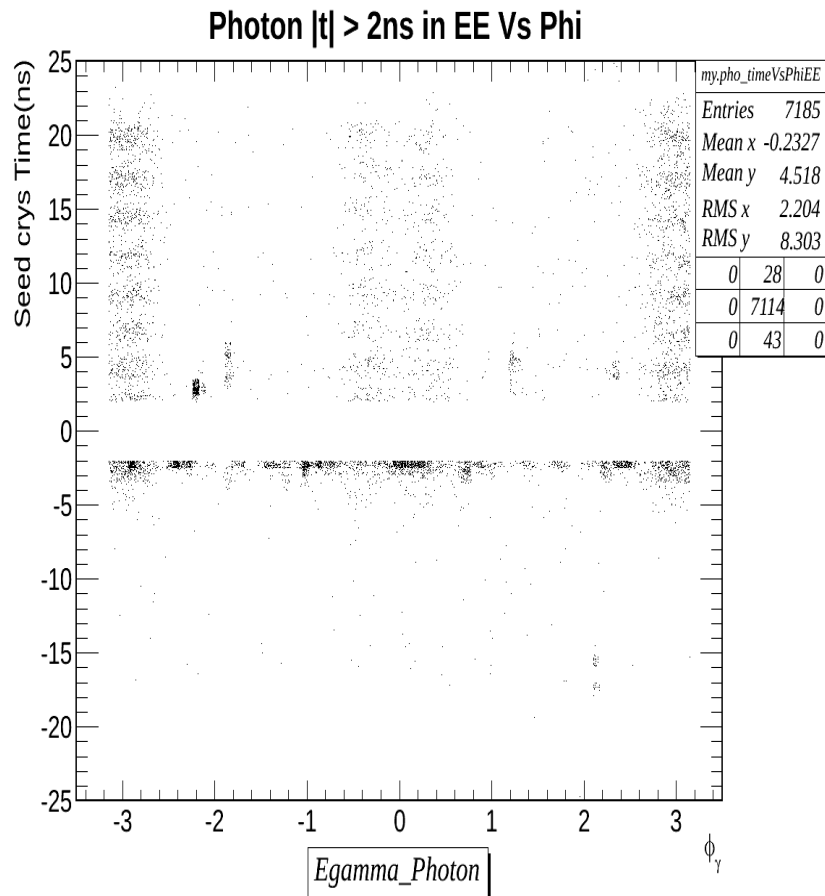
Franke E. Norbert(UMN US)

Halo photons.

20



# Region of Interest: EE



Few early time Halo photon than in EB

02/14/2012

Egamma photons

Tambe E. Norbert(UMN US)

Halo photons.

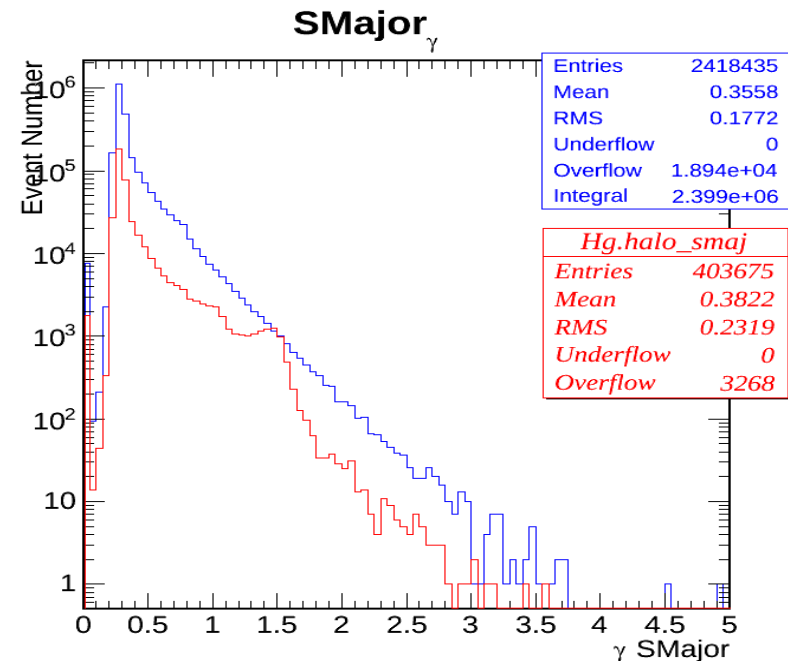
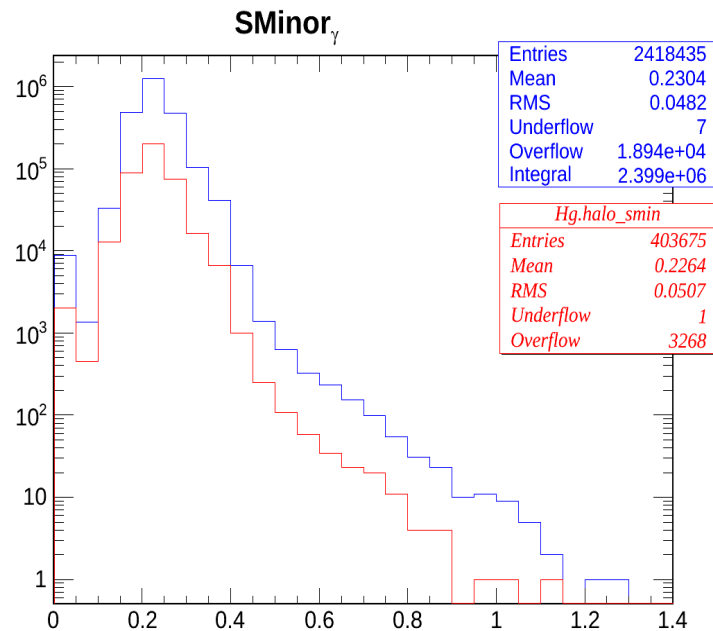
21

# *Halo Tagging & Egamma Photon Selection Efficiency.*

$ t  > 2\text{ns}$	Total Number of Photons	Egamma Photons	Selection Efficiency(%)	CSC Halo Tagged Photons	Tagging Efficiency
<b>EB</b>	25277	12040	<b>47.6</b>	13237	<b>52.36</b>
<b>EE (<math> \eta  &lt; 2.5</math>)</b>	9370	7185	<b>76.68</b>	2185	<b>23.3</b>

- For photons with time **outside 2ns window**:
  - Halo tagging efficiency : **EB(EE) = 52.36(23.3)%**
  - Egamma selection efficiency: **EB(EE) = 47.6(76.68)%**

# Photon Iso and Id variables.



SMinor

SMajor

Egamma and Halo have very similar isolation criteria.

Egamma(**Blue**) Halo(**Red**) Photon

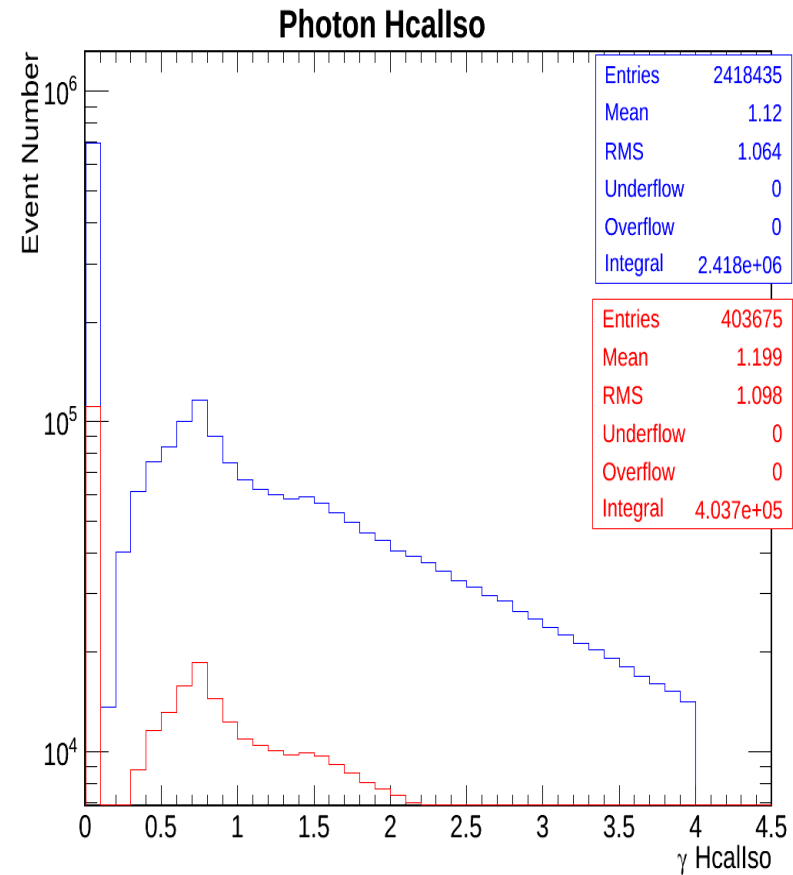
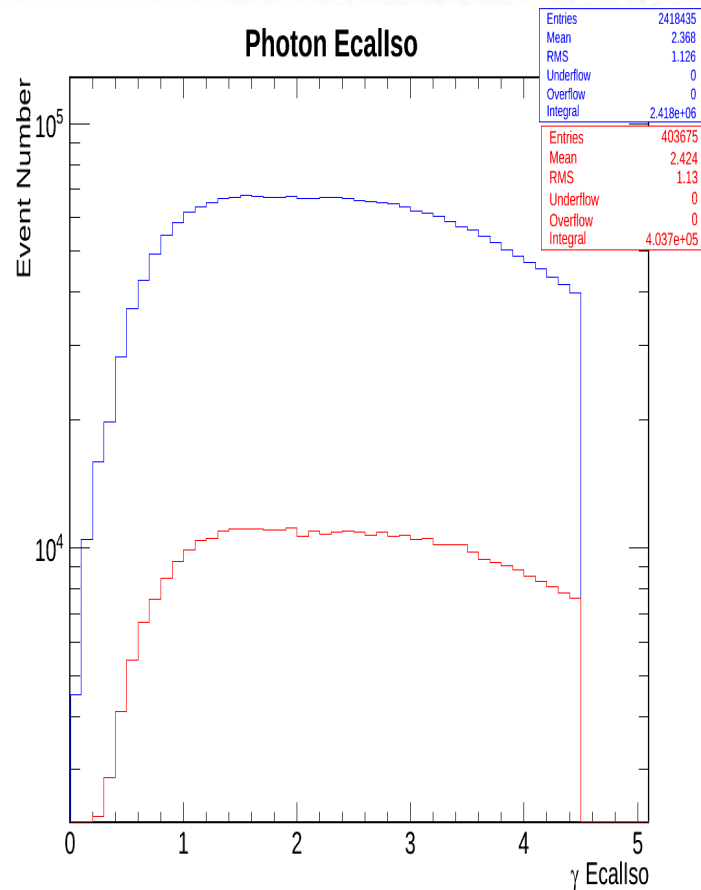


## *Summary*

- There are many non trivial background sources to delayed photon.
- CSC tagging can be use to reject beam halo photons with good efficiency, however not every background photon.
- Egamma photon Isolation variable cannot reject these background photons( see BU slide).
- With current background understanding, we are ready for Moriond although identifying the different sources is yet to come, nevertheless, we are working on it.

# *BACK UP*

# *Egamma Photon Id variables.*



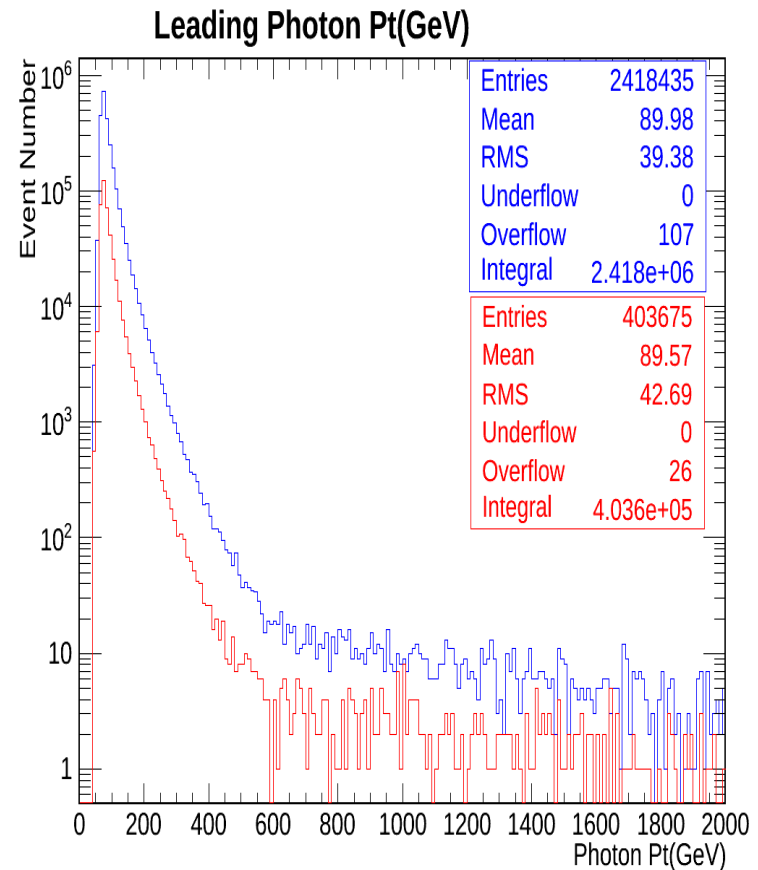
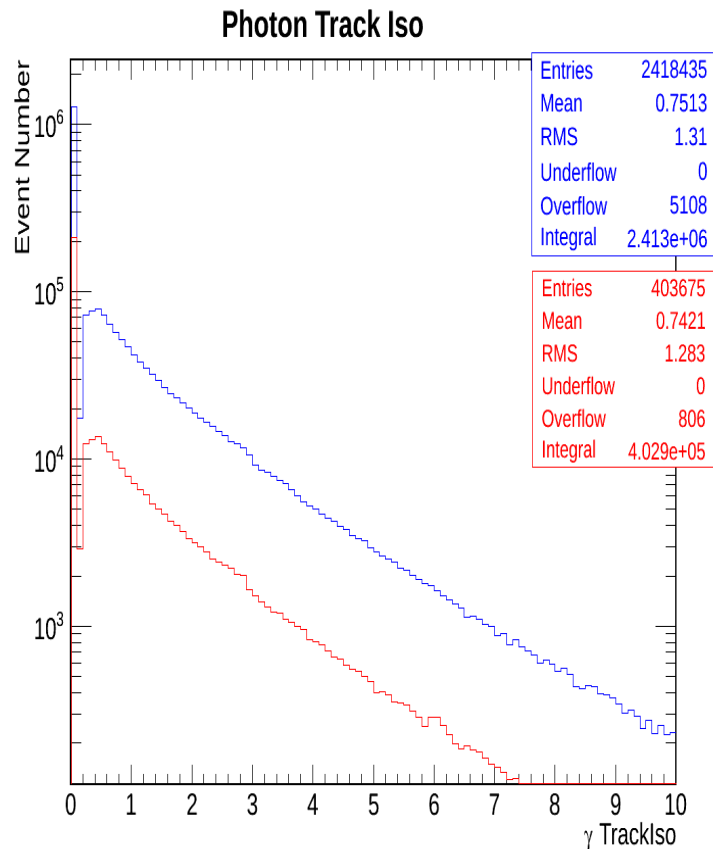
Ecal Isolation

Hcal Isolation

02/11/2012 Egamma(Blue) Halo(Red) Photon



# *Egamma Photon Id variables.*

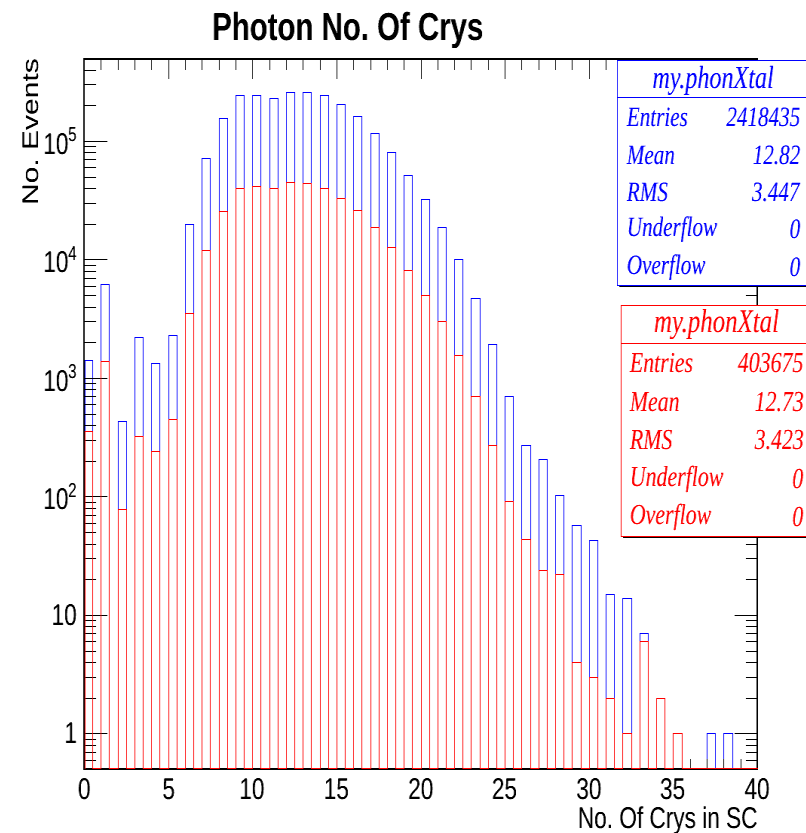
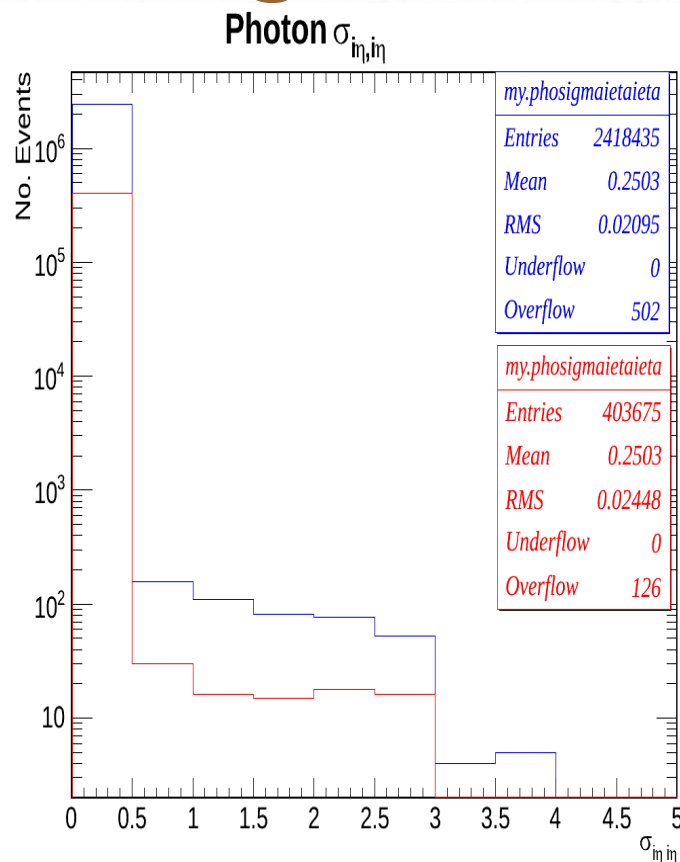


Track Isolation

Leading photon pt

02/11/2012 Egamma(Blue) Halo(Red) Photon

# Egamma Photon Id variables.



Sigma Ieta Ieta

Number of Crystals in BC

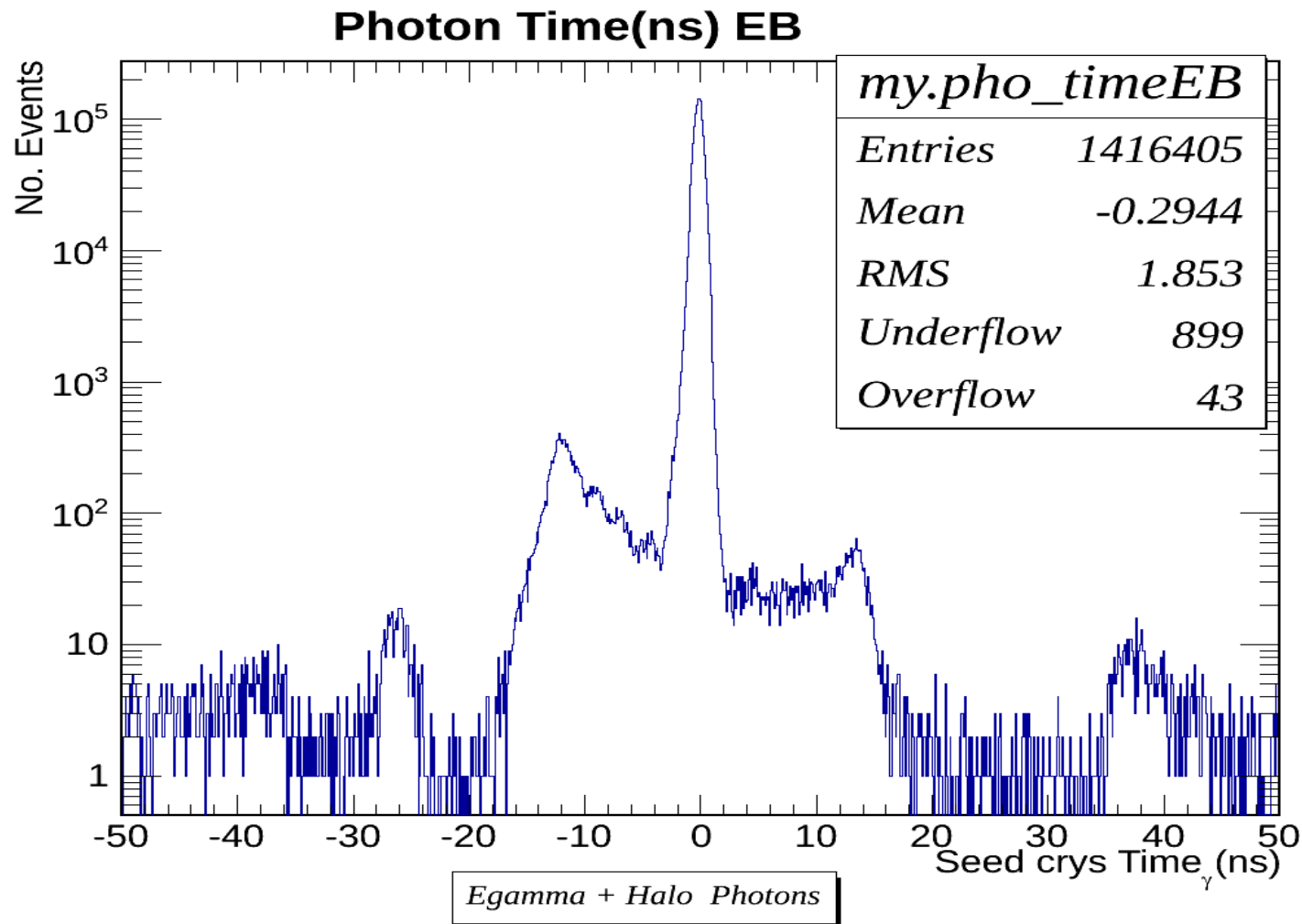
02/11/2012

Egamma(Blue) Halo(Red) Photon

Tamara E. Nelson (UMN/US)

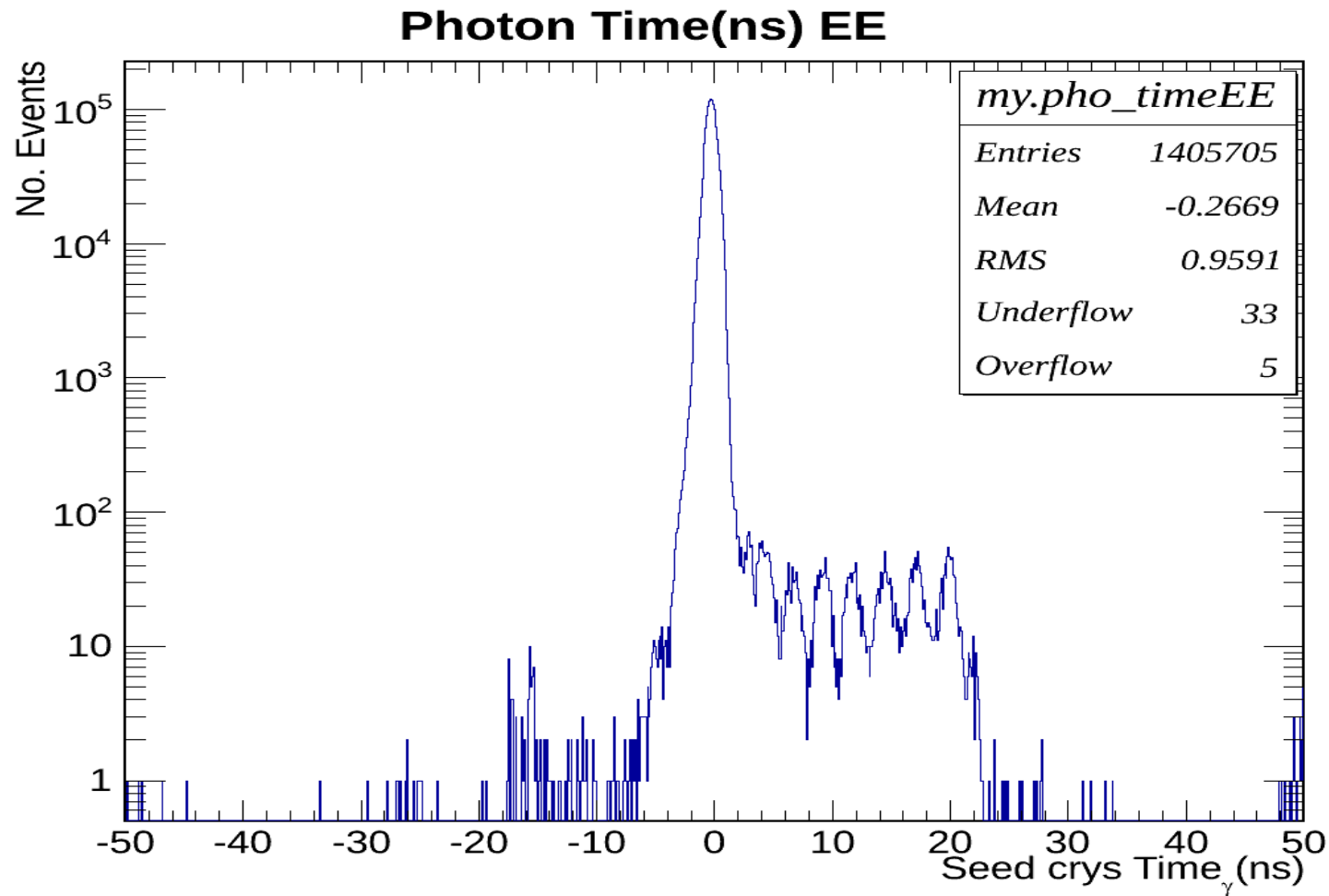
28

# *Halo + Egamma Photon time EB*





# *Halo + Egamma Photon time EE*



# Photon Time EE

