# Displaced Photon Background Studies Update LL Meeting Nov 02 2012

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# DataSet & Triggers

#### DataSets:

- /SinglePhoton/Run2012C/Prompt-Reco V1&2
- Int Lumi = 6.7 /fb

#### HLT Trigger:

 HLT\_DisplacedPhoton65\_CaloIdVL\_IsoL\_ PFMET25

# Signal from Background

• Ecal Time Calibration:

$$\langle t_{\gamma_{true}} \rangle \! \simeq \! 0$$
 but  $t_{\gamma_{Sig}} \! 
eq \! 0$ 

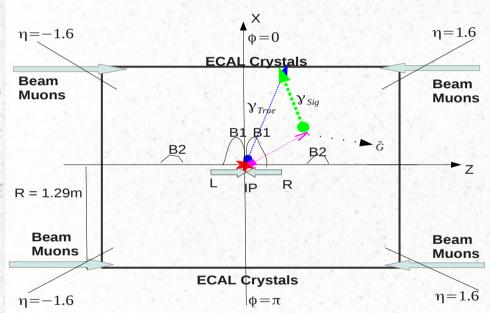
2 Types of Backgrounds



- Left B1 w/ Right B2
- Left B2 w/ Right B1
- LeftB2 w/ Right B2

#### Non-Collision:

 Beam halo (beam dump or P+Gas -> muons which Brem/shower in ECAL.



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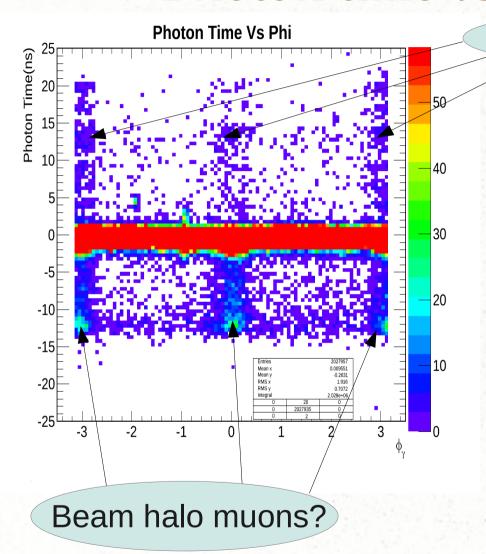
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Cosmic muons.

# **Event Selection**

- Selection :
  - $Gamma_Pt1(2) > 60(45)GeV$
  - |eta| < 2.5, Jet\_pt > 35 GeV
  - Egamma VL Iso cuts,
  - MET > 0 GeV
- Photon Tagging:
  - CSC Segment |eta| > 1.6
  - Halo tagged if dphi(cscSeg, gamma) < 0.6

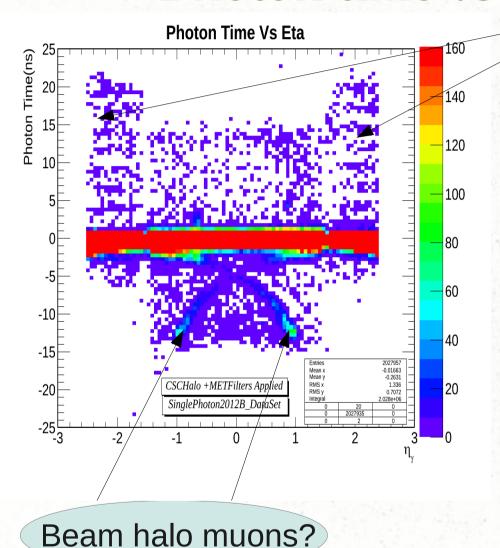
#### Photon time vs Phi



#### Cosmic muons?

- Observed in 2012B dataset
- High intensity inphi = 0, +/- pi
- Most photons arriving early in Ecal time indicate source could be beam halo muons.

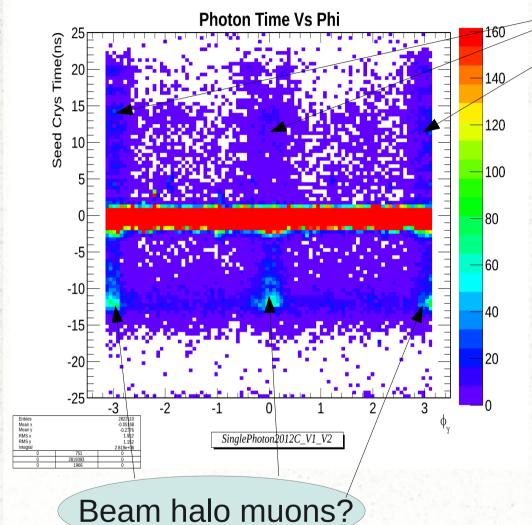
#### Photon time vs Eta



#### Cosmic muons?

- Intense at impact point in EB then slowly decreases towards IP.
- Eta dependence in early Ecal time.
- Surely photons do not all come from a unique source.

#### Photon time vs Phi



Cosmic muons?

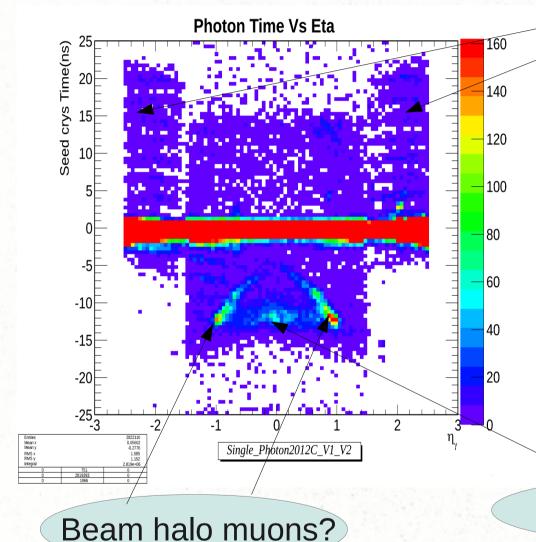
- Similar
   phenomenon in

   2012C dataset.
- Increased intensity with luminosity.
- Phi dependence in Ecal time of photons.

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#### Photon time vs Eta



Cosmic muons?

- 2012C dataset.
- Increased intensity with luminosity.
- Spikes with time between -10 and -15ns.

Spikes failing spike cleaning.

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# Definition

• Egamma Photon:

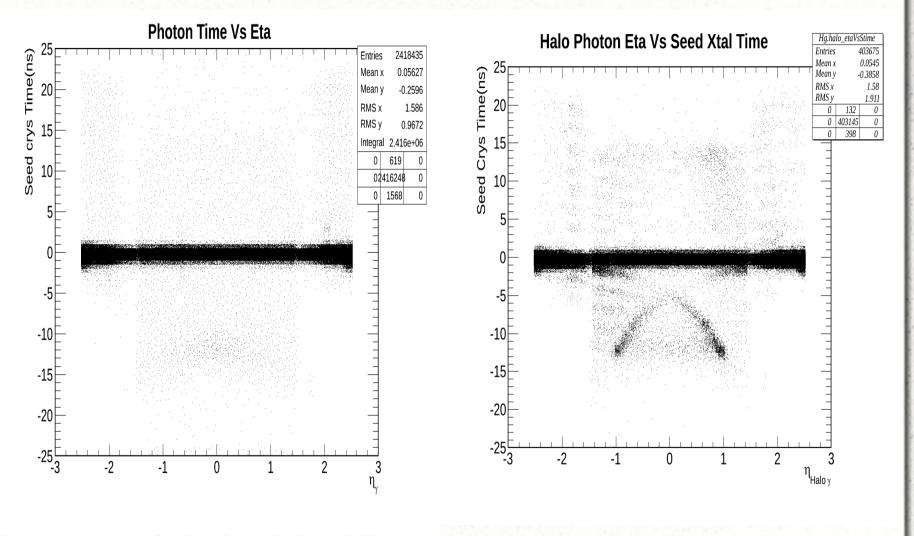
\* Egamma + VL Iso selection criteria + Sminor.

Halo Photon :

Tag as Halo photon if:

\* CSC Segment matching: dphi (cscsegment, gamma) < 0.6

#### Photon Time Vs Eta

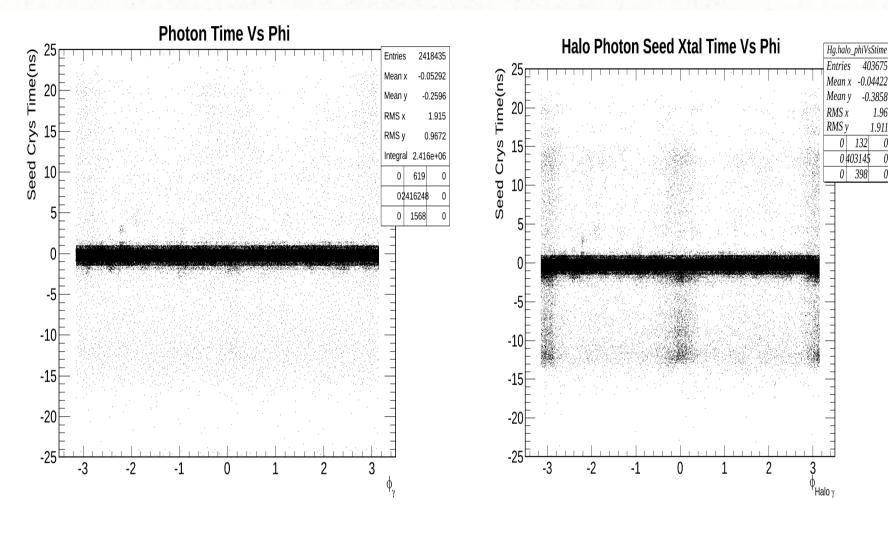


 $\overset{\scriptscriptstyle{02/11/2012}}{Egamma}\ Photon$ 

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Halo Photon

#### Photon Time Vs Phi



Egamma Photon Tambe E. Norbert(UMN US)

Halo Photon

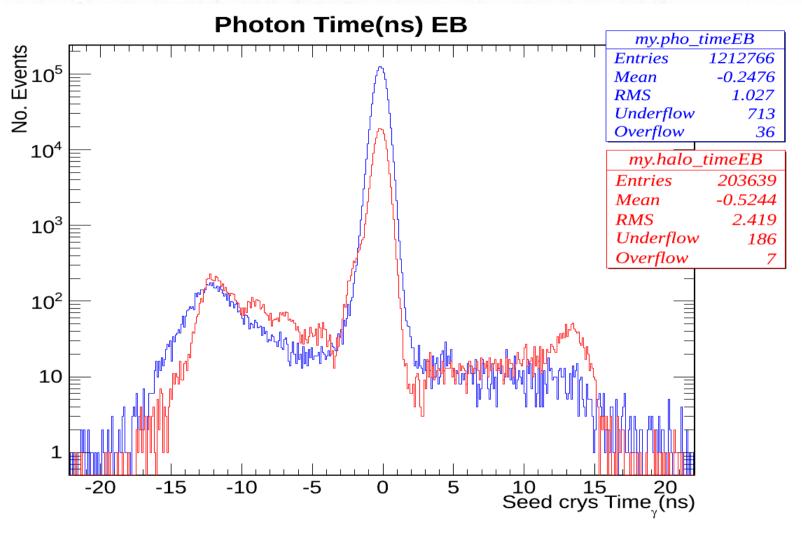
11

-0.3858

1.96

1.911

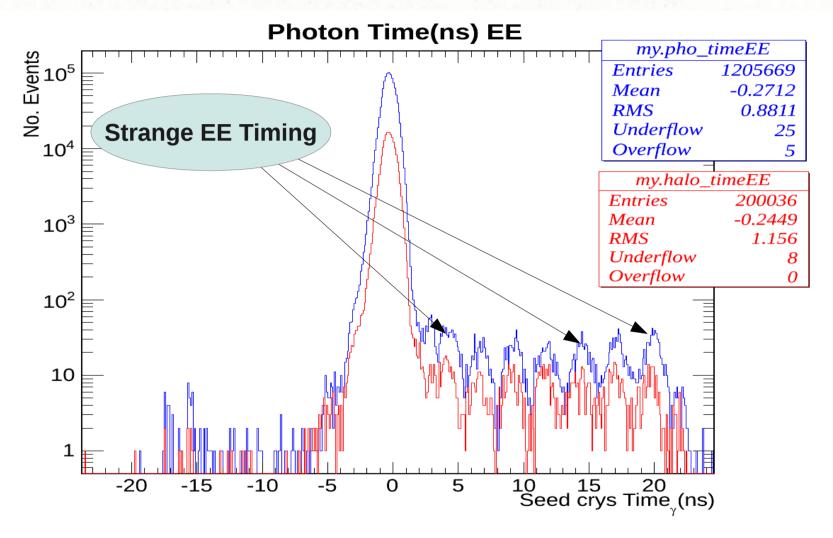
#### Photon Time EB



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#### Photon Time EE



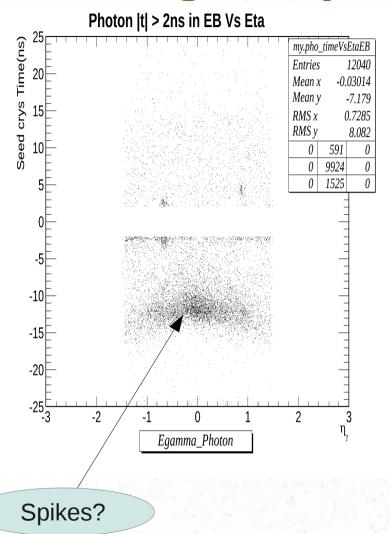
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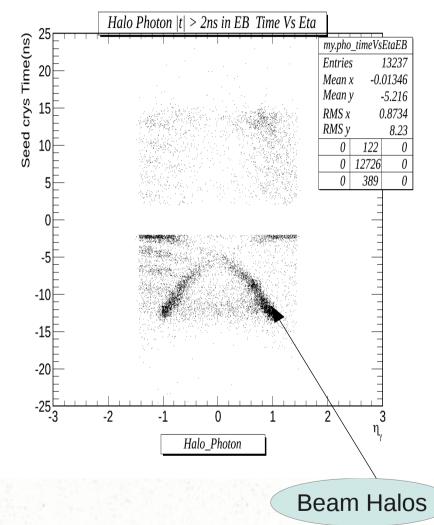
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Egamma(Blue) Halo(Red) Photon Ecal time

# Region of Interest: EB





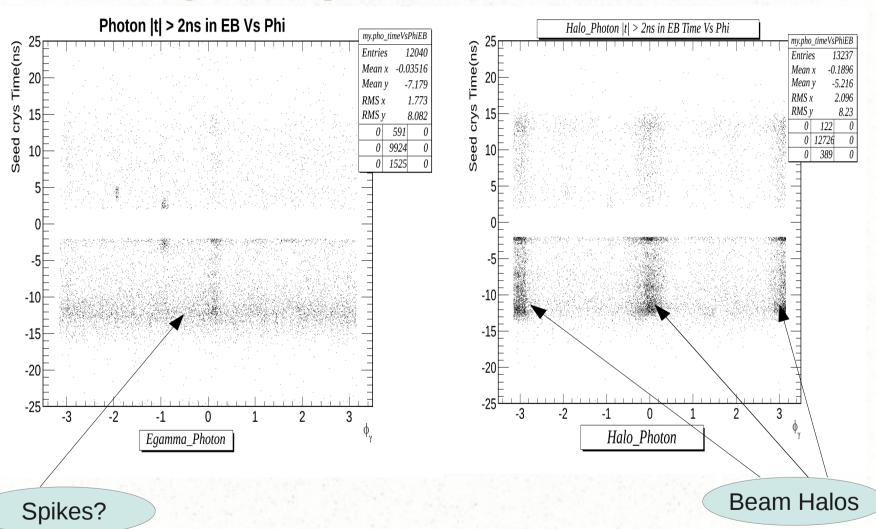
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Egamma photons

Halo photons.

# Region of Interest: EB

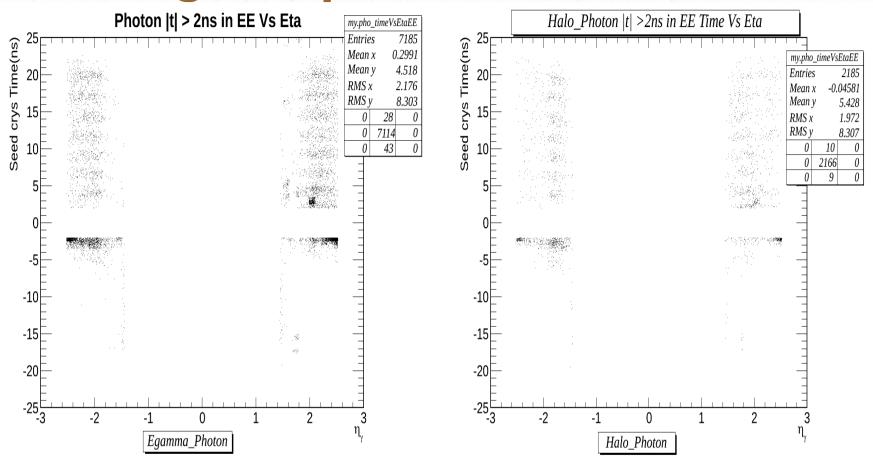


Egamma photons

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Halo photons.

# Region of Interest: EE



Few early time Halo photon than in EB

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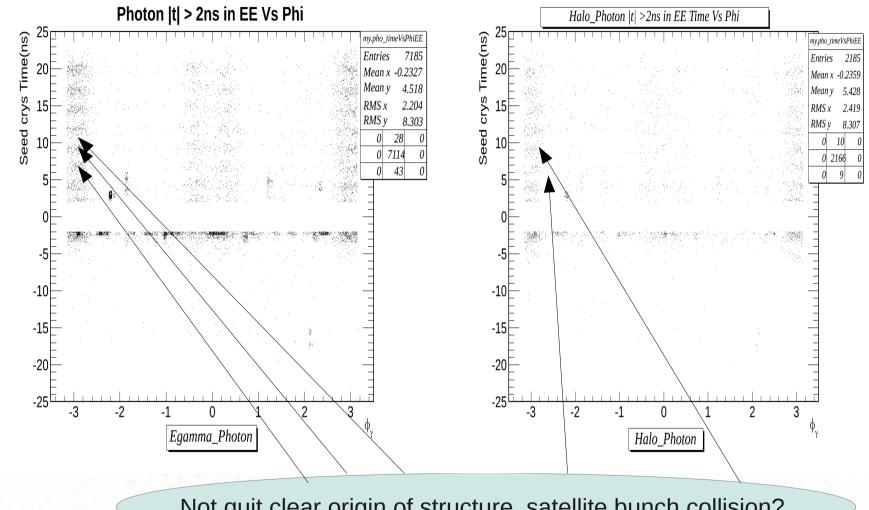
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Egamma photons

Halo photons.

# Region of Interest: EE



Not quit clear origin of structure, satellite bunch collision?

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Halo photons.

Egamma photons

# Halo Tagging & Egamma Photon Efficiency.

t  > 2ns	Total Number of Photons	Egamma Photons	% of Non- Halo- candidates	CSC Halo Tagged Photons	% of Halo candidates.
EB	25277	12040	48	13237	52
EE ( eta  < 2.5 )	9370	7185	77	2185	23

- For photons with time outside 2ns window:
  - Halo tagging efficiency : EB(EE) = 52(23)%
  - Egamma non-Halo photon % : EB(EE) =

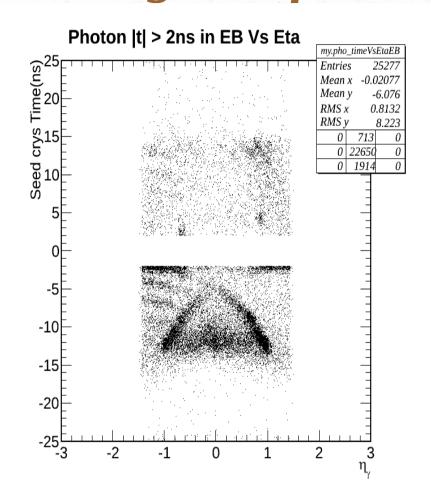
48 (77)%

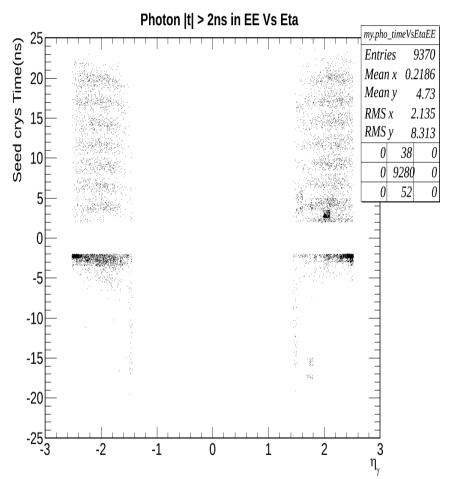
# **Summary**

- There are many non trivial background sources to delayed photon.
- CSC tagging can be use to reject beam halo photons with good efficiency, however not every background photon.
- With current background understanding, we are ready for Moriond although identifying the different sources is yet to come, nevertheless, we are working on it.

# BACK UP

# Region of Interest EB & EE

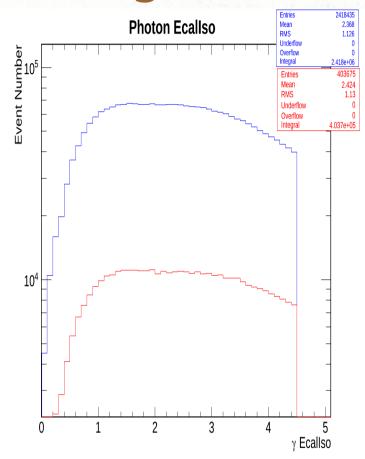


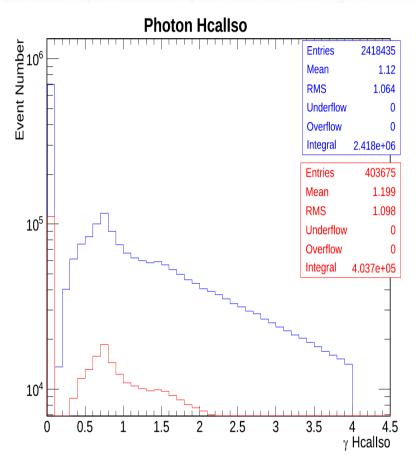


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# Egamma Photon Id variables.





**Ecal Isolation** 

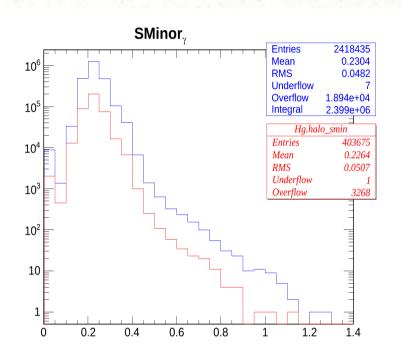
**Hcal Isolation** 

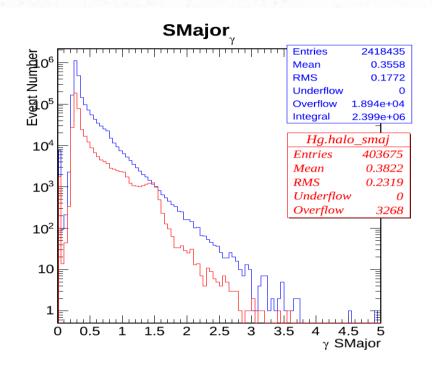
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#### Photon Iso and Id variables.



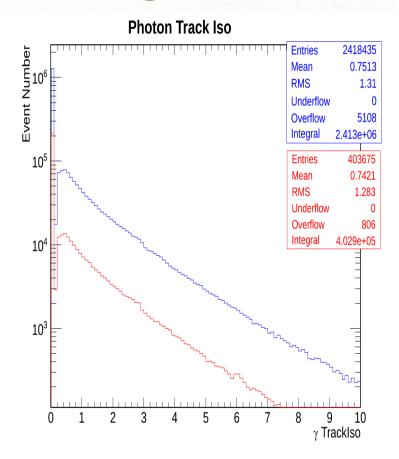


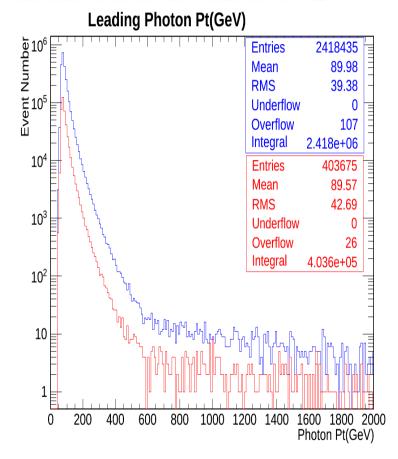
**SMinor** 

**SMajor** 

Egamma and Halo have very similar isolation criteria.

# Egamma Photon Id variables.





Track Isolation

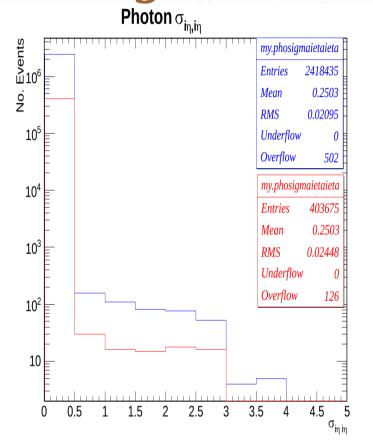
Leading photon pt

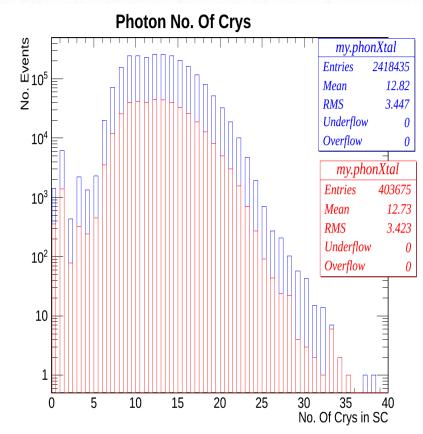
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# Egamma Photon Id variables.





Sigma Ieta Ieta

Number of Crystals in BC

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