

Phonon decay in 1D atomic Bose quasicondensates via Beliaev-Landau damping

Décroissance des phonons dans les quasi-condensats de Bose atomiques 1D via l'amortissement de Beliaev-Landau

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In a Bose gas, there is no non-trivial scattering channel involving three Bogoliubov modes that conserves both energy and momentum. Nevertheless, we show that such processes (Beliaev and Landau damping) account for their decay via interactions. Within an appropriate time window where the Fermi Golden Rule is applicable, the occupation number of the initially occupied mode decays exponentially in time, in an analytic form. The result is shown to compare favorably with simulations based on the Wigner Approximation. It is also shown that the same processes govern the growth of phonons induced by a parametric oscillation.

Dans un condensat de Bose 1D, il n'y a pas de canal de diffusion non trivial impliquant trois modes de Bogoliubov qui conserve à la fois l'énergie et la quantité de mouvement. Néanmoins, nous montrons que de tels processus de mélange à trois ondes (amortissement de Beliaev et Landau) expliquent leur décroissance via des interactions avec les fluctuations thermiques. Dans une fenêtre appropriée où la règle d'or de Fermi est censée s'appliquer, le nombre d'occupation du mode initialement occupé décroît de manière exponentielle et le taux prend une forme analytique. Le résultat se compare favorablement avec les simulations basées sur l'approximation de Wigner. Il est également montré que les mêmes processus ralentissent la croissance des phonons induite par une oscillation paramétrique.

I. INTRODUCTION

Introduction

Ultracold gases have proved to be a fruitful arena for both theoretical and experimental research. In particular, the propagation of elementary excitations on top of a macroscopic condensed background provides an accessible realisation of a quantum field in an effective curved spacetime [1, 2]. This can be exploited to, *e.g.*, mimic a black-hole horizon so as to induce the analogue of Hawking radiation (as recently achieved in [3]). It can also be used as a platform for realising an analogue of preheating, or the Dynamical Casimir Effect, which is a topic of current interest [4–6]. The degree of experimental control, combined with the intrinsically quantum nature of ultracold gases, makes them well-suited for such experiments. Les gaz ultrafroids se sont révélés être un domaine très fécond pour la recherche théorique et expérimentale. En particulier, la propagation des excitations élémentaires sur un fond condensé macroscopique constitue une réalisation accessible d'un champ quantique dans un espace-temps courbe effectif [1, 2]. Cela peut être exploité pour, par exemple, imiter un horizon

de trou noir afin d'induire l'analogue de la radiation de Hawking (comme cela a été récemment réalisé dans [3]). Cela peut également servir de plateforme pour réaliser un analogue du préchauffage ou de l'effet Casimir dynamique, un sujet d'intérêt actuel [4–6]. Le degré de contrôle expérimental, combiné à la nature intrinsèquement quantique des gaz ultrafroids, les rend particulièrement adaptés à de telles expériences.

In studies of this kind, an important issue concerns the effect of dissipation on the expected signal. Dissipation arises in closed systems as an effective phenomenon due to quasiparticle interactions [7, 8]. This entails the existence of an intrinsic quasiparticle decay. In 3D Bose gases, the principle mechanisms behind this decay are the 3-wave mixing processes of Beliaev and Landau damping [9, 10]. However, in 1D Bose gases with only two-body contact interactions (*i.e.*, of the Lieb-Liniger model [11]), two objections have been raised against the possibility of such processes. The first is that the integrability of the model prevents quasiparticle decay in principle, and that integrability-breaking perturbations must therefore be included before damping can occur [12–14]. The second is that quasiparticle decay requires the existence of non-trivial scattering channels conserving both momentum and energy, a criterion that has been routinely applied in many systems of all dimensionalities [15–19]. As there are no such channels involving three collinear Bogoliubov excitations due to their gapless and convex spectrum, it has been concluded that 3-wave mixing in 1D cannot induce decay [12, 18, 20]. Dans les études de ce type, une question importante concerne l'effet de la dissipation sur le signal attendu. La dissipation apparaît dans les systèmes fermés comme un phénomène effectif dû aux interactions entre quasiparticules [7, 8]. Cela implique l'existence d'une désintégration intrinsèque des quasi-particules. Dans les gaz de Bose 3D, les mécanismes principaux derrière cette désintégration sont les processus de mélange à trois ondes, tels que l'amortissement de Beliaev et de Landau [9, 10]. Cependant, dans les gaz de Bose 1D avec uniquement des interactions de contact entre deux corps (*i.e.*, du modèle de Lieb-Liniger [11]), deux objections ont été soulevées quant à la possibilité de tels processus. La première objection est que l'intégrabilité du modèle empêche en principe la désintégration des quasi-particules, et que des perturbations rompant l'intégrabilité doivent donc être incluses avant que l'amortissement puisse se produire [12–14]. La seconde objection est que la désintégration des quasi-particules nécessite l'existence de canaux de diffusion non triviaux conservant à la fois la quantité de mouvement et l'énergie, un critère qui a été systématiquement appliqué dans de nombreux systèmes de toutes les dimensions [15–19]. Étant donné qu'il n'existe pas de tels canaux impliquant trois excitations de Bogoliubov colinéaires en raison de leur spectre sans gap et convexe, il a été conclu que le mélange à trois ondes en 1D ne peut pas induire de désintégration [12, 18, 20].

In this paper, we wish to push back a little against these conclusions. First, integrability does not seem to prevent relaxation within a physical model, but only thermalization: the system tends towards an equilibrium state with rather more structure than a thermal state [21], described by a generalized

Gibbs ensemble [22], see [23] for a recent application to the Lieb-Liniger gas. In addition, in the context of 1D Bose gases, this relaxation could even be necessary in order to comply with the Mermin-Wagner-Hohenberg theorem [24, 25], as it provides a mechanism by which the long-range order induced by a sufficiently narrowband excitation spectrum is washed away. Dans cet article, nous souhaitons remettre en question quelque peu ces conclusions. Tout d'abord, l'intégrabilité ne semble pas empêcher la relaxation dans un modèle physique, mais seulement la thermalisation : le système tend vers un état d'équilibre avec une structure plutôt plus complexe qu'un état thermique [21], décrit par un ensemble de Gibbs généralisé [22], voir [23] pour une application récente au gaz de Lieb-Liniger. De plus, dans le contexte des gaz de Bose 1D, cette relaxation pourrait même être nécessaire pour se conformer au théorème de Mermin-Wagner-Hohenberg [24, 25], car elle fournit un mécanisme par lequel l'ordre à longue portée induit par un spectre d'excitation suffisamment étroit est effacé.

Second, we wish to show and emphasise that the apparent absence of an elastic scattering channel does not necessarily preclude any quasiparticle decay. In the Fermi Golden Rule (FGR), the Dirac delta enforcing energy conservation is typically interpreted in a binary way: either an energy-conserving channel exists and quasiparticle decay occurs, or there is no such channel and the quasiparticle is stable. However, the Dirac delta is an idealization of a narrow distribution with a small but finite width, so we must consider those final states that are *in the vicinity* of the exactly energy-conserving one. In particular, in 1D quasicondensates, the trivial 3-wave mixing channel involving the zero-energy mode is not physical itself. Yet the divergent thermal population of nearby infrared modes yields a well-defined contribution to the FGR decay rate. As anticipated, this mechanism leads to a broadening of the excitation spectrum and, within an appropriate time interval, the occupation of the initially occupied mode decays exponentially at a constant rate. The calculation of this rate is our main result. Deuxièmement, nous souhaitons montrer et souligner que l'absence apparente de canal de diffusion élastique ne préjuge pas nécessairement de l'absence de désintégration des quasiparticules. Dans la règle d'or de Fermi (FGR), le delta de Dirac imposant la conservation de l'énergie est généralement interprété de manière binaire : soit un canal conservant l'énergie existe et la désintégration des quasi-particules se produit, soit il n'existe pas de tel canal et les quasi-particules sont stables. Cependant, le delta de Dirac est une idéalisation d'une distribution étroite avec une largeur petite mais finie, donc nous devons considérer les états finaux qui sont *dans la proximité* de l'état exactement conservateur en énergie. En particulier, dans les quasi-condensats 1D, le canal de mélange à trois ondes trivial impliquant le mode à énergie nulle n'est pas physique en soi. Pourtant, la population thermique divergente des modes infrarouges voisins donne une contribution bien définie au taux de désintégration selon la FGR. Comme anticipé, ce mécanisme conduit à un élargissement du spectre d'excitation et, dans un intervalle de temps approprié, l'occupation du mode initialement occupé décroît exponentiellement à un taux constant. Le

calcul de ce taux est notre résultat principal.

The paper is organized as follows. In Sec. II, we recall the Lieb-Liniger model of a 1D Bose gas, including its hydrodynamical description in the weakly-interacting regime, and define the quasiparticles whose decay we are interested in. In Sec. III, we present a derivation of the intrinsic quasiparticle decay rate, as well as the outline of a more precise formulation that includes the main deviations from exponential decay. In Sec. IV, we show that the decay rate extracted from numerical simulations is in good agreement with the prediction of Sec. III, both in the relaxation of an initial injection of phonons and in the slowing of exponential growth induced by a parametric oscillation (the latter example being inspired by experiments [4]). We conclude and make links with other works in Sec. V, while some further details are relegated to Appendices. L'article est organisé comme suit. Dans la section II, nous rappelons le modèle de Lieb-Liniger pour un gaz de Bose 1D, y compris sa description hydrodynamique dans le régime de faible interaction, et définissons les quasi-particules dont nous étudions la désintégration. Dans la section III, nous présentons une dérivation du taux de désintégration intrinsèque des quasi-particules, ainsi que les grandes lignes d'une formulation plus précise incluant les principales déviations par rapport à une décroissance exponentielle. Dans la section IV, nous montrons que le taux de désintégration extrait des simulations numériques est en bon accord avec la prédiction de la section III, tant dans la relaxation d'une injection initiale de phonons que dans le ralentissement de la croissance exponentielle induite par une oscillation paramétrique (cet exemple étant inspiré par des expériences [4]). Nous concluons et établissons des liens avec d'autres travaux dans la section V, tandis que certains détails supplémentaires sont reportés en annexes.

L'article explore la désintégration des quasi-particules dans les gaz de Bose 1D, en contournant les conclusions précédentes qui affirmaient que l'intégrabilité du modèle et l'absence de canaux de diffusion élastique excluaient la désintégration. Voici un résumé des points principaux :

1. ****Contexte Théorique**** : Les gaz ultrafroids, notamment les gaz de Bose 1D, sont étudiés pour leur capacité à simuler des phénomènes quantiques dans un espace-temps courbe, comme l'analogie de la radiation de Hawking ou l'effet Casimir dynamique. Cependant, la dissipation dans ces systèmes, due aux interactions entre quasi-particules, peut influencer les résultats expérimentaux.
2. ****Problématique de la Dissipation**** : Dans les gaz de Bose 1D avec interactions de contact entre deux corps, l'intégrabilité du modèle semble empêcher la désintégration des quasi-particules, et les canaux de diffusion nécessaires pour cette désintégration sont jugés inexistantes. Cela a conduit à la conclusion que le mélange à trois ondes ne peut pas induire de désintégration.
3. ****Arguments Contre cette Conclusion**** :

- **Relaxation vs Thermalisation** : L'intégrabilité n'empêche pas la relaxation, mais seulement la thermalisation. Les gaz de Bose 1D peuvent tendre vers un état d'équilibre avec une structure plus complexe que celle d'un état thermique, ce qui est décrit par un ensemble de Gibbs généralisé.
- **Canaux de Diffusion** : Même en l'absence de canaux de diffusion élastiques clairs, la désintégration des quasi-particules peut se produire si l'on considère les états finaux proches de l'état exact de conservation de l'énergie. La population thermique des modes voisins infrarouges peut contribuer au taux de désintégration.

4. **Organisation de l'Article** :

- **Section II** : Présentation du modèle de Lieb-Liniger pour un gaz de Bose 1D et description hydrodynamique dans le régime de faible interaction, ainsi que définition des quasi-particules étudiées.
- **Section III** : Dérivation du taux de désintégration intrinsèque des quasi-particules et formulation plus précise incluant les déviations de la décroissance exponentielle.
- **Section IV** : Vérification numérique montrant que le taux de désintégration est en accord avec les prédictions théoriques, à la fois pour la relaxation des phonons et le ralentissement de la croissance exponentielle due à une oscillation paramétrique.
- **Section V** : Conclusion et mise en relation avec d'autres travaux, avec des détails supplémentaires dans les annexes.

II. SETTINGS

Réglages

Here we recall the necessary preliminaries for describing a 1D Bose gas. The quasicondensate nature, and the appropriateness of the hydrodynamical description, are introduced. The quasiparticles are defined in the hydrodynamical framework. A brief description of the numerical procedure adopted for simulating the dynamics of the gas is also given, and some preliminary numerical observations of the quasiparticle decay are presented. Nous rappelons ici les éléments préliminaires nécessaires pour décrire un gaz de Bose 1D. La nature de quasi-condensat et la pertinence de la description hydrodynamique sont introduites. Les quasi-particules sont définies dans le cadre hydrodynamique. Une brève description de la procédure numérique adoptée pour simuler la dynamique du gaz est également fournie, et quelques observations numériques préliminaires sur la désintégration des quasi-particules sont présentées.

A. Quasicondensate description

Description du quasi-condensat

Consider a 1D gas of identical bosons of mass m with only two-body contact interactions and no atom losses, described by the Hamiltonian

$$\hat{H} = \int_0^L dx \left\{ \frac{\hbar^2}{2m} \partial_x \hat{\Psi}^\dagger \partial_x \hat{\Psi} + \frac{g}{2} \hat{\Psi}^{\dagger 2} \hat{\Psi}^2 \right\}, \quad (1)$$

where $\hat{\Psi}$ is the atomic field and g is the 1D interaction constant. We impose periodic boundary conditions, so that the gas effectively lives on a torus of length L . The gas is “trapped” by the finiteness of the torus, but its dynamics and statistics are completely homogeneous. This allows an equivalent description using Fourier modes, each characterized by a wave number $k = 2\pi n/L$ for $n \in \mathbb{Z}$. Each Fourier amplitude $\hat{\Psi}_k$ is the annihilation operator for atoms of momentum $\hbar k$. The state of the system is fully characterized by the expectation values of the $\hat{\Psi}_k^{(\dagger)}$ and all their products. Considérons un gaz 1D de bosons identiques de masse m avec uniquement des interactions de contact entre deux corps et sans perte d’atomes, décrit par l’Hamiltonien

$$\hat{H} = \int_0^L dx \left\{ \frac{\hbar^2}{2m} \partial_x \hat{\Psi}^\dagger \partial_x \hat{\Psi} + \frac{g}{2} \hat{\Psi}^{\dagger 2} \hat{\Psi}^2 \right\},$$

où $\hat{\Psi}$ est le champ atomique et g est la constante d’interaction 1D. Nous imposons des conditions aux limites périodiques, de sorte que le gaz vit effectivement sur un tore de longueur L . Le gaz est « piégé » par la finitude du tore, mais sa dynamique et ses statistiques sont complètement homogènes. Cela permet une description équivalente en utilisant des modes de Fourier, chacun caractérisé par un nombre d’onde $k = 2\pi n/L$ pour $n \in \mathbb{Z}$. Chaque amplitude de Fourier $\hat{\Psi}_k$ est l’opérateur d’annihilation pour les atomes de moment $\hbar k$. L’état du système est entièrement caractérisé par les valeurs d’attente des $\hat{\Psi}_k^{(\dagger)}$ et de tous leurs produits.

A condensate occurs when one of the states of the system (typically the lowest-energy state $k = 0$) contains a macroscopic fraction of the total atom number N_{at} or, equivalently, when the gas demonstrates long-range order: the one-particle density matrix $g_1(x - x') = \langle \hat{\Psi}^\dagger(x) \hat{\Psi}(x') \rangle$ has a finite limit when $|x - x'| \rightarrow \infty$ [26]. However, the Mermin-Wagner-Hohenberg theorem [24, 25] precludes the apparition of such long-range order in a 1D system, essentially because the excitations of the system induce large fluctuations in the relative phase of $\hat{\Psi}$ between widely separated points. The one-particle density matrix then decays exponentially: $g_1(x - x') \approx \exp(-|x - x'|/r_0)$, where the coherence length r_0 is given by [25]

$$\frac{r_0}{\xi} = \left(\frac{k_B T}{mc^2} \right)^{-1} 2\rho_0 \xi. \quad (2)$$

Un condensat se produit lorsque l’un des états du système (typiquement l’état de plus basse énergie $k = 0$) contient

une fraction macroscopique du nombre total d'atomes N_{at} , ou, de manière équivalente, lorsque le gaz montre un ordre à longue portée : la matrice de densité à une particule $g_1(x - x') = \langle \hat{\Psi}^\dagger(x) \hat{\Psi}(x') \rangle$ a une limite finie lorsque $|x - x'| \rightarrow \infty$ [26]. Cependant, le théorème de Mermin-Wagner-Hohenberg [24, 25] exclut l'apparition d'un tel ordre à longue portée dans un système 1D, essentiellement parce que les excitations du système induisent de grandes fluctuations dans la phase relative de $\hat{\Psi}$ entre des points largement séparés. La matrice de densité à une particule décroît alors exponentiellement : $g_1(x - x') \approx \exp(-|x - x'|/r_0)$, où la longueur de cohérence r_0 est donnée par [25]

$$\frac{r_0}{\xi} = \left(\frac{k_B T}{mc^2} \right)^{-1} 2\rho_0 \xi.$$

A *quasi*-condensed state can however be achieved in a 1D Bose gas at sufficiently low temperature [27]. As indicated above, the quasicondensate is characterized by large relative phase fluctuations over large distances, while the density fluctuations remain small. It is therefore appropriate to adopt these as the field variables. They are related to the atomic field via the Madelung transformation: $\hat{\Psi} = e^{i\hat{\theta}} \sqrt{\hat{\rho}}$. The representation in terms of the density and phase fields is known as the *hydrodynamical* description, since $\hat{\theta}$ acts like a potential for the flow velocity: $\hat{v} = \frac{\hbar}{m} \partial_x \hat{\theta}$. At the classical level where these are all *c*-numbers, this is an exact canonical transformation, with ρ and θ being conjugate variables. At the quantum level, this is approximately true as long as the discreteness of the atoms can be neglected, which requires a sufficiently weak interaction and coarse-graining over sites containing many atoms [28]. We may then write the hydrodynamical version of Eq. (1), up to some irrelevant term coming from normal ordering:

$$\hat{H} = \int_0^L dx \left\{ \frac{\hbar^2}{2m} \frac{\partial \hat{\theta}}{\partial x} \hat{\rho} \frac{\partial \hat{\theta}}{\partial x} + \frac{\hbar^2}{8m\hat{\rho}} \left(\frac{\partial \hat{\rho}}{\partial x} \right)^2 + \frac{g}{2} \hat{\rho}^2 \right\}, \quad (3)$$

while imposing the canonical commutation relation

$$[\hat{\rho}(x), \hat{\theta}(x')] = i \delta(x - x'). \quad (4)$$

Un état *quasi*-condensé peut cependant être atteint dans un gaz de Bose 1D à une température suffisamment basse [27]. Comme indiqué ci-dessus, le quasi-condensat est caractérisé par de grandes fluctuations de phase relative sur de grandes distances, tandis que les fluctuations de densité restent faibles. Il est donc approprié d'adopter ces variables comme champs. Elles sont reliées au champ atomique via la transformation de Madelung : $\hat{\Psi} = e^{i\hat{\theta}} \sqrt{\hat{\rho}}$. La représentation en termes des champs de densité et de phase est connue sous le nom de description *hydrodynamique*, puisque $\hat{\theta}$ agit comme un potentiel pour la vitesse d'écoulement : $\hat{v} = \frac{\hbar}{m} \partial_x \hat{\theta}$. Au niveau classique où ces variables sont des nombres *c*, il s'agit d'une transformation canonique exacte, avec ρ et θ étant des variables

conjuguées. Au niveau quantique, cela est approximativement vrai tant que la discrétisation des atomes peut être négligée, ce qui nécessite une interaction suffisamment faible et un lissage sur des sites contenant de nombreux atomes [28]. Nous pouvons alors écrire la version hydrodynamique de l'équation (1), à l'exception d'un terme négligeable venant de l'ordonnancement normal :

$$\hat{H} = \int_0^L dx \left\{ \frac{\hbar^2}{2m} \frac{\partial \hat{\theta}}{\partial x} \hat{\rho} \frac{\partial \hat{\theta}}{\partial x} + \frac{\hbar^2}{8m\hat{\rho}} \left(\frac{\partial \hat{\rho}}{\partial x} \right)^2 + \frac{g}{2} \hat{\rho}^2 \right\} ,$$

tout en imposant la relation de commutation canonique

$$\left[\hat{\rho}(x) , \hat{\theta}(x') \right] = i \delta(x - x') .$$

Un état ****quasi-condensé**** peut être réalisé dans un gaz de Bose 1D à des températures suffisamment basses. Ce quasi-condensat se distingue par de grandes fluctuations de phase relative sur de longues distances, tandis que les fluctuations de densité restent faibles. Pour décrire cet état, on utilise les champs de densité et de phase, obtenus à partir du champ atomique par la transformation de Madelung : $\hat{\Psi} = e^{i\hat{\theta}} \sqrt{\hat{\rho}}$. Cette représentation est appelée ****description hydrodynamique**** car le champ de phase $\hat{\theta}$ fonctionne comme un potentiel pour la vitesse d'écoulement : $\hat{v} = \frac{\hbar}{m} \partial_x \hat{\theta}$. À un niveau classique, cette transformation est exacte avec ρ et θ comme variables conjuguées. Quant au niveau quantique, cette approximation est valable si la discrétisation des atomes est négligée, ce qui nécessite une interaction faible et un lissage sur des sites contenant de nombreux atomes.

La version hydrodynamique de l'Hamiltonien du système, en négligeant un terme d'ordonnancement normal, est :

$$\hat{H} = \int_0^L dx \left\{ \frac{\hbar^2}{2m} \frac{\partial \hat{\theta}}{\partial x} \hat{\rho} \frac{\partial \hat{\theta}}{\partial x} + \frac{\hbar^2}{8m\hat{\rho}} \left(\frac{\partial \hat{\rho}}{\partial x} \right)^2 + \frac{g}{2} \hat{\rho}^2 \right\} ,$$

avec la relation de commutation canonique :

$$\left[\hat{\rho}(x) , \hat{\theta}(x') \right] = i \delta(x - x') .$$

B. Perturbative expansion of Hamiltonian

Développement perturbatif de l'Hamiltonien

We wish to study elementary excitations, which requires a well-defined splitting of the total field into a background plus perturbations. The background is defined as the homogeneous solution of the classical equation associated to the Hamiltonian (1) (*i.e.*, the Gross-Pitaevskii equation [25]), working in the rest frame of the gas. The density ρ_0 is then constant (the total number of atoms is $N_{\text{at}} = \rho_0 L$) and the phase $\theta_0 = -g\rho_0 t/\hbar$. The density fluctuations $\delta\hat{\rho}$ around this background are assumed small while only the spatial variation of the phase fluctuations $\partial_x \delta\hat{\theta}$ is assumed small. We then expand the Hamiltonian

in $(\delta\hat{\rho}, \partial_x \delta\hat{\theta})$

$$\hat{H} = E_0 \hat{\mathbb{1}} + \hat{H}_2 + \hat{H}_3 + \sum_{i \geq 0} \hat{H}_{4+i}, \quad (5)$$

where the zeroth-order term $E_0 = g\rho_0^2 L/2$ is the energy of the homogeneous background, before any fluctuations are included. Since the background is an exact solution of the classical equation of motion, the first-order term vanishes identically. The higher orders are given by

$$\begin{aligned} \hat{H}_2 &= \int_0^L \left[\frac{\hbar^2 \rho_0}{2m} \left(\frac{\partial \delta\hat{\theta}}{\partial x} \right)^2 + \frac{\hbar^2}{8m\rho_0} \left(\frac{\partial \delta\hat{\rho}}{\partial x} \right)^2 + \frac{g}{2} \delta\hat{\rho}^2 \right] dx, \\ \hat{H}_3 &= \int_0^L \left[\frac{\hbar^2}{2m} \frac{\partial \delta\hat{\theta}}{\partial x} \delta\hat{\rho} \frac{\partial \delta\hat{\theta}}{\partial x} - \frac{\hbar^2}{8m\rho_0} \left(\frac{\partial \delta\hat{\rho}}{\partial x} \right)^2 \frac{\delta\hat{\rho}}{\rho_0} \right] dx, \\ \hat{H}_{4+i} &= (-1)^i \int_0^L \frac{\hbar^2}{8m\rho_0} \left(\frac{\partial \delta\hat{\rho}}{\partial x} \right)^2 \left(\frac{\delta\hat{\rho}}{\rho_0} \right)^{(2+i)} dx, \end{aligned} \quad (6)$$

for $i \geq 0$. Nous souhaitons étudier les excitations élémentaires, ce qui nécessite une séparation bien définie du champ total en un fond plus des perturbations. Le fond est défini comme la solution homogène de l'équation classique associée à l'Hamiltonien (1) (c'est-à-dire, l'équation de Gross-Pitaevskii [25] en travaillant dans le référentiel du gaz. La densité ρ_0 est alors constante (le nombre total d'atomes est $N_{\text{at}} = \rho_0 L$) et la phase $\theta_0 = -g\rho_0 t/\hbar$. Les fluctuations de densité $\delta\hat{\rho}$ autour de ce fond sont supposées petites, tandis que seule la variation spatiale des fluctuations de phase $\partial_x \delta\hat{\theta}$ est supposée petite. Nous développons alors l'Hamiltonien en $(\delta\hat{\rho}, \partial_x \delta\hat{\theta})$:

$$\hat{H} = E_0 \hat{\mathbb{1}} + \hat{H}_2 + \hat{H}_3 + \sum_{i \geq 0} \hat{H}_{4+i},$$

où le terme d'ordre zéro $E_0 = g\rho_0^2 L/2$ est l'énergie du fond homogène, avant l'inclusion de toute fluctuation. Étant donné que le fond est une solution exacte de l'équation classique du mouvement, le terme d'ordre un est nul identiquement. Les ordres supérieurs sont donnés par :

$$\begin{aligned} \hat{H}_2 &= \int_0^L \left[\frac{\hbar^2 \rho_0}{2m} \left(\frac{\partial \delta\hat{\theta}}{\partial x} \right)^2 + \frac{\hbar^2}{8m\rho_0} \left(\frac{\partial \delta\hat{\rho}}{\partial x} \right)^2 + \frac{g}{2} \delta\hat{\rho}^2 \right] dx, \\ \hat{H}_3 &= \int_0^L \left[\frac{\hbar^2}{2m} \frac{\partial \delta\hat{\theta}}{\partial x} \delta\hat{\rho} \frac{\partial \delta\hat{\theta}}{\partial x} - \frac{\hbar^2}{8m\rho_0} \left(\frac{\partial \delta\hat{\rho}}{\partial x} \right)^2 \frac{\delta\hat{\rho}}{\rho_0} \right] dx, \\ \hat{H}_{4+i} &= (-1)^i \int_0^L \frac{\hbar^2}{8m\rho_0} \left(\frac{\partial \delta\hat{\rho}}{\partial x} \right)^2 \left(\frac{\delta\hat{\rho}}{\rho_0} \right)^{(2+i)} dx, \end{aligned} \quad (7)$$

pour $i \geq 0$.

A couple of remarks are in order here. Notice first that the quasicondensate perturbative scheme clearly differs from the

standard Bogoliubov treatment since the standard non-linear term $\hat{\Psi}^{\dagger 2}\hat{\Psi}^2$ is fully included in the quadratic Hamiltonian, while the infinite series of perturbative corrections comes entirely from the kinetic term. Second, Eq. (7) indicates that each order is suppressed by an additional factor of $\delta\hat{\rho}/\rho_0$ which suggests that $\langle\delta\rho^2\rangle/\rho_0^2$ can be used as a measure of the importance of taking these higher orders into account. For the typical values of parameters used in this work we have $\langle\delta\rho^2\rangle/\rho_0^2 \sim 10^{-3}$ so that we will consider only the second-order term \hat{H}_2 and the first perturbation Hamiltonian \hat{H}_3 . Quelques remarques s'imposent ici. Remarquez d'abord que le schéma perturbatif du quasi-condensat diffère clairement du traitement standard de Bogoliubov, puisque le terme non-linéaire standard $\hat{\Psi}^{\dagger 2}\hat{\Psi}^2$ est entièrement inclus dans l'Hamiltonien quadratique, tandis que la série infinie des corrections perturbatives provient entièrement du terme cinétique. Deuxièmement, l'équation (7) indique que chaque ordre est supprimé par un facteur supplémentaire de $\delta\hat{\rho}/\rho_0$, ce qui suggère que $\langle\delta\rho^2\rangle/\rho_0^2$ peut être utilisé comme une mesure de l'importance de la prise en compte de ces ordres supérieurs. Pour les valeurs typiques des paramètres utilisés dans ce travail, nous avons $\langle\delta\rho^2\rangle/\rho_0^2 \sim 10^{-3}$, de sorte que nous ne considérerons que le terme du second ordre \hat{H}_2 et le premier Hamiltonien de perturbation \hat{H}_3 .

Dans cette section, nous développons l'Hamiltonien du gaz de Bose 1D autour d'un fond homogène défini par une solution de l'équation de Gross-Pitaevskii. La densité du fond, ρ_0 , est constante et la phase, θ_0 , varie linéairement avec le temps. Nous supposons que les fluctuations de densité $\delta\hat{\rho}$ et les variations spatiales des fluctuations de phase $\partial_x\delta\hat{\theta}$ sont petites. L'Hamiltonien est ainsi développé en termes de ces petites fluctuations :

$$\hat{H} = E_0\hat{1} + \hat{H}_2 + \hat{H}_3 + \sum_{i\geq 0} \hat{H}_{4+i},$$

où E_0 est l'énergie du fond homogène, \hat{H}_2 est le terme quadratique en $\delta\hat{\rho}$ et $\partial_x\delta\hat{\theta}$, \hat{H}_3 est le premier terme de perturbation, et les termes d'ordre supérieur sont de plus en plus petits. Il est important de noter que ce développement diffère du traitement standard de Bogoliubov car le terme non linéaire $\hat{\Psi}^{\dagger 2}\hat{\Psi}^2$ est entièrement inclus dans l'Hamiltonien quadratique, tandis que les corrections perturbatives proviennent du terme cinétique. Chaque ordre supérieur est supprimé par un facteur de $\delta\hat{\rho}/\rho_0$, indiquant que l'importance des ordres supérieurs peut être mesurée par $\langle\delta\rho^2\rangle/\rho_0^2$. Dans ce travail, nous trouvons que ce ratio est très petit ($\sim 10^{-3}$), nous limitons donc notre analyse aux termes \hat{H}_2 et \hat{H}_3 .

C. Quasiparticle definition

Définition des quasi-particules

Working in the canonical ensemble, the atom number N_{at} is a fixed parameter, and therefore so is the background density

ρ_0 . The zero mode of the density fluctuations thus vanishes identically: $\delta\hat{\rho}_{k=0} = 0$. Consequently, the conjugate variable $\delta\hat{\theta}_{k=0}$ is non-dynamical and can be ignored. The fluctuations $\delta\hat{\rho}$ and $\delta\hat{\theta}$ on top of this background are then composed of the non-zero Fourier modes $\delta\hat{\rho}_k$ and $\delta\hat{\theta}_k$:

$$\delta\hat{\rho}(x) = \sqrt{\frac{\rho_0}{L}} \sum_{k \neq 0} e^{ikx} \delta\hat{\rho}_k, \quad \delta\hat{\theta}(x) = \frac{1}{\sqrt{\rho_0 L}} \sum_{k \neq 0} e^{ikx} \delta\hat{\theta}_k. \quad (8)$$

With this writing $\delta\hat{\rho}_k$ and $\delta\hat{\theta}_k$ are dimensionless, and they satisfy $[\delta\hat{\rho}_k, \delta\hat{\theta}_{k'}] = i \delta_{k, -k'}$. Since $\delta\hat{\rho}(x)$ and $\delta\hat{\theta}(x)$ are Hermitian operators, the Fourier components satisfy $\delta\hat{\rho}_{-k} = \delta\hat{\rho}_k^\dagger$ and $\delta\hat{\theta}_{-k} = \delta\hat{\theta}_k^\dagger$. \hat{H}_2 can be diagonalized into normal modes, called *phonons*¹, represented by operators $\hat{\varphi}_k^{(\dagger)}$ such that

$$\hat{H}_2 = \sum_{k \neq 0} \hbar \omega_k \left(\hat{\varphi}_k^\dagger \hat{\varphi}_k + \frac{1}{2} \right), \quad (9)$$

where the phonon frequency

$$\omega_k = c |k| \sqrt{1 + \frac{1}{4} k^2 \xi^2}, \quad (10)$$

with $c = \sqrt{g\rho_0/m}$ the speed of sound and $\xi = \hbar/mc$ the healing length. In the limit $k\xi \rightarrow 0$, we have an exactly linear dispersion relation like that of the Luttinger liquid. The phononic operators are related to the density and phase fluctuations by

$$\hat{\varphi}_k = \frac{1}{\sqrt{2}} \left(C_k^{-1} \delta\hat{\rho}_k + i C_k \delta\hat{\theta}_k \right), \quad (11)$$

where $C_k^2 = \hbar k^2 / (2m\omega_k)$. The use of inverse coefficients C_k and C_k^{-1} ensures that the transformation is canonical and hence that the phonon operators satisfy the bosonic commutation relation $[\hat{\varphi}_k, \hat{\varphi}_{k'}^\dagger] = \delta_{k, k'}$. Travaillant dans l'ensemble canonique, le nombre d'atomes N_{at} est un paramètre fixe, et donc la densité de fond ρ_0 l'est aussi. Ainsi, le mode zéro des fluctuations de densité s'annule identiquement : $\delta\hat{\rho}_{k=0} = 0$. Par conséquent, la variable conjuguée $\delta\hat{\theta}_{k=0}$ n'est pas dynamique et peut être ignorée. Les fluctuations $\delta\hat{\rho}$ et $\delta\hat{\theta}$ sur ce fond sont alors composées des modes de Fourier non nuls $\delta\hat{\rho}_k$ et $\delta\hat{\theta}_k$:

$$\delta\hat{\rho}(x) = \sqrt{\frac{\rho_0}{L}} \sum_{k \neq 0} e^{ikx} \delta\hat{\rho}_k, \quad \delta\hat{\theta}(x) = \frac{1}{\sqrt{\rho_0 L}} \sum_{k \neq 0} e^{ikx} \delta\hat{\theta}_k. \quad (12)$$

Avec cette écriture, $\delta\hat{\rho}_k$ et $\delta\hat{\theta}_k$ sont sans dimension, et elles satisfont $[\delta\hat{\rho}_k, \delta\hat{\theta}_{k'}] = i \delta_{k, -k'}$. Puisque $\delta\hat{\rho}(x)$ et $\delta\hat{\theta}(x)$ sont des opérateurs hermitiens, les composantes de Fourier satisfont

¹ In this paper, we use the term “phonon” to refer to a quasiparticle of any wavelength, and *not* just to those well within the linear part of the dispersion relation (14).

$\delta\hat{\rho}_{-k} = \delta\hat{\rho}_k^\dagger$ et $\delta\hat{\theta}_{-k} = \delta\hat{\theta}_k^\dagger$. \hat{H}_2 peut être diagonalisé en modes normaux, appelés *phonons*², représentés par les opérateurs $\hat{\varphi}_k^{(\dagger)}$ tels que

$$\hat{H}_2 = \sum_{k \neq 0} \hbar \omega_k \left(\hat{\varphi}_k^\dagger \hat{\varphi}_k + \frac{1}{2} \right), \quad (13)$$

où la fréquence du phonon est

$$\omega_k = c |k| \sqrt{1 + \frac{1}{4} k^2 \xi^2}, \quad (14)$$

avec $c = \sqrt{g\rho_0/m}$ la vitesse du son et $\xi = \hbar/mc$ la longueur de guérison. Dans la limite $k\xi \rightarrow 0$, nous avons une relation de dispersion exactement linéaire comme celle du liquide de Luttinger. Les opérateurs phononiques sont reliés aux fluctuations de densité et de phase par

$$\hat{\varphi}_k = \frac{1}{\sqrt{2}} \left(C_k^{-1} \delta\hat{\rho}_k + i C_k \delta\hat{\theta}_k \right), \quad (15)$$

où $C_k^2 = \hbar k^2 / (2m\omega_k)$. L'utilisation des coefficients inverses C_k et C_k^{-1} garantit que la transformation est canonique et donc que les opérateurs de phonon satisfont la relation de commutation bosonique $[\hat{\varphi}_k, \hat{\varphi}_{k'}^\dagger] = \delta_{k,k'}$.

ξ and the associated healing time, $t_\xi = \xi/c$, provide natural units in which to express quantities adimensionally. There are three dimensionless parameters describing the system: the 1D density $\rho_0\xi$, the length L/ξ , and the temperature $k_B T/mc^2$ (where $mc^2 = \hbar/t_\xi$ is the chemical potential $\partial E_0/\partial N_{\text{at}}$). The interaction strength is characterised by $\gamma_{\text{LL}} = 1/(\rho_0\xi)^2$, the dimensionless Lieb-Liniger constant [11, 27]. Numerical simulations presented in this work typically have $\gamma_{\text{LL}} \sim 10^{-5} - 10^{-3}$, placing us firmly in the weakly-interacting regime. We also choose a grid spacing Δx such that $\rho_0\Delta x \sim 20$ atoms per site, justifying our use of the hydrodynamical description. ξ et le temps de guérison associé, $t_\xi = \xi/c$, fournissent des unités naturelles pour exprimer les quantités de manière adimensionnelle. Il y a trois paramètres sans dimension décrivant le système : la densité 1D $\rho_0\xi$, la longueur L/ξ , et la température $k_B T/mc^2$ (où $mc^2 = \hbar/t_\xi$ est le potentiel chimique $\partial E_0/\partial N_{\text{at}}$). La force d'interaction est caractérisée par $\gamma_{\text{LL}} = 1/(\rho_0\xi)^2$, la constante adimensionnelle de Lieb-Liniger [11, 27]. Les simulations numériques présentées dans ce travail ont typiquement $\gamma_{\text{LL}} \sim 10^{-5} - 10^{-3}$, nous plaçant fermement dans le régime des interactions faibles. Nous choisissons également un espacement de grille Δx tel que $\rho_0\Delta x \sim 20$ atomes par site, justifiant ainsi notre utilisation de la description hydrodynamique.

Just as for the atom operators $\hat{\Psi}_k$, the phonon operators $\hat{\varphi}_k$ provide a complete description of the system, whose state is

² Dans cet article, nous utilisons le terme "phonon" pour désigner une quasi-particule de n'importe quelle longueur d'onde, et *non* seulement celles qui se trouvent bien dans la partie linéaire de la relation de dispersion (14).

fully characterized by the expectation values of the $\hat{\varphi}_k^{(\dagger)}$ and all their products. Since the $\hat{\varphi}_k$ come close to diagonalizing the full Hamiltonian \hat{H} , the phonons are close to the exact normal modes of the system and their mutual interactions are relatively weak. The phonons therefore provide the most natural basis in which to examine the state of the system and interpret its dynamical behavior. However, the simplicity of the full Hamiltonian (1) makes the atom basis more convenient for numerical treatments of the evolution. Tout comme pour les opérateurs d'atomes $\hat{\Psi}_k$, les opérateurs de phonons $\hat{\varphi}_k$ fournissent une description complète du système, dont l'état est entièrement caractérisé par les valeurs d'attente des $\hat{\varphi}_k^{(\dagger)}$ et de tous leurs produits. Étant donné que les $\hat{\varphi}_k$ se rapprochent de la diagonalisation complète de l'Hamiltonien \hat{H} , les phonons sont proches des modes normaux exacts du système, et leurs interactions mutuelles sont relativement faibles. Les phonons constituent donc la base la plus naturelle pour examiner l'état du système et interpréter son comportement dynamique. Cependant, la simplicité de l'Hamiltonien complet (1) rend la base des atomes plus pratique pour les traitements numériques de l'évolution.

Pour étudier les excitations élémentaires dans un gaz de Bose 1D, il est nécessaire de diviser le champ total en un fond homogène et des perturbations. Le fond est défini comme la solution homogène de l'équation classique associée à l'Hamiltonien (1), c'est-à-dire l'équation de Gross-Pitaevskii, dans le cadre de repos du gaz. La densité de fond ρ_0 est constante, et la phase de fond est $\theta_0 = -g\rho_0 t/\hbar$. Les fluctuations de densité $\delta\hat{\rho}$ autour de ce fond sont supposées petites, tout comme les variations spatiales de la phase $\partial_x \delta\hat{\theta}$. L'Hamiltonien est développé en séries de perturbation autour de ce fond :

$$\hat{H} = E_0 \hat{\mathbb{1}} + \hat{H}_2 + \hat{H}_3 + \sum_{i \geq 0} \hat{H}_{4+i}, \quad (16)$$

où $E_0 = g\rho_0^2 L/2$ est l'énergie du fond homogène, et les termes de perturbation \hat{H}_2 , \hat{H}_3 , etc., représentent les corrections de plus haut ordre. Les expressions pour ces termes sont :

$$\begin{aligned} \hat{H}_2 &= \int_0^L \left[\frac{\hbar^2 \rho_0}{2m} \left(\frac{\partial \delta\hat{\theta}}{\partial x} \right)^2 + \frac{\hbar^2}{8m\rho_0} \left(\frac{\partial \delta\hat{\rho}}{\partial x} \right)^2 + \frac{g}{2} \delta\hat{\rho}^2 \right] dx, \\ \hat{H}_3 &= \int_0^L \left[\frac{\hbar^2}{2m} \frac{\partial \delta\hat{\theta}}{\partial x} \delta\hat{\rho} \frac{\partial \delta\hat{\theta}}{\partial x} - \frac{\hbar^2}{8m\rho_0} \left(\frac{\partial \delta\hat{\rho}}{\partial x} \right)^2 \frac{\delta\hat{\rho}}{\rho_0} \right] dx, \\ \hat{H}_{4+i} &= (-1)^i \int_0^L \frac{\hbar^2}{8m\rho_0} \left(\frac{\partial \delta\hat{\rho}}{\partial x} \right)^2 \left(\frac{\delta\hat{\rho}}{\rho_0} \right)^{(2+i)} dx, \end{aligned} \quad (17)$$

pour $i \geq 0$. Dans le cadre de l'ensemble canonique, le nombre d'atomes N_{at} est un paramètre fixe, et donc la densité de fond ρ_0 l'est aussi. Le mode zéro des fluctuations de densité $\delta\hat{\rho}_{k=0}$ disparaît identiquement, et la variable conjuguée $\delta\hat{\theta}_{k=0}$ est non-dynamique et peut être ignorée. Les fluctuations $\delta\hat{\rho}$ et

$\delta\hat{\theta}$ autour de ce fond sont alors composées des modes de Fourier non nuls $\delta\hat{\rho}_k$ et $\delta\hat{\theta}_k$:

$$\delta\hat{\rho}(x) = \sqrt{\frac{\rho_0}{L}} \sum_{k \neq 0} e^{ikx} \delta\hat{\rho}_k, \quad \delta\hat{\theta}(x) = \frac{1}{\sqrt{\rho_0 L}} \sum_{k \neq 0} e^{ikx} \delta\hat{\theta}_k.$$

Les opérateurs de phonons $\hat{\varphi}_k$ sont introduits en diagonalissant \hat{H}_2 . Ils sont liés aux fluctuations de densité et de phase par

$$\hat{\varphi}_k = \frac{1}{\sqrt{2}} \left(C_k^{-1} \delta\hat{\rho}_k + i C_k \delta\hat{\theta}_k \right),$$

où $C_k^2 = \hbar k^2 / (2m\omega_k)$. Les opérateurs de phonons satisfont les relations de commutation bosoniques, ce qui rend leur base la plus naturelle pour examiner l'état du système et interpréter son comportement dynamique. Cependant, pour les traitements numériques de l'évolution, la base des atomes reste plus pratique.

D. Quasiparticle interactions

Interactions des quasi-particules

Turning now to \hat{H}_3 , and neglecting terms of the form $\varphi^\dagger \hat{\varphi}^\dagger \hat{\varphi}^\dagger$ ($\hat{\varphi} \hat{\varphi} \hat{\varphi}$) which cause the unbalanced appearance (disappearance) of three phonons typically associated with strong violation of energy conservation, the relevant part of the interaction Hamiltonian takes the form

$$\hat{V}_3 = \frac{1}{\sqrt{N_{\text{at}}}} \sum_{\substack{p, q \neq 0 \\ p+q \neq 0}} \hbar V_3(p, q) \left\{ \hat{\varphi}_p^\dagger \hat{\varphi}_q^\dagger \hat{\varphi}_{p+q} + \hat{\varphi}_{p+q}^\dagger \hat{\varphi}_p \hat{\varphi}_q \right\}, \quad (18)$$

where

$$V_3(p, q) = \sqrt{\frac{\hbar}{32m}} \sqrt{\frac{|pq(p+q)|}{v_p^{\text{ph}} v_q^{\text{ph}} v_{p+q}^{\text{ph}}}} \left\{ - \left(\frac{\hbar}{2m} \right)^2 [p^2 + pq + q^2] \right. \\ \left. + v_p^{\text{ph}} v_q^{\text{ph}} + v_q^{\text{ph}} v_{p+q}^{\text{ph}} + v_p^{\text{ph}} v_{p+q}^{\text{ph}} \right\}, \quad (19)$$

and $v_k^{\text{ph}} = \omega_k/k$ is the phase velocity at k . En considérant maintenant \hat{H}_3 et en négligeant les termes de la forme $\varphi^\dagger \hat{\varphi}^\dagger \hat{\varphi}^\dagger$ ($\hat{\varphi} \hat{\varphi} \hat{\varphi}$) qui entraînent l'apparition (ou la disparition) déséquilibrée de trois phonons, typiquement associée à une forte violation de la conservation de l'énergie, la partie pertinente de l'Hamiltonien d'interaction prend la forme suivante :

$$\hat{V}_3 = \frac{1}{\sqrt{N_{\text{at}}}} \sum_{\substack{p, q \neq 0 \\ p+q \neq 0}} \hbar V_3(p, q) \left\{ \hat{\varphi}_p^\dagger \hat{\varphi}_q^\dagger \hat{\varphi}_{p+q} + \hat{\varphi}_{p+q}^\dagger \hat{\varphi}_p \hat{\varphi}_q \right\},$$

où

$$V_3(p, q) = \sqrt{\frac{\hbar}{32m}} \sqrt{\frac{|pq(p+q)|}{v_p^{\text{ph}} v_q^{\text{ph}} v_{p+q}^{\text{ph}}}} \left\{ - \left(\frac{\hbar}{2m} \right)^2 [p^2 + pq + q^2] \right. \\ \left. + v_p^{\text{ph}} v_q^{\text{ph}} + v_q^{\text{ph}} v_{p+q}^{\text{ph}} + v_p^{\text{ph}} v_{p+q}^{\text{ph}} \right\}, \quad (20)$$

et $v_k^{\text{ph}} = \omega_k/k$ est la vitesse de phase à k .

\hat{V}_3 describes the decomposition of a single phonon into two phonons, as well as the inverse process where two phonons combine into one. Note that it is momentum-conserving, reflecting the homogeneity of the system, and that if p is held fixed and $q \rightarrow 0$, $V_3(p, q)$ vanishes as \sqrt{q} . A similar writing of \hat{V}_3 can be found in [13], in a form that is equivalent to the one given here. \hat{V}_3 décrit la décomposition d'un phonon unique en deux phonons, ainsi que le processus inverse où deux phonons se combinent en un seul. Il est à noter qu'il conserve la quantité de mouvement, reflétant l'homogénéité du système, et que si p est maintenu fixe et $q \rightarrow 0$, $V_3(p, q)$ tend vers zéro comme \sqrt{q} . Une écriture similaire de \hat{V}_3 peut être trouvée dans [13], sous une forme équivalente à celle donnée ici.

Nous nous tournons maintenant vers \hat{H}_3 , en négligeant les termes de la forme $\hat{\varphi}^\dagger \hat{\varphi}^\dagger \hat{\varphi}^\dagger$ ($\hat{\varphi} \hat{\varphi} \hat{\varphi}$) qui causent l'apparition (ou la disparition) déséquilibrée de trois phonons, typiquement associée à une forte violation de la conservation de l'énergie. La partie pertinente de l'Hamiltonien d'interaction prend la forme

$$\hat{V}_3 = \frac{1}{\sqrt{N_{\text{at}}}} \sum_{\substack{p, q \neq 0 \\ p+q \neq 0}} \hbar V_3(p, q) \left\{ \hat{\varphi}_p^\dagger \hat{\varphi}_q^\dagger \hat{\varphi}_{p+q} + \hat{\varphi}_{p+q}^\dagger \hat{\varphi}_p \hat{\varphi}_q \right\},$$

où

$$V_3(p, q) = \sqrt{\frac{\hbar}{32m}} \sqrt{\frac{|pq(p+q)|}{v_p^{\text{ph}} v_q^{\text{ph}} v_{p+q}^{\text{ph}}}} \left\{ - \left(\frac{\hbar}{2m} \right)^2 [p^2 + pq + q^2] \right. \\ \left. + v_p^{\text{ph}} v_q^{\text{ph}} + v_q^{\text{ph}} v_{p+q}^{\text{ph}} + v_p^{\text{ph}} v_{p+q}^{\text{ph}} \right\}, \quad (21)$$

et $v_k^{\text{ph}} = \omega_k/k$ est la vitesse de phase à k . L'Hamiltonien \hat{V}_3 décrit la décomposition d'un phonon unique en deux phonons, ainsi que le processus inverse où deux phonons se combinent pour former un seul phonon. Il est à noter qu'il conserve le moment, ce qui reflète l'homogénéité du système, et que si p est fixé et $q \rightarrow 0$, $V_3(p, q)$ tend vers zéro comme \sqrt{q} . Une formulation similaire de \hat{V}_3 peut être trouvée dans [13], sous une forme équivalente à celle donnée ici.

E. Numerical simulations

Simulations numériques

The system is modeled numerically using the truncated Wigner approximation (TWA) [29] (see Appendix C for more details). The operators $\hat{\Psi}$ are replaced by classical variables Ψ , and products of these variables are identified with the corresponding fully symmetrized quantum operators. A series of *ab initio* Monte Carlo simulations are performed, with quantum indeterminacy appearing through the statistical ensemble describing the initial state. The field is then evolved according to the dynamics of Hamiltonian (1). This is repeated for a large number of independent initial realisations, so as to get good statistics when computing averages.

The phenomenon of interest is illustrated in Fig. 1, which shows the typical evolution observed in numerical simulations. Starting from a thermal state, the occupation number of a mode is increased by δn , to be considered as the initial number of phonons in the probe. Indeed, throughout this paper we shall adopt the following decomposition:

$$n_k = n_k^{\text{th}} + \delta n_k, \quad (22)$$

where $n_k = \langle \hat{\varphi}_k^\dagger \hat{\varphi}_k \rangle$ is the full phonon spectrum and n_k^{th} is the (thermal) spectrum in the absence of any probe. Further details of the simulations are given in Sec. IV. The figure shows n_k at a series of different times. We see that n_k^{probe} , the population of the initially occupied mode, decays. The observed behaviour is essentially linear, in the sense that the relative change $\delta n_k(t)/\delta n$ is independent of δn , the number of phonons injected in the probe. Moreover, it is clear that the spectrum of the probe has broadened. Summing over nearby modes (those within the vertical dotted lines), the total n is found to be constant in time. The phonons seem not to have been lost, but rather to have been kicked into neighboring modes. Note that the broadening is essentially symmetric: the phonons are just as likely to be kicked towards a higher momentum as a lower one. The small shifts in momentum suggest that the evolution is primarily driven by interactions with infrared phonons from the thermal bath.

Le système est modélisé numériquement en utilisant l'approximation de Wigner tronquée (TWA) [29] (voir l'Annexe C pour plus de détails). Les opérateurs $\hat{\Psi}$ sont remplacés par des variables classiques Ψ , et les produits de ces variables sont identifiés avec les opérateurs quantiques symétrisés correspondants. Une série de simulations Monte Carlo *ab initio* est effectuée, avec l'indétermination quantique apparaissant à travers l'ensemble statistique décrivant l'état initial. Le champ est ensuite évolué selon la dynamique de l'Hamiltonien (1). Cela est répété pour un grand nombre de réalisations initiales indépendantes, afin d'obtenir de bonnes statistiques lors du calcul des moyennes. Le phénomène d'intérêt est illustré dans la Fig. 1, qui montre l'évolution typique observée dans les simulations numériques. En partant d'un état thermique, le nombre de particules d'un mode est augmenté de δn , considéré comme le nombre initial de phonons dans la sonde. En effet, tout au long de cet article, nous adopterons la décomposition suivante :

$$n_k = n_k^{\text{th}} + \delta n_k,$$

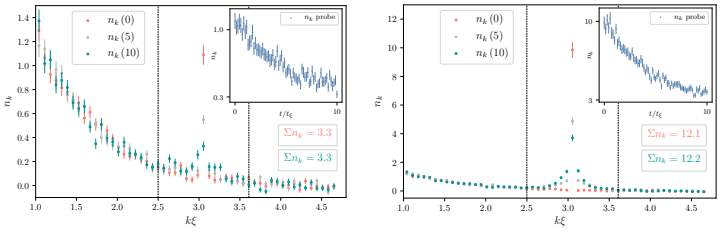


Figure 1. Snapshots of the phonon number spectrum at times $t/t_\xi = 0, 5, 10$ assuming an initial thermal state at temperature $k_B T / mc^2 = 2$ on top of which an additional $\delta n = 1$ (top) or $\delta n = 10$ (bottom) phonons are added in the mode $k\xi = 3.1$. The size of the system is $L/\xi = 90.5$ and its atomic density $\rho_0 \xi = 49.9$. The averages are calculated from an ensemble of 400 independent realisations, while the error bars represent the standard deviation. There are 256 points on the grid. The inset shows a more time-resolved evolution of the population in the probe mode. The total number of phonons within the vertical dotted lines is given for the initial and final times, and is seen to be conserved.

où $n_k = \langle \hat{\varphi}_k^\dagger \hat{\varphi}_k \rangle$ est le spectre complet des phonons et n_k^{th} est le spectre (thermique) en l'absence de toute sonde. Des détails supplémentaires sur les simulations sont donnés dans la Sec. IV. La figure montre n_k à une série de temps différents. Nous voyons que n_k^{probe} , la population du mode initialement occupé, diminue. Le comportement observé est essentiellement linéaire, dans le sens où le changement relatif $\delta n_k(t)/\delta n$ est indépendant de δn , le nombre de phonons injectés dans la sonde. De plus, il est clair que le spectre de la sonde s'est élargi. En sommant sur les modes voisins (ceux à l'intérieur des lignes verticales pointillées), le total n est constant dans le temps. Les phonons ne semblent pas avoir été perdus, mais plutôt avoir été transférés vers des modes voisins. Notez que l'élargissement est essentiellement symétrique : les phonons ont autant de chances d'être transférés vers un moment plus élevé qu'un moment plus bas. Les petits décalages en moment suggèrent que l'évolution est principalement entraînée par les interactions avec les phonons infrarouges du bain thermique.

Des instantanés du spectre du nombre de phonons aux temps $t/t_\xi = 0, 5, 10$, en supposant un état thermique initial à une température $k_B T / mc^2 = 2$, auquel sont ajoutés $\delta n = 1$ (en haut) ou $\delta n = 10$ (en bas) phonons dans le mode $k\xi = 3.1$. La taille du système est $L/\xi = 90.5$ et sa densité atomique est $\rho_0 \xi = 49.9$. Les moyennes sont calculées à partir d'un ensemble de 400 réalisations indépendantes, tandis que les barres d'erreur représentent l'écart-type. Il y a 256 points sur la grille. L'encadré montre une évolution plus résolue dans le temps de la population dans le mode de la sonde. Le nombre total de phonons à l'intérieur des lignes verticales pointillées est donné pour les temps initial et final, et il est observé qu'il est conservé.

Le système est modélisé numériquement en utilisant l'approximation de Wigner tronquée (TWA). Les opérateurs quantiques $\hat{\Psi}$ sont remplacés par des variables classiques Ψ , et les produits de ces variables sont identifiés avec les opérateurs quantiques

symétrisés correspondants. Des simulations Monte Carlo *ab initio* sont réalisées, avec l'indétermination quantique apparaissant à travers l'ensemble statistique décrivant l'état initial. Le champ est ensuite évolué selon la dynamique de l'Hamiltonien (??). Ce processus est répété pour un grand nombre de réalisations initiales indépendantes afin d'obtenir une bonne statistique pour les moyennes. Les résultats de ces simulations sont illustrés dans la Figure ??, montrant l'évolution typique observée. En partant d'un état thermique, le nombre de phonons dans un mode donné est augmenté de δn , ce qui est considéré comme le nombre initial de phonons dans la sonde. On utilise la décomposition suivante :

$$n_k = n_k^{\text{th}} + \delta n_k, ,$$

où $n_k = \langle \hat{\varphi} k^\dagger \hat{\varphi} k \rangle$ est le spectre complet des phonons et n_k^{th} est le spectre thermique en l'absence de sonde. Les détails des simulations sont donnés en Sec. ??. La figure montre n_k à différents moments. On observe que n_k^{probe} , la population du mode initialement occupé, décroît. Le comportement observé est essentiellement linéaire, en ce sens que le changement relatif $\delta n_k(t)/\delta n$ est indépendant de δn , le nombre de phonons injectés dans la sonde. De plus, il est clair que le spectre de la sonde s'est élargi. En sommant les modes voisins (ceux à l'intérieur des lignes verticales pointillées), le nombre total de phonons est trouvé constant dans le temps. Les phonons semblent ne pas avoir été perdus, mais plutôt avoir été transférés vers des modes voisins. Il est à noter que l'élargissement est essentiellement symétrique : les phonons sont tout aussi susceptibles d'être transférés vers un moment plus élevé qu'un moment plus bas. Les petits décalages de moment suggèrent que l'évolution est principalement entraînée par les interactions avec les phonons infrarouges du bain thermique.

III. THEORETICAL DERIVATION

Développement théorique

In this section, we provide an analytical description of the dissipative process at play. In Sec. III A, we derive the decay rate using the FGR; as this applies to a discrete eigenstate in interaction with a continuum, we expect it to describe the decay of the population of a singularly occupied mode (*i.e.*, when the spectrum is sufficiently narrow). In Sec. III B, we sketch a formalism which takes us beyond such an approximation, using response functions to model the dynamics of the spectrum more accurately and to describe the deviations from purely exponential behavior. A fuller treatment of this formalism is given in Appendix A, while a comprehensive analysis of the main deviations observed numerically is given in Appendix B. Dans cette section, nous fournissons une description analytique du processus dissipatif en jeu. Dans la section III A, nous dérivons le taux de décroissance en utilisant la règle d'or de Fermi (FGR) ; comme cela s'applique à un état propre discret en interaction avec un continuum, nous nous attendons à ce qu'elle décrive la décroissance de la population d'un mode singulièrement occupé

(c'est-à-dire lorsque le spectre est suffisamment étroit). Dans la section III B, nous esquissons un formalisme qui va au-delà de cette approximation, en utilisant des fonctions de réponse pour modéliser la dynamique du spectre plus précisément et pour décrire les déviations par rapport à un comportement purement exponentiel. Un traitement plus complet de ce formalisme est donné en Annexe A, tandis qu'une analyse détaillée des principales déviations observées numériquement est fournie en Annexe B.

A. Using the Fermi Golden Rule

Utilisation de la règle d'or de Fermi

Given an initial eigenstate $|i\rangle$ that couples to a continuum of final eigenstates $|\nu_f\rangle$ via a time-independent weak perturbation Hamiltonian \hat{H}' , the FGR gives the transition rate into the continuum [30]:

$$dW_{i \rightarrow f} = \frac{2\pi}{\hbar} \left| \langle \nu_f | \hat{H}' | i \rangle \right|^2 \delta(E_f - E_i) d\nu_f, \quad (23)$$

where E_i and E_f are the energies of the initial and final states, the latter being labeled by the dimensionless number ν_f . The Dirac delta imposes energy conservation, *i.e.* the rate is evaluated as a matrix element between states with the same (unperturbed) energy. The total rate is found by integrating over ν_f . Étant donné un état propre initial $|i\rangle$ qui couple à un continuum d'états propres finaux $|\nu_f\rangle$ via un Hamiltonien de perturbation faible et indépendant du temps \hat{H}' , la règle d'or de Fermi (FGR) donne le taux de transition dans le continuum [30] :

$$dW_{i \rightarrow f} = \frac{2\pi}{\hbar} \left| \langle \nu_f | \hat{H}' | i \rangle \right|^2 \delta(E_f - E_i) d\nu_f,$$

où E_i et E_f sont les énergies des états initial et final, ce dernier étant étiqueté par le nombre sans dimension ν_f . Le delta de Dirac impose la conservation de l'énergie, c'est-à-dire que le taux est évalué comme un élément de matrice entre des états ayant la même énergie (non perturbée). Le taux total est trouvé en intégrant sur ν_f .

Consider a singularly occupied phonon mode k , which decays due to interactions with the thermal population of phonons. Each available momentum is a multiple of $2\pi/L$. For sufficiently large kL the thermal distribution of phonons, as well as the states available to decay to, can be approximated as a continuum from the point of view of the mode k . We thus expect the FGR (23) to be applicable. We need only determine the relevant perturbation and corresponding initial and final states. Considérons un mode de phonon singulièrement occupé k , qui se désintègre en raison des interactions avec la population thermique de phonons. Chaque moment disponible est un multiple de $2\pi/L$. Pour un kL suffisamment grand, la distribution thermique de phonons, ainsi que les états disponibles pour la désintégration, peuvent être approximatés comme un continuum du point de vue du mode k . Nous nous attendons donc à ce que

la règle d'or de Fermi (FGR) (23) soit applicable. Il nous suffit de déterminer la perturbation pertinente et les états initiaux et finaux correspondants.

In the phonon basis, the role of the perturbation entering the FGR is played by \hat{V}_3 of Eq. (18). It describes two distinct processes involving the annihilation of a probe phonon at wave number k :

- Phonons of wave-numbers k and q combine to produce a single phonon with wave number $k + q$. The relevant term is $\hat{\varphi}_{k+q}^\dagger \hat{\varphi}_q \hat{\varphi}_k$, taking $|i\rangle = |n_q, n_k, n_{k+q}\rangle$ to $|f\rangle = |n_q - 1, n_k - 1, n_{k+q} + 1\rangle$, and the corresponding squared matrix element is $\frac{1}{N_{\text{at}}} \hbar^2 |2 V_3(k, q)|^2 \times n_k n_q (n_{k+q} + 1)$. The factor 2 comes from the $p \leftrightarrow q$ symmetry of the sum in Eq. (18). The factors of n come from the action of the phonon operators on the state $|i\rangle$, with the 'n' and '+1' terms encoding stimulated and spontaneous processes, respectively. The energy difference is $E_f - E_i = \hbar \delta \omega_L = \hbar (\omega_{k+q} - \omega_q - \omega_k)$; at small q , this gives $\delta \omega_L \approx v_k^{\text{gr}} q - c|q| = q(v_k^{\text{gr}} \mp c)$, where $v_k^{\text{gr}} = d\omega_k/dk$ is the group velocity at k .
- k decays to two phonons, with wave numbers $k - q$ and q . The relevant term is $\hat{\varphi}_{k-q}^\dagger \hat{\varphi}_q^\dagger \hat{\varphi}_k$, taking $|i\rangle = |n_q, n_{k-q}, n_k\rangle$ to $|f\rangle = |n_q + 1, n_{k-q} + 1, n_k - 1\rangle$, and the corresponding squared matrix element is $\frac{1}{N_{\text{at}}} \hbar^2 |2 V_3(k - q, q)|^2 \times n_k (n_q + 1) (n_{k-q} + 1)$. The energy difference is $E_f - E_i = \hbar \delta \omega_B = \hbar (\omega_{k-q} + \omega_q - \omega_k)$; at small q , this gives $\delta \omega_B \approx -q(v_k^{\text{gr}} \mp c) = -\delta \omega_L$.

The two scattering processes described above lead to the well-known Landau and Beliaev damping of phonons in BECs: the former is associated with the absorption of a thermal phonon, the latter with a splitting into two phonons.³ The associated frequency differences are thus labeled $\delta \omega_L$ and $\delta \omega_B$. Dans la base des phonons, le rôle de la perturbation entrant dans la règle d'or de Fermi (FGR) est joué par \hat{V}_3 de l'équation (??). Il décrit deux processus distincts impliquant l'annihilation d'un phonon de la sonde au nombre d'onde k :

- Les phonons des nombres d'onde k et q se combinent pour produire un seul phonon avec le nombre d'onde $k + q$. Le terme pertinent est $\hat{\varphi}_{k+q}^\dagger \hat{\varphi}_q \hat{\varphi}_k$, passant de $|i\rangle = |n_q, n_k, n_{k+q}\rangle$ à $|f\rangle = |n_q - 1, n_k - 1, n_{k+q} + 1\rangle$. L'élément de matrice au carré correspondant est $\frac{1}{N_{\text{at}}} \hbar^2 |2 V_3(k, q)|^2 \times n_k n_q (n_{k+q} + 1)$. Le facteur 2 provient de la symétrie $p \leftrightarrow q$ de la somme dans l'équation (??). Les facteurs de n proviennent de l'action des opérateurs de phonons sur l'état $|i\rangle$, les termes n et '+1' encodant respectivement les processus stimulés et spontanés. La différence d'énergie est $E_f - E_i = \hbar \delta \omega_L = \hbar (\omega_{k+q} - \omega_q - \omega_k)$;

³ While Beliaev damping often occurs spontaneously (as it is a $1 \rightarrow 2$ -phonon process), we are concerned here with a regime of sufficiently high temperature in which it is stimulated by the presence of a large number of thermal infrared phonons.

pour q petit, cela donne $\delta\omega_L \approx v_k^{\text{gr}} q - c|q| = q(v_k^{\text{gr}} \mp c)$, où $v^{\text{gr}}k = d\omega k/dk$ est la vitesse de groupe en k .

- k se désintègre en deux phonons, avec les nombres d'onde $k - q$ et q . Le terme pertinent est $\hat{\phi}k - q^\dagger \hat{\phi}q^\dagger \hat{\phi}k$, passant de $|i\rangle = |nq, n_{k-q}, n_k\rangle$ à $|f\rangle = |n_q + 1, n_{k-q} + 1, n_k - 1\rangle$, et l'élément de matrice au carré correspondant est $\frac{1}{N_{\text{at}}} \hbar^2 |2, n_k(n_q + 1)(n_{k-q} + 1)|$. La différence d'énergie est $E_f - E_i = \hbar, \delta\omega_B = \hbar(\omega_{k-q} + \omega_q - \omega_k)$; pour q petit, cela donne $\delta\omega_B \approx -q(v^{\text{gr}}k \mp c) = -\delta\omega_L$.

Les deux processus de diffusion décrits ci-dessus conduisent aux atténuations bien connues de Landau et de Beliaev des phonons dans les BECs : le premier est associé à l'absorption d'un phonon thermique, le second à une scission en deux phonons.⁴ Les différences de fréquence associées sont ainsi étiquetées $\delta\omega_L$ et $\delta\omega_B$.

Imposing also energy conservation $\delta\omega_{L,B}(q; k) = 0$, as demanded by the Dirac delta in Eq. (23), we encounter the problem anticipated above: the only exact solution respecting both momentum *and* energy conservation is the trivial one, $q = 0$.⁵ In the vicinity of this channel, the interaction vanishes as $|V_3(k, q)|^2 \propto q$. However, the frequency ω_q also vanishes as $\omega_q \rightarrow c|q|$, and therefore the thermal population $n_q \approx k_B T / \hbar \omega_q$ simultaneously diverges. It therefore makes sense to refer to the product $|V_3(k, q)|^2 n_q$ as an effective interaction strength for the corresponding channel, as it has a finite limit as $q \rightarrow 0$, and it is this finite limit that is picked up by the Dirac delta in Eq. (23). Although the delta is centered at $q = 0$, it does not actually matter that the trivial elastic channel is unphysical and thus removed from the dynamics: the Dirac delta is a placeholder for a steadily narrowing distribution of final states in the vicinity of this channel (see Appendix A). As long as there are sufficiently many modes within this distribution, the single removed mode at $q = 0$ has a relative measure of zero. This allows the application of the FGR, as if the limiting processes at $q \rightarrow 0^\pm$ were physically allowed.⁶ En imposant également la conservation de l'énergie $\delta\omega_{L,B}(q; k) = 0$, comme l'exige la fonction delta de Dirac dans l'équation (??), nous rencontrons le problème anticipé ci-dessus : la seule solution exacte respectant à la fois la conservation de la quantité de mouvement *et* de l'énergie est celle triviale, $q = 0$.⁷ À proximité de ce canal, l'interaction s'annule comme $|V_3(k, q)|^2 \propto q$. Cependant, la fréquence ω_q s'annule également comme $\omega_q \rightarrow c|q|$, et donc la population thermique $n_q \approx k_B T / \hbar \omega_q$ diverge simultanément. Il est donc logique de se référer au produit $|V_3(k, q)|^2 n_q$ comme une force

⁴ Alors que l'atténuation de Beliaev se produit souvent spontanément (car il s'agit d'un processus $1 \rightarrow 2$ phonons), nous nous intéressons ici à un régime de température suffisamment élevée où il est stimulé par la présence d'un grand nombre de phonons infrarouges thermiques.

⁵ Notice that for a Luttinger liquid where $\xi = 0$ the problem is exactly the opposite as there is an infinite number of elastic scattering channels leading to a divergence of the FGR decay rate [31, 32].

⁶ One important modification with respect to the standard case arises because the effective interaction strength $|V_3(k, q)|^2 n_q$ is discontinuous across the resonant channel $q = 0$. This obliges us to consider separately the coupling with positive and negative q , leading to two applications of

d'interaction effective pour le canal correspondant, car il a une limite finie lorsque $q \rightarrow 0$, et c'est cette limite finie qui est captée par la fonction delta de Dirac dans l'équation (??). Bien que la delta soit centrée sur $q = 0$, il importe peu que le canal élastique trivial soit non physique et donc supprimé de la dynamique : la fonction delta de Dirac est un substitut pour une distribution de plus en plus étroite d'états finaux à proximité de ce canal (voir l'Annexe ??). Tant qu'il y a suffisamment de modes dans cette distribution, le mode unique supprimé à $q = 0$ a une mesure relative de zéro. Cela permet l'application de la règle d'or de Fermi, comme si les processus limitants à $q \rightarrow 0^\pm$ étaient physiquement autorisés.⁸

The Dirac delta in Eq. (23) also serves to multiply by the density of available states with respect to the energy. Since the available states are evenly spaced in momentum, in a window of size Δq the number of states is $\Delta N = L/2\pi \times \Delta q \approx L/2\pi\hbar \times |dq/d(\delta\omega_{L/B})| \times \Delta E$. Recalling from above that, for small q , we have $\delta\omega_L \approx -\delta\omega_B \approx q(v_k^{\text{gr}} \mp c)$, this gives the following density of states:

$$\rho_E(k) \equiv \frac{dN}{dE} = \frac{1}{2\pi\hbar} \frac{L}{v_k^{\text{gr}} \mp c}. \quad (24)$$

For ease of notation we restrict to $k > 0$, so that $v_k^{\text{gr}} \mp c$ is always positive. Due to isotropy, $-k$ will behave in exactly the same way as k . We recall that the sign in the denominator is related to the sign of q , the momentum of the infrared phonons with which the interaction takes place. The abrupt change in the velocity as q crosses zero means that there are two distinct limits to be taken: $q \rightarrow 0$ from above and from below. Putting everything together, we thus derive two decay rates, corresponding to the coupling with positive and negative q :

$$\Gamma_\pm(k) = \frac{\lim_{q \rightarrow 0^\pm} \left[\frac{L}{N_{\text{at}}} |2V_3(k, q)|^2 n_q \right]}{(v_k^{\text{gr}} \mp c)}. \quad (25)$$

Inserting the explicit form of $V_3(k, q)$ yields:

$$\Gamma_\pm t_\xi = \frac{k_B T}{mc^2} \frac{1}{\rho_0 \xi} f_\pm(k\xi) \quad (26)$$

where

$$f_\pm(k\xi) = \frac{1}{2} \frac{(k\xi)^2}{(v_k^{\text{ph}}/c)^2} \frac{(v_k^{\text{ph}}/c \pm 1/2)^2}{v_k^{\text{gr}}/c \mp 1}. \quad (27)$$

The total decay rate presented in the figures is $\Gamma_k = \Gamma_+ + \Gamma_-$. La fonction delta de Dirac dans l'équation (??) sert également

⁸ Une modification importante par rapport au cas standard survient car la force d'interaction effective $|V_3(k, q)|^2 n_q$ est discontinue de part et d'autre du canal résonant $q = 0$. Cela nous oblige à considérer séparément le couplage avec q positif et négatif, conduisant à deux applications de la règle d'or de Fermi où chaque canal capte la "moitié" de la fonction delta de Dirac autour de $q = 0$. Ce facteur est annulé par la co-occurrence des canaux de Landau et de Beliaev, qui contribuent de manière égale au taux de décroissance.

à multiplier par la densité d'états disponibles par rapport à l'énergie. Puisque les états disponibles sont régulièrement espacés en quantité de mouvement, dans une fenêtre de taille Δq , le nombre d'états est $\Delta N = L/2\pi \times \Delta q \approx L/2\pi\hbar \times |dq/d(\delta\omega_{L/B})| \times \Delta E$. En rappelant ci-dessus que, pour de petits q , nous avons $\delta\omega_L \approx -\delta\omega_B \approx q(v^{\text{gr}}k \mp c)$, cela donne la densité d'états suivante :

$$\rho E(k) \equiv \frac{dN}{dE} = \frac{1}{2\pi\hbar} \frac{L}{v^{\text{gr}}k \mp c}.$$

Pour simplifier la notation, nous restreignons à $k > 0$, de sorte que $v^{\text{gr}}k \mp c$ est toujours positif. En raison de l'isotropie, $-k$ se comportera exactement de la même manière que k . Nous rappelons que le signe dans le dénominateur est lié au signe de q , la quantité de mouvement des phonons infrarouges avec lesquels l'interaction a lieu. Le changement brusque de la vitesse lorsque q traverse zéro signifie qu'il y a deux limites distinctes à prendre : $q \rightarrow 0$ par le haut et par le bas. En rassemblant tout, nous dérivons ainsi deux taux de décroissance, correspondant au couplage avec des q positifs et négatifs :

$$\Gamma_{\pm}(k) = \frac{\lim_{q \rightarrow 0^{\pm}} \left[\frac{L}{N_{\text{at}}} |2, V_3(k, q)|^2 n_q \right]}{(v^{\text{gr}}k \mp c)},.$$

En insérant la forme explicite de $V_3(k, q)$, nous obtenons :

$$\Gamma_{\pm} t_{\xi} = \frac{k_B T}{mc^2} \frac{1}{\rho_0 \xi} f_{\pm}(k\xi)$$

où

$$f_{\pm}(k\xi) = \frac{1}{2}, \frac{(k\xi)^2}{(v^{\text{ph}}k/c)^2}, \frac{(v^{\text{ph}}k/c \pm 1/2)^2}{v^{\text{gr}}k/c \mp 1},.$$

Le taux de décroissance total présenté dans les figures est $\Gamma k = \Gamma_+ + \Gamma_-$.

Equation (III A) constitutes our main result: an explicit expression for the rate at which the population of a singularly populated mode will decay, due to interactions with thermal phonons. L'équation (??) constitue notre principal résultat : une expression explicite du taux auquel la population d'un mode singulièrement peuplé va décroître en raison des interactions avec les phonons thermiques.

Since at small q we have $\delta\omega_L \approx -\delta\omega_B \approx v_k^{\text{gr}}q - c|q|$, at fixed $|q|$ the magnitude of the frequency difference is smaller when q has the same sign as k . We therefore expect that such modes provide the dominant contribution to the decay, while the coupling to modes of the opposite sign induces a subdominant correction. So the Landau channel tends to kick a phonon at k into a higher-energy mode, while the Beliaev channel acts in the opposite direction. Since the two channels contribute equally, there is no preference for a phonon at k to be kicked towards either a higher or lower momentum. This explains the approximate symmetry of the broadening observed in Fig. 1. Puisque pour de petites valeurs de q nous avons $\delta\omega_L \approx -\delta\omega_B \approx v_k^{\text{gr}}q - c|q|$, à $|q|$ fixe, l'amplitude de la

différence de fréquence est plus faible lorsque q a le même signe que k . Nous nous attendons donc à ce que ces modes fournissent la contribution dominante à la décroissance, tandis que le couplage avec les modes de signe opposé induit une correction subordonnée. Ainsi, le canal de Landau tend à propulser un phonon en k vers un mode d'énergie supérieure, tandis que le canal de Beliaev agit dans la direction opposée. Étant donné que les deux canaux contribuent de manière égale, il n'y a aucune préférence pour qu'un phonon en k soit propulsé vers un momentum supérieur ou inférieur. Cela explique la symétrie approximative de l'élargissement observée dans la Fig. 1.

Nous considérons un mode phonon singulièrement occupé k , qui se désintègre en raison des interactions avec la population thermique de phonons. Pour un grand kL , la distribution thermique des phonons ainsi que les états disponibles pour la désintégration peuvent être approximés comme un continuum du point de vue du mode k . La règle d'or de Fermi (FGR) est donc applicable. Nous déterminons la perturbation pertinente ainsi que les états initiaux et finaux correspondants. Dans la base des phonons, le rôle de la perturbation dans la FGR est joué par \hat{V}_3 , qui décrit deux processus distincts :

- Les phonons de nombres d'ondes k et q se combinent pour produire un seul phonon avec le nombre d'onde $k + q$. Le terme pertinent est $\hat{\varphi}_{k+q}^\dagger \hat{\varphi}_q \hat{\varphi}_k$, avec une différence d'énergie $\delta\omega_L$, et l'élément de matrice au carré est $\frac{1}{N_{\text{at}}} \hbar^2 |2 V_3(k, q)|^2 \times n_k n_q (n_{k+q} + 1)$.
- k se désintègre en deux phonons, avec les nombres d'ondes $k - q$ et q . Le terme pertinent est $\hat{\varphi}_{k-q}^\dagger \hat{\varphi}_q^\dagger \hat{\varphi}_k$, avec une différence d'énergie $\delta\omega_B$, et l'élément de matrice au carré est $\frac{1}{N_{\text{at}}} \hbar^2 |2 V_3(k - q, q)|^2 \times n_k (n_q + 1) (n_{k-q} + 1)$.

En imposant la conservation de l'énergie $\delta\omega_{L,B}(q; k) = 0$, nous rencontrons un problème : la seule solution exacte respectant à la fois la conservation de l'énergie et du momentum est le cas trivial $q = 0$. Cependant, ce canal élastique trivial est retiré de la dynamique car l'interaction efficace $|V_3(k, q)|^2 n_q$ a une limite finie lorsque $q \rightarrow 0$. La règle d'or de Fermi est ainsi appliquée en considérant que les processus limites à $q \rightarrow 0^\pm$ sont physiquement permis. La fonction delta de Dirac dans la FGR sert également à multiplier par la densité des états disponibles en énergie. Comme les états disponibles sont uniformément espacés en momentum, la densité des états est :

$$\rho_E(k) \equiv \frac{dN}{dE} = \frac{1}{2\pi\hbar} \frac{L}{v_k^{\text{gr}} \mp c}.$$

En combinant tout, nous dérivons deux taux de désintégration, correspondant aux couplages avec des q positifs et négatifs :

$$\Gamma_\pm(k) = \frac{\lim_{q \rightarrow 0^\pm} \left[\frac{L}{N_{\text{at}}} |2 V_3(k, q)|^2 n_q \right]}{(v_k^{\text{gr}} \mp c)}.$$

En insérant la forme explicite de $V_3(k, q)$, nous obtenons :

$$\Gamma_\pm t_\xi = \frac{k_B T}{mc^2} \frac{1}{\rho_0 \xi} f_\pm(k\xi)$$

où

$$f_{\pm}(k\xi) = \frac{1}{2} \frac{(k\xi)^2}{\left(v_k^{\text{ph}}/c\right)^2} \frac{\left(v_k^{\text{ph}}/c \pm 1/2\right)^2}{v_k^{\text{gr}}/c \mp 1}.$$

Le taux de désintégration total est donné par $\Gamma_k = \Gamma_+ + \Gamma_-$.

B. Using response functions

Utilisation des fonctions de réponse

Here, we give the outline of a more precise description of how a perturbation to the phonon number spectrum δn_k evolves. For brevity, we shall avoid details here, though they can be found in Appendix A. However, the writing will allow us to point out significant deviations from the exponential decay rate of Eq. (III A), which are studied in more detail in Appendix B. Dans cette section, nous donnons un aperçu d’une description plus précise de l’évolution d’une perturbation dans le spectre de nombre de phonons δn_k . Par souci de concision, nous éviterons les détails ici, bien qu’ils puissent être trouvés dans l’Annexe ??.

Cependant, cette présentation nous permettra de souligner des écarts significatifs par rapport au taux de décroissance exponentielle de l’équation (??), qui sont étudiés plus en détail dans l’Annexe ??.

The time-derivative of δn_k depends both on δn_k and on phonon correlations, the most important of which are the 3-point correlations induced by the Beliaev and Landau processes described above: $C_{p,q}^{(3)} = \langle \hat{\phi}_p^\dagger \hat{\phi}_q^\dagger \hat{\phi}_{p+q} \rangle$. Neglecting other connected correlation functions, the equations of motion can be written entirely in terms of n_k , though they become integro-differential equations that include response functions, which encode how the system “remembers” and responds to its past behaviour. Neglecting any self-interaction of the perturbation, we linearize the equations in δn_{k+q} :

$$\begin{aligned} \partial_t (\delta n_k) = & - \int_0^t dt' D_k(t-t') \delta n_k(t') \\ & + \int_0^t dt' \sum_{q \neq -k} M_{k,k+q}(t-t') \delta n_{k+q}(t'). \end{aligned} \quad (28)$$

La dérivée temporelle de δn_k dépend à la fois de δn_k et des corrélations des phonons, dont les plus importantes sont les corrélations à trois points induites par les processus de Beliaev et Landau décrits ci-dessus : $C^{(3)}_{p,q} = \langle \hat{\phi}_p^\dagger \hat{\phi}_q^\dagger \hat{\phi}_{p+q} \rangle$. En négligeant d’autres fonctions de corrélation connectées, les équations du mouvement peuvent être écrites entièrement en termes de n_k , bien qu’elles deviennent des équations intégrales-différentielles qui incluent des fonctions de réponse, lesquelles codifient comment le système “se souvient” et réagit à son comportement passé. En négligeant toute auto-interaction de la

perturbation, nous linéarisons les équations en δn_{k+q} :

$$\begin{aligned} \partial_t (\delta n_k) = & - \int_0^t dt' D_k(t-t') \delta n_k(t') \\ & + \int_0^t dt' \sum_{q \neq -k} M_{k,k+q}(t-t') \delta n_{k+q}(t') . \end{aligned}$$

The first term on the right-hand side of (28), governed by D_k , describes the essential behavior of δn_k , when back-reaction from its near-neighbors can be neglected and finite-size effects can also be ignored (see below). $D_k(t-t')$ decays to zero within a time scale t_{crit} , and its integral approaches Γ_k of Eq. (III A). As long as $t \gg t_{\text{crit}}$ and δn_k does not vary much on time scales of order t_{crit} , its value at t can be taken outside the integral, and the first term becomes $-\Gamma_k \delta n_k$. This is the regime in which the analysis of the previous subsection applies. We can thus expect deviations when δn_k varies significantly over times of order t_{crit} . This can happen either when $\partial_t (\delta n_k)$ is particularly large, or for a time t_{crit} after a sudden injection of phonons. The latter case will be relevant for some of our numerical simulations. Le premier terme du côté droit de (??), régi par D_k , décrit le comportement essentiel de δn_k , lorsque la réaction des voisins proches peut être négligée et que les effets de taille finie peuvent également être ignorés (voir ci-dessous). $D_k(t-t')$ décroît vers zéro dans un temps caractéristique t_{crit} , et son intégrale tend vers Γ_k de l'Eq. (??). Tant que $t \gg t_{\text{crit}}$ et que δn_k ne varie pas beaucoup sur des échelles de temps de l'ordre de t_{crit} , sa valeur à t peut être prise en dehors de l'intégrale, et le premier terme devient $-\Gamma_k \delta n_k$. C'est le régime dans lequel l'analyse de la sous-section précédente s'applique. Nous pouvons donc nous attendre à des écarts lorsque δn_k varie de manière significative sur des périodes de l'ordre de t_{crit} . Cela peut se produire soit lorsque $\partial_t (\delta n_k)$ est particulièrement grande, soit après un temps t_{crit} suivant une injection soudaine de phonons. Ce dernier cas sera pertinent pour certaines de nos simulations numériques.

The second term of Eq. (28), governed by $M_{k,k+q}$, describes the influence of other modes on the evolution of n_k . These typically act to slow down the decay of n_k , because some of the phonons in neighboring modes will be kicked into the mode of interest. This is particularly relevant in situations where k is at the center of a peak with a finite width, and in such a scenario we expect the net decay rate to be smaller than that predicted by Eq. (III A). Note that this is not in contradiction with the FGR, which applies when a single discrete mode loses energy to a continuum of modes; this picture becomes less applicable when the mode losing energy is itself part of a continuum. Generally, the effect of this back-reaction term is difficult to take fully into account, but in certain situations (particularly in the case of parametric resonance) we can make approximations and derive the expected qualitative behavior for the decay rate as a function of the peak width. Le deuxième terme de l'Eq.(??), régi par $M_{k,k+q}$, décrit l'influence des autres modes sur l'évolution de n_k . Ceux-ci agissent généralement pour ralentir la décroissance de n_k , car certains des phonons

dans les modes voisins seront transférés vers le mode d'intérêt. Cela est particulièrement pertinent dans les situations où k se trouve au centre d'un pic ayant une largeur finie ; dans un tel scénario, nous nous attendons à ce que le taux de décroissance net soit plus faible que celui prédit par l'Eq.(??). Notez que cela n'est pas en contradiction avec la règle d'or de Fermi (FGR), qui s'applique lorsqu'un seul mode discret perd de l'énergie au profit d'un continuum de modes ; cette description devient moins applicable lorsque le mode perdant de l'énergie fait lui-même partie d'un continuum. En général, il est difficile de prendre pleinement en compte l'effet de ce terme de rétroaction, mais dans certaines situations (en particulier dans le cas de la résonance paramétrique), nous pouvons faire des approximations et dériver le comportement qualitatif attendu pour le taux de décroissance en fonction de la largeur du pic.

Dans cette section, nous donnons un aperçu de la description plus précise de l'évolution d'une perturbation du spectre de nombre de phonons δn_k . Nous éviterons les détails ici, qui peuvent être trouvés en A, mais nous soulignerons les écarts significatifs par rapport au taux de décroissance exponentiel donné par l'Eq. (III A), étudiés en B. La dérivée temporelle de δn_k dépend à la fois de δn_k et des corrélations des phonons, notamment les corrélations à trois points induites par les processus de Beliaev et Landau. En négligeant d'autres fonctions de corrélations connectées, les équations de mouvement peuvent être écrites en termes de n_k , mais deviennent des équations intégral-différentielles incluant des fonctions de réponse qui codifient la manière dont le système « se souvient » et répond à son comportement passé. En négligeant toute auto-interaction de la perturbation, nous linéarisons les équations en δn_{k+q} :

$$\begin{aligned} \partial_t (\delta n_k) = & - \int_0^t dt' D_k(t-t') \delta n_k(t') \\ & + \int_0^t dt' \sum_{q \neq -k} M_{k,k+q}(t-t') \delta n_{k+q}(t') . \end{aligned} \quad (29)$$

Le premier terme sur le côté droit de l'Eq. (28), régi par D_k , décrit le comportement essentiel de δn_k , lorsque la rétroaction de ses voisins proches peut être négligée et que les effets de taille finie peuvent également être ignorés. $D_k(t-t')$ décroît à zéro dans un temps caractéristique t_{crit} , et son intégrale approche Γ_k de l'Eq. (III A). Tant que $t \gg t_{\text{crit}}$ et que δn_k ne varie pas beaucoup sur des échelles de temps de l'ordre de t_{crit} , sa valeur à t peut être prise en dehors de l'intégrale, et le premier terme devient $-\Gamma_k \delta n_k$. C'est le régime dans lequel l'analyse de la section précédente s'applique. Nous pouvons donc nous attendre à des écarts lorsque δn_k varie considérablement sur des temps de l'ordre de t_{crit} . Cela peut se produire soit lorsque $\partial_t (\delta n_k)$ est particulièrement grand, soit pour un temps t_{crit} après une injection soudaine de phonons. Ce dernier cas sera pertinent pour certaines de nos simulations numériques. Le deuxième terme de l'Eq. (28), régi par $M_{k,k+q}$, décrit l'influence des autres modes sur l'évolution de n_k . Ceux-ci agissent généralement pour ralentir la décroissance de n_k , car certains des phonons dans les modes voisins seront transférés vers le mode d'intérêt. Cela est

particulièrement pertinent dans les situations où k se trouve au centre d'un pic ayant une largeur finie ; dans un tel scénario, nous nous attendons à ce que le taux de décroissance net soit plus faible que celui prédit par l'Eq. (III A). Notez que cela n'est pas en contradiction avec la règle d'or de Fermi (FGR), qui s'applique lorsqu'un seul mode discret perd de l'énergie au profit d'un continuum de modes ; cette description devient moins applicable lorsque le mode perdant de l'énergie fait lui-même partie d'un continuum. En général, il est difficile de prendre pleinement en compte l'effet de ce terme de rétroaction, mais dans certaines situations (en particulier dans le cas de la résonance paramétrique), nous pouvons faire des approximations et dériver le comportement qualitatif attendu pour le taux de décroissance en fonction de la largeur du pic.

C. Suppression due to finite size of system

Suppression due à la taille finie du système

The second term of Eq. (28) includes a contribution from $q = 0$, which will be present even when neighboring modes are not significantly populated. This $q = 0$ term slows down the decay, but it is proportional to $1/L$ and vanishes entirely in the limit $L \rightarrow \infty$ where we have a continuum in k -space. The primary function of this term is to account for the finite resolution in k by keeping track of those phonons which, in the continuum limit, would be lost to nearby modes within $1/L$ of the main decaying mode. With the finite resolution induced by the finiteness of L , these phonons remain in the same bin as the decaying mode. Therefore, their contribution to the variation of n_k is removed, and the net decay is effectively slowed down. Le deuxième terme de l'équation (??) inclut une contribution provenant de $q = 0$, qui sera présente même lorsque les modes voisins ne sont pas significativement peuplés. Ce terme en $q = 0$ ralentit la décroissance, mais il est proportionnel à $1/L$ et disparaît complètement dans la limite où $L \rightarrow \infty$, où nous avons un continuum en espace k . La fonction principale de ce terme est de prendre en compte la résolution finie en k en tenant compte de ces phonons qui, dans la limite du continuum, seraient perdus vers des modes voisins dans une fenêtre de $1/L$ du mode principal en décroissance. Avec la résolution finie induite par la finitude de L , ces phonons restent dans le même compartiment que le mode en décroissance. Par conséquent, leur contribution à la variation de n_k est éliminée, et la décroissance nette est effectivement ralentie.

In Appendix B we derive the slowing down of the decay rate induced by this contribution. On sufficiently short time scales, it enters as a quadratic correction to the exponential decay:

$$\delta n_k \approx \delta n_k(0) e^{-\Gamma_k t + \frac{1}{2} \gamma_k t^2}, \quad (30)$$

where

$$\gamma_k t_\xi^2 = \frac{k_B T}{mc^2} \frac{1}{\rho_0 L} g(k\xi) \quad (31)$$

and where we have defined

$$g(k\xi) = \frac{(k\xi)^2}{v_{\text{ph}}^2(k)/c^2} \left\{ \left(\frac{v_{\text{ph}}(k)}{c} - \frac{1}{2} \right)^2 + \left(\frac{v_{\text{ph}}(k)}{c} + \frac{1}{2} \right)^2 \right\}. \quad (32)$$

The finite-size effect grows in importance as time progresses, on a time scale t_{fs} such that $\gamma_k t_{\text{fs}}^2 = \Gamma_k t_{\text{fs}}$ (where the decay in Eq. (III C) slows to zero). Its relevance therefore depends on how t_{fs} compares to the typical time-scale of decay Γ_k^{-1} , *i.e.*, the relevant quantity is

$$\begin{aligned} \Gamma_k t_{\text{fs}} &= \frac{\Gamma_k^2}{\gamma_k} = \frac{2L}{r_0} \frac{f^2(k\xi)}{g(k\xi)} \\ &\approx \frac{L}{r_0} \quad \text{when} \quad k\xi \gg 1. \end{aligned} \quad (33)$$

In the second line we have taken the large wavenumber limit, where our exponential prediction for the decay is most accurate (see Sec. IV below). r_0 is the (one-body) coherence length of the quasicondensate defined in Eq. (2). So, when $L/r_0 \ll 1$, the decay proceeds hardly at all before it is stopped by the finite-size effect. On the other hand, when $L/r_0 \gg 1$, the decay is significant before the slowing-down kicks in, and the finite-size effect becomes irrelevant. Dans l'Appendice B, nous dérivons le ralentissement du taux de décroissance induit par cette contribution. Sur des échelles de temps suffisamment courtes, il entre comme une correction quadratique à la décroissance exponentielle :

$$\delta n_k \approx \delta n_k(0) e^{-\Gamma_k t + \frac{1}{2} \gamma_k t^2},$$

où

$$\gamma_k t_{\xi}^2 = \frac{k_B T}{mc^2} \frac{1}{\rho_0 L} g(k\xi),$$

et où nous avons défini

$$g(k\xi) = \frac{(k\xi)^2}{v_{\text{ph}}^2(k)/c^2} \left\{ \left(\frac{v_{\text{ph}}(k)}{c} - \frac{1}{2} \right)^2 + \left(\frac{v_{\text{ph}}(k)}{c} + \frac{1}{2} \right)^2 \right\}.$$

L'effet de taille finie croît en importance au fur et à mesure que le temps progresse, sur une échelle de temps t_{fs} telle que $\gamma_k t_{\text{fs}}^2 = \Gamma_k t_{\text{fs}}$ (où la décroissance dans l'équation (III C) ralentit à zéro). Sa pertinence dépend donc de la comparaison entre t_{fs} et l'échelle de temps typique de décroissance Γ_k^{-1} , *c'est-à-dire*, la quantité pertinente est

$$\begin{aligned} \Gamma_k t_{\text{fs}} &= \frac{\Gamma_k^2}{\gamma_k} = \frac{2L}{r_0} \frac{f^2(k\xi)}{g(k\xi)} \\ &\approx \frac{L}{r_0} \quad \text{lorsque} \quad k\xi \gg 1. \end{aligned}$$

Dans la seconde ligne, nous avons pris la limite des grands nombres d'onde, où notre prédiction exponentielle pour la décroissance est la plus précise (voir Sec. IV ci-dessous). r_0 est la longueur de cohérence (à un corps) du quasi-condensat

définie dans l'équation (2). Donc, lorsque $L/r_0 \ll 1$, la décroissance progresse à peine avant d'être arrêtée par l'effet de taille finie. En revanche, lorsque $L/r_0 \gg 1$, la décroissance est significative avant que le ralentissement ne prenne effet, et l'effet de taille finie devient alors négligeable.

As mentioned in Sec. II A, when considering a quasicondensate over distances larger than r_0 the long-range order characterising condensation is lost due to large thermal fluctuations in the phase. Conversely, if we restrict attention to distances much shorter than r_0 , the one-body correlation is preserved and the gas looks like a “true” condensate. Equation (34) then shows that if we consider a small enough system that appears as a true condensate, the decay will be strongly suppressed by finite-size effects. In effect, the decay processes we have identified manifest in position space as an x -dependent drift in the phase of the excited phonon, whose variance becomes significant only over distances larger than r_0 . Therefore, the decay is only effective when we reach the quasicondensate regime $L \gg r_0$. Comme mentionné dans la section ??, lorsqu'on considère un quasi-condensat sur des distances supérieures à r_0 , l'ordre de longue portée caractéristique de la condensation est perdu en raison des grandes fluctuations thermiques de la phase. Inversement, si nous nous limitons à des distances beaucoup plus petites que r_0 , la corrélation unitaire est préservée et le gaz ressemble à un “vrai” condensat. L'équation (??) montre alors que si nous considérons un système suffisamment petit pour apparaître comme un vrai condensat, la décroissance sera fortement supprimée par les effets de taille finie. En effet, les processus de décroissance que nous avons identifiés se manifestent dans l'espace des positions comme un dérive dépendante de x dans la phase du phonon excité, dont la variance devient significative uniquement sur des distances supérieures à r_0 . Par conséquent, la décroissance n'est efficace que lorsque nous atteignons le régime du quasi-condensat $L \gg r_0$.

This correspondence is further explored in Appendix B, where it is also shown that binning modes in momentum space yields an evolution that is very similar to that on a shorter torus. We thus conclude that the system size L appearing in Eq. (34) can be generalized to the size of any subsection of a larger system. If this size is small compared to r_0 , local measurements will be insensitive to the decay. Cette correspondance est explorée plus en détail dans l'Appendice ??, où il est également montré que la binning des modes dans l'espace des moments donne une évolution très similaire à celle observée sur un torus plus court. Nous concluons donc que la taille du système L apparaissant dans l'équation (??) peut être généralisée à la taille de toute sous-section d'un système plus grand. Si cette taille est petite par rapport à r_0 , les mesures locales seront insensibles à la décroissance.

La taille finie du système influence significativement le taux de décroissance des phonons. Le terme $q = 0$ dans l'équation (28) ralentit la décroissance, mais il est proportionnel à $1/L$ et disparaît lorsque la taille du système tend vers l'infini, où l'espace des moments devient un continuum. Ce terme ajuste la résolution finie en k et ralentit la décroissance en maintenant les phonons proches du mode principal de

décroissance dans la même "bin". Ainsi, la contribution des phonons voisins est modifiée, ce qui ralentit la décroissance nette. L'Appendice B analyse comment ce ralentissement de la décroissance se manifeste sur des échelles de temps courtes. La correction quadratique à la décroissance exponentielle est donnée par :

$$\delta n_k \approx \delta n_k(0) e^{-\Gamma_k t + \frac{1}{2} \gamma_k t^2},$$

où

$$\gamma_k t_\xi^2 = \frac{k_B T}{m c^2} \frac{1}{\rho_0 L} g(k\xi).$$

et où nous avons défini

$$g(k\xi) = \frac{(k\xi)^2}{v_{\text{ph}}^2(k)/c^2} \left\{ \left(\frac{v_{\text{ph}}(k)}{c} - \frac{1}{2} \right)^2 + \left(\frac{v_{\text{ph}}(k)}{c} + \frac{1}{2} \right)^2 \right\}.$$

L'effet de taille finie devient significatif à une échelle de temps t_{fs} telle que $\gamma_k t_{\text{fs}}^2 = \Gamma_k t_{\text{fs}}$ (où la décroissance dans l'équation (III C) ralentit à zéro). Sa pertinence dépend donc de la comparaison entre t_{fs} et l'échelle de temps typique de décroissance Γ_k^{-1} , c'est-à-dire, la quantité pertinente est

$$\begin{aligned} \Gamma_k t_{\text{fs}} &= \frac{\Gamma_k^2}{\gamma_k} = \frac{2L}{r_0} \frac{f^2(k\xi)}{g(k\xi)} \\ &\approx \frac{L}{r_0} \quad \text{lorsque } k\xi \gg 1. \end{aligned}$$

Dans la seconde ligne, nous avons pris la limite des grands nombres d'onde, où notre prédiction exponentielle pour la décroissance est la plus précise (voir Sec. IV ci-dessous). r_0 est la longueur de cohérence (à un corps) du quasi-condensat définie dans l'équation (2). Ainsi, lorsque $L/r_0 \ll 1$, la décroissance est pratiquement inexistante avant d'être stoppée par l'effet de taille finie. En revanche, lorsque $L/r_0 \gg 1$, la décroissance est significative avant que le ralentissement dû à la taille finie ne devienne pertinent. En résumé, lorsqu'un système est de taille inférieure à r_0 , il peut apparaître comme un condensat "vrai", et l'effet de taille finie peut fortement supprimer la décroissance. Pour des systèmes plus grands, avec $L \gg r_0$, les effets de taille finie deviennent négligeables, et le système se comporte plus comme un condensat quasi-condensé. L'Appendice B montre également que la binning des modes dans l'espace des moments produit une évolution similaire à celle observée sur un torus plus court, confirmant que la taille du système L peut être généralisée à celle de toute sous-section d'un système plus grand. Si cette taille est petite par rapport à r_0 , les mesures locales seront insensibles à la décroissance.

IV. NUMERICAL CONFIRMATION

We have run several numerical simulations in order to test our prediction for the phonon damping rate. The essentials of

the numerical method have already been described in Sec. II E, and more details can be found in Appendix C.

The simulations fall into two types. First, we perform a series of simulations like that of Fig. 1, where a number of phonons is injected into a single mode k , and the evolution of the phonon number spectrum is followed in time. This setup allows us to study the response of a thermal system to a single perturbation, the addition of phonons in the mode k . We use it to demonstrate that the thermal processes discussed in the previous section are indeed the main culprit involved in the scattering of injected phonons, and to check that the scaling properties of the decay rate match those predicted by Eq. (III A). Moreover, this controlled scenario enables us to track precisely how phonons scatter to other modes (as shown in Fig. 1), informing our study of the decay process and its generalization to a peak of finite width (see Sec. IV B below, and Appendix B).

The second series of simulations is inspired by the parametric resonance experiments of [4] and theoretically studied in [8, 33]: starting from a thermal state, a sinusoidal modulation of the 1D atomic interaction strength is applied, inducing exponential growth of the phonon occupation number within a resonance window. (This can be achieved experimentally by modulating the transverse stiffness of the trap [4, 7, 8].) The phonon damping is then observed as a reduction in the rate of exponential growth. As suggested above, the finite width of the resonant peak induces a deviation from the decay rate (III A), which strictly speaking is only applicable in the limit of a singularly occupied mode. However, this deviation occurs in a controlled fashion, and is found to be consistent with the back-reaction term appearing in Eq. (28).

A. Initial injection of phonons

In the first run of simulations, the initial state is taken to be a thermal state for the quadratic Hamiltonian \hat{H}_2 , with the possible addition of a probe in a single phonon mode of wave vector k . When the probe is present, it is given an initial occupation number δn on top of the thermal distribution by simply multiplying the amplitude of the relevant mode by a constant factor. This ensures that the phonon modes are initially independent of each other, and separately exhibit Gaussian statistics. Each realisation is then evolved twice, with and without the probe, to take account of the slight degree of time-dependence that occurs even when the probe is absent. The spectra are calculated independently at different times in the interval $t/t_\xi \in [0, 10]$, and the probe spectrum $\delta n_k(t)$ is defined as the difference between the two, in accordance with Eq. (22). $\delta n_k(t)$ is then fitted to the template $A \exp(-\Gamma t + \gamma t^2/2)$ of Eq. (III C), the fitted value of Γ being the extracted decay rate. As mentioned in Sec. III C above, the t^2 correction is related to the limited resolution in k -space induced by the finite length of the condensate, coming from the $q = 0$ term in the second line of Eq. (28). It is included here as it makes a small but noticeable difference to the fit. More details on the extraction of the γ term are given

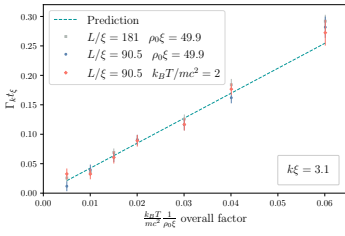


Figure 2. Best fit values for $\hbar\Gamma_k$ extracted from the TWA simulations, as a function of $k_B T / \rho_0 \xi$. All points are extracted from simulations performed using an addition of $\delta n = 3$ phonons on average. We used 400 realisations, a grid spacing $\Delta x / \xi = 0.35$, and a time window $t/t_\xi \in [0, 10]$ with 140 time steps. For the grey and blue points the temperature varies $k_B T / mc^2 \in [0.25, 3]$, while for the red points the density varies in $\rho_0 \xi \in [33, 399]$. The parameters that are fixed in each run are listed in the legend.

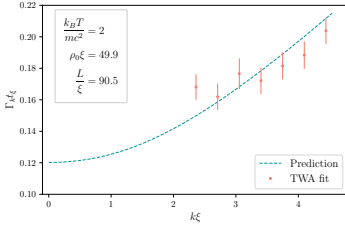


Figure 3. Best fit values for $\hbar\Gamma_k$ extracted from the TWA simulations, as a function of $k\xi$. All other physical parameters are fixed at the values shown, while numerical parameters are the same as in Fig. 2. Numerically extracted values are shown only for $k\xi \gtrsim 2.5$. For lower $k\xi$, significant deviations appear due to the longer response time. (A fuller treatment of this regime is given in Appendix B.)

in Appendix B.

In Fig. 2 we demonstrate the linearity of the best-fit values of Γ in the overall prefactor in Eq. (III A), at fixed $k\xi$. The numerical results are in good agreement with the prediction. We illustrate that varying T at fixed $\rho\xi$ (and *vice versa*) yields the expected behaviour, and moreover that the fitted decay rate is unaffected by a change in L , as expected from prediction (III A). Any L -dependence of the observed behavior would then appear in the other fitting parameters like γ .

In Fig. 3 we show instead the dependence of these best-fit values for Γ on the probe mode $k\xi$, with the prefactor in Eq. (III A) fixed. The numerical observations agree well with the predicted behaviour at modestly high $k\xi \gtrsim 2.5$. At lower $k\xi$, we expect significant deviations, for as noted in Sec. III B, the FGR becomes valid only after a critical time t_{crit} (see also [30, 34]). We show in Appendix B that t_{crit} diverges like $1/(k\xi)^3$ as $k\xi \rightarrow 0$, and for $k\xi \lesssim 2.5$ it is not reached within the sampled time frame. In this regime, the decay proceeds quadratically in time rather than exponentially, with a lifetime that is significantly longer than predicted by (III A). This early-time behavior at small $k\xi$ is examined in more detail in Appendix B.

B. Slowing of exponential growth

In our second set of simulations, the initial state is simply a thermal state for the quadratic Hamiltonian \hat{H}_2 , with no additional component added by hand. Instead, during the evolution, parametric resonance is induced by a sinusoidal modulation of the 1D atomic interaction parameter:

$$g(t) = g(1 + a \sin(\omega_p t)) . \quad (34)$$

As $c^2(t) \propto g(t)$, this translates into a sinusoidal modulation of the squared phonon frequencies:

$$\begin{aligned} \omega_k^2(t) &= c^2 k^2 \left[1 + a \sin(\omega_p t) + \frac{1}{4} k^2 \xi^2 \right] \\ &= \omega_k^2 (1 + A_k \sin(\omega_p t)) , \end{aligned} \quad (35)$$

where $A_k = a / (1 + k^2 \xi^2 / 4)$. This is exactly the situation modeled in [33]. The result is an exponential growth of the number of phonons within a resonant frequency window centered at $\omega_p/2$. In the absence of any damping mechanism, the analysis in Appendix A of [33] shows that, for the exactly resonant mode at $\omega_k = \omega_p/2$, the occupation number is parametrically amplified according to

$$n_k(t) \approx n_k^{\text{in}} + (2n_k^{\text{in}} + 1) \sinh^2 \left(\frac{1}{2} G_k t \right) , \quad (36)$$

where the growth rate $G_k = A_k \omega_k / 2$. Since the initial state is thermal, the initial occupation number n_k^{in} is simply the thermal population of the mode.

When including the effects of phonon interactions, the Beliaev-Landau scattering with the thermal population acts simultaneously with the parametric amplification, kicking phonons out of the resonant mode as they are being produced. The damping mechanism thus acts much like the phenomenological damping introduced in [33], and is therefore expected to reduce the growth rate. This is illustrated in Fig. 4, where $n_k(t)$ of the exactly resonant mode is extracted from the numerical simulations and shown for two different values of $\rho_0 \xi$. We clearly see that the exponential growth rate is lower than predicted by Eq. (36), and that it is further reduced as $\rho_0 \xi$ is reduced, so that Γ_k of Eq. (III A) is accordingly increased.

We may use the difference between the observed growth rate and the “pure” growth rate of Eq. (36) as a measure of the damping.⁹ Assuming that the damping acts straightforwardly as a reduction of the growth rate, we make a slight generalization of Eq. (36) and propose the following ansatz for the occupation number of the resonant mode:

$$n_k(t) = n_k^{\text{in}} + (2n_k^{\text{in}} + 1) \sinh^2 \left[\frac{1}{2} (G_k - \Gamma_k) t \right] . \quad (37)$$

⁹ Similar simulations were performed in [8] where, although the focus was on nonlinear effects at large n_k , hints of an n_k -independent damping rate were seen at sufficiently early times (see in particular Fig. 6 and footnote 10). Given the parameters used ($k\xi \sim 1$, $k_B T / mc^2 \sim 1/2$, $\rho_0 \xi \sim 400$), this intrinsic decay rate has $\hbar \Gamma_k / mc^2$ equal to a few times 10^{-3} , in agreement with prediction (III A).

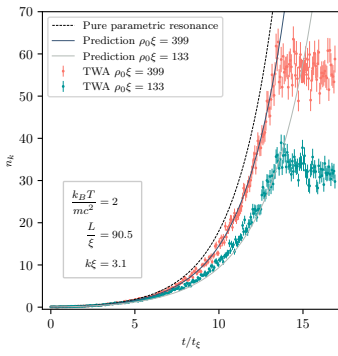


Figure 4. Mean occupation of the resonant mode as a function of time, as extracted from TWA simulations. We take an initial thermal state at temperature $k_B T / m c^2 = 2$, and a modulation of amplitude $a = 0.5$ at frequency $\omega_p = 2\omega_k$ where $k\xi = 3.1$ (so that the modulation of ω_k^2 has amplitude $A_k = 0.15$) lasting 24 periods, *i.e.*, $13.5 t_\xi$. The size of the system is $L/\xi = 90.5$ and its atomic density is $\rho_0\xi = 399$ for the red dots and $\rho_0\xi = 133$ for the green ones. The dashed black line is the estimate Eq. (36), where n_{in} has been set to the thermal value. The solid lines correspond to the corrected estimate Eq. (37), with Γ_k as predicted by Eq. (III A). Each data point is calculated independently, averaged over 400 realizations.

Figure 4 shows that this ansatz, taken together with the assumption that Γ_k is as predicted by Eq. (III A), accounts very well for the reduced growth observed in TWA simulations with respect to that of Eq. (36). We stress that this reduction rapidly leads to a sizeable change in the number of produced phonons. For example, considering the red data points in Fig. 4 (corresponding to $k\xi = 3.1$, $k_B T / m c^2 = 2$, and $\rho_0\xi = 399$), the relative damping is $\Gamma_k / G_k = 5\%$, yet the reduction with respect to the non-damped case is very clear, and we are thus able to extract quite precise values for the damping rate.

We now perform a more systematic examination of the numerically observed reduction of the growth rate. Proceeding as for the first set of simulations, we fit $n_k(t)$ for the exactly resonant mode to Eq. (37), where Γ_k is now treated as a fitting parameter. This is done for two different wave vectors, one in the high- k regime ($k\xi = 3.1$) where the first set of simulations worked reasonable well, and one in the low- k regime ($k\xi = 1.0$) where they did not. The temperature is fixed at $k_B T / m c^2 = 2$, and the 1D density $\rho_0\xi$ is varied. The extracted values of Γ_k are shown in Fig. 5 alongside the prediction of Eq. (III A).

A few remarks are in order concerning these results. We begin by focusing on the high- k mode ($k\xi = 3.1$), the first and third data points of which correspond to the simulations shown in Fig. 4. While the extracted decay rate tends towards prediction (III A) at large $\rho_0\xi$, there is a clear trend for it to fall further away from this prediction as $\rho_0\xi$ is decreased. We attempt to explain this behavior by appealing to the finite width of the resonant peak. Examining the evolution of the number spectrum shows that, to a good approximation, the shape of the resonant peak saturates such that $n_k \sim R_k e^{Gt}$ at sufficiently late time, for some profile R_k and some growth rate

G . This is in contrast to the results of the “pure” parametric resonance with no phonon interactions, where the growth rate is k -dependent and largest at exact resonance [33], so that the width of the peak approaches zero as $t \rightarrow \infty$. In the present case, in addition to the saturation of the profile, we observe that larger interaction strengths are associated with wider peaks, see Fig. 6. It thus seems likely that n_k is also being fed by phonons in neighboring modes through the last term on the right-hand side of Eq. (28). We investigate this effect by adopting a Lorentzian ansatz for the profile of δn_k . Figure 6 shows how our ansatz compares with the numerically observed number spectrum. Under this assumption, we solve Eq. (28) self-consistently to extract the net growth rate. The details of this calculation are given in Appendix B, but the corrected predictions for the decay rate are shown by the blue dots in Fig. 5. For $k\xi = 3.1$ this corrected prediction is found to be in very good agreement with the extracted rate.

For the low- k mode ($k\xi = 1.0$) there remains a clear discrepancy even when accounting for the finite width of the peak. Part of the explanation lies in the non-Lorentzianity of the profile visible in Fig. 6, making the correction less valid. However, note that the extracted decay rate does not tend well to prediction (III A) at large $\rho_0\xi$, as it seems to approach a line with a different slope. We believe this to be a consequence of the large critical time in the low- k regime: in effect, the occupation number is growing a little too fast for the system to have time to react, and the amount of damping is thus lower than expected.

The parametric resonance simulations described here yield results that corroborate and complement those found using straightforward phonon injection. In the high- k regime where the critical response time is sufficiently short, the observed deviations from prediction (III A) are well described via corrections due to the finite width of the peak, which only appear in the case of parametric resonance. On the other hand, at low k where the critical time is long, the phonon injection method yields a non-exponential behavior that has not been shown here (see instead Appendix B), whereas the parametric resonance approach still gives a damping rate that has the same qualitative behavior as in the high- k regime (see the near-linear behavior of the red dots in the lower panel of Fig. 5). Moreover, the method of parametric resonance for exciting phonons has some interesting advantages over that of phonon injection. Most notably, the damping mechanism manifests in a rather more dramatic way, as is evident from the difference in the final phonon numbers shown in Fig. 4. It is also of considerable experimental relevance: while the injection of phonons is conceptually simple, it is not very practical, whereas parametric resonance is a way of exciting phonons that has already been implemented in experiments [4]. On the conceptual side, the method of parametric resonance bypasses the finite-size effect. In Sec. IV A, it arises due to the discretization of a time-dependent continuous profile, the “binned” mode of interest containing both the exponentially decaying mode and the nearby modes within π/L that are growing in time. Here, the profile is fixed (the only time-dependence being an overall exponential factor e^{Gt}), and

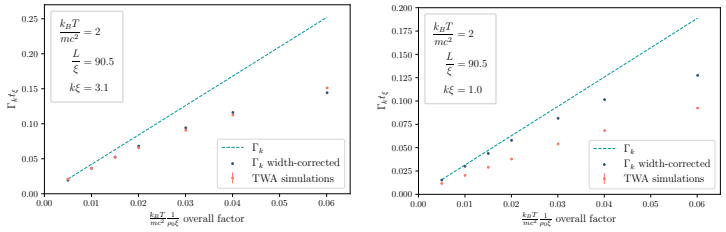


Figure 5. Best fit values for $\hbar\Gamma_k$ extracted from the TWA simulations as a function of $k_B T / \rho_0 \xi$. (We divide by mc^2 to adimensionalize.) The temperature is kept fixed to $k_B T / mc^2 = 2$ and only the density is varied $\rho_0 \xi \in [33, 399]$. The parameters that are fixed in each run are listed in the legend. All points are extracted from simulations performed using a continuous modulation of amplitude $a = 0.5$ at frequency $\omega_p = 2\omega_k$ where $k\xi = 3.1$ (top) and $k\xi = 1.0$ (bottom); the corresponding amplitudes for the modulation of ω_k^2 are $A_k = 0.15$ (top) and $A_k = 0.39$ (bottom). We use $n_r = 400$ realisations, a spatial grid with spacing $\Delta x / \xi = 0.35$. The fit is performed using the template Eq. (37) over a time window $t/t_\xi \in [0, 13.5]$ (top) and $t/t_\xi \in [0, 22.5]$ (bottom) with 224 (top) and 280 (bottom) time steps. The green dashed line represents the FGR prediction Eq. (III A) while the blue dots are the corrected predictions taking into account the effect of a finite width of the resonant peak, see Appendix B. The peak is assumed to be of Lorentzian shape and its parameters are extracted by a procedure described in Fig. 6.

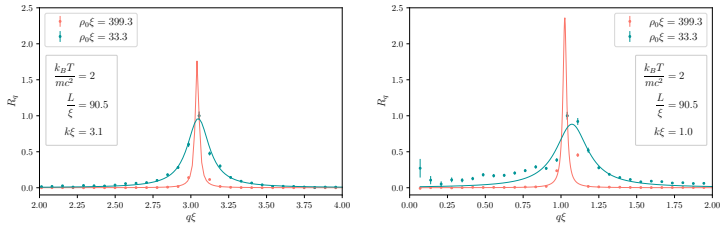


Figure 6. Snapshots of the ratio $R_q = \frac{\delta n_q}{\delta n_k}$ as a function of $q\xi$ at time $t/t_\xi = 13.5$ (top) and $t/t_\xi = 22.5$ (bottom). The spectrum is plotted for $\rho_0 \xi = 399$ (red) and $\rho_0 \xi = 33$ (green). They are obtained by a continuous modulation of the type Eq. (34) with $a = 0.5$ and at the appropriate frequency so that $k\xi = 3.1$ (top) or $k\xi = 1.0$ (bottom) are exactly at the resonance. A Lorentzian is fitted to the distribution taking into account the first five neighbours on each side of the central resonant mode letting the overall amplitude, width and center be free fitting parameters. The best-fit value of the width is then used to correct the prediction of the decay rate, blue dots in Fig. 5. The averages are calculated from an ensemble of 400 independent realisations, while the error bars represent the standard deviation. There are 256 points on the grid.

the discretization is therefore irrelevant.

V. CONCLUDING REMARKS

We have identified a mechanism whereby interactions with a thermal bath of phonons, through Landau and Beliaev damping mechanisms, yields exponential decay of a singularly occupied

phonon mode. It is effectively described by an application of the FGR with the resonant channel being the trivial one, in the limit $q \rightarrow 0$. The result was numerically verified using TWA simulations, following both the decay of a mode excited “by hand” and the reduced growth of a parametrically resonant mode. The prediction works particularly well in the higher- k regime, where the critical response time is relatively short and the discreteness of the thermal population is invisible to the mode in question. We also observe a reduction of the damping rate due to the finite width of the peak.

While this limit yields a non-vanishing rate only in 1D (since in higher dimensions the IR divergence of the thermal population is tamed by the volume element in k -space), it seems generally applicable to 1D systems with an approximately linear excitation spectrum in the limit of small k . We thus believe that it could appear as an additional contribution to the decay of excitations in such systems, such as the 1D dipolar gas considered in [19].

Comparing with higher dimensionality D [35], there are some clear similarities: prediction (III A) exhibits an increasing dependence on $k\xi$, is proportional to $\gamma_{\text{LL}}^{D/2}$, and is linear in T (as found in higher dimensions when $k_B T/mc^2 \gtrsim 1$).

The experimental relevance of this decay mechanism was exemplified by numerically studying the parametric growth of phonons in the gas, as experimentally tested in [4]. We have demonstrated that the numerically observed reduction of the growth rate can be quantitatively explained assuming that the phonons generated by the parametric process simultaneously decay by thermal Landau-Beliaev processes.

These thermal scatterings of phonons are expected to occur generally in 1D quasicondensate experiments. They can produce important deviations to the dynamics of phonons and should therefore be taken into account in designing and analysing analogue gravity experiments on this platform.

Finally, let us comment on the absence of entanglement between the induced peaks in the experimental observations of [4]. In their phenomenological treatment of dissipation, the authors of [33] showed that a dissipative rate $\Gamma/\omega_k \sim 4.2\%$ ¹⁰ would be sufficient to explain this negative result. Using the relevant parameters ($k\xi \sim 1$, $k_B T/\hbar\omega_k \sim 1$, $\rho_0\xi \sim 60$), Eq. (III A) gives $\Gamma/\omega_k \sim 5\%$. The decay mechanism identified here thus provides a possible microphysical basis to this scenario. The precise dynamics of the entanglement will be the subject of a future work.

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¹⁰ Since we focus directly on the phonon number n_k rather than the amplitudes $\hat{\varphi}_k$, there is a difference of a factor 2 between the dissipative rates defined in [33] and the one defined here.

was responsible for the genesis of this work and whose guidance and input in its initial stages proved invaluable. We would like to dedicate the paper to his memory. This work was supported by the French National Research Agency via Grant No. ANR-20-CE47-0001 associated with the project COSQUA (Cosmology and Quantum Simulation).

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Appendix A FULLER DERIVATION OF EVOLUTION OF δn_k

In this appendix we derive Eq. (28), a more precise equation of motion for the phonon spectrum, and we show how the FGR emerges despite there being no exactly elastic scattering channel. We then analyze more fully the deviations with respect to the FGR result.

A Equations of motion for phonon operators

We start by computing the Heisenberg equation of motion for $\hat{\varphi}_k$. The one for $\hat{\varphi}_k^\dagger$ is easily deduced by taking the adjoint. We consider only the dynamics under the quadratic (13) and cubic (18) Hamiltonians, neglecting higher orders. We have :

$$\partial_t \hat{\varphi}_k = -i \left\{ \omega_k \hat{\varphi}_k + \frac{1}{\sqrt{N_{\text{at}}}} \sum_{q \neq 0, -k} 2V_3(k, q) \hat{\varphi}_q^\dagger \hat{\varphi}_{k+q} + \frac{1}{\sqrt{N_{\text{at}}}} \sum_{q \neq 0, k} \right. \quad (38)$$

Considering then the full number operator $\hat{n}_k = \hat{\varphi}_k^\dagger \hat{\varphi}_k$, we have

$$\begin{aligned} \partial_t \hat{n}_k &= \hat{\varphi}_k^\dagger \cdot \partial_t \hat{\varphi}_k + \partial_t \hat{\varphi}_k^\dagger \cdot \hat{\varphi}_k, \\ &= -\frac{i}{\sqrt{N_{\text{at}}}} \sum_{q \neq 0, -k} 2V_3(k, q) \hat{\varphi}_k^\dagger \hat{\varphi}_q^\dagger \hat{\varphi}_{k+q} - \frac{i}{\sqrt{N_{\text{at}}}} \sum_{q \neq 0, k} V_3(k - q, q) \hat{\varphi}_k^\dagger \hat{\varphi}_q^\dagger \hat{\varphi}_{k+q} \end{aligned} \quad (39)$$

On the right-hand side of the equation of motion for \hat{n}_k appears the momentum preserving 3-phonon operator $\hat{\varphi}_k^\dagger \hat{\varphi}_q^\dagger \hat{\varphi}_{k+q}$, the same that appears in \hat{V}_3 . We are thus compelled to consider the dynamics of this operator as well. It is useful to define $\hat{\varphi}_p^\dagger \hat{\varphi}_q^\dagger \hat{\varphi}_{p+q} = \hat{c}_3(p, q) \exp[-i(\omega_{p+q} - \omega_p - \omega_q)t]$, where the oscillations are made explicit. These operators obey

$$\begin{aligned} e^{-i(\omega_{p+q} - \omega_p - \omega_q)t} \partial_t \hat{c}_3(p, q) &= \partial_t \hat{\varphi}_p^\dagger \cdot \hat{\varphi}_q^\dagger \hat{\varphi}_{p+q} + \hat{\varphi}_p^\dagger \cdot \partial_t \hat{\varphi}_q^\dagger \cdot \hat{\varphi}_{p+q} + \hat{\varphi}_p^\dagger \cdot \hat{\varphi}_q^\dagger \cdot \partial_t \hat{\varphi}_{p+q} \\ &= \frac{2i}{\sqrt{N_{\text{at}}}} V_3(p, q) \hat{\varphi}_{p+q}^\dagger \hat{\varphi}_{p+q} \\ &+ \frac{2i}{\sqrt{N_{\text{at}}}} \sum_{\lambda \neq 0, -p} V_3(p, \lambda) \hat{\varphi}_{p+\lambda}^\dagger \hat{\varphi}_q^\dagger \hat{\varphi}_\lambda \hat{\varphi}_{p+q} + \sum_{\lambda \neq 0, -q} V_3(q, \lambda) \hat{\varphi}_p^\dagger \hat{\varphi}_{q+\lambda}^\dagger \hat{\varphi}_{p+q} \\ &+ \frac{i}{\sqrt{N_{\text{at}}}} \sum_{\lambda \neq 0, p} V_3(p - \lambda, \lambda) \hat{\varphi}_\lambda^\dagger \hat{\varphi}_{p-\lambda}^\dagger \hat{\varphi}_q^\dagger \hat{\varphi}_{p+q} + \sum_{\lambda \neq 0, q} V_3(q - \lambda, \lambda) \hat{\varphi}_p^\dagger \hat{\varphi}_{q-\lambda}^\dagger \hat{\varphi}_{p+q} \end{aligned} \quad (40)$$

The extra term on the first line involving the amplitudes for the $p + q$ mode comes from rearranging the one term which is not initially in normal order.

B Equations of motion for average values

We want to take average values in both equations of motion above and derive the evolution of $n_k = \langle \hat{n}_k \rangle$. However, in order

to get a closed system we have to make some approximations. We assume that the state is initially homogeneous; it will remain so as our Hamiltonian is momentum-conserving. We also assume that the initial state is Gaussian, and that the only deviation from Gaussianity to evolve is a non-vanishing value of $\hat{c}_3(p, q)$, *i.e.*, every connected correlation function of order higher than three is negligible. Therefore, when taking average values on the right hand-side of Eq. (40), the 4-point functions reduce to products of 2-point functions which have to respect the homogeneity of the state:

$$\langle \hat{\varphi}_p^\dagger \hat{\varphi}_q^\dagger \hat{\varphi}_{p+q-\lambda} \hat{\varphi}_\lambda \rangle = n_p n_q \delta_{q,\lambda} + n_p n_q \delta_{p,\lambda} + c_q^* c_\lambda \delta_{p,-q}, \quad (41)$$

where $c_p = \langle \hat{\varphi}_p \hat{\varphi}_{-p} \rangle$. The quantities $n_{\pm p}$ and c_p – respectively, the population and 2-mode correlation of the modes $\pm p$ – are the only non-vanishing 2-point functions in a homogeneous state. We shall further assume, for simplicity, that the c_p are all vanishing.¹¹ The equation of motion for n_p and $c_3(p, q) = \langle \hat{c}_3(p, q) \rangle$ then reads :

$$\begin{aligned} \partial_t n_k &= \frac{1}{\sqrt{N_{\text{at}}}} \sum_{q \neq 0, -k} 4V_3(k, q) \Im \left[c_3(k, q) e^{-i\delta\omega_L(q; k)t} \right] - \frac{1}{\sqrt{N_{\text{at}}}} \\ \partial_t c_3(p, q) &= 2i \frac{V_3(p, q)}{\sqrt{N_{\text{at}}}} [n_{p+q} (n_p + n_q + 1) - n_p n_q] e^{i(\omega_{p+q} - \omega_p - \omega_q)t} \end{aligned} \quad (42)$$

We can reduce this to a single equation of motion for n_k by writing $c_3(p, q)$ explicitly in terms of n_k :

$$c_3(p, q)(t) = 2i \frac{V_3(p, q)}{\sqrt{N_{\text{at}}}} \int_0^t dt' N_{p, q}(t') e^{i(\omega_{p+q} - \omega_p - \omega_q)t'}, \quad (43)$$

where $N_{p, q} = n_{p+q} (n_p + n_q + 1) - n_p n_q$. Equation (43) can now be substituted directly into the equation of motion for n_k :

$$\begin{aligned} \partial_t n_k &= 8 \sum_{q \neq 0, -k} \frac{|V_3(k, q)|^2}{N_{\text{at}}} \int_0^t dt' N_{k, q}(t') \cos[(\omega_{k+q} - \omega_k - \omega_q)t] \\ &- 4 \sum_{q \neq 0, k} \frac{|V_3(k - q, q)|^2}{N_{\text{at}}} \int_0^t dt' N_{k-q, q}(t') \cos[(\omega_k - \omega_{k-q} - \omega_q)t] \end{aligned} \quad (44)$$

Note that the terms depending on the various populations can be rewritten as the difference between the direct process mentioned in the text below Eq. (18) and the corresponding reverse process. For instance in the first sum, corresponding to the Landau scattering, we have

$$n_{p+q} (n_p + n_q + 1) - n_p n_q = (n_p + 1) (n_q + 1) n_{p+q} - (n_{p+q} + 1) n_p \quad (45)$$

¹¹ This is not actually the case for the 2-mode squeezed state generated by the parametric resonance considered in the second set of simulations. Yet, we achieve a quantitative agreement with the TWA simulations for large enough value of $k\xi$. We relegate the analysis of the influence of the correlation on the decay to a future work dedicated to the dynamics of entanglement in this context.

On the right-hand side the $+1$ terms are associated only with the decay products and allow for spontaneous processes, while the n_s encode the stimulated part. This generalizes the matrix elements given in the main text, which only include one direction where an excitation at p is removed by the interaction.

Equations (42) and (44) are the key equations governing the system, given our simplifying approximations. They are entirely equivalent if $c_3(p, q)$ is set to zero at $t = 0$, though Eqs. (44) can be straightforwardly modified in a more general case. Since Eqs. (42) are Markovian, they are much more suitable for numerical integration. On the other hand, Eq. (44) is a nonlinear, non-Markovian equation for the full phonon spectrum n_k . However, if analyzed using appropriate approximations, we will show that it encodes both the exponential decay of the phononic population, its first deviations and the corrections to the growth of population in a parametric resonance process.

C Dynamics of a probe on top of a quasicondensate

Up to this point, we have assumed homogeneity and quasi-Gaussianity of the state but worked with the full phonon spectrum n_q . As explained in the main text, the presence of a near-thermal population in the IR modes is instrumental in the decay process that we study. Therefore, as done in Eq. (22), we split n_q into a thermal background $n_q^{\text{th}} = 1/[\exp(\hbar\omega_q/k_B T) - 1]$ plus a perturbation δn_q . Physically this setup allows us to analyze the response of a quasicondensate at temperature T to the addition of phonons around a certain mode k . This can either be done “by hand”, as we have done in the first set of simulations presented in the text, or by a parametric amplification as in the second set, see Sec. IV. As illustrated by Fig. 1 of the main text, we expect these δn probe phonons to redistribute across the system and want to compute the ensuing decay rate.

Notice that the background thermal population n_q^{th} is not strictly stationary due to the addition of the interaction term \hat{V}_3 , but as a first approximation we will assume it to be a solution of the equation of motion. Since we expect the relevant interactions to be between the peak and the thermal population, and not of the peak on itself, we insert $n_k = n_k^{\text{th}} + \delta n_k$ in Eq. (44) and linearise in δn . We have

$$\begin{aligned} N_{p,q} &= N_{p,q}^{\text{th}} + \delta n_{p+q} (n_p^{\text{th}} + n_q^{\text{th}} + 1) + \delta n_p (n_{p+q}^{\text{th}} - n_q^{\text{th}}) + \delta n_q (n_{p+q}^{\text{th}} - n_p^{\text{th}}) \\ &= N_{p,q}^{\text{th}} + \delta N_{p,q} + O(\delta n^2), \end{aligned} \quad (46)$$

with $N_{p,q}^{\text{th}}$ the specific combination of populations evaluated with thermal populations and

$$\delta N_{k,q} = \delta n_k (n_{k+q}^{\text{th}} - n_q^{\text{th}}) + \delta n_{k+q} (n_k^{\text{th}} + n_q^{\text{th}} + 1) + \delta n_q (n_{k+q}^{\text{th}} - n_k^{\text{th}}) \quad (47)$$

$$\delta N_{k-q,q} = \delta n_k (n_{k-q}^{\text{th}} + n_q^{\text{th}} + 1) + \delta n_{k-q} (n_k^{\text{th}} - n_q^{\text{th}}) + \delta n_q (n_{k-q}^{\text{th}} - n_k^{\text{th}}) \quad (48)$$

Inserting back in the equations of motion we get

$$\begin{aligned}
\partial_t \delta n_k &= 8 \sum_{q \neq 0, -k} \frac{|V_3(k, q)|^2}{N_{\text{at}}} \int_0^t dt' \delta N_{k, q} \cos[(\omega_{k+q} - \omega_k - \omega_q)(t - t')] \\
&\quad - 4 \sum_{q \neq 0, k} \frac{|V_3(k - q, q)|^2}{N_{\text{at}}} \int_0^t dt' \delta N_{k-q, q} \cos[(\omega_k - \omega_{k-q} - \omega_q)(t - t')] \\
&= - \int_0^t dt' D_k(t - t') \delta n_k(t') + \int_0^t dt' \sum_{q \neq -k} M_{k, k+q}(t - t') \delta n_{k+q}(t')
\end{aligned} \tag{49}$$

In the last line we have split the RHS into two terms, defining a “diagonal” response function $D_k(t - t')$ that acts only on $\delta n_k(t')$, and a “matrix” response function that includes contributions from other modes $\delta n_{k+q}(t')$. Notice that when defining the response functions we have included the term $q = 0$, by extending the summand to its finite limit as q goes to 0, in both the diagonal and the matrix function; one can check that they exactly cancel out. This inclusion allows to separate clearly the exponential decay from its deviations, as laid out below. The last equality in Eq. (49) gives Eq. (28) of the main text.

D Diagonal response function and exponential decay

Considering first the diagonal response function, we have explicitly:

$$\begin{aligned}
D_k(\tau) &= 8 \sum_{q \neq -k} |V_3(k, q)|^2 (n_q^{\text{th}} - n_{k+q}^{\text{th}}) \cos[(\omega_{k+q} - \omega_k - \omega_q)\tau] \\
&\quad + 8 \sum_{q \leq k/2} |V_3(k - q, q)|^2 (n_q^{\text{th}} + n_{k-q}^{\text{th}} + 1) \cos[(\omega_k - \omega_{k-q} - \omega_q)\tau]
\end{aligned}$$

where we have used the $q \rightarrow k - q$ symmetry in the second sum to fold it on $q \leq k/2$ adding a factor 2. We want to show that $D_k(\tau)$ reduces to a Dirac delta, and for this purpose it is convenient to first consider its integral. We define

$$I_k(\tau) = \int_0^\tau d\tau' D_k(\tau') \quad \Longleftrightarrow \quad D_k(\tau) = I'_k(\tau), \quad I_k(0) = 0 \tag{51}$$

Since $D_k(\tau)$ is an even function ($D_k(-\tau) = D_k(\tau)$), we find that $I_k(\tau)$ is odd:

$$I_k(-\tau) = \int_0^{-\tau} d\tau' D_k(\tau') = \int_0^\tau d(-\tau') D_k(-\tau') = - \int_0^\tau d\tau' D_k(\tau') \tag{52}$$

Explicitly, we note that $\int_0^\tau \cos(\Omega\tau') d\tau' = \tau \text{sinc}(\Omega\tau)$. Therefore,

$$I_k(\tau) = 8 \sum_{q \neq -k} |V_3(k, q)|^2 (n_q^{\text{th}} - n_{k+q}^{\text{th}}) \tau \text{sinc}[(\omega_{k+q} - \omega_k - \omega_q)\tau] \\ + 8 \sum_{q \leq k/2} |V_3(k - q, q)|^2 (n_q^{\text{th}} + n_{k-q}^{\text{th}} + 1) \tau \text{sinc}[(\omega_k - \omega_{k-q} - \omega_q)\tau]$$

It is also useful to take the continuous (large L) limit, replacing $\frac{1}{\Delta q} \sum_q \Delta q \rightarrow \frac{L}{2\pi} \int dq$

$$I_k(\tau) = \frac{4L}{\pi} \int_{-\infty}^{+\infty} dq |V_3(k, q)|^2 (n_q^{\text{th}} - n_{k+q}^{\text{th}}) \tau \text{sinc}[(\omega_{k+q} - \omega_k - \omega_q)\tau] \\ + \frac{4L}{\pi} \int_{-\infty}^{k/2} dq |V_3(k - q, q)|^2 (n_q^{\text{th}} + n_{k-q}^{\text{th}} + 1) \tau \text{sinc}[(\omega_k - \omega_{k-q} - \omega_q)\tau]$$

Considering the large τ limit of Eq. (54), the sinc functions in the integrand become highly peaked around $q = 0$, and to a good approximation can be replaced by Dirac deltas proportional to $\delta(q)$. However, the limit must be taken with care since the general term has a discontinuity as q goes to 0. We split the integrals for negative and positive momenta and in the limit of large τ we have $\int_0^\epsilon f(x) \tau \text{sinc}(\tau \delta\omega_{L/B}(q)) dq \rightarrow f(0^+) \int_0^\infty \tau \text{sinc}(\tau \delta\omega'_{L/B}(0^+) q) dq = \frac{\pi}{2} f(0^+) / |\delta\omega'_{L/B}(0^+)|$, and similarly on the other side. Explicitly, for $\tau \rightarrow \infty$

$$\int_{-\epsilon}^\epsilon |V_3(k, q)|^2 (n_q^{\text{th}} - n_{k+q}^{\text{th}}) \tau \text{sinc}[\delta\omega_L(q)\tau] dq = \frac{\pi}{2} \left(\lim_{q \rightarrow 0^-} + \lim_{q \rightarrow 0^+} \right) |V_3(k, q)|^2 (n_q^{\text{th}} - n_{k+q}^{\text{th}}) \\ \int_{-\epsilon}^\epsilon |V_3(k - q, q)|^2 (n_{k-q}^{\text{th}} + n_q^{\text{th}} + 1) \tau \text{sinc}[\delta\omega_B(q)\tau] dq = \frac{\pi}{2} \left(\lim_{q \rightarrow 0^-} + \lim_{q \rightarrow 0^+} \right) |V_3(k - q, q)|^2 (n_{k-q}^{\text{th}} + n_q^{\text{th}} + 1) \quad (55)$$

In both integrals, in the limit q to 0, $n_{k \pm q}^{\text{th}}$ and 1 are negligible compared to n_q^{th} that diverges as $1/q$. In addition, we have $|\delta\omega'_{L/B}(0^\pm)| = |v_{\text{gr}}(k) \mp c|$. Combining the above two limits we get the formula (III A) for the decay rate of phonon of momentum k . Therefore

$$I_k(\tau) = \int_0^\tau D_k(\tau') d\tau' \rightarrow \pm \Gamma_k \quad \text{as} \quad \tau \rightarrow \pm\infty, \quad (56)$$

where the opposite limit is obtained by anti-symmetry. I_k is asymptotically constant in both directions and only varies in the vicinity of $\tau = 0$. As a consequence D_k is peaked around $\tau = 0$. We now assume that D_k is sufficiently peaked compared to the variation of δn_k so that the integral in Eq. (49) only picks out its instantaneous value $\delta n_k(t)$

$$\int_0^t dt' D_k(t - t') \delta n_k(t') \approx \delta n_k(t) \int_0^t dt' D_k(t - t') , \quad (57) \\ = \delta n_k(t) I_k(t) .$$

The equation of motion becomes

$$\partial_t \delta n_k = -\delta n_k(t) I_k(t) + \int_0^t dt' \sum_{q \neq -k} M_{k,k+q}(t-t') \delta n_{k+q}(t'). \quad (58)$$

Let us now consider the ideal situation where at initial time every mode but a single mode k is exactly thermal: $n_q(0) = n_q^{\text{th}} + \delta n_{\delta k,q}$ where δn is the number of phonons added in mode k on top of the quasicondensate. This corresponds to the situation analyzed in Sec. IV A. Then we may neglect the matrix response function (since $\delta n_q \ll \delta n_k$):

$$\partial_t \delta n_k = -\delta n_k(t) I_k(t). \quad (59)$$

Finally we take a large time limit to have $I_k(t) \rightarrow \pm \Gamma_k$ so that in this limit n_k obeys

$$\partial_t \delta n_k = -\delta n_k(t) \Gamma_k, \quad (60)$$

i.e., the population of the mode k decays exponentially at the rate predicted by Eq. (III A).

E Matrix response function: the slowing effect of a finite width

It remains to give the form of $M_{k,k+q}$ and describe its effects. It reads:

$$\begin{aligned} M_{k,k+q}(t-t') = & 8 |V_3(k, q)|^2 (n_q^{\text{th}} + n_k^{\text{th}} + 1) \cos[(\omega_{k+q} - \omega_k - \omega_q)(t-t')] \\ & + 8 |V_3(k+q, -q)|^2 (n_q^{\text{th}} - n_k^{\text{th}}) \cos[(\omega_k - \omega_{k+q} - \omega_q)(t-t')] \\ & + 8 |V_3(k, k+q)|^2 (n_{2k+q}^{\text{th}} - n_k^{\text{th}}) \cos[(\omega_{2k+q} - \omega_k - \omega_q)(t-t')] \end{aligned} \quad (61)$$

Firstly, a couple of technical remarks. Notice that there is an indicator function in the last term, stating that this piece should not be evaluated at $q = -2k$. Also, the $q = 0$ term has been included to compensate for the inclusion of the opposite term in $D_k(t-t')$. Therefore, even in the case where only a single mode is significantly occupied, the matrix response function can never be completely neglected and (60) has to be amended. This first limitation is dealt with in Appendix B below.

$M_{k,k+q}$ represents the indirect interaction of phonons in the modes k and $k+q$ through the thermal population. Equation (61) shows that there are three such processes. The first term corresponds to the conversion between phonons of wavevectors k and q and that of wavevector $k+q$, which we write symbolically as $(k, q) \leftrightarrow k+q$. Similarly, the second term encodes processes of the form $(k+q, -q) \leftrightarrow k$, and the third term $(k+q, k) \leftrightarrow 2k+q$. Even though the perturbation is initially localised in the mode k , the decay process will generate a non-zero δn_q in the vicinity of both $q = 0$ and $q = k$. We shall assume that the perturbation spectrum is relatively narrow around these two points. Equation (58) then implies

that we need only consider the values of $M_{k,k+q}$ for $k+q \approx k$ and $k+q \approx 0$.

Consider first $k+q \approx 0$. In the second and third terms of (61), the factor of $|V_3|^2$ and the combination of thermal populations independently vanish in the limit $k+q \rightarrow 0$, making these terms negligible. In the first term, only the factor of $|V_3|^2$ tends to zero, but the rapid oscillations of the cosine function at frequencies close to $2\omega_k$ will greatly suppress its contribution to the integral of (58). Therefore, the back-reaction from the decay products at very low momenta is expected to be negligible.

On the other hand, for $k+q \approx k$ only the contribution of the third term in (61) is expected to be negligible, for the frequency of the cosine function is large (roughly $\omega_{2k} - 2\omega_k$) and suppresses its contribution. We are then left with contributions coming from the first and second terms. Their frequency differences are small, preventing any averaging out due to rapid oscillations, and can be written to first order in q . Their coefficients have non-vanishing limit since the vanishing of $|V_3|^2$ is compensated by the divergence of the thermal population, and can be approximated by their low- q limits. Finally, we get ¹²

$$M_{k,k+q}(t-t') \approx \frac{8}{N_{\text{at}}} \lim_{q \rightarrow 0^+} \left[|V_3(k,q)|^2 n_q^{\text{th}} \right] \cos[q(v_{\text{gr}}(k) - c)(t-t')] + \frac{8}{N_{\text{at}}} \lim_{q \rightarrow 0^-} \left[|V_3(k,q)|^2 n_q^{\text{th}} \right] \cos[q(v_{\text{gr}}(k) + c)(t-t')] . \quad (62)$$

Equation (62) will be used to compute the correction to the FGR prediction of the decay rate due to the finite width of the peak, *i.e.*, to the back-reaction of neighbouring modes.

Appendix B DEVIATIONS FROM FGR RESULT

We turn in this appendix to the deviations with respect to the simple exponential decay predicted by the FGR. Each of these deviations is described by equation of motion (49).

A Non-exponential decay of IR phonons

1 Numerical observations

It was shown in the main text that, in the first set of simulations where an initial probe is simply injected into a single phonon mode, the numerically observed decay of modes with a large enough momentum ($k\xi \geq 2.5$) is very well described by an exponential decay at a rate given by Eq. (III A). In Fig. 7,

¹² Notice that we dropped the terms n_k^{th} and $+1$ compared to n_q^{th} which is divergent as $q \rightarrow 0$. However, if we work in the discrete setting and consider a small $k \sim j2\pi/L$, then one could argue that q can never be less than $2\pi/L$ so that n_q^{th} is always finite and of the order of n_k^{th} . Still, in the limit $k \approx 0$, these terms come with opposite signs and practically cancel out unlike the ones in n_q^{th} .

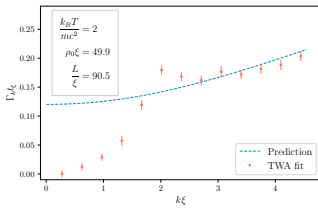


Figure 7. Best fit values for $\hbar\Gamma_k$ extracted from the TWA simulations as a function of $k\xi$. The window shown here is larger than that of Fig. 3 while all physical parameters are fixed at the same values.

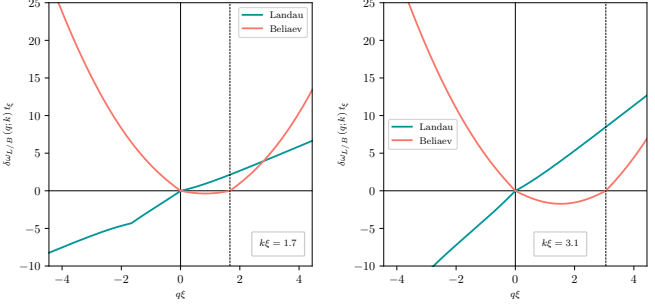


Figure 8. Plots of the Rabi frequencies associated to Landau and Beliaev damping processes $\delta\omega_{L,B}(q; k)$ as a function of q for $k\xi = 1.7$ (left.) and $k\xi = 3.1$ (right.). The value of k is shown by the vertical dashed line.

a more complete set of results is shown, in which the template $\delta n_k(t) = A \exp(-\Gamma t + \gamma t^2/2)$ is fitted to the behavior of a larger set of initial momenta. This figure demonstrates that there are significant deviations at lower $k\xi$, with the values of Γ_k extracted from the simulations going to zero as $k\xi \rightarrow 0$ instead of the predicted finite limit. These deviations are due to the fact that the FGR result is only valid within a certain time window [30, 34].

We review the last approximation made in the passage from Eq. (59) to Eq. (60), which is that t be large enough for $I_k(t)$ to be equal to its asymptotic value, or equivalently (as shown in Eqs. (55)-(56)), being able to replace $t \text{sinc}(\delta\omega_{L/B}t)/\pi$ by a δ function. This requires that $\delta\omega_{L/B}(k, q)t$, which vanishes for the elastic scattering channel $q = 0$, should nevertheless reach a large enough value in its vicinity so that the most significant part of the sinc is squeezed into the region of constant effective interaction strength. We may introduce a critical response time, t_{crit} , which marks the time after which this condition is satisfied. This critical time depends on k , so that, for a fixed time window $t/t_\xi \in [0, 10]$ over which the fit is performed, those k for which $t_{\text{crit}}/t_\xi \ll 10$ will be in the FGR regime, while those k for which $t_{\text{crit}}/t_\xi \gg 10$ will be in an early-time regime long before the FGR can be applied.

Examples of $\delta\omega_{L/B}(q, k)t_\xi$ are plotted in Fig. 8, for $k\xi = 1.7$ and 3.1. The key observation is that $|\delta\omega_B|$ is the most significantly constrained, especially in the window $q \in [0, k]$ where the decay processes are expected to be most efficient (as the effec-

tive interaction strength is largest there). It is this restriction on $|\delta\omega_B|$ that is most clearly responsible for the failure of the approximation. $|\delta\omega_B|$ reaches a maximum $\delta\omega_B^{(\max)} = \omega_k - 2\omega_{k/2}$ at $q = k/2$. At lowest order in $k\xi$, we have $\delta\omega_B^{(\max)} t_\xi \approx \frac{3}{32} (k\xi)^3$. The critical response time can be (somewhat arbitrarily) defined via $\delta\omega_B^{(\max)} t_{\text{crit}} = 2\pi$, but the key point is that it increases quickly at low momentum: $t_{\text{crit}}/t_\xi \propto (k\xi)^{-3}$. Since the fitting window of Fig. 7 extends only up to $t/t_\xi = 10$, we require at least $10 \delta\omega_B^{(\max)} t_\xi \approx (k\xi)^3 \geq 2\pi$, which imposes $k\xi \geq 2$. This estimate is in close correspondence with the onset of deviations seen for $k\xi \sim 2$ in Fig. 7. Moreover, as the total duration of the simulation is varied, the estimated threshold for $k\xi$ will vary only slowly, as $(t/t_\xi)^{-1/3}$.

At lower $k\xi$, we are well outside the validity regime of the FGR, in an early-time regime where a t^2 behavior is expected [34] (see below). Furthermore, for very low-lying modes $k \sim 2\pi/L$, the range $q \in [0, k]$ will be poorly sampled. We thus expect a strong suppression of the Beliaev component of the decay as $k\xi \rightarrow 0$ due to the small number of available modes to decay to, in accordance with similar remarks made in [9, 10]. While such low-lying k mode can still decay via a Landau process we expect that in this regime the system becomes sensitive to the discreteness of excitations and this process would thus require a separate analysis (see, *e.g.*, Ref. [36]).

2 Approximate analytical description of behavior

We expect that these low- k modes can still be well described by Eq. (44). However, we shall adopt an early-time approximation, assuming that $N_{p,q}(t)$ varies sufficiently slowly for it to be taken out of the time-integral as an overall prefactor. We get

$$\begin{aligned} \partial_t n_k = & 8 \sum_{q \neq 0, -k} \left\{ \frac{|V_3(k, q)|^2}{N_{\text{at}}} [n_{k+q} (n_k + n_q + 1) - n_k n_q] t \text{sinc}[\delta\omega_B t] \right. \\ & \left. - 8 \sum_{0 < q \leq k/2} \left\{ \frac{|V_3(k - q, q)|^2}{N_{\text{at}}} [n_k (n_{k-q} + n_q + 1) - n_{k-q} n_q] t \text{sinc}[\delta\omega_B t] \right\} \right. \end{aligned} \quad (63)$$

where we have used the symmetry $k \rightarrow k - q$ in the second term to sum over only half the momenta, compensating with the inclusion of a factor 2. This equation is then numerically solved twice: once using only the thermal population as an initial state, and a second time using the thermal population plus the δn phonons added in the mode k . We then take the difference to obtain the green curves in Fig. 9. This procedure correct for small variations of the background thermal population. A good agreement is found with the TWA simulations even for large values of $k\xi$.

Let us try to get an analytical estimate for the behavior of the low-lying modes. Their critical time being very large, it is appropriate to consider the early-time limit of Eq. (59) where $\delta n_k(t) \approx \delta n_k(0)$ and $\text{sinc}(\delta\omega \tau) \sim 1$, therefore $I_k(\tau) = \alpha_k \tau$ for

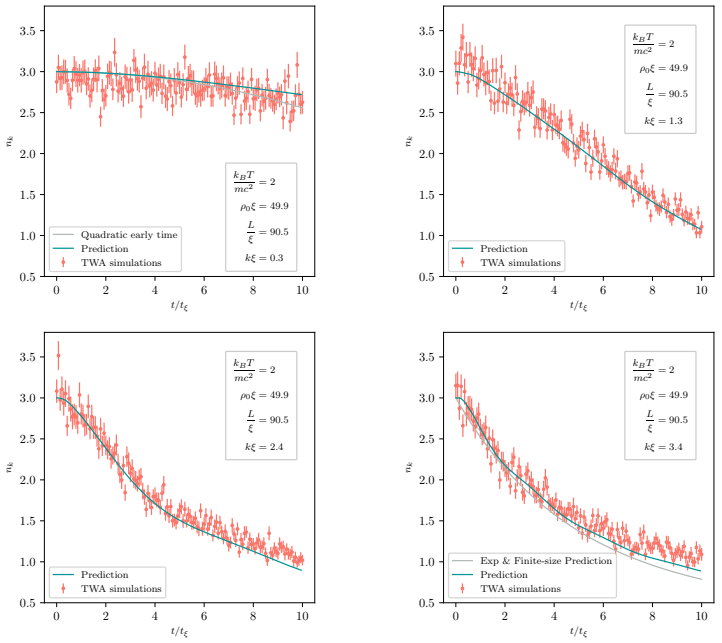


Figure 9. Number of phonons n_k in the mode $k\xi = 0.3$ (top left), $k\xi = 1.3$ (top right), $k\xi = 2.4$ (bottom left), $k\xi = 3.4$ (bottom right) as a function of time t/t_ξ for $k_B T/mc^2 = 2$, $\rho_0 \xi = 49.9$, $L/\xi = 90.5$ and $n_r = 400$ realisations. Each plot is comprised of $n_t = 140$ points. The red dots correspond to the result of the TWA simulations. The green line is obtained by taking the difference of the prediction of Eq. (63) for the background thermal population and the full population. The gray line in the top left panel corresponds to the prediction for an early-time quadratic decay given by Eq. (64), while in the bottom right panel it corresponds to the prediction of an exponential decay at the rate given by Eq. (III A) of the text with corrections given by Eq. (III C).

a certain constant α_k . This gives

$$\delta n_k(t) = \delta n_k(0) (1 - \alpha_k t^2/2) , \quad (64)$$

where we now have to calculate α_k .

We cannot simultaneously set all the sinc functions in Eq. (54) to 1, as the $+1$ term in the second sum would then lead to a divergence. Indeed, for large q the frequency difference diverges and the sinc decay accordingly quickly. We typically consider the evolution of the system over time scales of the order of t_ξ . Considering $k\xi \sim 0.3$, we want to sort every mode q in two categories. Either $\delta\omega_{L/B}(q, k)t_\xi \ll 1$, then this mode can be considered to experience an early time behaviour and the related sinc can be set to 1, or $\delta\omega_{L/B}(q, k)t_\xi \gg 1$ then the sinc should be set to 0. We simply exclude the latter modes from the sum.

On the one hand, as noted above for the Beliaev-type channels it is clear from Fig. 8 that the modes $q \in [0, k]$ oscillate at frequencies smaller than the others and should be the one kept in the sum. On the other hand there is no clear separation for the Landau-type scatterings. However, the terms associated to $\delta\omega_L$ in Eq. (54) are already suppressed at large q by an

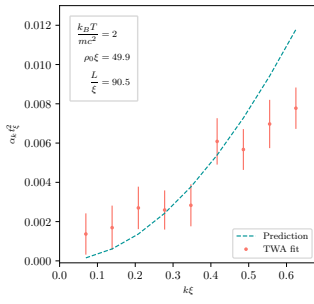


Figure 10. Red dots are the best fit values for $\alpha_k t_\xi^2$ extracted from the TWA simulations using a template of the form of Eq. (64). The green dashed line is the prediction for $\alpha_k t_\xi^2$ of equation Eq. (65). The simulation parameters are $k_B T / m c^2 = 2$, $\rho_0 \xi = 49.9$ and $L / \xi = 90.5$ with $n_r = 400$. The fits are performed over a time-window $t/t_\xi \in [0, 5]$, which is half the time-window used in for the fits of Γ_k in Figs. 2 and 3, but with the same number of points $n_t = 140$.

exponentially decaying thermal population n_q in factor. The inclusion of them in the sum should then be irrelevant for the resulting numerical prediction and we include all the modes in the first sum. We then get:

$$\alpha_k = \frac{8}{N_{\text{at}}} \left\{ \sum_{q \neq 0, -k} |V_3(k, q)|^2 (n_q^{\text{th}} - n_{k+q}^{\text{th}}) + \sum_{0 < q \leq k/2} |V_3(k - q, q)|^2 (n_{k-q}^{\text{th}} - n_q^{\text{th}}) \right\} \quad (65)$$

In Fig. 10 we compare this prediction with the best-fit value of α for the template $A(1 - \alpha t^2/2)$ applied to the TWA simulations with A and α as fitting parameters. We considered the smallest values of k of Fig. 3 in the text and used the same time-window $t/t_\xi \in [0, 10]$. The agreement is good for very small k and deviations appear around $k\xi = 0.5$. This can be understood using Fig. 9. For $k\xi = 0.3$ the decay is satisfyingly described by the quadratic prediction Eq. (65), while for $k\xi = 3.4$ it is well described by the exponential prediction of Eq. (III A) plus the finite size-correction examined in the next section. On the other hand for the intermediate values $k\xi = 1.4$ and $k\xi = 2.4$, corresponding to the second and third panels, the decay from $t/t_\xi = 0$ to $t/t_\xi = 10$ is neither quadratic nor exponential all the way. $k\xi = 2.4$ is precisely the value around which our prediction of exponential decay seems to break down in Fig. 7.

B Finite-size effect

1 Description of effect

The inequivalent curves in Fig. 11 plotted for $L/\xi = 90.5$ and $L/\xi = 181$ demonstrate that the dynamics of n_k in the TWA simulations is not completely insensitive to the size of the system L . Therefore, the L -independent exponential decay rate of Eq. (III A) cannot fit exactly the result of the simulations.

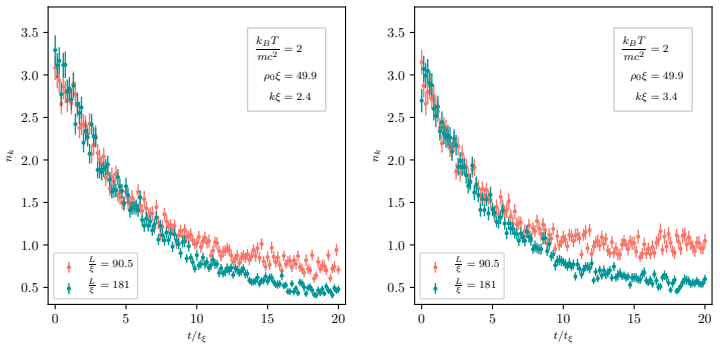


Figure 11. Number of phonons n_k in the mode $k\xi = 2.4$ (left.), $k\xi = 3.1$ (right.) as a function of time t/t_ξ for $k_B T / mc^2 = 2$, $\rho_0 \xi = 49.9$ and $n_r = 400$ realisations. The red dots correspond to $L/\xi = 90.5$ and the green ones to $L/\xi = 181$.

The key to understanding this effect is the $q = 0$ contribution to the matrix part of the response function. For a singularly occupied mode, this is the only term in the matrix part that plays any significant role. Recall that it must be included simultaneously in both the diagonal and matrix parts, so that there is no net change in the total response function. If these $q = 0$ terms were not included, the effect would be related to the *absence* of the $q = 0$ term in the *diagonal* response function.

What effect does this term have? Considering the case where a certain number of phonons are injected at $t = 0$, and working in the regime where the diagonal response function reduces to a Dirac delta, we have

$$\begin{aligned} \partial_t (\delta n_k) &= -\Gamma_k \delta n_k + \int_0^t dt' M_{k,k}(t-t') \delta n_k(t') \\ &= -\Gamma_k \delta n_k + \gamma_k \int_0^t dt' \delta n_k(t'), \end{aligned} \quad (66)$$

where we have made it manifest that $\gamma_k \equiv M_{k,k}$ does not depend on $t - t'$:

$$\gamma_k = 8 \left(\lim_{q \rightarrow 0^+} + \lim_{q \rightarrow 0^-} \right) \frac{|V_3(k, q)|^2}{N_{\text{at}}} n_q^{\text{th}}. \quad (67)$$

At fixed density, N_{at} is proportional to L , so $\gamma_k \propto 1/L$. Performing computations explicitly we get Eq. (III C) of the main text.

It is clear from Eq. (66) that there is some push-back on the decay. The simplest solution is when δn_k does not vary much over the duration of interest, so that we can pull it out of the integral:

$$\partial_t (\delta n_k) \approx (-\Gamma_k + \gamma_k t) \delta n_k, \quad (68)$$

with solution given by Eq. (III C) of the main text. In this approximation, the γ_k term simply provides a constant deceleration to the decay rate.¹³ In numerical simulations, we have

¹³ Here a time dependence of the rate appears through the L -dependence

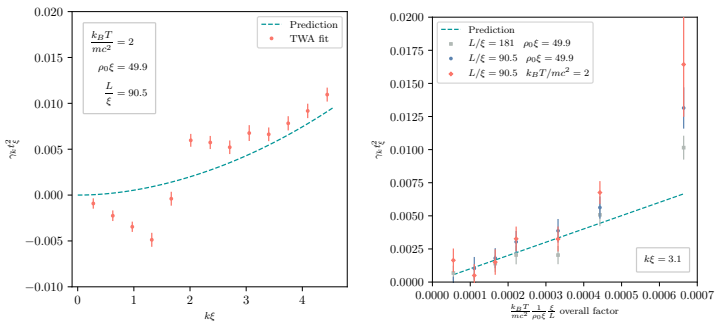


Figure 12. (Left) Plot of $\gamma_k t_\xi^2$ as a function of $k\xi$. The parameters and data used are the same as Fig. 3. (Right) Plot of $\gamma_k t_\xi^2$ as a function of $k_B T / mc^2$, $\rho_0 \xi$ and L/ξ . The parameters and data used are the same as Fig. 2. In both panels the red dots are the best fit values extracted from the TWA simulations using Eq. (III C) as a template. The green dashed line is the prediction of equation Eq. (III C).

used Eq. (III C) as a template to extract the fitted values of Γ_k shown in the figures of the main text. The values of γ_k extracted from the same fits in the time window $t/t_\xi \in [0, 10]$ are plotted in Fig. 12, along with the prediction of Eq. (III C). It demonstrates a reasonable agreement in the same regime of validity as the one of exponential decay ($k\xi \leq 2.5$) shown in Fig. 3. At smaller $k\xi$, the system exhibits the t^2 behavior described above, and this becomes reflected in the fitting parameters. In particular, γ_k moves over into the t^2 coefficient $-\alpha_k$, and we see indeed that its sign changes.

That said, there is a difference between how γ_k and α_k enter the expression of $n_k(t)$, respectively, linearly and exponentially. In particular, while the early t^2 behavior is described by the single fitting parameter α_k , the inclusion of γ_k as a correction to the exponential decay means that it is one of two fitting parameters. The time window of fit is large enough so that the difference between these two templates (applied on the same data) is visible, as illustrated by a direct comparison of Fig. 10 with Fig. 12.

2 Relation to finite size and resolution in k -space

Is there an intuitive way to link this apparent deceleration to the finite size of the system? As our calculations are based in

i.e. the finite size effect. The reader might wonder why we never mentioned possible deviations to the exponential decay solely due to the finiteness of t and hence the sinc being not exactly a Dirac delta in Eq. (54). The finite t deviation can be obtained by expanding the effective interaction strength around $q = 0$. Indeed, a Dirac delta would pick up only the value at $q = 0$, first term in the expansion, while a sinc is still sensitive to the deviations from the value at $q = 0$ which are encoded by the higher order terms. We can then proceed by integration by part in Eq. (54) to find that each term in q^n corresponds to a term decaying as $1/t^n$. A careful analysis shows that by combining both integrals the next to leading order term in the expansion around $q = 0$ is in

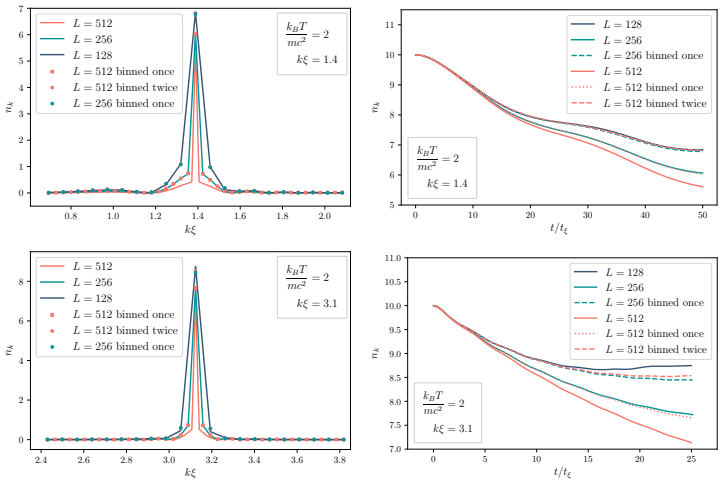


Figure 13. Number of non-thermal phonons δn_k as a function of $k\xi$ at $t/t_\xi = 50$ (top left.), and $t/t_\xi = 25$ (bottom left.), when $\delta n = 10$ phonons were initially added in the mode $k\xi = 1.4$ (top left.), or $k\xi = 3.1$ (bottom left.). This number is shown in full line for different values of density of the grid $\Delta k = 2\pi/L$ where $L = 128$ (blue), $L = 256$ (green) and $L = 512$ (red). The number of modes in the simulation for $L = 512$ is then halved by merging nearby modes into a single bin via Eq. (69) and shown in red squares. Equation (69) is then applied on the set of modes for $L = 256$ and repeated via Eq. (70) on the new set of modes to have the same set of modes as $L = 128$ shown respectively the red and green dots. They are almost everywhere superimposed so that they can hardly be distinguished. The evolution as a function of t/t_ξ of the population in the probe mode is shown in full lines on the right figures for $k\xi = 1.4$ (top right.) and $k\xi = 3.1$ (bottom right.). The result of the binning procedure from $L = 512$ to the set of modes of $L = 256$ is shown in red dotted line. The result of the second binning and the binning of $L = 256$ to the set of modes of $L = 128$ is shown in red and green dashed lines. The initial thermal state used corresponds to $k_B T / mc^2 = 2$. The evolution is performed using the same numerical strategy as described for the green curves of Fig. 9 for $\rho_0 \xi = 49.9$.

Fourier space, the size of the system enters through the discreteness of the available modes. To some extent, the discretized description should simply provide a finite-resolution view of the continuous-mode (infinite- L) case.

We have investigated this by numerically solving Eqs. (42) directly. We inject a number of phonons in a single mode, and allow the system to evolve in time, for three different values of L . We then compare the evolution by appropriately binning the phonon number spectrum n_k when the k -space resolution is higher, so that each simulation is ultimately represented using the same set of discrete modes. This procedure can also be seen as restricting attention to a section of length D of the system so that the relevant momenta are $2\pi n/D$ rather than $2\pi n/L$, for $n \in \mathbb{Z}$. The precise relationship between the two sets of modes is equivalent to a choice of window in position space, over which the Fourier transform of the field is taken. We take a more heuristic approach here.

For example, to map the data for a simulation of a system of length L onto the set of modes applicable to a system or section of length $L/2$, the phonons in every second mode must be reallocated to neighbouring modes. We do this by dividing them symmetrically into their two nearest neighbours, half the phonons going into the mode above, half into the mode below. So:

$$n_k^{\text{binned}} = \frac{1}{2}n_{k-\Delta k} + n_k + \frac{1}{2}n_{k+\Delta k}. \quad (69)$$

Similarly, to map onto a set of modes applicable to a system of length $L/4$, we adopt the following binning procedure:

$$n_k^{\text{binned}} = \frac{1}{2}n_{k-2\Delta k} + n_{k-\Delta k} + n_k + n_{k+\Delta k} + \frac{1}{2}n_{k+2\Delta k}. \quad (70)$$

Results are shown in Fig. 13, for two different values of k . Interestingly, the binning procedure applied to larger- L data gives a very good approximation to the data for smaller- L . This suggests that there is no new (relevant) physics due to the discretization in k -space, in the sense that we can solve for a continuous k -space (*i.e.*, in the limit of infinite L) and then simply apply a suitable binning procedure to see how the discretized system behaves. Similarly, the (relevant) physics on a section of length L of a system of infinite size is the same as that of a finite-size system of the same length.

This binning procedure provides an intuitive explanation for the apparent deceleration of the decay encoded in Eq. (66). For, while the singularly occupied mode decays exponentially, the lost phonons are kicked into nearby modes, whose occupation numbers grow in time. This is what is captured by the second term of Eq. (66): it represents the growth of those modes in the continuous spectrum that are very near k , but which, due to the finite resolution in k -space, are included in the same bin. By local conservation of the number of phonons described in the text, if we bin all modes within the width of the peak into a single mode, we effectively suppress the decay, *i.e.*, if we consider a sufficiently short section we will not witness the decay of the phononic excitations. The time-scale comparison of Eq. (34) shows that the relevant length scale is the coherence length of the quasicondensate r_0 . For shorter distances we do not expect to be able to resolve the decay of the phonons and the broadening of the peak in momentum space.

Of course, we do not expect the good correspondence shown in Fig. 13 between simulations of different L to last indefinitely. Eventually, the relevant components of the system – namely, the thermal spectrum and the probe – will be able to tell that they are on a finite torus, rather than a finite section of an infinite-size system, and we can then expect the different simulations to diverge significantly. The critical time can be conceptualized as the re-crossing of the relevant components, which would not re-cross if the system were truly infinite in length. There are two such times, corresponding to the re-crossing of the probe with the positive and negative wave vector components of the thermal spectrum, which propagate and speeds c

and $-c$, respectively. Then the re-crossing times are

$$t_+^{\text{rec}} = \frac{L}{v_{\text{gr}}(k) - c}, \quad t_-^{\text{rec}} = \frac{L}{v_{\text{gr}}(k) + c}. \quad (71)$$

Since these are simply proportional to L , it makes sense that the simulation with smallest L should be the first to show deviations, as we see in the lower-right panel of Fig. 13. (In this example, $t_-^{\text{rec}}/t_\xi = 30.1$.) However, the deviations remain small, likely due to the weaker coupling between the probe and the negative part of the thermal spectrum. We expect more significant differences after time t_+^{rec} ; indeed, we have observed some recurrence of δn_k on this time scale.

C Finite width of the resonant peak

Here we will discuss the corrections to the decay rate induced by a finite peak width. In particular, we will show how the approximate corrections shown in Fig. 5 were calculated.

The reduction of the decay rate is induced by the off-diagonal part of the matrix response function. Phonons next to the mode of interest also decay due to interaction with the thermal population, and some of these are thus transferred into the mode of interest. Since they can only contribute positively to the occupation number of the main mode, they decelerate the decay of n_k , thereby effectively reducing the decay rate Γ_k .

In our first set of simulations, we impose a very narrow initial peak, and this effect does not have a chance to build up significantly. By contrast, in the second set of simulations where the peak is induced via parametric resonance, it naturally has a finite width which is observed to depend on the interaction strength. We thus see signs of the overall decay rate being smaller than the predicted value at larger interaction strengths, as seen in Fig. 5.

We wish to predict the expected reduction in Γ_k given the observed peak width. To this end we employ Eq. (28) with two key assumptions. First, we modify the equation to include a source term that induces exponential growth, modeling the parametric resonance. This is achieved by adding $G_k \delta n_k$ to the right-hand side, where G_k is the predicted growth rate in the absence of any phonon-phonon interactions:

$$G_k = \frac{1}{2} A_k \omega_k, \quad (72)$$

which is consistent with Eq. (36) for sufficiently large t . Second, we assume that a steady state is reached where the occupation numbers of all relevant nearby modes grow at the same net rate, G_{net} . Therefore, the shape of the peak is assumed constant, and evolves in time simply according to

$$\delta n_k(t) = f_k e^{G_{\text{net}} t}. \quad (73)$$

This appears to be consistent with numerical observations see Fig. 14.

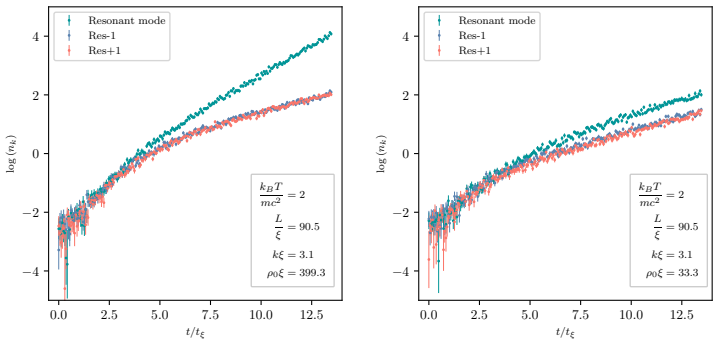


Figure 14. Logarithm of the mean occupation of the resonant mode (green) and its nearest neighbours on either sides (red, blue) as a function of time extracted from the TWA simulations. The parameters for the initial state, the amplitude and the duration of the modulation are the same as in Fig. 4. We see a quicker convergence of the growth rates within the peak on the right plot, which is due to a smaller atomic density resulting in a larger effective interaction strength.

Incorporating these assumptions in Eq. (28) for the occupation number of the resonant mode yields a consistency equation that determines the expected net growth rate, G_{net} :

$$\begin{aligned} \partial_t (\delta n_k) &= G_{\text{net}} \delta n_k = (G_k - \Gamma_k) \delta n_k + \int_0^t dt' \sum_{q \neq -k} M_{k,k+q}(t-t') \delta n_{k+q} \\ &\approx (G_k - \Gamma_k) \delta n_k + \int_0^t dt' \int dq \frac{L}{2\pi} M_{k,k+q}(t-t') \delta n_{k+q} \\ &= (G_k - \Gamma_k) \delta n_k + \delta n_k(t) \int_0^t dt' \int dq \frac{L}{2\pi} M_{k,k+q}(t-t') \end{aligned}$$

where in the second line we have taken the continuum limit to replace the sum over wave vectors q by an integral. The exponential suppression in $t - t'$ allows us to replace the lower limit of the t' integral by $-\infty$, which in turn allows us to write the integral independently of t and to divide through the whole equation by $\delta n_k(t)$:

$$G_{\text{net}} = G_k - \Gamma_k + \int_0^\infty d\tau \int dq \frac{L}{2\pi} M_{k,k+q}(\tau) R_{k+q} e^{-G_{\text{net}}\tau}. \quad (75)$$

This self-consistency equation can be solved for G_{net} , given that we know G_k , Γ_k , and the profile shape R_{k+q} . If we assume that the peak is narrow enough so that only small q are relevant, we may replace $M_{k,k+q}$ with its small- q limit, given in Eq. (62). This gives

$$\begin{aligned} G_{\text{net}} &\approx G_k - \Gamma_k + \int_0^\infty d\tau \int dq \{ f_+ \cos[q(v_{\text{gr}}(k) - c)\tau] + f_- \cos[q(v_{\text{gr}}(k) + c)\tau] \} \\ &= G_k - \Gamma_k + \int dq \left\{ f_+ \frac{G_{\text{net}}}{G_{\text{net}}^2 + q^2(v_{\text{gr}}(k) - c)^2} + f_- \frac{G_{\text{net}}}{G_{\text{net}}^2 + q^2(v_{\text{gr}}(k) + c)^2} \right\} \end{aligned}$$

where we have defined

$$f_{\pm} = \frac{4}{\pi} \frac{L}{N_{\text{at}}} \lim_{q \rightarrow 0^{\pm}} \left\{ |V_3(k, q)|^2 n_q^{\text{th}} \right\}. \quad (77)$$

All that remains is to plug in a suitable profile R_{k+q} for the profile of the peak. Over a significant range of parameters, we observe numerically that the profile is fairly well described by a Lorentzian:

$$R_{k+q} = \frac{1}{(q/\sigma)^2 + 1}. \quad (78)$$

Adopting this as an ansatz and plugging it into the integral above, we find

$$G_{\text{net}} - \Gamma_+ g\left(\frac{G_{\text{net}}}{(v_{\text{gr}}(k) - c)\sigma}\right) - \Gamma_- g\left(\frac{G_{\text{net}}}{(v_{\text{gr}}(k) + c)\sigma}\right) = G_k - \Gamma_k, \quad (79)$$

where Γ_{\pm} are the predicted decay rates given in Eq. (III A) (recall that $\Gamma_k = \Gamma_+ + \Gamma_-$), and where we have defined

$$g(x) = \frac{1}{1 + x}. \quad (80)$$

The effective decay rate Γ_{eff} is then defined as $\Gamma_{\text{eff}} = G - G_{\text{net}}$, thus capturing the reduction of the growth rate due to phonon-phonon interactions. Note that Eq. (79) yields the expected behavior in the limits of very narrow and very broad peaks:

- When the peak is very narrow, we take $\sigma \rightarrow 0$ so that the argument of $g(x)$ becomes very large, and we can set $g(x)$ to 0. Then we just get $G_{\text{net}} = G_k - \Gamma_k$, so $\Gamma_{\text{eff}} = \Gamma_k$; this is exactly our prediction for the decay rate of a narrow peak.
- Conversely, when the peak is very broad so that we can send $\sigma \rightarrow \infty$, the argument of $g(x)$ becomes very small and we can replace $g(x)$ by 1. Since $\Gamma_k = \Gamma_+ + \Gamma_-$, Eq. (79) tells us that $G_{\text{net}} = G_k$, *i.e.*, there is no reduction of the growth rate. This corroborates our expectation that the broadness of the peak tends to suppress the decay rate.

To determine the corrected decay rates shown in Fig. 5, we first perform a fit of a Lorentzian profile to the occupation numbers around the resonant peak (including five data points on either side of the resonant mode), in order to extract the width σ . This is then used in Eq. (79) on order to determine the expected net growth rate G_{net} , and thereby the effective decay rate Γ_{eff} .

Appendix C MONTE-CARLO SIMULATIONS : TRUNCATED WIGNER APPROXIMATION (TWA)

In order to assess the validity of our predictions we compare them to the results of *ab initio* Monte-Carlo simulations of the system. To simulate the evolution of our quasicondensate we use the Truncated Wigner Approximation (TWA) also known as the classical field approximation. This method is based on the description of the state of the system by means of a quasi-probability distribution, the Wigner function. The TWA has

been repeatedly used to describe Bose gas [29, 37], specifically one-dimensional quasicondensate [38, 39], as well as a wide variety of other systems (polaritons, spins etc.) [40, 41]. We will restrict our presentation of the TWA to the necessary minimum and we refer to [29] for further details.

A TWA in a nutshell

The Wigner function $W(\Psi, \Psi^*)$ is a quasi-probability distribution in phase space defined by a bijective transformation of the density matrix $\hat{\rho}$. Under this transformation the von Neumann equation of motion on the density matrix then translates into a partial differential equation for W [42]. For the Hamiltonian (1) the equation reads

$$i\hbar\dot{W}(\Psi, \Psi^*) = - \int_{-\infty}^{+\infty} \left\{ \frac{\delta}{\delta\Psi} \left[\frac{1}{2m} \partial_x^2 \Psi + g(|\Psi|^2 - 1) \Psi \right] - \frac{1}{4} \frac{1}{\delta^2} \right\} \quad (81)$$

where the derivatives act on the element inside of the brackets multiplied by $W(\Psi, \Psi^*)$, see Eq. (23) of [29]. The truncation giving its name to the Truncated Wigner Approximation consists in neglecting terms with three derivatives or more. The resulting equation is then solved by using the method of characteristics. Practically, the Wigner function at time t is found by first sampling the Wigner function at initial time i.e. drawing a set of values $\Psi_i(x, t=0)$ according to the initial Wigner function and evolving these realizations under an equation of motion which is simply the classical counterpart of the Heisenberg equation of motion of $\hat{\Psi}$ under the Hamiltonian (1)

$$i\hbar\dot{\Psi} = -\frac{1}{2m}\partial_x^2\Psi + g|\Psi|^2\Psi. \quad (82)$$

The resulting $\Psi_i(x, t)$ represent a sampling of the Wigner function of the state at time t . This sample can be used to compute average values of observables $\langle A(\varphi_k) \rangle_{\text{TWA}}$ built from φ_k . It can be shown that expectation values computed treating the Wigner function as a *bona fide* probability distribution are equal to quantum expectation values of the associated operator when the expression is completely symmetrized in $\hat{\varphi}_k$ and $\hat{\varphi}_k^\dagger$. We will not discuss here the conditions of validity of this truncation and of its numerical implementation; some considerations can be found in [18, 28, 29].

B Numerical implementation

The numerical implementation of the TWA is a straightforward application of the above program. The evolution is performed using a discretized version of Eq. (82) and a split-step Fourier algorithm [43]. The initial state is taken to be thermal at a temperature T , up to the addition of a few “probe” phonons in the mode k for the first set of simulations. The exact thermal state of the system is approximated by the one associated to the quadratic Hamiltonian \hat{H}_2 i.e., the phonon modes

are completely uncorrelated to each other and their number spectrum is set equal to the Bose-Einstein distribution at temperature T . The realisations of the atomic field are built from n_r realisations of the phonon modes φ_k . These are themselves built drawing independently $\Re[\varphi_k]$ and $\Im[\varphi_k]$ according to a centered Gaussian distribution with variance

$$\sigma_k^2 = \frac{1}{2} \left(n_k + \frac{1}{2} \right) = \frac{1}{4} \coth \left(\frac{\hbar \omega_k}{2k_B T} \right). \quad (83)$$

It is then straightforward to check that the TWA averages reproduce the averages of the symmetrized quantum operators in a thermal state:

$$\langle \varphi_k \rangle_{\text{TWA}} = \langle \varphi_k^* \rangle_{\text{TWA}} = 0, \quad (84)$$

$$\langle \varphi_k^* \varphi_k \rangle_{\text{TWA}} = n_k^{\text{th}} + \frac{1}{2} = \frac{1}{2} \left\langle \left\{ \hat{\varphi}_k, \hat{\varphi}_k^\dagger \right\} \right\rangle, \quad (85)$$

where $\{\hat{A}, \hat{B}\} = \hat{A}\hat{B} + \hat{B}\hat{A}$ is the anti-commutator. Since this is not the exact thermal state we let the system evolve for a certain duration to be as close as possible to a stationary state that we use as an initial state. For the first set of simulations phonons are then injected in the mode k by transforming its amplitude $\varphi_k(0)$ according to

$$\varphi_k(0) \rightarrow \sqrt{1 + \frac{\delta n}{n_k^{\text{th}} + \frac{1}{2}}} \varphi_k(0), \quad (86)$$

so that $\varphi_k(0)$ still has Gaussian statistics but with a variance corresponding to $n_k = n_k^{\text{th}} + \delta n$. We then evolve under Eq. (82) both the realisations with and without the addition of probe phonons using the same code. We repeat this process for the n_r realisations. Finally we compute the average values for n_q with and without the probe at any time and take the difference to get the evolution of the average value $\delta n_q(t)$.

This method closely mirrors the step of the derivation of the decay rate laid out around Eq. (49) where we subtract the evolution of the background thermal population to single out the evolution of the probe. However the TWA encodes also the deviations to the linear case considered in our equations. For the second set of simulations we simply evolve the initial quasi-stationary state according to (82) where g is modulated according to (35), and proceed to the same type of averaging as in the first set. We want to stress that in any figure of this work, for both set of simulations, a data point at time $t + dt$ is *not* obtained simply by evolving the realisations at t time by an extra-step dt as this would result in strongly correlated data points. We rather start the whole evolution process from new realisations of the initial state and evolve them until $t + dt$.