

**Search for New Massive Long-Lived Neutral Particles
decaying to Photons in pp Collisions at $\sqrt{S} = 8$ TeV**

**A DISSERTATION
SUBMITTED TO THE FACULTY OF THE GRADUATE SCHOOL
OF THE UNIVERSITY OF MINNESOTA
BY**

Tambe Ebai Norbert

**IN PARTIAL FULFILLMENT OF THE REQUIREMENTS
FOR THE DEGREE OF
Doctor of Philosophy**

Prof. Yuichi Kubota

October, 2014

© Tambe Ebai Norbert 2014
ALL RIGHTS RESERVED

Acknowledgements

There are many people that have earned my gratitude for their contribution to my time in graduate school.

Dedication

To those who held me up over the years

Abstract

Dark matter particles are believed to be neutral, stable, weakly interacting with ordinary matter and maybe massive. The hunt for dark matter particles is on! There are theoretical models which predict the existence of dark matter particles that can be produced in a proton-proton collider like the Large Hadron Collider provided there is sufficient center of mass energy. Gauge Mediating Supersymmetric Models are examples of such models. These models describe the production and decay into isolated energetic photons, new, massive, neutral, weakly interacting particles known as supersymmetric particles. The **lightest supersymmetric particle, the Neutralino ($\tilde{\chi}_1^0$)**, is a prime example and its decay into a photon is accompanied by a light weakly interacting and stable supersymmetric particle, the gravitino (\tilde{G}) which is considered to be a very good candidate for dark matter particles. The resulting photon from such a decay, is understood to be delayed in its arrival time at the detector. This timing delay is due to inherent dynamics understood and well described by models beyond the standard model. The signature of a delayed photon is not specific to only supersymmetric models but could be the result of probably some new kind of physics unrelated to the standard model. Using the Compact Muon Solenoid detector at the LHC, we have performed a search for delayed photons produced from proton-proton collisions at the center of mass energy, $\sqrt{S} = 8$ TeV. We did not find any excess of events over standard model background. Consequently, we compute an upper limit on the cross section $\sigma_{\tilde{\chi}_1^0} > \chi\chi$ on the possible production and decay of a lightest neutralino with mass and proper lifetime; $m_{\tilde{\chi}_1^0} \geq \chi\chi$ GeV/ c^2 and $\tau_{\tilde{\chi}_1^0} \geq \chi\chi$ ns respectively, as described in the gauge mediated supersymmetric models. We also show that using only the timing information of the electromagnetic calorimeter as an observable, the CMS detector is sensitive to neutralinos with **life time** up to 30 ns and mass, $m_{\tilde{\chi}_1^0} = 255$ GeV/ c^2 which no previous experiment had shown. We provide hints on possible improvements which might help discover delayed photons in future search analysis.

Contents

Acknowledgements	i
Dedication	ii
Abstract	iii
List of Tables	viii
List of Figures	ix
1 Introduction	1
2 Model and Phenomenology of Long-Lived Particles	5
2.1 The Standard Model of Particle Physics	5
2.1.1 Main Components	5
2.1.2 Spontaneous Symmetry Breaking	10
2.1.3 Decay Rate and Life Time	13
2.1.4 Limitations Of the Standard Model	17
2.2 Beyond Standard Model Physics	19
2.2.1 Supersymmetry	21
2.2.2 Minimal Supersymmetric Standard Model	24
2.2.3 Soft Supersymmetry Breaking	26
2.2.4 Gauge Mediated Supersymmetry Breaking and Phenomenology .	27
2.3 Long-Lived Particles in GMSB Models	33
2.3.1 Production of supersymmetric particles at Hadron Colliders . . .	34

2.3.2	Decay of supersymmetric particles in CMS detector	35
2.3.3	Why is the search for neutral long-lived particles important? . .	38
2.4	Previous Experiments and Results	41
3	Hadron Collider and Detector	42
3.1	Large Hadron Collider	42
3.1.1	Overview	42
3.1.2	Colliding Energy	43
3.1.3	Luminosity	44
3.1.4	Superconducting Electromagnets	46
3.1.5	Timing	47
3.2	Compact Muon Solenoid	52
3.2.1	Overview	52
3.2.2	Tracker	56
3.2.3	Calorimeter	58
3.2.4	Muon Chambers	63
3.2.5	Particle Detection	64
3.2.6	Triggering	65
4	Timing Reconstruction and Calibration	67
4.0.7	Electromagnetic Calorimeter Readout Chain	68
4.0.8	Timing Extraction	71
4.0.9	Timing Resolution	73
4.0.10	Timing Calibration Procedure	75
4.0.11	Electromagnetic Calorimeter Timing Performance	85
5	Physics Objects Reconstruction and Identification in CMS	90
5.1	Physics Objects Reconstruction	90
5.1.1	Supercluster Reconstruction	91
5.1.2	Vertex and Track Reconstruction	94
5.1.3	Photon or Electron Identification	96
5.1.4	Muon Reconstruction	101
5.1.5	Jet Reconstruction	103

5.1.6	Missing Transverse Energy Reconstruction	104
5.2	Anomalous Signals	106
5.3	Using Timing for Event Cleaning	109
6	Search Analysis for Long-Lived Particles	110
6.1	Analysis Strategy	110
6.1.1	Signal and Background Modelling	112
6.1.2	Datasets	113
6.2	Event Selection	114
6.2.1	Trigger	115
6.2.2	Offline Selection	116
6.2.3	ECAL Time	118
6.3	Background Estimation	121
6.3.1	Non-Collision Backgrounds	123
6.3.2	Collision Backgrounds	127
6.3.3	Event Cleaning	128
6.3.4	Background Estimation Cross Check	132
6.4	Results	136
6.4.1	Systematics Studies	136
7	Statistical Analysis	138
7.1	Limit Computation	138
7.1.1	CLs Technique	139
7.1.2	Statistical Test Formalism	140
7.1.3	Test Statistics and p -values	142
8	Limit Interpretation	146
8.1	Signal Efficiency and Acceptance	146
8.2	Future Improvements	150
8.2.1	Beam Halo Monitoring Detector	150
8.2.2	Back-end Electronics upgrade HCAL	150
9	Conclusion	151

Bibliography	152
Appendix A. Glossary and Acronyms	157
A.1 Glossary	157
A.2 Acronyms	157

List of Tables

3.1	The LHC operation parameter conditions during RUN 1:2010-2013 . . .	49
3.2	CMS Detector Material and Resolution(Time resolution: $N \approx 35$ ns, $\bar{C} \approx 0.020$ ns [?])	54
4.1	Table Comparing Timing Resolution performance of 2011 with 2012 . .	89
5.1	Simple Cut-Based criteria for High energy electron and photon identifi- cation in CMS	100
6.1	The dataset name and corresponding integrated luminosity of the data used in the analysis	113
6.2	The signal GMSB MC samples used in this analysis	114
6.3	The γ + jets samples used in this analysis	114
6.4	The photon ID selection as used in this analysis	117
6.5	The Jet ID selection used in this analysis	118
6.6	Final number of events estimated for each background and the number of events passing out event selection and acceptance criteria.	136
6.7	Summary of systematic uncertainties on the signal (top), background (mid- dle) and machine (bottom) as used in the σ_{UL} calculation.	137
A.1	Acronyms	158

List of Figures

2.1	Pdfs.	10
2.2	Higgs boson “Mexican hat” potential which leads to spontaneous symmetry breaking.	11
2.3	SM “particle periodic table” and their interactions through bosons as described in the SM.	13
2.4	Higgs self energy diagrams showing how the higgs boson mass is computed from both Higgs field and supersymmetric partner particle contributions.	20
2.5	SUSY Mass spectra in the mGMSB SPS8 model (left) and GGM Model (right) with mass of gluino ($M_{\tilde{g}} = 1.0$ TeV)	27
2.6	Feynman diagrams of of Gravitino interactions with superpartner pairs (ψ, ϕ) (a) and (λ, A)(b).	32
2.7	Parton distribution function (PDF) for partons against energy fraction on the horizontal axis for a particular momentum transfer Q value.	34
2.8	Feynman diagrams of single(top) and di(bottom) photon production from cascade decays of gluino and squark at LHC.	35
2.9	Neutralino transverse momentum distribution(top left) and proper decay length(top right) with its decayed photon transverse momentum distribution(Bottom right) and time of arrival at ECAL(Bottom right) for GMSB SPS8 model.	38
2.10	Neutralino lifetime and mass upper limit from ATLAS(left) and CMS(right) 7 TeV analysis with non-pointing photons and MET.	41
3.1	Schematic diagram showing the full Large hadron Collider. Image taken from [22]	43

3.2	Cumulative luminosity versus day delivered to (blue), and recorded by CMS (orange) during stable beams and for p-p collisions at 8 TeV centre-of-mass energy in 2012.	46
3.3	Longitudinal Profile taken with LDM detector showing definition of Ghost/Satellite bunches with respect to main bunches.	50
3.4	(left) Arrival time distribution(red) of ATLAS MBTS for LHC fill 1533 during 2010 Pb-Pb run and LDM profile(black) for Beam2(same for Beam1). (Right) Timing of Clusters in the CMS endcap calorimeters for fill 1089:Left: EEP detector(left side of IP $z < 0$) Right: EEM detector(right side of IP, $z > 0$). NB: Plots taken from [30] and [31]	51
3.5	CMS Detector showing the different subdetectors and their material. . .	53
3.6	Schematic diagram of CMS detector view showing definition of coordinates as used by CMS.	55
3.7	Schematic diagram of CMS Tracker showing the silicon pixel detector region (inner closer to LHC beam) and silicon strip region (outer). . . .	56
3.8	Schematic diagram of CMS calorimetry system with HCAL enclosing ECAL in the Barrel and Endcap regions.	59
3.9	Layout of the CMS electromagnetic calorimeter showing the arrangement of crystal modules, supermodules in the barrel with the preshower in front of endcap with supercrystals.	61
3.10	Longitudinal view of CMS showing the coverage range of its sub-detectors.	64
3.11	Transverse slice of the CMS detector showing how different types of particles interact and hence identified using this detector.	65
4.1	Schematic showing ReadOut Chain.	69
4.2	Typical pulse shape of a given signal showing signal amplitude and time.	71
4.3	left: A measured ECAL pulse shape for each channel. Right: $T - T_{MAX}$ Vs $R(T)$ showing the distribution of $T(R)$. Solid line is reference shape or shape from testbeam while dots correspond to a 10 discrete samples corresponding to signal from a single event in a single channel or crystal.	72

4.4	Deviation of the timing difference as a function of A_{eff}/σ_n between two crystals sharing an energy in the same electromagnetic shower obtained during electron testbeam measurements. The single crystal energy scales for barrel (EB) and endcap (EE) is overlaid. The fitted results give $N = (35.1 \pm 0.2)$ ns and $\bar{C} = (20 \pm 4)$ ns.	75
4.5	Top: Timing calibration maps showing the distribution of average time for each channel/PbWO ₄ crystal in EB (top) and EE (below: EE-(left), EE+(right)) before calibration. Bottom: Timing calibration maps showing the distribution of average time for each channel/PbWO ₄ crystal in EB (top) and EE (below: EE-(left), EE+(right)) after calibration. After calibration most crystals have an average time of zero.	79
4.6	Top: Crystal mean time distribution for crystals in readout electronics EB \pm 8. Crystal time is obtained from Laser data. Bottom: Distribution of CCU timing shift in readout electronics due to hardware intervention for EB (top) and EE (bottom). The adjustment for global timing shift per FED due to difference in light source for each CCU can be seen to reduce the possibility of CCU showing false timing shift. The figures show the distributions of the Δt_{CCU} with corresponding RMS values before (left) and after (right) the global shift has been subtracted.	81
4.7	Top: Distribution of mean time as a function of crystal energy for EB prior (left) and after (right) timing bias corrections depending on amplitude developed have been applied. Bottom: Distribution of timing resolution as a function of crystal energy for EB prior (left) and after (right) timing bias corrections depending on amplitude developed have been applied.	83
4.8	Top: Distribution of mean time as a function of amplitude (left) and Resolution as a function of amplitude (right) for different pseudo-rapidity regions in the barrel. Bottom: All modules in EB combined timing resolution as a function against η crystals in the same readout electronics in barrel (EB).	85

4.9	Ecal absolute time of a single reconstructed electron in $Z \rightarrow e^-e^+$ decay. The electron time is the seed (crystal with highest energy deposit)time of the electron.(a) in in EB and (b) in EE	87
4.10	Ecal time difference between the two reconstructed electrons in $Z \rightarrow e^-e^+$ decay. The electron time is the seed (crystal with highest energy deposit) time with additional correction due to the time of flight of the electron.(a) in in EB and (b) in EE	88
4.11	Resolution of time difference between two most energetic crystals in the same readout unit of an ECAL cluster as a function of effective amplitude- The Neighbouring crystal method(left), Resolution of time difference between the two reconstructed electrons in $Z \rightarrow e^-e^+$ decay as a function of effective amplitude($A_{eff} = A_1A_2/\sqrt{A_1^2 + A_2^2}$), for both electrons in EB; The Z method (right). The noise term N is the same as test beam for bother methods while the constant term \bar{C} is better for the Neighbouring crystal method (70 ps) than for the Z method (150 ps) indicating the effect due to different intercalibrations between two different readout of front end electronics.	89
5.1	Superclustering algorithm in ECAL for both hybrid (EB) and island (EE) clustering algorithms.	92
5.2	Superclustering in ECAL for hybrid clustering algorithm in barrel. . . .	92
5.3	Superclustering in ECAL for Island clustering algorithm in barrel. . . .	93
5.4	$Z \rightarrow e^+e^-$ mass plot showing resolution and energy scale that is obtained from applying energy scale corrections to account for intrinsic spread in crystal and photo-detector response and time-dependent corrections to compensate for channel response loss for EB (right) and EE (left)	97
5.5	Super-clusters showing resolution and energy scale that is obtained from applying energy scale corrections for EB (right) and EE (left)	97

5.6	Illustration of the differences between proton-proton collision muons, cosmic and halo muons. (a) Muons from collision propagating from the center and moving outwards in a well defined pattern, (b) Cosmic muons penetrating the detector and leaving signals in opposite hemispheres of the muon system, (c) Cosmic muons leaving signals in the tracker and opposite hemispheres, (d) cosmic muons entering and leaving the detector without passing through the muon detector layers, (e) beam halo muons penetrating the detector and leaving signals in the endcaps and (f) Cosmic muons entering the detector through the endcap and leaving through the barrel and which can happen in a vice-versa manner.	102
6.1	Trigger efficiency turn-on curves for photon p_T and $E_T^{\text{miss}} > 25$ GeV (left) and for E_T^{miss} with photon $pt > 80$ GeV/c (right).	116
6.2	Schematic diagram showing $\tilde{\chi}_1^0 \rightarrow \gamma + \tilde{G}$ decay topology within the ECAL volume of the CMS detector.	117
6.3	Pulse shape profile showing a spike and a real photon time from data. .	119
6.4	Timing distribution of photons with $p_T > 80$ GeV showing timing measurements using seed (left) and that using cluster time (right). Resolution from seed time is much better compared to that for cluster time. . . .	119
6.5	Timing distribution of photons with $p_T > 80$ GeV showing timing of data and MC γ + jets samples before (left) and after (right) smearing of MC is applied.	120
6.6	ECAL timing distribution of photons with $p_T > 80$ GeV from data showing contributions from main proton-proton collision in EB (left), EE (right) and all of ECAL combined (below). A 2.5 ns timing structure is clearly seen in EE compared to EB	120
6.7	ECAL time Vs η (left) and ECAL time Vs ϕ (right) for photons with $p_T > 80$ GeV from data. The lower plot show the photon timing distribution for events with different jet multiplicity.	122

6.8	ECAL time Vs η (left) and ECAL time Vs ϕ (right) and $CSC(Seg, \gamma)\Delta\phi$ for photons with $p_T > 80$ GeV from data. Halo photons show a clear matched between CSC segments and ECAL cluster in $\Delta\phi$ with their distribution peaking at $\phi = 0, \pm\pi$ and also the shape of their expected time.	125
6.9	Two dimensional plot showing $DT\Delta\phi(Seg, \gamma)$ against $DT\Delta\eta(Seg, \gamma)$ for photons with $p_T > 80$ GeV, ECAL Time > 2 ns and ECAL Time < -3 ns in proton-proton collision data (left) and non-proton-proton collision or cosmic data (right). Small $\Delta\eta$ and $\Delta\phi$ are cosmic-ray photon candidates.	126
6.10	Plot showing Number of crystals in photon super cluster for photons with $p_T > 80$ GeV, ECAL Time < -3 ns (blue), control region ($ t < 1.0$ ns (black) and spike control sample (red).	127
6.11	Figure showing E_T^{miss} distributions for events with out-of-time and in-time photons with $p_T > 80$ GeV. $E_T^{\text{miss}}_1$ and $E_T^{\text{miss}}_2$ definitions are given in context.	128
6.12	Diagrams showing background estimation technique.	131
6.13	ECAL time Vs η and ECAL time Vs ϕ (left) for photons from SinglePhoton dataset (left) similar plots from the DoubleElectron dataset (right). Photons at $\phi = 0, \pm\pi$ which are mostly halo photons not observed in the Z boson candidate sample.	133
6.14	Di-electron candidate mass distribution and the time of both electrons for the signal $71 < m_{\ell_1, \ell_2} < 100 \text{ GeV}/c^2$ Z boson sample(left) and similar distributions from the Control (outside signal region) sample (right). Candidates events from the DoubleElectron dataset.	133
6.15	<i>Top</i> : Control sample (left) and signal sample (right) of di-electron candidate mass distribution. <i>Bottom</i> : Figure showing definition of scale factor use in estimating the contributions from control sample in signal sample.	135
6.16	Timing distribution of genuine Z bosons after background contribution has been subtracted.	135
7.1	Sampling distributions for $f(t_\mu \mu)$ showing how one extracts the p -vlaues. left: is the using a analytic of the Asymptotic method and right: is from the HybridNew method.	143

7.2	Distribution of p -values showing how upper limit on μ is extracted for a given threshold probability.	145
8.1	Signal efficiency \times Acceptance for signal events passing our events selection for the SPS8(left) and GGM(right) models. The acceptance are photons with $t > 3$ ns.	147
8.2	Neutralino production cross section against proper delay length upper limit interpretation in SPS8 model.	148
8.3	Neutralino production cross section against proper delay length upper limit at 95% confidence levels interpretation in SPS8 model.	149
8.4	Neutralino two dimensional exclusion limit of neutralino mass (Λ) against proper delay length upper limit interpretation in SPS8 model in the decay $\tilde{\chi}_1^0 \rightarrow \gamma + \tilde{G}$ with limits from previous experiments shown.	150

Chapter 1

Introduction



We have performed a **model-independent** search for Neutral Massive Long-Lived Particles (NMLLP) decaying to photons using timing information. This analysis uses dataset recorded by the CMS detector from proton-proton (pp) collisions by the Large Hadron Collider (LHC) with a center of mass energy $\sqrt{S} = 8$ TeV.

Particles from NMLLP decay are considered **candidate particles for Dark Matter (DM)**. Matter observed around us make up only 4.5% of our total universe or multiverse. It is best described with unmatched precision by simple symmetries known as gauge symmetries. The mathematical formulation of gauge symmetries used in our understanding of the visible or baryonic universe is best implemented in the Standard Model (SM). The SM cannot describe non-visible or non-baryonic matter (*Dark Matter* (DM)) which make up the larger percentage of matter content in our universe. Although a direct detection of DM is yet to be presented, indirect detection experiments in Cosmology and Astrophysics support speculation that non-visible matter is made up of particles which may be very stable or have long lifetime collectively called *Long-Lived* (LL) particles. In general, LL particles are either charged (electromagnetically charged i.e interact with light (photons) or color charged) or neutral (cannot interact with light in the context of the SM).

Of particular interest to the scientific community are neutral LL particles, since DM is understood to not interact directly with light and could very weakly interact with visible matter. Recent negative search results is indicating that dark matter particles, if they

exists, could be very light i.e having very small mass of about a few eV to keV. These are known as Warm Dark Matter (WDM). DM particles could also be heavy with mass around GeV to TeV mass range called Cold Dark Matter (CDM). A common property is that they are stable.

The phenomenon of interest is a delayed photon produced in the decay of a meta-stable next-to-lightest supersymmetric particle (NLSP). The NLSP is the NMLLP. A classic example of a NLSP is the neutralino ($\tilde{\chi}_1^0$). It decays into a photon and the lightest supersymmetric particle (LSP) called the gravitino (\tilde{G}). In R-parity conserving (RPC) models, supersymmetric particles like the neutralino are produced in pairs at a collider. The neutralinos are also produced in a cascade decay of higher massive supersymmetric particles produced from pp collision. The gravitino being the LSP is stable, light in mass, neutral and weakly interacting with ordinary matter. This makes it a good candidate particle for DM. The photons from neutralino decay are energetic, isolated and delayed and can be detected using the electromagnetic calorimeter. High transverse momentum (p_T) spray of hadronic particles collectively called jets and missing transverse energy due to the weakly interacting nature of the gravitino as it leaves the detector undetected, accompany this decay. The photon arrival time at the electromagnetic calorimeter measured is delayed **due to the long lifetime of the neutralino**. This combination of jets, missing transverse momentum and delayed photon is a clear signal for a new kind of physics beyond the standard model (SM). An event with the decay of a neutralino, produced in the LHC pp collider would be recorded using the Compact Muon Solenoid (CMS) detector. The CMS detector is located at one of the beam crossing or collision points (also known as Interact**ing** Point (IP)) at Point 5 in Cessy, France. Relying on the excellent timing and energy resolution of Electromagnetic Calorimeter (ECAL) sub detector, of the CMS detector, we can distinguish between high energy photons from NMLLP decay and photons produced in interactions precisely and well described by the SM. Finding such a LL particle would address a lot of important questions in modern physics like: Why do we observe so much matter than anti-matter in our universe? Is there a reason why particles as currently observed in the SM have very different masses and can be classified to exists in 3 generations? What is the origin

and existence of Dark Matter (DM) and what is it made of? Do all fundamental forces behave the same at some higher energy scale? Answers to these questions will provide a clear understanding and direction towards studying physics beyond the standard model.

We have described in this thesis, a search analysis for delayed photons with results. We arrange this description according to the following chapters:

- Chapter 1 gives an introduction and general outline of this thesis.
- In Chapter 2, we give a brief description of the current standard model highlighting its strengths and weaknesses. We also described supersymmetric models beyond the standard model (BSM) showing the prediction of NMLLP. The physics of long-lived particle is also described. This chapter also presents compelling hints from theory, experiment as well as cosmological observation supporting the existence of NMLLPs which motivates our search. The phenomenology of NMLLP in gauge mediating supersymmetric models is used as a benchmark model in our search. Results from previous search analysis are also given.
- In Chapter 3, describes the experimental setup particularly the LHC collider and CMS detectors and also a detail description of the sub-detectors of the CMS which have been used in our search analysis.
- Timing reconstruction and calibration is described in chapter 4, detailing the method of extraction and calibration procedure used by CMS.
- In chapter 5, the reconstruction of physics objects such as superclusters, photons, jets and missing transverse energy E_T^{miss} according the CMS standards is described here. The presence of anomalous signals in the electromagnetic calorimeter is also mentioned.
- The search analysis is described here in chapter 6 detailing triggers used, dataset, our event selection and background estimation in our for neutral massive long-live particles presenting the result of the search. Sources of systematic considered in this experiment are also described.

- Chapter 7 presents the statistical analysis and methods providing clear meaning of p -values as used in this search analysis.
- Using the minimal gauge mediating supersymmetric model with snowmass signal point 8 (SPS8) as our benchmark model, chapter 8 provides an interpretation of our results in terms of exclusion regions reached by our analysis. Possible improvement for future analysis is briefly mentioned.
- In chapter 9, we present our conclusion from performing the search for delayed photons.

Chapter 2

Model and Phenomenology of Long-Lived Particles

2.1 The Standard Model of Particle Physics

The Standard Model (SM) provides a thorough and experimentally valid description of the fundamental constituents of all the matter and **it** interactions (except gravity) in our universe. Predictions by this model of elementary particle physics agrees with almost all of the available experimental data with unmatched precision. However, there are some theoretical and experimental difficulties with this model such as the existence of Dark Matter (DM) and neutrino masses which point to an extension of the SM to a much general model to which the SM is embedded within. An example of such a parent model could be based on the idea of supersymmetry.

This section briefly describes the SM **providing in summary its main building blocks, its strength as well as its limitations.**

2.1.1 Main Components

The mathematics used to formulate the SM is known as relativistic quantum field theory. Particles are represented as quantum fields and their dynamics and interaction is expressed using the Lagrangian formalism with a Lagrangian density \mathcal{L} describing its main building blocks being:

- *Fermions*: All of Matter around us can be described by fermion fields.
- *Interactions*: Matter fields interact with one another with these interactions mediated by vector bosons of a particular interaction based on a given symmetry.
- *Higgs Mechanism*: These matter fields are massless and obtain mass through their interaction with another field called the Higgs field. This process is known as Higgs mechanism. Massive matter fields are allowed to mix with each other to form new states of matter.

Fermions

Fundamental particles are characterized in terms of 3 quantities: Mass, Charge and Spin, where spin is a non-spatial or internal quantum number unlike mass and charge. The spin of a particle can be integer or half-integer. Fermions are ordinary matter, half-integer spin ($\frac{1}{2}\hbar$) particles distributed in a *Fermi-Dirac* distribution, meaning no two identical fermions can occupy the same state. They can be massive or massless, charged or be neutral. Anti-fermions have the same mass as fermions but have an opposite charge.

In the SM, the dynamics of fermions and their possible interactions is described by the *Dirac* equation given as:

$$\mathcal{L}(\bar{\Psi}, \Psi, G^\mu) = \bar{\Psi} (i\gamma^\mu \mathcal{D}_\mu - m) \Psi \quad (2.1)$$

Our current understanding favors the existence of two types of fermions: Leptons and Quarks. In the SM, these come in pairs of 3 generations but the SM itself provides **no reason for this classification**. Leptons participate in electromagnetic and weak interactions but not strong interactions while quarks can participate in all three interactions. In the SM, leptons have integer charge values and come in three families or flavors of pairs arranged in a certain mass hierarchy with the third generation being the most massive. The third generation flavor leptons can decay into the lower generation leptons through weak interactions. Each charged lepton of a particular flavor has its neutral pair partner known as its neutrino. For example, the pair partner of an electron

is the electron neutrino. In the SM, neutrinos are considered massless however, numerous experiments have confirmed neutrinos have a very tiny mass and can oscillate from one flavor into another under sufficiently large distances.

Quarks also come in pairs of three flavors with the most massive being the third generation flavor capable of decaying to its lower generation flavors through electro-weak interactions. Each pair flavor or generation of quarks consists of an “up-type” and a “down-type” quark. Quarks **have have** an electric charge as well as a color charge since they can participate in strong interactions. **Up-type** quarks like **up** (u), **charm** (c), **top** (t) have charge of $+\frac{2}{3}$ and **down-type** quarks such as **down** (d), **strange** (s), **bottom** (b) have a charge of $-\frac{1}{3}$. Charges are expressed in units of elementary charge e . Quarks make up the contents of less fundamental particles like the proton **use in** hadron collisions **and their distribution inside a proton is modeled using *parton distribution functions* (PDF).**

Electro-weak interactions allow for the distinction of fermions as being either “**Left**” or “**Right**” handed. From the SM, we understand that one generation of leptons –the electron and the electron neutrino (e, ν) and one generation of quarks –the up-quark and the down quark (u, d) is enough to describe all the visible matter **observe** around us.

Interactions

The interaction between matter (fermions) is mediated by force-mediating particles called *bosons*. Bosons have an integer spin ($n\hbar$, where n is an integer). The SM describes three different forces and their carriers. The electromagnetic force is described under the mathematical framework of *Quantum Electrodynamics* (QED), the force carrier is a massless boson called the *photon* (γ). The two nuclear forces; the weak force which was later in the 1960s developed in a combined electro-weak framework by Sidney Glashow, Abdus Salam and Steven Weinberg [7] have 3 massive vector bosons W^\pm , Z^0 discovered at CERN in 1983 as the force-mediator and the strong force described under the framework of *Quantum Chromodynamics* (QCD), not unified with the other two forces, is mediated by massless *gluon* (g).

At the heart of the formulation of the SM, is the concept of *symmetry* and *conserved quantum numbers*. Symmetry is the invariance of the dynamics (Lagrangian density,

\mathcal{L}), under a particular set of transformations. In the SM, these set of transformations are called local transformation or *gauge* transformations (symmetry groups) because they depend on space-time coordinates. The symmetries involved in the SM is:

$$SU(3)_C \otimes SU(2)_L \otimes U(1)_Y \quad (2.2)$$

$SU(3)_C$ is the symmetry group with quantum number *color* (C) charge for the strong interactions. There are 8 massless self interacting gluons. The color charge ensures there are three different color quarks for each quark flavor. Anti-quarks carry opposite color charges. The nature of this force ensures that quarks always exist in nature as bound states called *hadrons* consisting of 2 or 3 quarks such as protons although recent experiments have observed bound states consisting of 4 quarks.

$SU(2)_L \otimes U(1)_Y$ is the symmetry group with conserved quantum number *weak isospin* (T) for the electroweak interaction. There are four corresponding gauge massless bosons $W_\mu^{1,2,3}, B_\mu$ which later mix with each other to give the physical electroweak bosons; charged W^\pm and neutral Z^0 and γ . The W^\pm and Z^0 become massive after the Higgs mechanism. These bosons couple using the “charge” of the weak interaction called weak isospin T and the hypercharge Y to all matter. The W^\pm only interacts with left-handed fermions and right-handed anti fermions. This leads to a phenomenon called *parity* violation. The third component of the weak isospin, T_3 combines with the hyper charge Y , according to the relation:

$$Q = T_3 + \frac{Y}{2} \quad (2.3)$$

Left handed fermions have $T^3 = \pm \frac{1}{2}$ and are represented as *multiplets* or in this case of the SM, isospin *doublets*, while, right-handed fermions have $T^3 = 0$ and are isospin *singlets*. The doublet or singlet nature of a given fermion representation in the SM is summarized in table ??.

The $SU(2)_L \otimes U(1)_Y$ group is a combination of two symmetry groups with coupling strengths g and g' connected to the electric charge of each fermion as $e = g \sin \theta_w = g' \cos \theta_w$.

θ_w is the *Weinberg angle*, $\sin^2 \theta_w \approx 0.231$ measured from experiments. This angle allows for the rotation of the gauge bosons from their *weak eigen-* states to their

physically observed states.

$$W^{(\pm)\mu} = \frac{W_1^\mu \pm iW_2^\mu}{\sqrt{2}}, \quad \begin{pmatrix} A^\mu \\ Z^\mu \end{pmatrix} = \begin{pmatrix} \cos \theta_w & \sin \theta_w \\ -\sin \theta_w & \cos \theta_w \end{pmatrix} \begin{pmatrix} B^\mu \\ W_3^\mu \end{pmatrix} \quad (2.4)$$



Its also allows for the mixing of different quark flavors. There is no such mixing for leptons in the SM, however, recent neutrino experiments show mixing between different neutrino flavors or family. This mixing allows for the W^\pm to be able to change the flavor of a given quark, a typical interaction happening inside the core of the sun. The complete changing of quark flavors is described by the *Cabibbo-Kobayashi-Maskawa* (CKM) 3 by 3 matrix whose elements are all parameters measured from experiment and not predicted by the SM.

$$TABLE OF SM multiplets and their Quantum Numbers. \quad (2.5)$$

Quantum Chromodynamics and Parton Distribution Functions

Hadrons like protons and neutrons consists of quarks and massless gluons collectively referred to as *partons*. The strength of parton interaction is determined by the strong coupling constant (α_S). The value of α_S depends on the amount of momentum transferred, Q^2 , between the interacting partons. The distribution of partons inside a proton is described by a *Parton Distribution Functions* (PDF). PDF provides the probability of finding a parton with momentum fraction x of the total proton momentum p . PDFs are measured from electron-proton accelerator experiments such as HERA in Germany due to the difficulty of computations in Quantum Chromodynamics (QCD). Their measurement allows for the introduction of *uncertainty* in the use of PDFs in measuring other quantities such parton-parton collision cross sections. PDFs are expressed as a function of the fraction of the parton's momentum to the total proton momentum, $x = p/P$, and the momentum transfer Q^2 from a given electron or parton interacting with the parton, $f(x, Q^2)$. Figure 2.1 shows an example of the PDFs for a few quarks and gluons with momentum fractions x_q and x_g respectively a momentum transfer of Q^2 .

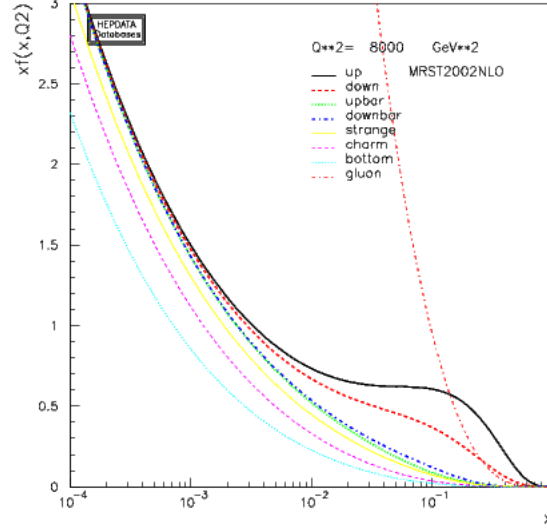


Figure 2.1: Pdfs.

It is imperative to know the uncertainty in the measurement of PDFs as this is usually a source of uncertainty in ~~a related experiment~~.

2.1.2 Spontaneous Symmetry Breaking

Early attempts prior to the 1960s to constructing a gauge theory of weak interactions failed because the gauge bosons were massless while experimental evidence proved otherwise.

In the Higgs-Brout-Englert [6] mechanism, consist of introducing an extra weak isospin complex scalar doublet ϕ .

The $SU(2)_L \otimes U(1)_Y$ symmetry is spontaneously broken into a $U(1)$ symmetry which describes the electromagnetic interaction. Figure 2.2 shows a picture of the potential of the spin 0 complex Higgs doublet field ϕ .

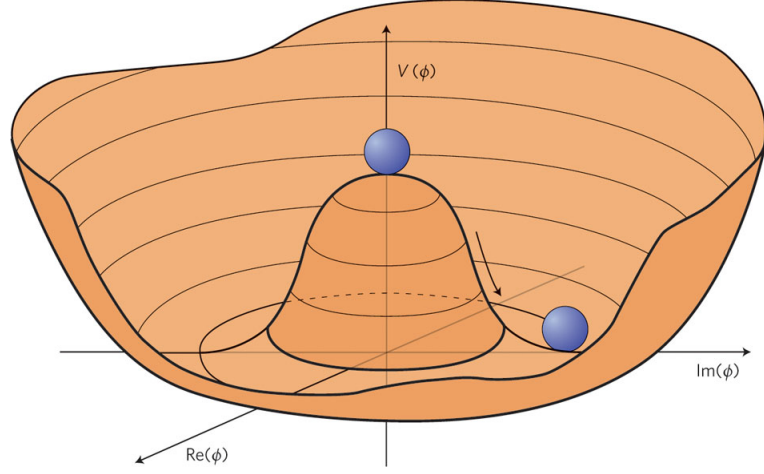


Figure 2.2: Higgs boson “Mexican hat” potential which leads to spontaneous symmetry breaking.

The choice of the minimum value of the potential,

$$|\phi_0| = \sqrt{\frac{-\mu^2}{\lambda}} = \nu$$

where μ and λ are parameters, is chosen to spontaneously break the $SU(2)_L \otimes U(1)_Y$ symmetry to $U(1)$ symmetry thus giving mass to matter and gauge fields (particles). This process is referred to as *Higgs-Brout-Englert mechanism* or *Higgs mechanism* for simplicity.

Matter particles such as quarks and leptons, get their mass, through their interaction with the Higgs field. The amount of their obtained mass, m_f , is proportional to the strength of their interaction or Yukawa coupling λ_f ~~with~~ with the Higgs field. Electro-weak interaction mediating gauge bosons, Z^0 and W^\pm obtain their masses m_Z and m_{W^\pm} respectively, by engulfing or “*eating*” the massless components (Nambu-Goldstone bosons) of the complex Higgs doublet. Out of the four scalar fields, only a physically massive Higgs boson remains.

$$m_f = \lambda_f \frac{\nu}{\sqrt{2}}, \quad \frac{m_{W^\pm}}{m_Z} = \frac{\frac{1}{2}\nu g}{\frac{1}{2}\nu \sqrt{g^2 + g'^2}} = \cos \theta_w \quad (2.6)$$

The search for the Higgs boson was the main purpose for the construction of the LHC at CERN. On ~~June~~ ^S 04, 2012, a Higgs-like candidate particle was found whose mass

$m_H = 125 \pm \text{GeV}/c$ which gives the parameter $\nu \approx 246 \text{ GeV}$.

Figure 2.3 shows a complete summary of all the particles in the SM and their interactions.

It is important to note that there is no fundamental reason given by the SM why there should be only one type of the Higgs field to which all fermions couple to obtain their mass, nor any prediction from the SM for the choice of parameters. As we will see in other models such as supersymmetry, that, there could be more than one Higgs field.

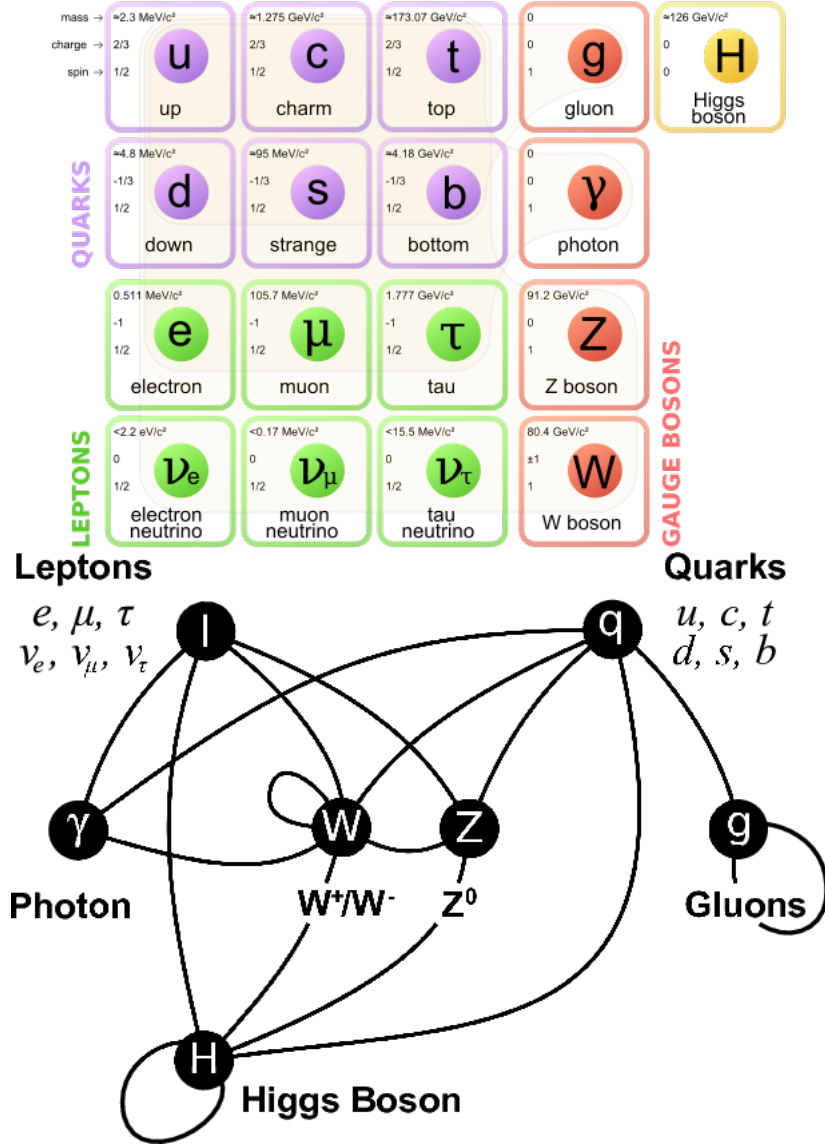


Figure 2.3: SM “particle periodic table” and their interactions through bosons as described in the SM.

2.1.3 Decay Rate and Life Time

Fundamental particles in the SM with the exception of the electron (e), **up** (u) and **down** (d) quarks, can transform from one particle or generation to the other, either through disintegration or oscillation. As a result most particles are understood to be

meta-stable. This stability is measured with respect to the age of the universe which is understood to be about 13.7 billion years old. Composite particles are even more unstable and easily decay into lower generation particles. The proton consisting of two u and one d valence quarks seemingly very stable can also disintegrate with some theories beyond the SM predicting the proton to remain stable in a time period of about 10^{33} years. Particle disintegration is understood as the interaction between a given particle and its daughter particle(s). This interaction could be electromagnetic, weak, strong, any pair or maybe all three types of interactions.

The stability of a particle is related to the conservation of a quantum number or conserved quantity such as energy, spin, angular momentum and charge. Other factors may also play an important role such as phase space (availability of lighter particles to decay into), violation of some property such as strangeness or the mediating particle usually a boson being very massive compared to the momentum transfer between the parent and the daughter particle.

The decay rate, Γ , can be computed in some theory and can also be measured from experiment. Its measurement provide access to the underlying type of interaction or mediating particles involved in the decay and thus can be used as a direct tool to search for other interactions beyond the current known ones. A single particle can decay into more than one type of particles. As long as the conditions for decay are set, a decay into any potential particle whose mass is less than the mass of the parent particle is possible. The decay into a particular set of particle(s) is known as a *decay channel*. Different decay channels can be quantified with respect to the overall *total decay rate* of the parent particle. This quantification is expressed as a *branching ratio (BR)*. Thus the BR gives quantitative estimate of the possibility of a parent particle decaying to specific daughter particles or through that channel.

The inverse of a decay rate is the *proper life time*, denoted as τ .

$$\tau = \frac{\hbar}{\Gamma} \quad (2.7)$$

A convenient way to express life time is in distance traveled. This distance traveled $c\tau$, where c is the speed of light in vacuum and τ is the proper life time is called the *decay length*. $c\tau$ is the decay length as measured in the frame with respect to the

center of mass of a moving particle. It is called the *proper decay length*. However, laboratory measurements are not done in the center of mass of the moving particle as this would be near impossible to do. So since the particle is moving with some speed with respect to a stationary laboratory, we have to take into consideration this difference in motion of frames (the moving particle frame and the stationary Laboratory frame) in understanding our measurement in the laboratory of the decay length of the moving particle and how this is translated into the true decay length of the particle. The effect we need to consider is known in special relativity as *time dilation* and the decay length is a measure of the distance between the position where the particle was produced to where it decayed. Our laboratory measurement is expressed as:

$$L = \beta\gamma c\tau \quad (2.8)$$

where $\beta = \frac{v}{c}$ with v being the speed of the moving particle and $\gamma = \frac{1}{\sqrt{1-\frac{v^2}{c^2}}}$ is the ratio accounting for time dilation effect and from this one can translate back into the true decay length of a particle $c\tau$. The decay length can also be expressed in terms of the momentum p of the particle where $p = \beta\gamma m$ and thus $L = \frac{p}{m}c\tau$. We see that the decay length of a particle is proportional to the momentum and inversely proportional to the mass of the particle.

The decay rate depends on quite a number of particle properties as we've mentioned earlier. As a result, the decay length for electromagnetic, can be very different to that of weak and strong interactions. The decay length of strong interactions having the shortest decay length due to the strong nature of the interaction leading high decay rates. This is followed by the weak and then electromagnetic interactions. There are some exceptions to this due to other factors playing a key role than interactions alone as we mentioned. The table and graph below show the mass and decay length of SM particles and their interaction.

TABLE SHOWING SM particle decay rates and interactions in vol as well as mass V_s decay length. (2.9)

In particle physics experiments it is very challenging to measure the life time or decay

length **or a particle** by measuring the time it travels from where it was produced to where it decayed. Rather, the number of events present initially and **that observed after a time period t** is used to measure the lifetime of a particle. The decay rate (or life time) of a particle is related to the number of particles observed though the equation:

$$N(t) = N_0 \exp\left(\frac{-t}{\tau}\right) = N_0 \exp\left(\frac{-\Gamma t}{\hbar}\right) \quad (2.10)$$

where $N(t)$ is the number of particles observed at an arbitrary time t and N_0 is the number of particles observed at an initial time where it is assumed no particle has decayed yet usually at $t = 0$. **A distribution of the observed number of particles(usually a Poisson distribution) can be plotted with time measured. The resulting distribution if fitted with a Poisson distribution faction and the parameter of the Poisson distribution function extracted to give us the decay rate or life time of a particle.**

Particles with large decay length or long life time are commonly referred to as *Long-Live* (LL) particles. Many models beyond the SM predict the existence of such particle. **They are also understood to be prime candidates for particles making up DM.** Before we dive into such models, it is necessary to understand in detail the decay of particles and factors which determine a particle's decay length as well as the kind of LL particles considered detectable in a multi-particle physics detector such as those at the Large Hadron Collider (LHC) CERN pursued in this thesis.

Particles described by the SM come in different types of long-lived. First, we have the stable elementary (as we currently believe) such as the electron and neutrinos. Second, we have the meta-stable elementary such as the muon and finally the (very) long-lived composite particles such as the neutrons and protons. By referring to the different classes of particles according to their life time, we can asked the question, what properties of a particle makes it stable, meta-stable or long-lived?

There are three possible answers to this question:

- A particle could be the lightest state carrying a conserved quantum number and as such remain entirely stable e.g the electron and proton.
- The decay of a particle to another lighter particle could only be made possible

through some **suppressed or effective coupling** and as a result ends up being meta-stable e.g the muon

- If the mass of a particle is relatively close in quantity to the particle it is decaying into such that their difference in mass is quite small, the decay will be ~~eventually~~ suppressed. This goes by the name "lack of phase space" for decay e.g decay of neutron ($n \rightarrow p + e^- + \bar{\nu}_e$). In this scenario the difference in mass between the neutron (n) and the proton (p) is ≈ 1.293 MeV and as a result **determines the type of associated particle produced in this decay as observed.**

In this thesis, we will only be interested in Meta-Stable (MS) particles, and in **particular focus on Massive Neutral Meta-Stable particles which we refer to as Neutral Massive Long-Lived Particles (NMLLP).** A rather more descriptive name would be Neutral Massive Meta-Stable Particles(NMMP) since these particles are not LL in the real sense but might decay into other elementary particles which are observable at detectors and how long-lived you refer to them depends on the possible lifetime of this particle your detector is sensitive to. As their lifetime might range from a few nanoseconds (10^{-9} s) to the age of the universe (13.7 billion years) likewise from a few μm to billions of km. Thus, our interpretation of LL particles will be those whose decay length range from a few μm to few meters or detectable size of the LHC detectors and in particular electromagnetic sector of LHC detectors.

We have restricted ourselves to electromagnetic (local U(1) gauge symmetry) neutral particles as their charge counterparts ~~can be~~ studied using ~~conventional~~ magnetic spectrometer and ionization methods as shown in **this** studies for Heavy Stable Charge Particles (HSCP).

2.1.4 Limitations Of the Standard Model

The SM (Glashow, 1961; Weinberg, 1967; Salam, 1968) currently describes almost entirely all of the observed phenomena and fundamental particles of nature with unmatched precision. However, as indicated from previous sections, there is more to be understood such as:

- **General Formalism:** There are way too many parameters in the SM which

are not derived within the theory but rather measured experimentally such as the Weinberg angle and the CKM matrix elements. The SM does not account for the multiplet structure of fields as well as why there are only three observed generations of particles. Currently observed neutrino masses are not required in the SM as neutrinos are considered massless in the SM. Even the Electroweak symmetry breaking is not very well understood.

- **Astrophysical:** Why is there so much matter and not anti-matter? If the Big Bang is suppose to be the correct theory about the universe existence, then matter and anti-matter must be observed in equal composition, however, Baryon asymmetry ratio show that there is more matter than anti-matter, where is all the expected anti-matter? This could be explained as charge-parity violation in weak interactions of the SM but this is quite small(not strong enough) to explain for all the observed discrepancy. Baryonic Acoustic Oscillation results also known as CMB as well as WMAP results all indicate the presence of excess matter which does not interact with light and is collision less called Dark Matter (DM) as well as increase energy density responsible for the rapid accelerated expansion of the universe called Dark Energy (DE). All these observations cannot be explained within the SM.
- **Theory:** The SM is seen as some low energy theory of some much deeper underlying theory(SM as an effective theory) due to the fact that it cannot describe gravity. In Addition to this, the coupling constants in the SM all vary with energy and so the definite question if whether there is a much higher scale where all these couplings become a single coupling(Unification of forces). Supersymmetric extensions of the SM show unification of forces at Grand Unified Theories (GUT) energy scale of $\approx 10^{15}$ GeV.
- **Mass Hierarchy and Fine-tuning:** Why are particles masses in the standard model arranged in such a hierarchy? from neutrino masses of few eV to top mass of 173 GeV. In terms of energy scale, from the electro-weak symmetry breaking scale of ≈ 100 GeV to Planck scale, $M_p = 10^{19}$ GeV, there are no other particles and especially scalars or any known interactions. The Higgs mass incredible precise contributions from higher order (loop) effects to its mass in order to maintained

its experimentally observed value of ≈ 125 GeV. Understanding this precise contributions and cancellations to arrived at the expected value is referred to as the fine-tuning problem and extensions of the SM like supersymmetry provide a very natural answer as we will see later.

2.2 Beyond Standard Model Physics

The Higgs boson's mass as understood from theoretical calculations should receive contributions from higher order (loop interactions as is known) in order to observe the experimentally measured mass of possibly 125 GeV or of order $O(100 \text{ GeV})$. However, all these corrections are said to cancel out such that the experimentally observed mass is as it is measured. The question of where these additional loop corrections disappeared into cannot be understood within the context of the SM. However, theories beyond the SM such as supersymmetry provide a natural understanding of how these loop effects cancel out to arrive at the observed mass. To go a step further with this; the mass of a particle can be expressed as

$$m_{physical}^2 = m_{Bare}^2 + \delta m_1^2 \quad (2.11)$$

where $m_{Physical}^2$ is the true mass of the particle measured in the laboratory; in the case of the Higgs boson of order $O(100 \text{ GeV})$ while m_{Bare}^2 is the true mass of the particle which cannot be calculated or measured. δm_1^2 is the quantum one loop corrections to the true mass which can be calculated. Thus from the measured mass and the calculated one loop quantum effect mass, one can get the true mass of the particle. The quantum loop contributions can come from bosons as well as fermions. For the Higgs scenario, the Higgs can couple or interact with every particle through interactions like $\lambda_f H \bar{f} f$ for fermions and $\lambda_S |H|^2 S^2$ for scalar or bosons with λ_f and λ_S the coupling constants and not necessarily equal. The quantum one loop corrections as calculated from the following diagrams:

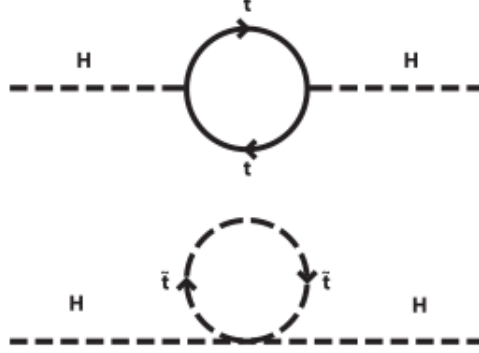


Figure 2.4: Higgs self energy diagrams showing how the higgs boson mass is computed from both Higgs field and supersymmetric partner particle contributions.

is given as:

$$\delta m_{1,f}^2 = \frac{1}{16\pi^2} |\lambda_f|^2 \left(-2\Lambda^2 + 6m_f^2 \ln \left(\frac{\Lambda}{m_f} \right) + \dots \right) \quad (2.12)$$

$$\delta m_{1,S}^2 = \frac{1}{16\pi^2} |\lambda_S|^2 \left(\Lambda^2 - 2m_S^2 \ln \left(\frac{\Lambda}{m_S} \right) + \dots \right) \quad (2.13)$$

where Λ is understood to be some cut-off scale where new kind of interaction at a much higher energy scale is needed to regulate the low energy behavior of the SM. Λ could be the Planck scale (10^{19} GeV) where this new kind of interaction is gravity. The two things to observe from this calculations is that first, corrections to the higgs bare mass m_{Bare}^2 are not proportional to the Higgs mass as is the case with other SM particles like the electron. Second, these corrections are of the order of $\approx 10^{38}$ GeV² with the signs reverse for fermions or scalar corrections. Despite this expected corrections to the true Higgs mass, the measured or physical Higgs mass squared $m_{H,Physical}^2$ is of the order $\approx 10^4$ GeV². The only way there is an agreement between this correction and physical Higgs mass is if the true Higgs mass, $m_{H,Bare}^2$ is fine-tuned with a precision of about 1 in 10^{17} . This enormous fine-tuning is considered a fundamental problem with the Higgs mechanism of the SM and is referred to as not *natural*. Infact, since in the SM, there is only once scalar particle which is the Higgs boson, this fine-tuning cannot be understood. However, in other models beyond the SM such as supersymmetry, the

fermion one loop quantum correction which comes with an opposite sign to the scalar one loop quantum correction with $\lambda_f = \lambda_S$, there is a cancellation and this fine-tuning can be understood. Another interpretation of this issue is through the question of why there is so much difference in energy scale between the electroweak scale $O(100 \text{ GeV})$ and the Planck energy scale $O(10^{19} \text{ GeV})$ where gravity effects to particle interaction becomes significant?. This is referred to as the *Hierarchy problem* stated above as one of the motivation to go beyond SM.

Another drawback with the SM is that of unification. It is believed that at a higher energy scale since SM seems to be describing the low energy behavior of some parent theory, the electromagnetic, weak and strong interactions all become one interaction just as the electromagnetic and weak interaction unified into the electro-weak interaction as the electro-weak energy scale $\approx 100 \text{ GeV}$. i.e

$$SU(3)_C \otimes SU(2)_L \otimes U(1)_Y \subset \mathcal{G} \quad (2.14)$$

where \mathcal{G} is some larger symmetry group. In SM, this does not happen at any higher energy scale. However, in supersymmetry, there is a clear unification of individual coupling constants or electromagnetic, weak and strong interactions at the GUT energy scale of $\approx 10^{15} \text{ GeV}$. This effect can be seen in the following figures as taken from [].

$$\textit{Plot showing Unification of Coupling Constants.} \quad (2.15)$$

2.2.1 Supersymmetry

In relativistic Quantum Field Theory (QFT), the idea of symmetry is used to provide a better understanding of a particle and its possible interaction with other particles. Symmetries can be classified into two broad categories: Space-Time or external symmetries known as Poincaré (rotational and translational) symmetries also known as groups and internal or gauge (which as we saw earlier; $SU(3)_C \otimes SU(2)_L \otimes U(1)_Y$ describing the quantum numbers— color, weak and hypercharge respectively) symmetries. In a quest to include gravitational interaction along with all the other forces of nature into a unique frame work called unification, it was thought that one could combine these two classes

of symmetries into a bigger class of symmetry. However, Coleman and Mandula in their so-called “*no-go*” theorem in 1967 [8] showed that the direct product nature of super groups; a direct approach to extending symmetries to bigger symmetries, is not possible. Thus these two class of symmetries cannot be combined into a bigger parent symmetry. This produced a challenge of finding a scenario where $[M^{\mu\nu}, T^a] \neq 0$, considering that the generators of these groups; P^μ , $M^{\mu\nu}$ and T^a corresponding to these symmetries have a direct product; Poincaré \times Gauge group, for which $[P^\mu, T^a] = [M^{\mu\nu}, T^a] = 0$. Because of this *no-go* theorem, such a parent symmetry group is not possible if one used generators of Lorentz tensors. The only way out is to find a symmetry which is generated by spinorial (particle’s spin) charges instead of tensorial (space-time) charges. Such a theorem was found in 1975 by Haag, Lapuszanski and Sohnius [?] called the *Haag-Lapuszanski-Sohnius* theorem with its corresponding algebra called the *Lie-superalgebra*. The generators of these *Lie-superalgebra*, Q^α , with $\alpha = 1, \dots, N$. N is the number of generators for the supersymmetry. In this Thesis we will only be considering the case where $N = 1$ i.e only one supersymmetry generator. This is known as the minimal supersymmetry. If you are interested in learning more about extensions of this minimal version see this excellent text [?]. The generators Q^α can be expressed in terms of Weyl Spinors Q_a where $a = 1, 2$ and must satisfy anti-commutation relation with its conjugate. This is the major aspect of introducing supersymmetry. Thus for these generators to be supersymmetric, they must satisfy the following relations:

$$\{Q_a, \bar{Q}_b\} = 2(\gamma^\mu)_{ab} P^\mu, \quad [Q_a, P^\mu] = 0, \quad [Q_a, M^{\mu\nu}] = \frac{1}{2}(\Sigma^{\mu\nu})_a^b Q_b \quad (2.16)$$

where γ^μ is define such that $\{\gamma^\mu, \gamma^\nu\} = 2g^{\mu\nu}$ and $\Sigma^{\mu\nu} = \frac{i}{2}[\gamma^\mu, \gamma^\nu]$ and \bar{Q}_a is the Hermitian conjugate to Q and is also a generator. From these relations, two very fundamental consequences arise:

- Particles in a given supermultiplet have the same mass but differ in their spin by half a unit.
- There is an equal number of fermionic and bosonic degree of freedom or states in every irreducible representation of supersymmetry.

From the above observation, we define supersymmetry as the symmetry which transforms particles from one spin into another. Hence these supersymmetry generators transform fermions into bosons or bosons into fermions with the same mass.

$$Q|\mathbf{Fermion}\rangle = |\mathbf{Boson}\rangle, \quad Q|\mathbf{Boson}\rangle = |\mathbf{Fermion}\rangle \quad (2.17)$$

This reveals to us that, in supersymmetry, every particle has a partner with the same mass with supermultiplets chosen such that every SM particle with spin $0, \frac{1}{2}, 1, 2$ have a partner with spin $\frac{1}{2}, 0, \frac{1}{2}, \frac{3}{2}$ respectively in the same supermultiplets. In supersymmetry, there are three kinds of supermultiplets referred to as *Chiral*, *Vector* and *gravity* multiplets. In building a minimal supersymmetric extension of SM only the Chiral and Vector supermultiplets are used and thus we will concentrate only on these two. The following table shows the different supermultiplets as encountered in supersymmetry.

TABLE of SUSY Multipletes and spin (2.18)

So far we have mostly seen the algebraic approach to supersymmetry. However in order to build models(supersymmetric Lagrangians) and make predictions, we have to return to the idea of fields. This brings us to the idea of the so-called *superfields* first proposed and realized by my hero Abdus Salam and Strathdee [?]. Superfields are fields defined on a superspace which is an ordinary MinKowski-space, x^μ and four anti-commuting Grassmann, numbers θ . For more on this see []. Thus a general superfield is an operator-valued function Φ on a superspace $\Phi(x^\mu, \theta, \bar{\theta})$ Its components consist of from ordinary scalar fields (real or complex), Lorentz vector fields and Left-handed or Right-Handed Weyl(2 degrees of freedom) spinor fields. The Table below shows an example of the components which make up the superfield or supermultiplets. Each component represents a SM particle and its super partner with the same mass.

TABLE OF A SUPERFIELD AND COMPONENTS (2.19)

In constructing the MSSM, only the Chiral and Vector supermultiplets or superfields

are used.

2.2.2 Minimal Supersymmetric Standard Model

The Minimal Supersymmetric Standard Model or MSSM is an extension of the SM to include its supersymmetric partner particles.

As a result the already 19 free parameters of the SM is increased with an additional 105 free parameters. These are a lot of parameters for any fundamental theory describing elementary particle interactions and thus undermines its predictive power. Thus, a generic or parent theory must be preferred in which the number of free parameters for the theory to predict is much reduced. Through this way, it is much easier to study the theory phenomenologically. For example, the supersymmetric (SUSY) extension of the SM theory with gravity is called mSUGRA and has only 6 parameters while a SUSY extension of SM with purely gauge interaction is called Gauge Mediated Supersymmetry Breaking (GMSB) and has only 5 parameters. Other SUSY theories such as Anomalous Supersymmetry Breaking (ASB) have 6 parameters.

In this thesis, we will only talked about GMSB theories as these describe and predict the existence of LL particles. A full table of SM particles and their SUSY counterparts as understood through the MSSM can be seen in the table below.

TABLE OF SM/SUSY PARTICLES IN MSSM. (2.20)

In SUSY, the particle dynamics and interaction Lagrangian is written in terms of supermultiplets are already seen in table above. However, the superpotential is what defines the phenomenology and hence mass spectrum of the model. In MSSM, the superpotential is given as thus[?]:

$$W_{\text{MSSM}} = \bar{u}\mathbf{y}_u\mathcal{Q}H_u - \bar{d}\mathbf{y}_d\mathcal{Q}H_d - \bar{e}\mathbf{y}_e\mathcal{L}H_d - \mu H_d H_u \quad (2.21)$$

Where the objects H_u , H_d , \mathcal{Q} , \mathcal{L} , \bar{u} , \bar{d} , \bar{e} are chiral superfields corresponding to the chiral supermultiplets given in table above. The dimensionless couplings $\mathbf{y}_u, \mathbf{y}_d, \mathbf{y}_e$ are 3×3 matrices known as the Yukawa couplings. One thing to note here is that instead of a single Higgs double as in the case with the SM, two Higgs double are present; H_u and

H_d to give mass to up-type and down-type quarks and leptons. The superpartners of these Higgs particles which are fermions and those of the gauge bosons called gauginos mix to produce new neutral and charged fermions called Neutralinos and Charginos respectively. An aspect which is not present in the SM.

In MSSM, a combination of baryon (B) and lepton number (L) symmetries is combined to give a more fundamental symmetry called *R-Parity*[?] or *Matter Parity*[?]. From a multiplicative combination of B and L numbers, we get R-parity expressed as:

$$R_P = (-1)^{3(B-L)+2S} \quad (2.22)$$

where S is the spin of the particle. This is a conserved quantum number which is from a discrete symmetry. This symmetry commutes with supersymmetry. Thus particles in a given supermultiplet do not have the same R parity. SM particles like quarks have an Even parity $R_P = 1$ while SUSY particles like squarks have Odd parity $R_P = -1$. This has a lot of important phenomenological consequences such as the following:

- In the decay of SUSY particles, the lightest SUSY particle (LSP) have odd parity $R_P = -1$ and thus it must be absolutely stable. In addition to its absolute stability, if it is neutral and interacts only and very weakly with ordinary matter, then it is a good candidate for non-baryonic dark matter as required by cosmology.[?]
- Every SUSY particle produced which is not the LSP, will eventually decay into the LSP or an odd number of LSPs.
- SUSY particles can only be produced in pairs in a collider experiment.

Thus in generic SUSY models with minimal particle content, where the superpotential include terms which violate Lepton (L) and baryon (B) numbers; R-parity conservation can be imposed giving rise to R-parity Conserving (RPC) models with the LSP stable while R-parity can be violated resulting to R-parity Violating (RPV) models where the LSP is unstable and decays to SM particles.

In this Thesis, we consider only RPC models since our motivation is to search for neutral stable particles motivated by them being candidates particles which make up dark matter (DM).

If SUSY is to become a theory which describes nature, then the observation of components within the same supermultiplets having the same mass, i.e $m_B = m_F$ must be unrealistic as until presently no experiment has found a selectron(SUSY partner of electron) with a mass of 0.512 MeV for example. Therefore, SUSY must be spontaneously broken. Spontaneous Supersymmetry Breaking (SSB) means that the vacuum expectation value of a scalar field(in SUSY an auxiliary field) must be non-zero. The manner in which this breaking occurs determines the phenomenology of any model. As one would imagine, there are many different ways of breaking SUSY(through gravity, gauge etc) resulting to many different SUSY models. However, in this thesis, will concentrate only on those for which gauge interactions is responsible for SUSY breaking. Such SUSY models are generally referred to as Gauge Mediated SUSY breaking models (GMSB) ranging from Pure to General Gauge Mediation (GGM) as one would easily find in the literature. We will also focus on models with Soft Breaking as we would like favour SUSY models with phenomenology within the reach of the LHC. Soft breaking would mean the SUSY breaking terms in the SUSY potential consists of only masses and terms whose couplings have positive mass dimension. This ensures the existence of sparticles with masses around a few TeV where they can possibly be produced at current particle colliders.

2.2.3 Soft Supersymmetry Breaking

The idea of being soft is such that the spontaneous breaking must be caused by couplings with positive mass dimension and not dimensionless coupling. Also this allows for the observed hierarchy between the electroweak energy scale 100 GeV and the Planck energy scale 10^{19} GeV. The Lagrangian for soft SUSY breaking terms can be written as thus:

$$\mathcal{L}_{\text{Soft}}^{\text{MSSM}} = -\frac{1}{2} \left(M_3 \tilde{g} \tilde{g} + M_2 \tilde{W} \tilde{W} + M_1 \tilde{B} \tilde{B} \right) + c.c \quad (2.23)$$

$$- m_{H_u}^2 H_u^* H_u - m_{H_d}^2 H_d^* H_d - (b H_u H_d + c.c.) \quad (2.24)$$

$$- m_{\tilde{Q}}^2 \tilde{Q}^+ \tilde{Q} - m_{\tilde{L}}^2 \tilde{L}^+ \tilde{L} - m_{\tilde{u}}^2 \tilde{u}^+ \tilde{u} - m_{\tilde{d}}^2 \tilde{d}^+ \tilde{d} - m_{\tilde{e}}^2 \tilde{e}^+ \tilde{e} \quad (2.25)$$

$$- \left(a_u \tilde{u} \tilde{Q} H_u - a_d \tilde{d} \tilde{Q} H_d - a_e \tilde{e} \tilde{L} H_d + c.c \right) \quad (2.26)$$

where M_1 , M_2 and M_3 are the superpartners of the gauge bosons of the SM symmetry

group. They are referred to as the Bino, Wino and gluinos(8 gluinos because there are 8 gluons in the SM). I have intentionally omitted scalar mass terms in this Lagrangian and rather included SUSY breaking contribution to the Higgs potential.

This results in the following mass spectrum for particles in the MSSM as seen in the figure below:

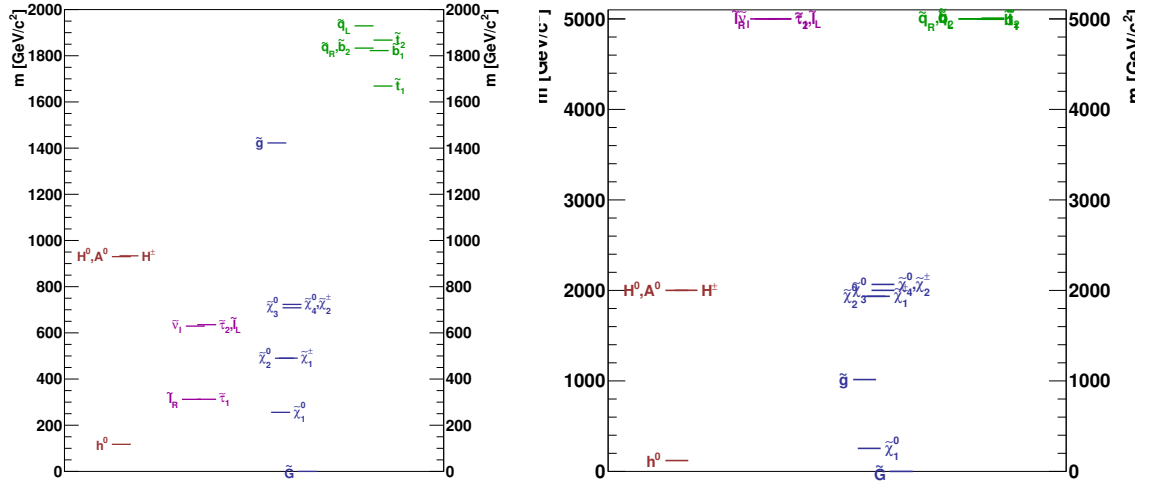


Figure 2.5: SUSY Mass spectra in the mGMSB SPS8 model (left) and GGM Model (right) with mass of gluino ($M_{\tilde{g}} = 1.0 \text{ TeV}$)

In summary, SUSY predicts in addition to SM particles, new particles whose spin (S) differ from their SM counterparts by half-integer. Bosons (fermions) in the SM have superpartners which are fermions (bosons) respectively. The superpartners of SM fermions are scalars comprising of sfermions (\tilde{l}), sneutrinos ($\tilde{\nu}$) and squarks (\tilde{q}) while gluinos (\tilde{g}) being fermions are the superpartners of the massless gauge bosons of strong interaction, gluons. The scalar Higgs (2 needed for obvious reasons) boson and the vector gauge bosons of electro-weak interaction have fermionic superpartners called higgsinos, winos and binos. These can mix to form a pair of mass eigenstates called charginos ($\tilde{\chi}_j^\pm, j = 1, 2$) and a quartet mass eigenstates neutralinos ($\tilde{\chi}_i^0, i = 1, \dots, 4$).

2.2.4 Gauge Mediated Supersymmetry Breaking and Phenomenology

General Model

Constructing a SUSY model requires the following items:

1. A gauge group describing the nature of particle interaction.
2. Specifying a superpotential.
3. Providing the method for SUSY breaking.

In the case of our Minimal Supersymmetric Standard Model (MSSM), we have already provided these as seen in equations 2.4, 2.38 and 2.40-2.43. These provide the particle content and interactions we need. However, we must account for the interactions needed in the mediation of SUSY breaking down to the energy scale of the MSSM or soft terms to be observed at the LHC.

SUSY breaking is realised through the existence of a *Hidden Sector* (Hidden because it couples only indirectly and very weakly to our “observable sector” of SM particles and their superpartners) whose dynamics manages to break SUSY. The nature of this breaking is not relevant for phenomenology but rather the “mediators” which communicate the effects of this breaking to the super partners of the SM particles. Thus, these mediators or agents must couple to this “Hidden Sector” as well as the “observable sector”. In gauge mediating SUSY breaking (GMSB), these agents have the usual SM gauge interactions and are called *Messenger fields*. These Messenger fields through loops (instead of normal tree level interaction) couple with the SM superpartners. As a result MSSM particle (gauginos and sfermions) get SUSY breaking masses at the loop level referred to as soft terms. The mass of this Messenger fields M_{mess} defines the energy scale of SUSY breaking. If $M_{\text{mess}} \ll M_{\text{Pl}}$ then induced SUSY breaking occurs at a much lower energy scale instead at the Planck energy scale where gravitational interactions become very significant and the effects of the breaking is first felt by these Messenger fields and later communicated to the observable sector through SM gauge interactions. In terms of energy scales, the picture is such that SSB happens at an energy scale \mathbf{F} which defines the mass of a gravity mediating superpartner particle, the gravitino as

$$m_{3/2} = \frac{\mathbf{F}}{\sqrt{3}M_{\text{Pl}}} \quad (2.27)$$

where $M_{\text{Pl}} = 1.3 \times 10^{19}$ GeV, then the energy scale \mathbf{F}_S , which is the *induced SUSY breaking* scale in the hidden sector. This along with the mass of the messenger particles

M_{mess} , defines the masses of the gauginos and sfermions of the MSSB or visible sector. If $\mathbf{F}_S < \mathbf{F}$ then the interaction between the hidden sector and the fundamental SUSY breaking is weak interaction other wise $\mathbf{F}_S \approx \mathbf{F}$ and the interaction is strong. The consequences of this is that one would no longer expect the mass of the gravitino $m_{3/2}$ to be given as in equation (2.45) but rather suppressed by $\frac{M_{\text{mess}}}{M_{Pl}}$ in GMSB models. In the mass spectrum of these models, the gravitino mass can be varied to a very small value only bounded by cosmological results, thus making it the lightest SUSY particle (LSP). Spanning the gravitino mass is expressed as a parameter c_{grav} which directly determines the lifetime of the SUSY next to lightest sparticle decaying to the gravitino. We will see more of this ahead.

In General GMSB, a simple model of the messenger sector can be chosen to consists of chiral supermultiplets of leptons and quark with the same quantum numbers $SU(3)_C \times SU(2)_L \times U(1)_Y$ as the SM gauge groups with representations given as:

$$\tilde{\ell} \sim (1, 2, 1) \quad \tilde{\ell}' \sim (1, 2^*, -1) \quad (2.28)$$

$$\tilde{q} \sim (3, 1, -\frac{2}{3}) \quad \tilde{q}' \sim (3^*, 1, \frac{2}{3}) \quad (2.29)$$

They couple to each other via a superpotential of a gauge singlet chiral supermultiplet S with an F-term as in the O'RAIFEARTAIGH model[?]. This messenger superpotential can be given as:

$$W_{\text{mess}} = \lambda_{\ell} S \tilde{q} \tilde{q}' + \lambda_q S \tilde{\ell} \tilde{\ell}' \quad (2.30)$$

We can thus obtain SUSY breaking by allowing vacuum expectation values VEV for both S and its auxiliary components F-term as $\langle S \rangle$ and $\langle \mathbf{F}_S \rangle$ where the \mathbf{F}_S does not have to coincided with \mathbf{F} as mentioned earlier and can be parametrised as:

$$\mathbf{F} = C_{grav} \cdot \mathbf{F}_S \quad (2.31)$$

This equation indicates that the non-zero VEV for the F-term is the main cause of fundamental SUSY breaking and this breaking is transferred to the messenger particles through radiative interactions as C_{grav} is a dimensionless parameter. Diagonalising the leptons and fermions masses with their scalar superpartners leads to mass terms for

messenger particles as:

$$m_{\tilde{\ell}\tilde{\ell}'}^2 = |\lambda_\ell \langle S \rangle|^2, \quad m_{\tilde{\ell} \text{ scalars}}^2 = |\lambda_\ell \langle S \rangle|^2 \pm |\lambda_\ell \langle F_S \rangle| \quad (2.32)$$

$$m_{\tilde{q}\tilde{q}'}^2 = |\lambda_q \langle S \rangle|^2, \quad m_{\tilde{q} \text{ scalars}}^2 = |\lambda_q \langle S \rangle|^2 \pm |\lambda_q \langle F_S \rangle| \quad (2.33)$$

From these messenger particle mass terms, we can defined a general scale for which these definations are can be trusted as:

$$M_{\text{mess}} = (\lambda_q, \lambda_\ell) \langle S \rangle \quad (2.34)$$

In pure gauge mediated SUSY breaking models (PGGM) it is required that $\lambda_q \neq \lambda_\ell$ [?] whereas in general gauge mediated (GGM) and minimal gauge mediated SUSY breaking (GMBS)[?] models, $\lambda_q \simeq \lambda_\ell \simeq \lambda$ and $M_{\text{mess}} = \lambda \langle S \rangle$

In the MSSM sector, the gauginos and scalars obtained their mass through 1-loop and 2-loop level corrections respectively. A simple diagrams for these corrections can be seen in figure .

In the minimal GBSM scenario, an additional parameter N_5 specitying the number of messenger vector-like supermultiplets transforming under $SU(5)$ so as to allow for unification of gauge couplings at the GUT energy scale($M_{GUT} \approx 10^{16}$ GeV. N_5 may not be too large so as to avoid gauge couplings diverging before GUT scale and as such the masses of MSSM gauge and scalars particles can be writen as:

$$M_a = \frac{\alpha_a}{4\pi} N_5 \Lambda \quad (2.35)$$

$$\phi_i^2 = 2\Lambda^2 N_5 \sum_{a=1}^3 C_a(i) \left(\frac{\alpha_a}{4\pi}\right)^2 \quad (2.36)$$

where $C_a(i)$ are some constants of $O(1)$, α_a are coupling constants and

$$\Lambda = \frac{F_S}{\lambda \langle S \rangle} = \frac{F_S}{M_{\text{mess}}} \quad (2.37)$$

Ofcourse for PGGM models we will have two seperate λ s defined as:

$$\Lambda_G = \frac{F_S}{\lambda_q \langle S \rangle} \quad (2.38)$$

$$\Lambda_S = \frac{F_S}{\lambda_\ell \langle S \rangle} \quad (2.39)$$

For complete description of PGGM and its parameters see [?]. Λ is called *the effective SUSY breaking scale* which defines the mass spectrum of MSSM gauginos and scalars. In GGM models [11, 12, 13], the mass of the gauginos $\mathbf{M}_a, a = 1, 2, 3$ defines the parameter space for these models. From equation we redefine the fundamental SUSY breaking scale F in terms of the effective SUSY breaking scale Λ as:

$$\mathbf{F} = C_{grav} \cdot \Lambda \cdot \mathbf{M}_{mess} \quad (2.40)$$

and hence from equation gravitino mass is re-written as:

$$m_{\tilde{G}} = C_{grav} \cdot \frac{\Lambda \mathbf{M}_{mess}}{\sqrt{3} \mathbf{M}_{pl}} \quad (2.41)$$

We observe from above equation that in GMSB models the gravitino can become very light compared to gravity mediating models. It should not be surprising when the gravitino is identified as least stable SUSY particle (LSP) in GMSB models and is seen as a candidate for DM. C_{grav} as mentioned earlier determines the decay rate and hence the lifetime of a SUSY particle decaying to the gravitino. For instantaneous decay $C_{grave} \approx 1$ whereas for non-instantaneous decays C_{grav} can be varied accordingly to achieve of-the- order-of a particle detector size lifetimes.

GMSB Phenomenology

GMSB models attract a lot of interests because of the existence of first, light gravitinos—unlike gravity mediating SUSY models which require the gravitino mass to be of the order of $\simeq 100$ GeV and as such does not play any role in collider phenomenology, GMSB models allow for gravitino masses as low as a few eV. Second, unique gravitino-scalar-chiral fermion and gravitino-gaugino-gauge boson interactions. Figure shows feynman diagrams for these interaction.

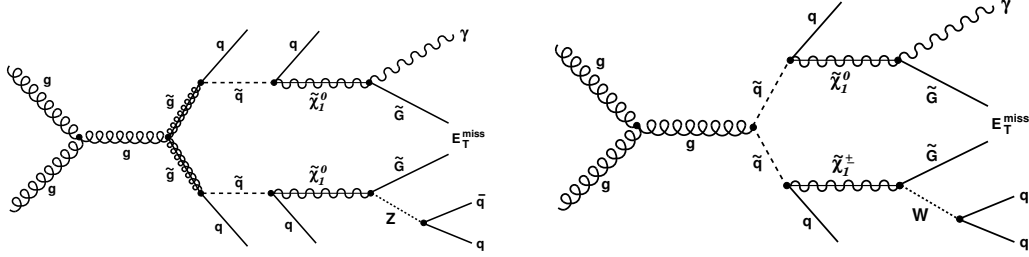


Figure 2.6: Feynman diagrams of of Gravitino interactions with superpartner pairs (ψ, ϕ) (a) and (λ, A) (b).

The presence of light gravitinos allows for the decay of any next-to-lightest MSSM particle to a gravitino as $\tilde{p} \rightarrow p\tilde{G}$ with the decay rate depending on the mass of the gravitino $m_{\tilde{G}}$ as long as R-parity is conserved. Thus every MSSM particle decay will eventually include a gravitino in its final states. This decay rate can be parametrised by C_{grav} . It is easy to see that $C_{grav} \geq 1$. Thus we define the parameter space for GMSB models as follows:

- For minimal GMSB (mGMSB): the parameter space is

$$\{\Lambda, \quad M_{\text{mess}}, \quad N_5, \quad \tan \beta, \quad \text{sgn}(\mu), \quad C_{grav}\} \quad (2.42)$$

- For General Gauge Mediation SUSY breaking (GGM): The parameter space to scan is [9, 10]

$$\{\mathbf{M}_3(\text{gluino mass}), \quad \mathbf{M}_2(\text{Wino mass}), \quad \mathbf{M}_1(\text{Bino mass}), \quad \tan \beta, \quad \text{sgn}(\mu), \quad c\tau_{NLSP}\} \quad (2.43)$$

The advantage with GGM models is that colored sparticles are not required to be heavier than their electroweak sparticles allowing for greater discovery potential at hadron collider[14]

- For Pure General Gauge Mediation SUSY breaking (PGGM): The parameter space to scan is:

$$\{\Lambda_G, \quad \Lambda_S, \quad M_{\text{mess}}\} \quad (2.44)$$

In these models, the Next-To-Lightest SUSY particle (NLSP) decays to the lightest SUSY particle (LSP), the gravitino and its SM partner. if \tilde{p} is the NLSP, then it will

decay is as follows:

$$\tilde{p} \rightarrow p + \tilde{G} \quad (2.45)$$

In mGMSB models \tilde{p} is the lightest neutralino (neutralinos come in four types and they are a mixture of Bino (\tilde{B}°), Wino (\tilde{W}°), higgsino ($\tilde{H}_u^\circ, \tilde{H}_d^\circ$) depending on the choice of parameters M_1, M_2, M_3 , or Λ , $\tan \beta$, and $sgn(\mu)$. and particle p is the photon (γ), SM(or new) Z boson (Z)(or Z') and the higgs (h). In this thesis, we will only focus on the parameter space for which the the particle $p = \gamma$ and $C_{grav} > 1$. This ensures that with the lifetime of our NLSP being finite, its decay happens *within the detector volume* and the resulting photon is delayed or non-prompt on detector time scales.

The decay rate for an NLSP to its SM partner and a gravitino goes like(details can be found in[7, ?]):

$$\Gamma(N\tilde{LSP} \rightarrow \gamma\tilde{G}) \approx \frac{m_{NLSP}^5}{\mathbf{F}^4} \quad (2.46)$$

This approximation is almost the same for the non-minimal GMSB models except that we add other parameters showing explicit dependence of how the neutralino life time can be made as long as expected in collider detectors. It is important to observe here that, the decay rate is large for smaller values of fundamental SUSY breaking scale or equivalently smaller gravitino mass provided the neutralino mass is kept fixed. Thus if m_{NLSP} is of the O(100 GeV) or more and $\mathbf{F} \ll 1000$ TeV, meaning $m_{\tilde{G}} \leq 1$ KeV, then the above decay rate is of the order than can be observed at hadron collider detectors.

2.3 Long-Lived Particles in GMSB Models

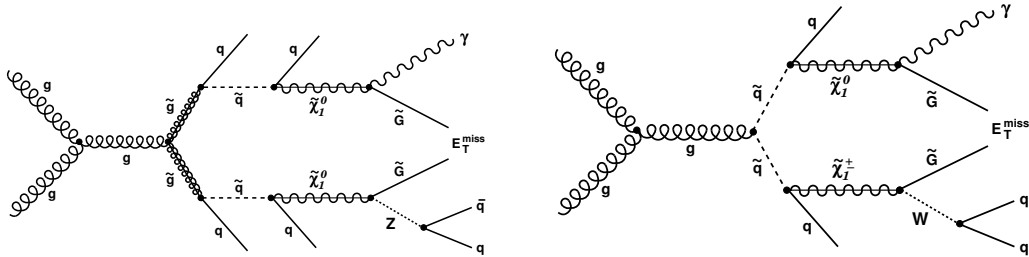
We have in previous discussions mentioned in passing some reasons why the study of LL particles is very important for uncovering new physics beyond the SM. In addition to mass, charge and spin being experimental handles for the search of new physics, a particle's life time or decay length is indispensable as importable parameters related to the underlying interaction type of the decay can be extracted from the decay rate and hence provide direct window towards new physics interactions beyond the SM.

2.3.1 Production of supersymmetric particles at Hadron Colliders

Production of a supersymmetric (SUSY) particle at a hadron collider is a probabilistic process. This process can be quantitatively expressed as the production cross-section. The production cross-section is basically the sum of all the different probabilities of a given process happening. This probability depends on the distribution of type of incoming particles inside the colliding protons with enough proton energy fraction to create a supersymmetric particle. Since the mass of the supersymmetric particle being create depends on the available center of mass energy of the colliding partons. In a hadron collider like the LHC, gluons are the one of these partons with the highest probability of proton energy fraction distribution. Sea quarks such as up and down quarks which make up the proton as well as valence quark which can exist inside the proton provide their energy fraction is large enough can also collide to create SUSY particles. The figure (2.3.1) show the parton distribution function against its energy fraction on the horizontal axis within a proton. The probability of creating a SUSY particle depends on the momentum transfer Q of these partons.

Figure 2.7: Parton distribution function (PDF) for partons against energy fraction on the horizontal axis for a particular momentum transfer Q value.

In order to be able to make comparisons between theoretical expectations and experimental observation or measurements, we have to be able to quantify these cross-sections of a given process happening at a particle collider. We do this by summing the probabilities of all the given ways by which a process can occur. The given ways by which a process can occur are represented by diagrams known as feynman diagrams. The fynman diagrams for the production of SUSY particles at LHC is given in (2.3.1).



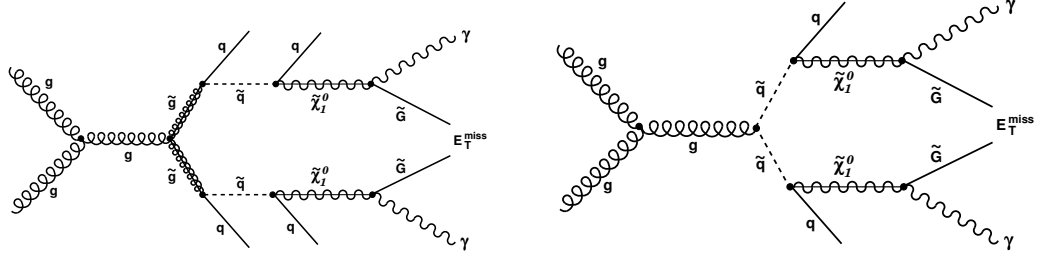


Figure 2.8: Feynman diagrams of single(top) and di(bottom) photon production from cascade decays of gluino and squark at LHC.

The list presented is not exhaustive at all but rather simplicity in computation. The sum or complete computations results in a single number known as the cross-section of the process occurring at the LHC for some amount of center of mass energy. This number called the cross-section is very small that it is express in term of a unit called a *barns* (b). ($1 b = 10^{-28} m^2$). Because this number is quite small it is usually further express with suffix in front of the unit barn such as *nano* (n) or *pico* (p) or *femto* (f). For example, the typical cross-section of producing a SUSY particle at the LHC is of the order of $1 pb = 10^{-12} \times 10^{-28} m^2$ or at times $1 fb = 10^{-15} \times 10^{-28} cm^2$ for some extremely rare SUSY processes. While a typical production of a standard model process such as the production of the Z or W bosons is of the order of a few $nb = 10^{-9} \times 10^{-28} m^2$. Thus, finding a SUSY particle in an environment such as the LHC is liken to *finding a needle in a haystack*, since most SUSY production cross-section are very small compared to an overwhelming high cross-section processes from the standard model.

2.3.2 Decay of supersymmetric particles in CMS detector

NLSP Decay Length

The SUSY particle produced will also decay instantaneously or stay a bit longer-before disintegrating. Usually this disintegration process occurs in a series of subsequent decays in a cascade manner until the final particle is stable and cannot decay. This decay is once again a probabilistic process which can also be expressed and quantified in a single real number. The probability depends on quite a number of factors like the availability of smaller particles for it to decay into and some others. Usually there must be a conservation of a fundamental quantum number or symmetry like the energy conservation for such a decay or decay chain to be possible. There are always many

options for a given particle to decay into and each of these options is quantified by a number known as the *Branching fraction or Ratio*. Summing all the branching ratio for all the different decay processes gives the total *Decay Width*. This total decay width, calculated in theory from the different probabilities of the given particle to decay into other particles is expressed as a single real number in units of GeV or MeV (GeV = giga electron volt). This number can be compared with measurements from experiments observing this different decays. In the case of the SUSY particles like the NLSP, the decay width depends on the nature of the interaction which in this case is different from the interactions currently understood within the SM.

The probability for a NLSP particle produced with an energy E having mass m to travel a distance x before decaying to a photon and gravitino in the laboratory frame is given as:

$$\mathcal{P}(x) = 1 - \exp\left(-\frac{x}{L}\right) \quad (2.47)$$

where

$$L = c\tau_{NLSP} \cdot (\beta\gamma)_{NLSP} [mm] \quad (2.48)$$

and

$$(\beta\gamma) = \frac{|p|}{m} = \sqrt{\left(\frac{E}{m}\right)^2 - 1} \quad (2.49)$$

with its proper decay length as given by as (we have used equation 2.23 and 2.64 to go from decay rate to lifetime):

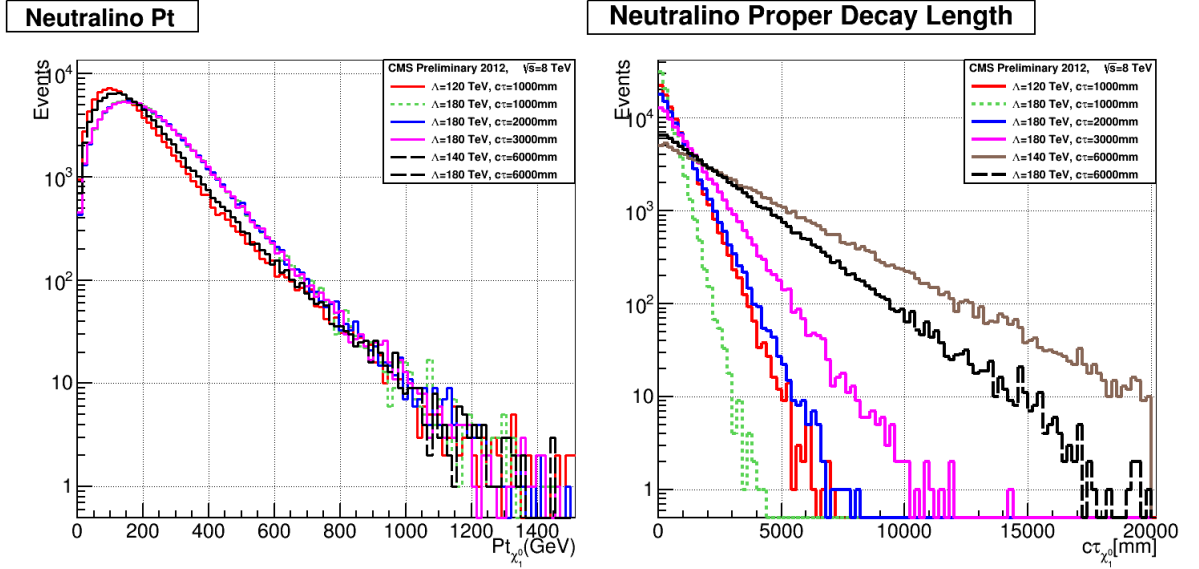
$$c\tau_{NLSP} \approx \left(\frac{m_{NLSP}}{\text{GeV}}\right)^{-5} \left(\frac{\sqrt{\mathbf{F}}}{\text{TeV}}\right)^4 \quad (2.50)$$

It is important to observe here that, varying \mathbf{F} changes the lifetime of the NLSP from being prompt to long-lived. This variation can be easily archived in GMSB models where the parameter C_{grav} is used so that the above proper decay length equation becomes:

$$c\tau_{NLSP} \approx C_{grav}^2 \left(\frac{m_{NLSP}}{\text{GeV}}\right)^{-5} \left(\frac{\sqrt{\mathbf{F}_S}}{\text{TeV}}\right)^4 \quad (2.51)$$

With C_{grav} , one could change the decay length of the NLSP such that

its decay occurs within the volume of the detector such that the resulting photon is delayed. This gives a unique signature for discovering SUSY in hadron colliders as photons produced from SM interactions are prompt. In a simple GMSB model such as the SPS8, where the neutralino is the NLSP and gravitino is LSP, figure (??) show the kinematic properties and neutralino proper decay length distribution for the neutralino and its decayed photon in different parameter choices. The HepMC class in CMS Software (CMSSW) is used to measure and an exponential distribution as the one in (2.47) is fitted on the decay length distribution to extract its proper decay length as produced in the MC generation.



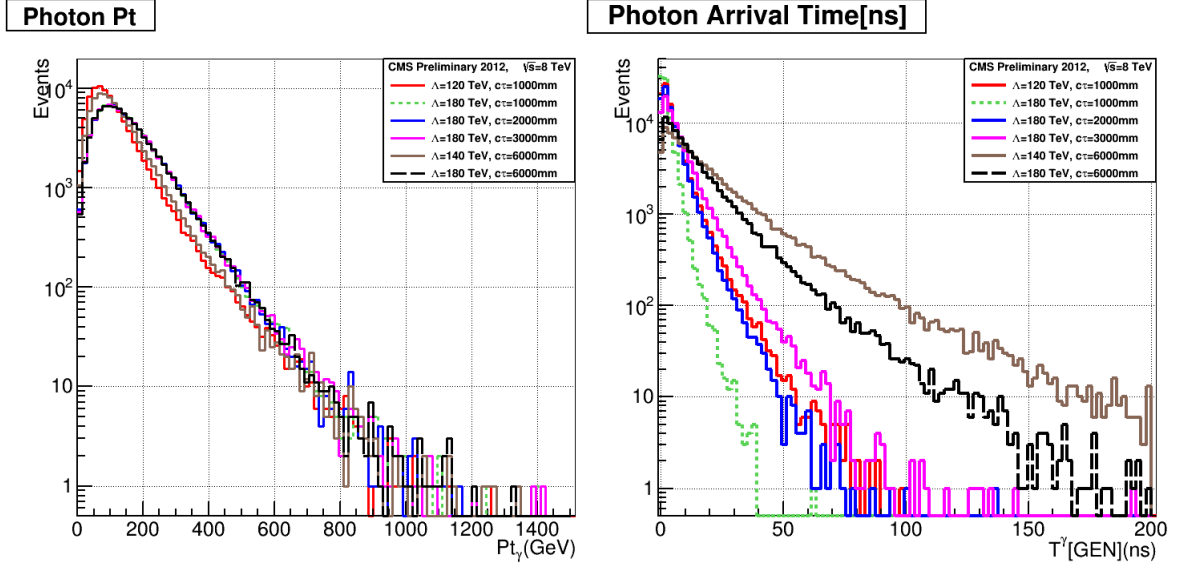


Figure 2.9: Neutralino transverse momentum distribution(top left) and proper decay length(top right) with its decayed photon transverse momentum distribution(Bottom right) and time of arrival at ECAL(Bottom right) for GMSB SPS8 model.

2.3.3 Why is the search for neutral long-lived particles important?

Finding answers to fundamental questions such as the following: What is the origin of neutrino masses, Why are there only 3 types of leptons and quarks? Where do all the parameters in SM come from? Other big questions include: What is Dark Matter (DM)? is DM made of particles? Can one detect these particles? Why is there so much asymmetry between matter and anti-matter in our universe? Is there some energy scale or early epoch in the evolution of the universe where all the fundamental forces behaved as a unique kind of force. Do baryons such as the proton exist forever? What is the lifetime of the proton? What is Dark Energy (DE)? Is the Universe expanding indefinitely? Are there other Universes? provide added impetus to search for particles BSM. We believe that the discovery of a long-live particle which are not known to exist within the current SM will provide unique access to understanding physics BSM and making measurements which can provide answers to most of the above questions. Searching for LL neutral particles decaying to photon using timing gives us a unique advantage compared to other experiments as we expect very limited background process

contribution from SM. Most of our background will be detector originated contributions. Infact, our search using lifetime gives us a wide range of search techniques depending on the lifetime of the LL particle ranging from quantitative measurements to statistics. The figure below shows the wide variety of techniques which can be used to search for LL particles in general.

ADD figure of techniques for Searches using LL (2.52)

Using equations (2.49), precise measurement of fundamental parameters in SUSY or new physics can be archived. Another advantage is that, the our search for neutral particles is unique in that a lot of previous searches been performed for charged particles but very limited for neutral LL particles since DM is speculated to be made of stable neutral particle(s) with long lifetime, we might as well go for DM. We also use lifetime because no particle with lifetime $\gtrsim 10^{-7}$ s and mass $\gtrsim 1.5$ GeV has been found and obviously because our detector has a very good timing resolution as can be seen in the section of the CMS detector in this thesis.

There have been previous attempts to search for quasi-stable neutral massive particles but all the search results show no evidence for neutral particles with long lifetime. The challenge with such an experiment is that neutral particles cannot be studied using conventional magnetic spectrometer as they are not affected by magnetic field because they are charge (local U(1) gauge symmetry) neutral. Nevertheless, there are countless theoretical as well as observational reasons why studying these particles using novel experimental techniques is very important in the field of particle physics. Some of these reasons will emerge naturally as we see in the subsequent sections below.

Theory Motivation.

Physics BSM can be summarized to answering three major theoretical questions: Is there a reliable explanation behind the ordering in mass of SM particles as observed? This is the Hierarchy problem. Is there a single theory which can provide a derivation for all the numerous parameters (19) in the SM and also unify all the fundamental forces of nature? Grand Unified Theories (GUT). What is DM and Dark Energy (DE)? (DE is the stuff that is responsible for the accelerated expansion of our universe). And finally

being a particle physicist it is only natural to ask if DM is made up of particles and if yes, Can one construct a model which can consistently describe DM as is already the case with visible matter in SM?

Most of the efforts in the last decades in theoretical particle physics has been to find answers to the above questions.

Experiment and Observation Motivation

As early as 1956 Reines and Cowan[?] observed that when a neutron decays into a proton and an electron, an elusive particle called neutrino is also produced. The observation of neutrinos was later incorporated into the SM. In the formulation of the SM, the neutrinos are considered massless. However, recent results from experiments [?] have shown that neutrinos of different flavours can oscillate or mix into one another. The only way they can do this is if they have a tiny but finite mass. Recent experiments measuring the different neutrino flavours and their mass difference point towards the existence of a much larger theory that can incorporate the existence of neutrino masses and the observed phenomenon of neutrino mixing in which the SM is embedded in it and can be understood as a low energy version of a much broader and deeper theory.

Galactic and supernovae observations using the Hubble and a host of other telescope as well as results from Baryonic Acoustic Oscillation (BAO) and WMAP reveal unique matter content of our universe. In addition to these cosmological observations including galaxy profiles, cluster formation, large scale structure formation and Cosmic Microwave background (CMB) power spectrum can be somehow explained by DM [?]. These observation reveals about 25% of our universe is made of DM with the current DM relic density is measured to be However, the question of “What is DM?” remains a very interesting one to both the particle physics astronomy society. Understanding DM and the rest of our universe will be crucial for future developments in high energy physics from both theory and experimental fronts. A possible property of DM is that they must be made of up long lived neutral particles. There are candidate particles from SUSY which have these properties. A few of these include lightest neutralino ($\tilde{\chi}_1^0$) and gravitino (\tilde{G}). From the GMSB point of view, because gravitinos are stable, neutral and very weakly interacting; they are seen as good candidates for the particles which

make up DM.

2.4 Previous Experiments and Results

The hunt for the discovery of DM and new particles, has led to several experimental search for neutral long-live particles decaying to photons. Obviously, negative results from these experiments has led to the putting limits of the lifetime, mass and cross section of possible existence of SUSY particles in different models of SUSY. Results from experiments(DO, CDF, CMS and ATLAS)[?, ?] in the search for Neutralino NLSP decaying to photon and gravitino and interpreted within the Snowmass Point and Slop (SPS)8 benchmark[?] scenario with parameter set be seen in figures 2.3(a) and (b). These results show that within the SPS8 model, neutralinos with mass $m_{\tilde{\chi}_1^0} \leq 245$ GeV and proper decay length $c\tau_{\tilde{\chi}_1^0} \leq 6000$ mm cannot exist at hadron colliders and thus their existence has been excluded.

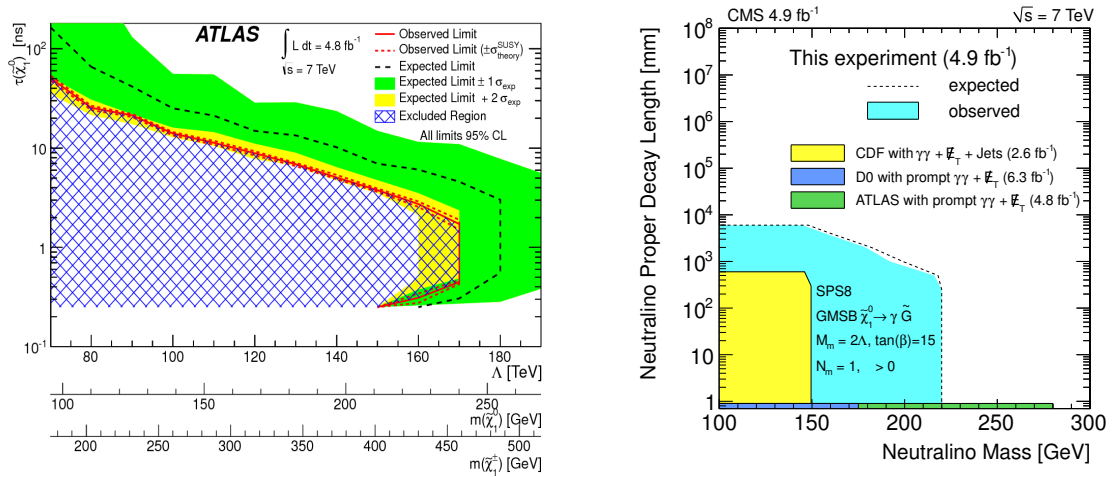


Figure 2.10: Neutralino lifetime and mass upper limit from ATLAS(left) and CMS(right) 7 TeV analysis with non-pointing photons and MET.

Chapter 3

Hadron Collider and Detector

This section describes the particle accelerator and detectors that are used to produce and detect particles at colliders. The first section describes the particle accelerator while the following section describes the CMS detectors with emphasis to those sections which are directly relevant to this analysis. A detailed description of the LHC and detectors can be found in [21] and [?].

3.1 Large Hadron Collider

3.1.1 Overview

The Large Hadron collider (LHC) is a proton-proton and heavy ion collider designed to achieve a center of mass \sqrt{S} energy of 14 TeV. It is hosted and controlled by the European Organisation for Nuclear Research (CERN). Unlike linear colliders, the LHC is a circular collider with nearly 27 km in circumference located at the border between France and Switzerland. It is designed to smash protons and ions against each other controlled by powerful magnets at officially four main locations. At each major collision point are multi-purpose particle detectors ranging from A Toroidal LHC Apparatus (ATLAS) and Compact Muon Solenoid (CMS) both non-fixed target detectors, A Large Ion Collider Experiment (ALICE) for colliding heavy ions and finally Large Hadron Collider beauty (LHCb), a fixed target experiment for investigating the properties of B-Hadrons. We give a full description of the important parts of the LHC in the following subsections,