

SIGNAL RECOVERY CHEATSHEET

Giacomo Tombolan giacomo.tombolan@mail.polimi.it

Notes

Follow the lessons: not everything is written here. This is just a recap with the most important things to know according to me. Most qualitative aspects and explanations are contained in the slides/videos. This is especially true for the photodetector section

This is just a brief guide when doing written exercises so do not expect to study theory here.

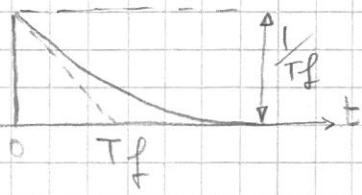
Do a lot of exercises so everything that is written here becomes natural and automatic.

There can be errors, do not take for granted every thing that is written here. Check slides/lessons/tutorials/book.

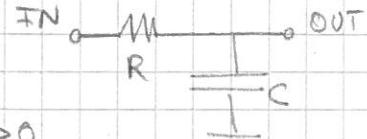
PART 1 - Filters

LPF - RC → Simplest filter

- Used in the 1st question to compute noise w/o filtering → preamplifier



$T_f = RC$ time constant → exp is zero after $\sim 5T_f$



$$h(t) = 1(t) \frac{1}{T_f} e^{-t/T_f} \text{ where } 1(t) = \begin{cases} 1 & \text{for } t \geq 0 \\ 0 & \text{for } t < 0 \end{cases}$$

Autocorrelation:

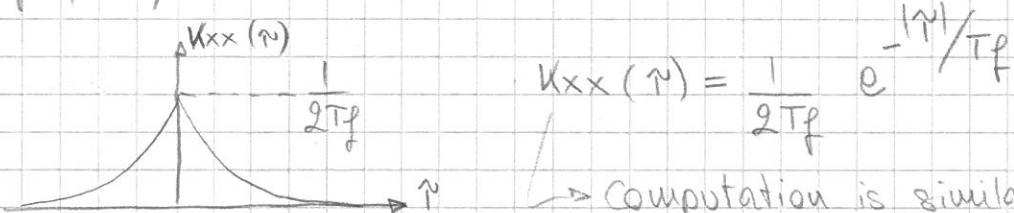
$$K_{xx}(0) = \int_{-\infty}^{+\infty} x(\tau)^2 d\tau = \frac{1}{T_f^2} \int_0^{+\infty} e^{-2\tau/T_f} d\tau = -\frac{1}{T_f^2} \cdot \frac{T_f}{2} \left[e^{-2\tau/T_f} \right]_0^{+\infty} = \frac{1}{2T_f}$$

because of $1(t)$

$$|W_{nn}(f)|^2 = \frac{1}{1 + (2\pi f T_f)^2}$$

Autocorr. can be computed in freq:

$$K_{xx}(0) = \int_{-\infty}^{+\infty} |W_{nn}(f)|^2 df = \frac{1}{2\pi T_f} \int_{-\infty}^{+\infty} \frac{1}{1 + (2\pi f T_f)^2} df = \frac{1}{2\pi T_f} \left[\arctan 2\pi f T_f \right]_{-\infty}^{+\infty} = \frac{\pi}{2\pi T_f} = \frac{1}{2T_f}$$



$$K_{xx}(\tau) = \frac{1}{2T_f} e^{-|\tau|/T_f}$$

Computation is similar to $K_{xx}(0)$

Signal filtering: heavily depends on signal shape

Noise filtering: suppose wide-band white noise.

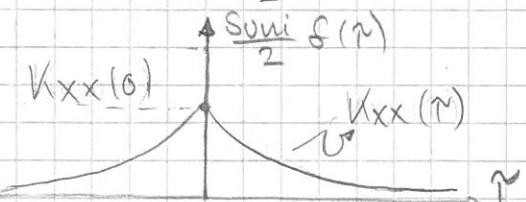
e.g.: preamp limits noise at 50 MHz and RC-LPF has a pole

$$\text{at } f_{RC} = \frac{1}{2\pi T_f} = 5 \text{ MHz} \quad T_{RC} = 10 \text{ TPA} \rightarrow f_{pole_{RC}} \ll f_{pole_{PA}}$$

In this case we can neglect PA pole, so $R_{nn}(\tau) = \frac{S_{uni}}{2} \cdot f(\tau)$

$$C_n^2 = \int_{-\infty}^{+\infty} R_{nn}(\tau) K_{xx}(\tau) d\tau = \frac{S_{uni}}{2} K_{xx}(0)$$

$$R_{nn}(\tau) = \frac{S_{uni}}{2} f(\tau)$$



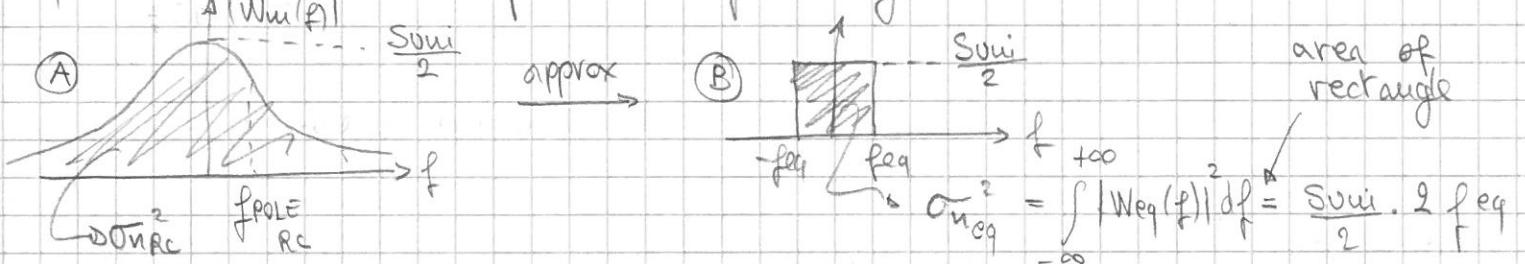
$$C_n^2 = \frac{S_{uni}}{2} K_{xx}(0) \rightarrow \text{delta narrows } -\infty/+ \infty \text{ integration to just the value in zero}$$

example: $T_{RC} = 10 \text{ TPA}$ $\sigma_n = \sqrt{S_{\text{uni}} \cdot \frac{1}{4 \text{ TPA}}} = \sqrt{S_{\text{uni}} \cdot \frac{1}{4 T_{RC}}}$

$\text{SNR} = \frac{V_p}{\sigma_n} \rightarrow$ Suppose signal isn't affected by RC-LPF

$\text{SNR}_{\text{RC}} = \sqrt{10} \text{ SNR}_{\text{PA}} \rightarrow \text{SNR increases by } \approx 3 \text{ times}$

Equivalent bandwidth for RC filtering



Noise integration in A has to be equal (equivalent) to case B

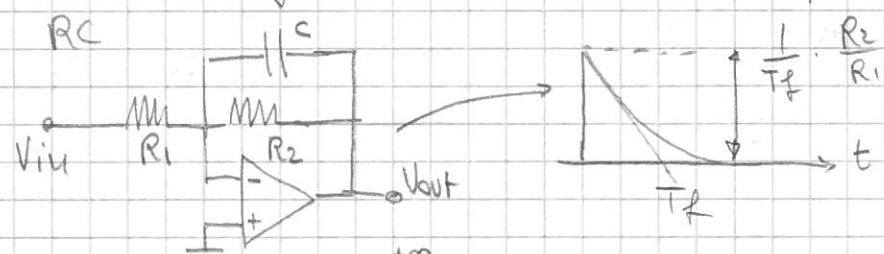
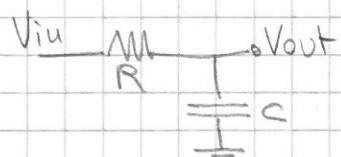
$$\sigma_{\text{RC}}^2 = \sigma_{\text{eq}}^2 \rightarrow \frac{S_{\text{uni}}}{2} \cdot \frac{1}{2T_f} = \frac{S_{\text{uni}}}{2} \cdot 2 \cdot f_{\text{eq}} \quad f_{\text{eq}} = \frac{1}{4T_f}$$

Knowing that, for a signal, $f_{\text{pole}} = \frac{1}{2\pi T_f}$ then we can do the following:

$$\frac{1}{4T_f} = f_{\text{eq}} = \frac{\pi}{2} f_{\text{pole}} \Big|_{\text{RC}} \quad \text{Therefore the equivalent noise BW is } \frac{\pi}{2}$$

higher than the RC pole frequency

Passive vs Active RC



$$\text{DC Gain} = \int_{-\infty}^{+\infty} h(t) dt = \frac{1}{j\omega} \cdot \frac{R_2}{R_1} (-j\omega) \int_0^{+\infty} e^{-\frac{\omega}{j\omega} t} dt = \frac{R_2}{R_1}$$

Note: DC gain does not depend on frequency and capacitor value
(This will not be true for Rateometer integrator)

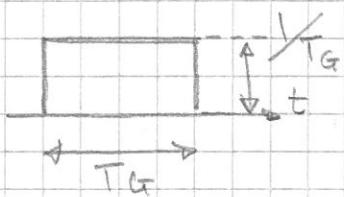
What about SNR? \rightarrow input signal is not affected by RC-LPF

$$\text{SNR} = \frac{V_p \cdot \frac{R_2}{R_1}}{\sqrt{\frac{S_{\text{uni}}}{2} K_{\text{noise}}(\omega)}} = \frac{V_p R_2 / R_1}{\sqrt{\frac{S_{\text{uni}} \cdot 1}{2T_f} \cdot \left(\frac{R_2}{R_1}\right)^2}} = \frac{V_p R_2 / R_1}{\sqrt{\frac{S_{\text{uni}} \cdot \frac{\pi}{2} f_{\text{pole}}}{2T_f}} \cdot \frac{R_2}{R_1}} = \text{SNR}_{\text{PASSIVE}}$$

SNR is not affected by DC gain.

When asked during the exam, answer is: Gain does not increase SNR but it helps to keep the noise of following stages low enough to be negligible (because signal will be increased thanks to the gain) Moreover, it exploits the full dynamic range of a possible ADC connected to the measurement system

GATED INTEGRATOR AND MOBILE MEAN FILTER

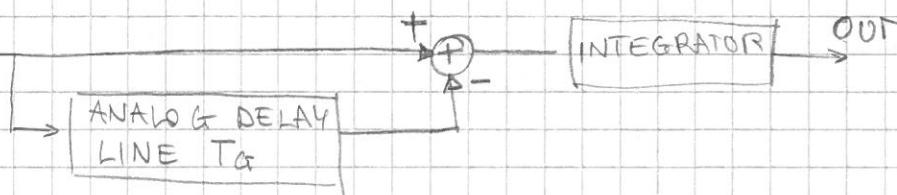


MMF makes use of analog delay lines \rightarrow
 \rightarrow be careful! T_G can't be really high

eg: $T_G = 20\text{ns}$ length of delay line = $c_0 \cdot 20\text{ns} = 6\text{m}$

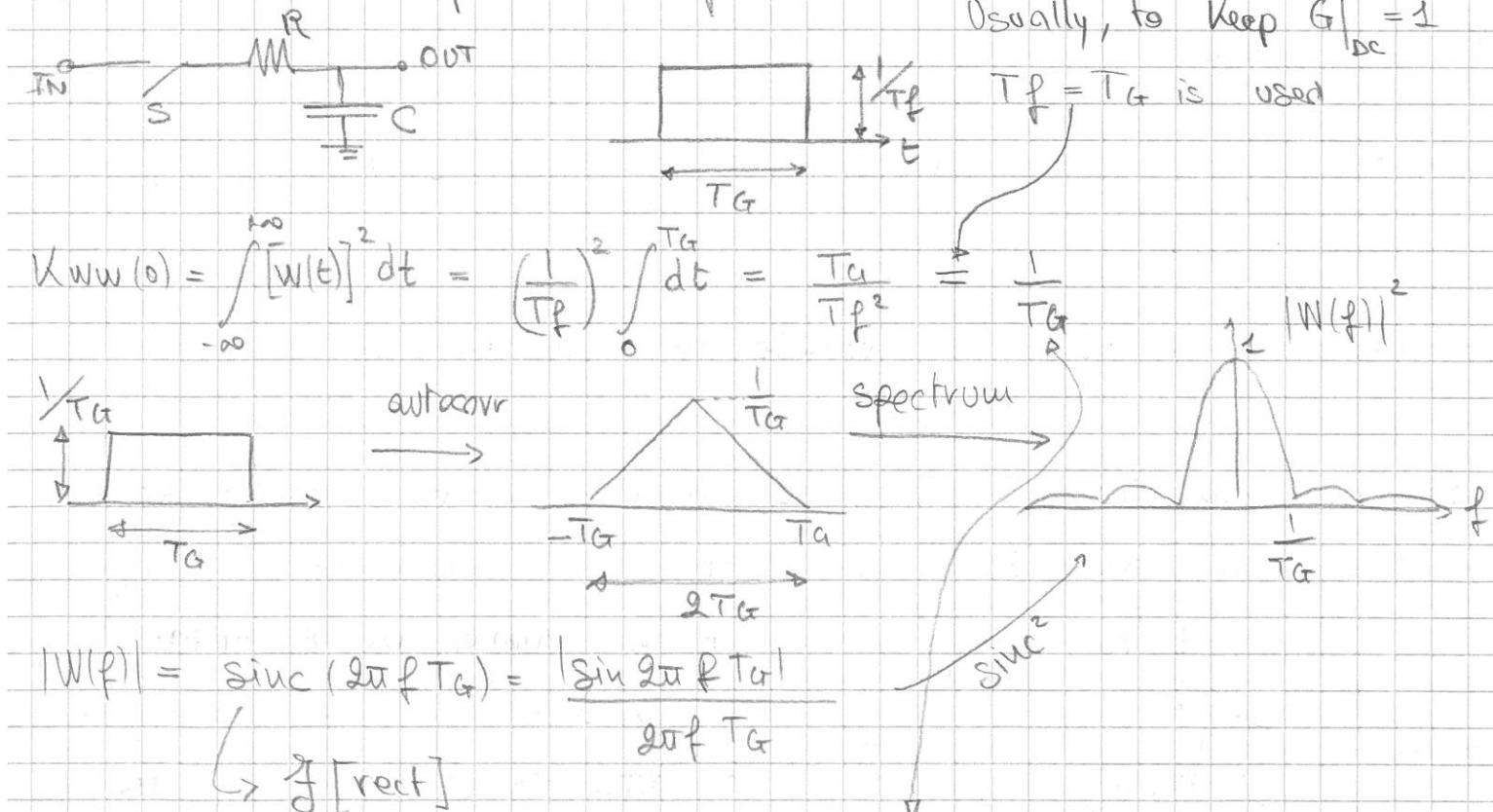
speed of light

Do not go above $10/20\text{ns} = T_C$ for MMF



When T_G has to be higher than $10/20\text{ns} \rightarrow$ GI filter

It's a switched parameter filter



Filtering noise: $\overline{\sigma_n^2} = \frac{\sum n_i}{2} \cdot K_{WW}(0) = \frac{\sum n_i}{2} \cdot \frac{1}{T_G} \rightarrow f_{eq}|_{GI} = \frac{1}{2T_G}$

Comparison between GI and RC

Taking into account noise filtering only:

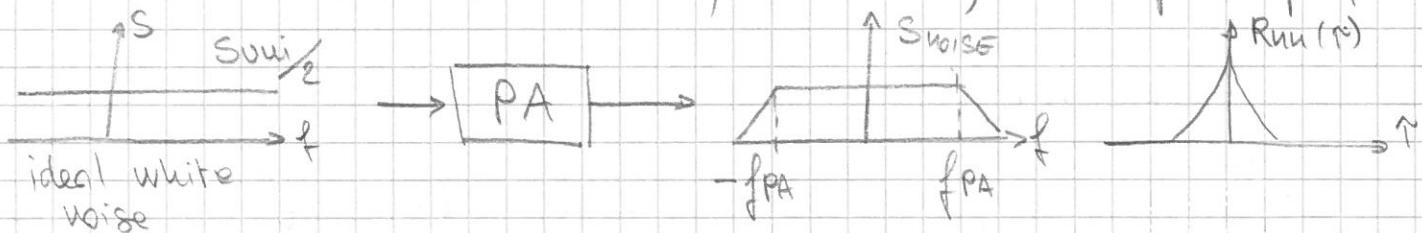
$$\sigma_{\text{RC}}^2 = \frac{S_{\text{out}}}{2} \cdot \frac{1}{2T_f} \quad \sigma_{\text{GI}}^2 = \frac{S_{\text{out}}}{2} \cdot \frac{1}{T_G} \quad \Rightarrow \quad T_{\text{RC}}|_{\text{eq}} = \frac{T_G}{2}$$

So the equivalent bandwidth of a GI seen from a RC-LPF perspective would be half of TG

Note on noise filtering

We assumed for the sake of simplicity that $R_{\text{PA}}(\tau) = \frac{S_{\text{out}}}{2} S(\tau)$

In reality deltas do not exist \rightarrow we have the shortest autocorrelation time that usually is set by the preamplifier.



$R_{\text{PA}}(\tau)$ resembles a delta but it really is an exponential.

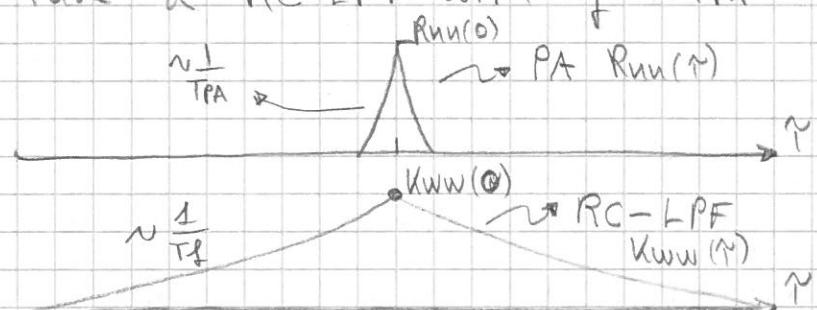
On the other hand, if we take a RC-LPF with $T_f = 10T_{\text{PA}}$

It will be:

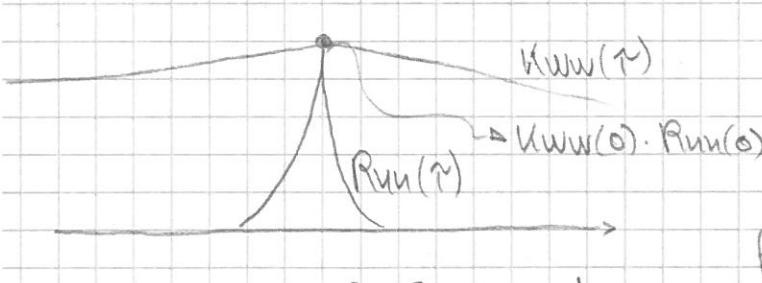
$$\sigma_n^2 = \int_{-\infty}^{+\infty} R_{\text{PA}}(\tau) K_{\text{WW}}(\tau) d\tau$$

$\xrightarrow{\text{RC-LPF}}$

$\xrightarrow{\text{exp from PA}}$



Suppose $T_f = 10T_{\text{PA}}$ then PA autocorrelation time ($n \frac{1}{T_{\text{PA}}}$) would narrow the integral from $-\infty$ to $+\infty$ to an interval that considers the very short autocorrelation of PA with respect to the autocorrelation of RC-LPF. That said, if we zoom:

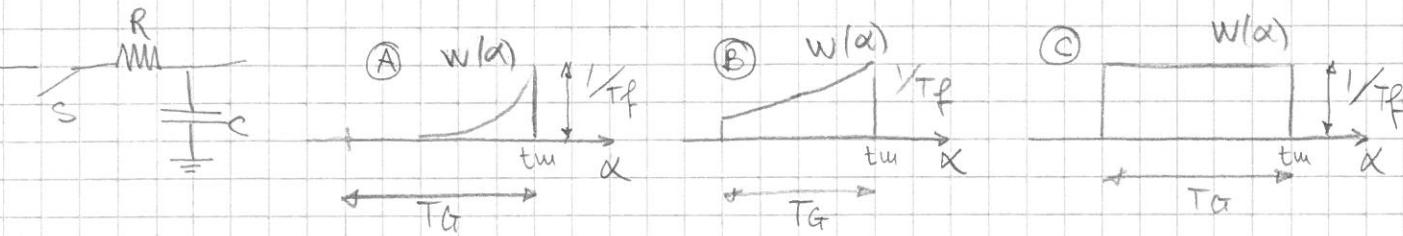


We can approximate (by accepting a small error) the $R_{\text{PA}}(\tau)$ to a delta \rightarrow integral narrows

from $\pm \infty$ to just zero evaluation:

$$\text{Therefore } \sigma_n^2 = \frac{S_{\text{out}}}{2} \cdot \frac{1}{2T_f}$$

OTHER SWITCHED PARAMETER FILTERS



(A) Sample & Hold: used to sample signals before an ADC by $T_G \gg T_f$ opening/closing the switch at acquisition.

$$W(t) \xrightarrow{T_G} \Delta W_{\text{NN}}(0) = \int_0^{T_G} [W(t)] dt \approx \int_0^{T_G} \left(\frac{1}{T_f} e^{\frac{t}{T_f}} \right)^2 dt = \frac{1}{2T_f}$$

Noise filtering is the same of a RC-LPF (T_G has no effect). Usually T_f is short so noise reduction is poor. It's never used to limit noise.

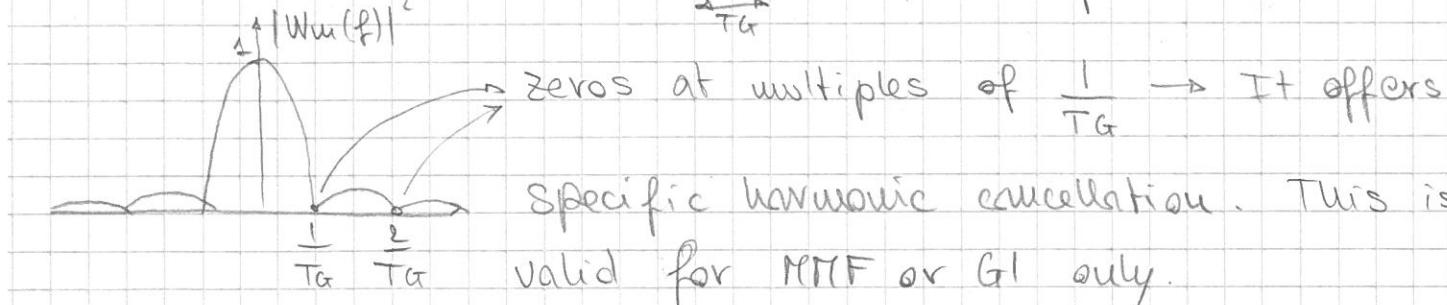
(B) Switched RC: mild filtering. $T_G \sim T_f$.

$$\Delta W_{\text{NN}}(0) = \int_0^{T_G} \left(\frac{1}{T_f} \right)^2 e^{-\frac{2t}{T_f}} dt = \frac{1}{2T_f} \left(1 - e^{-\frac{2T_G}{T_f}} \right)$$

Mild noise filtering. It is barely used because of the ease of calculations that a GI can give

(C) Gated Integrator: $T_f \gg T_G$ now noise filtering can be heavy if $\frac{T_G}{T_f^2}$ is a high number.

Note on GI: consider $\frac{1}{T_G} \frac{1}{1+T_f t}$, its spectrum will have



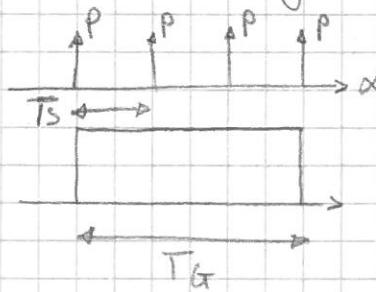
A RC-LPF can not exclude specific harmonics.

e.g.: signal has a 50Hz/100Hz/200Hz/... distort/binterference that needs to be cancelled out $\rightarrow \frac{1}{T_G} = 50\text{Hz} \rightarrow T_G = 20\text{ms}$

In this way it is possible to cancel out the disturb.

Digital filters

Discrete integrator



Note: if noise was ideal white \rightarrow no noise

Approx: take into account that noise is wideband white noise limited by the preamp (autocorrelation time of the noise is T_h).

Consider sampling frequency $f_s = \frac{1}{T_s}$. Hypothesis: $2T_h \ll T_s$ to keep uncorrelated the samples.

Input sequences: s_x is the input signal series while n_x is the input noise one.

Equivalent analog $G_1 \rightarrow T_G = NT_s$ DC gain = $N \cdot P$

It is called discrete time averager if $P = \frac{1}{N}$.

Sampling signal: $s_y = \sum_{k=1}^N P \cdot s_x = N \cdot P \cdot s_x$

Sampling noise $n_y = \sum_{k=1}^N P \cdot n_{xk} \rightarrow \overline{n_y^2} = P^2 (\overline{n_{x1}^2} + \overline{n_{x2}^2} + \overline{n_{x1} n_{x2}} + \dots)$

But if $2T_h \ll T_s$ (sampling freq is much lower than autocorrelation time) then all the samples are uncorrelated:

$\overline{n_{x1}} = \overline{n_{x2}} = \overline{n_x}$ because noise is stationary

$\overline{n_{x1} n_{x2}} = 0 \rightarrow$ uncorrelated noise samples

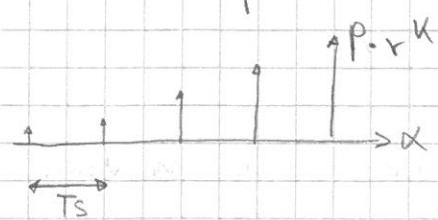
Therefore $\overline{n_y^2} = N \cdot P^2 \cdot \overline{n_x^2}$

$$\text{SNR} = \frac{s_y}{\sqrt{\overline{n_y^2}}} = \frac{N \cdot P \cdot s_x}{\sqrt{N P^2 \cdot \overline{n_x^2}}} = \frac{s_x}{\overline{n_x}} \cdot \sqrt{N} = \text{SNR}_{\text{IN}} \cdot \sqrt{N}$$

SNR_{IN} = ratio computed before the DI, so it will take into account the preamplifier filtering only. e.g.:

$$\text{SNR}_{\text{IN}} = \frac{V_p}{\sqrt{\frac{S_{\text{out}} \cdot 1}{2 \cdot 2T_{\text{PA}}}}} \approx \text{voltage amplitude of a DC signal}$$

Discrete exponential averager



Weight decreases with sample number

$$w_k = p \cdot r^k \text{ where } (1-r) \ll 1$$

\Rightarrow preamp autocorrelation time

As always $2T_u \ll T_s \rightsquigarrow$ uncorrelated noise samples

$$\text{Signal } S_y = S_x \cdot p \cdot \sum_{k=0}^{\infty} r^k = S_x \cdot \frac{p}{1-r} \quad \text{DC gain} = \frac{p}{1-r}$$

Integrated signal is the sum of the samples but averaged

$$\text{Noise } \bar{n}_y^2 = \bar{n}_x^2 p^2 \sum_{k=0}^{\infty} r^{2k} = \bar{n}_x^2 \frac{p^2}{1-r^2} \rightsquigarrow \text{Same reasoning of DI}$$

$$\left. \text{SNR} \right|_{\text{out}} = \left. \text{SNR} \right|_{\text{in}} \frac{\frac{1}{1-r}}{\sqrt{\frac{1}{1-r^2}}} = \sqrt{\frac{1+r}{1-r}} \rightsquigarrow 1-r \ll 1 \\ 1+r \approx 2$$

$$\text{SNR}_{\text{out}} = \left. \text{SNR} \right|_{\text{in}} \cdot \sqrt{\frac{2}{1-r}}$$

If we take $r = e^{-T_s/T_e}$ then $1-r \propto \frac{T_s}{T_e}$ therefore

$$\left. \text{SNR} \right|_{\text{out}} = \left. \text{SNR} \right|_{\text{in}} \cdot \sqrt{\frac{2 T_s}{T_e}}$$

Removing uncorrected noise samples hypothesis

$2T_n \ll T_s$ can also be not true \rightarrow this means that, for a DI, the \sqrt{N} improvement factor will change if N is noise samples are correlated ($\rightarrow T_s \sim T_n$)

Frequency demonstration

DI is a sampled GI \rightarrow Sampling in time \Leftrightarrow replica in spectrum

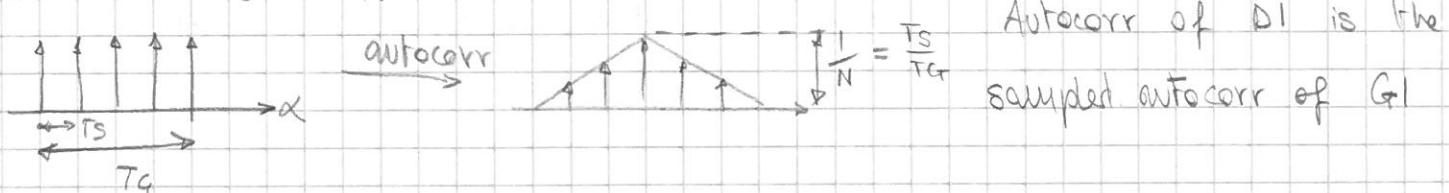


If $f_s \uparrow \rightarrow$ replicas are more spaced between each other

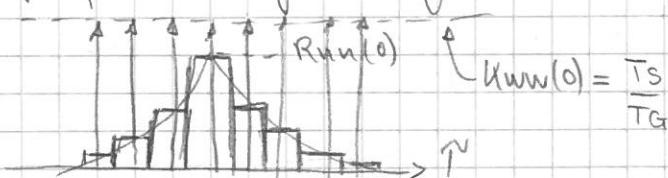
The more $f_s \uparrow$; the more replicas will fall out of noise band limit f_{PA} . Therefore we can conclude that for $f_s \rightarrow \infty$ only the center replica survives \rightarrow it's the analog GI

Conclusion: $SNR \sim \sqrt{N}$ for $2T_n \ll T_s$, for larger T_s the improvement factor \sqrt{N} starts to change, with an upper boundary set at its analog counterpart \rightarrow GI is the limit.

Time demonstration



If f_s is high enough, more deltas will cut Run(τ):



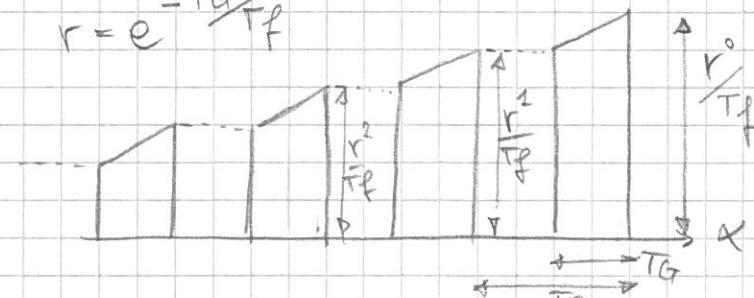
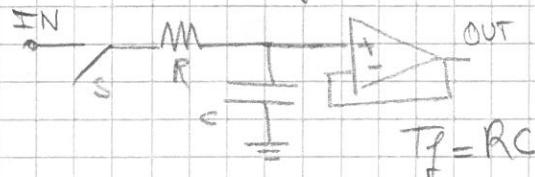
$$\sigma_n^2 = \frac{1}{T_G} (\text{area of the scaloid}) \approx \text{like in the analog domain}$$

It is possible to demonstrate that area of the scaloid is always higher than the area of $\text{Run}(\tau)$ \rightarrow noise will be higher than GI for $f_s \rightarrow \infty$. Exactly the analog GI is obtained

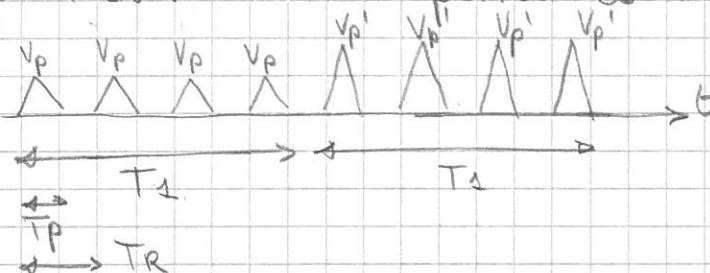
Exploit repetition rates

$$r = e^{-T_G/T_f}$$

Boxcar integrator



Boxcar is: G1 + exponential averager. A sync signal just integrates the time replicas of a pulse (e.g. train of pulses coming from a modulated laser). The replicas need to have a fairly stable peak over a time period so that repetition can be exploited. E.g.:



Input signal (duration T_p) is repeated each T_R and peak does not change for a time T_1 .

Exponential transient has to be lower than T_1 , so

$$\underbrace{5T_f}_{\text{transient}} \Big|_{B1} = \underbrace{\frac{T_1}{T_R} \cdot T_p}_{L} \rightarrow \text{duration of integration for each replica}$$

\rightarrow Number of replicas inside the wanted time period (5↑)

Note: T_R could be a statistical variable \rightarrow it can change over time.

Since $B1$ maintains its charge when switch is open, T_R can change (for example jitter or different arrival time) without issues.

This is because DC gain does not depend on frequency

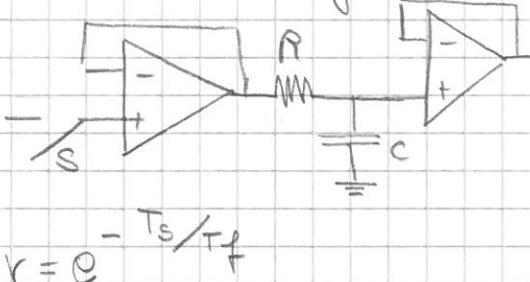
$$\text{DC gain} = \int_{-\infty}^{+\infty} w_{B1}(t) dt = 1 \rightarrow \text{it's the same of a RC-LPF}$$

\rightarrow G1 integration $T_G \ll T_f \rightarrow r \approx 1$

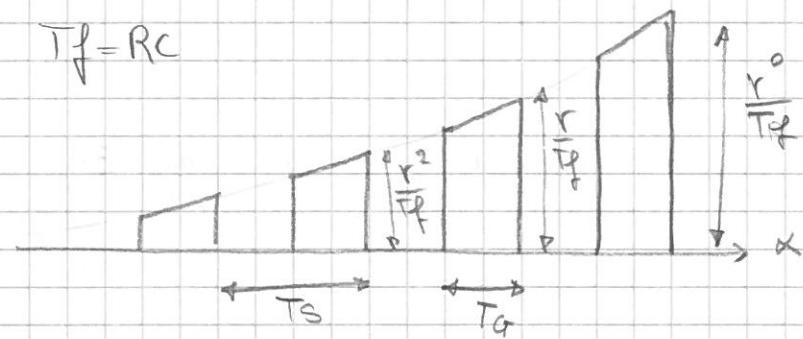
$$\left. \text{SNR} \right|_{\text{OUT}} = \left. \text{SNR} \right|_{\text{G1}} \frac{\sum_k r^k}{\sqrt{\sum k r^{2k}}} = \left. \text{SNR} \right|_{\text{G1}} \frac{\frac{1}{1-r}}{\sqrt{\frac{1}{1-r^2}}} = \left. \text{SNR} \right|_{\text{G1}} \sqrt{\frac{1+r}{1-r}}$$

$$\approx \left. \text{SNR} \right|_{\text{G1}} \sqrt{\frac{2T_f}{T_G}}$$

Ratemeter Integrator



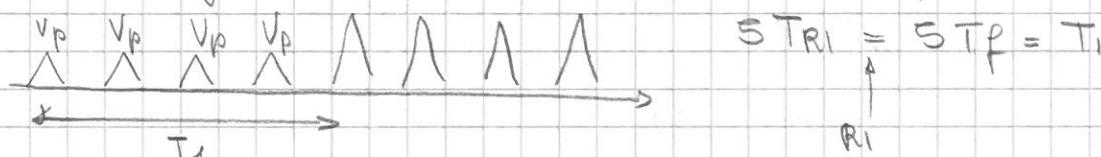
$$T_f = RC$$



→ Note: r depends on repetition T_s !!!

There is no hold state → capacitor (unlike BI) can now discharge.
We still have GI + exponential averaging.

Considering the same situation as before:



$$5TR_I = 5T_f = T_1$$

$$\text{SNR}_{\text{out}} = \text{SNR}_{\text{BI}} \cdot \frac{\sum \text{--}}{\sum \text{--}} = \text{SNR}_{\text{BI}} \cdot \sqrt{\frac{2T_f}{T_s}} = \text{SNR}_{\text{BI}} \sqrt{2T_f \cdot f_s}$$

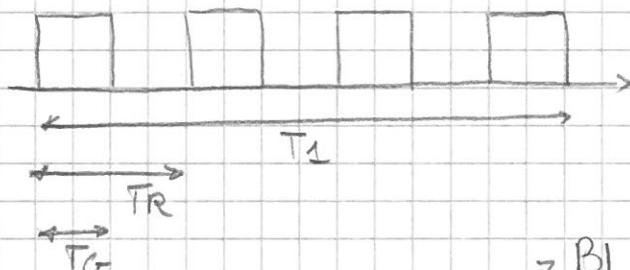
Where f_s = repetition frequency

This means that now SNR depends on frequency → if TR is affected by jitter SNR will change as well.

This can be a drawback compared with the BI (therefore RI can not be used in these kind of situations).

Ratemeter becomes useful when a DC signal is fed and switch is closed with f_s : a frequency-to-Voltage is obtained.

Note: if TR is fixed and T_g is the same, using a BI and/OR a RI will lead to the same improvement factor. Proof:



$$T_f_{\text{BI}} = \frac{T_1 T_g}{5TR}$$

$$T_f_{\text{RI}} = \frac{T_1}{S}$$

They are equivalent
unless TR has a σ
in time

$$\text{BI} = \sqrt{\frac{2T_f}{T_g}} = \sqrt{\frac{2T_1 T_g}{5TR}} = \sqrt{\frac{2T_1}{5TR}}$$

Improvement factor

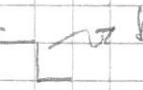
$$RI = \sqrt{\frac{2T_f}{TR}} = \sqrt{\frac{2T_1}{S} \cdot \frac{1}{TR}} = \sqrt{\frac{2T_1}{STR}}$$

$$\text{Ratemeter DC gain} = |W_{RI}(0)| = \int_{-\infty}^{+\infty} W_{RI}(t) dt \approx \frac{T_G}{T_R} = T_G \cdot f_p$$

As we can see, DC gain (unlike BI) now depends on repetition rate.

OPTIMUM FILTERING

NOTE: OF theory works for white noise only \rightarrow if noise spectrum is not white (e.g.: SI going into an integrator), a whitening filter will be needed.

Example signal:  $\sim V_p \cdot \frac{1}{T_g}$ 

Separate peak from signal shape $b(t)$

$$\text{Signal} = V_p \int_{-\infty}^{+\infty} b(\tau) w(\tau) d\tau = K_{bw}(0) \cdot V_p$$

$$K_{bw}(0) = \frac{\text{Sum}}{2} S(0)$$

$$\text{Noise} = \int_{-\infty}^{+\infty} R_{ww}(\tau) K_{ww}(0) d\tau = \frac{\text{Sum}}{2} K_{ww}(0)$$

$$\text{SNR} = \frac{V_p \cdot K_{bw}(0)}{\sqrt{\frac{\text{Sum}}{2} K_{ww}(0)}} = \text{maximum (because of OF theory) if } K_{bw}(0) = K_{bb}(0) \text{ and } K_{ww}(0) = K_{bb}(0)$$

filter has same signal shape

$$\text{SNR}_{OF} = \frac{V_p K_{bb}(0)}{\sqrt{\frac{\text{Sum}}{2} K_{bb}(0)}} = \frac{V_p}{\sqrt{\frac{\text{Sum}}{2}}} \sqrt{K_{bb}(0)}$$

Example:

$$\begin{array}{c} V_p \\ \hline \xleftarrow[T_G]{} \end{array} \rightarrow \begin{array}{c} b(t) \\ \hline \xleftarrow[T_G]{} \end{array} \quad K_{bb}(0) = T_G$$

$$\text{SNR}_{OF} = \frac{V_p}{\sqrt{\frac{\text{Sum}}{2}}} \sqrt{T_G}$$

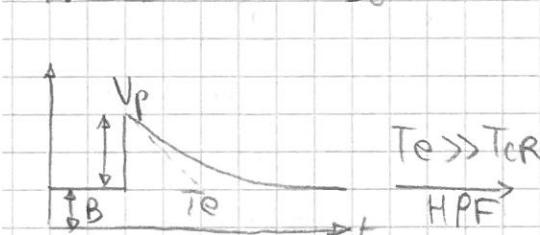
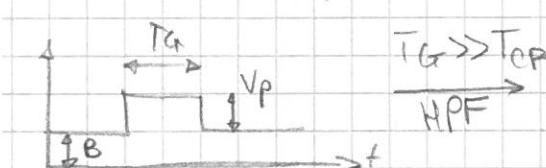
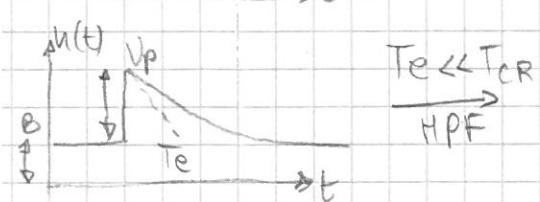
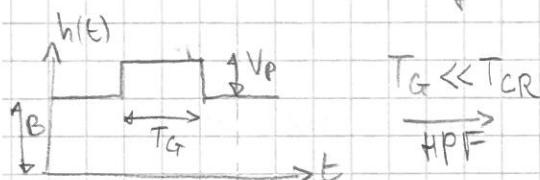
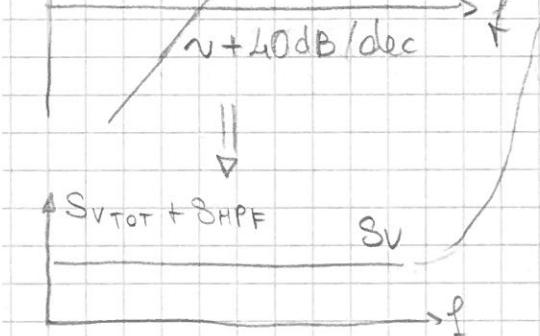
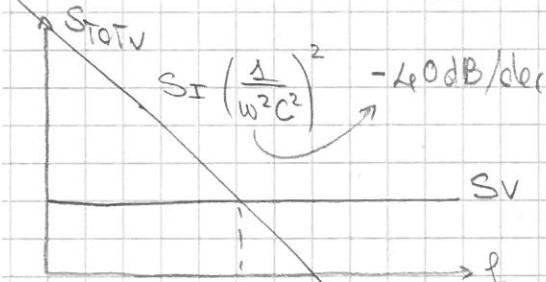
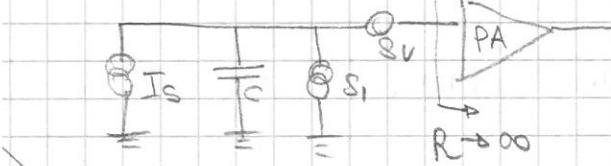
It also works for whatever $b(t)$ amplitude

$$\begin{array}{c} V_p \\ \hline \xleftarrow[T_G]{} \end{array} \rightarrow \begin{array}{c} b(t) \\ \hline \xleftarrow[T_G]{} \end{array} \quad K_{bb}(0) = \frac{1}{T_G}$$

$$\text{SNR}_{OF} = \frac{(V_p \cdot T_G)}{\sqrt{\frac{\text{Sum}}{2}}} \sqrt{\frac{1}{T_G}} = \frac{V_p \sqrt{T_G}}{\sqrt{\frac{\text{Sum}}{2}}}$$

$$\begin{array}{c} V_p \\ \hline \xleftarrow[T_G]{} \end{array} \rightarrow V_p \cdot \frac{1}{T_G} \cdot \begin{array}{c} b(t) \\ \hline \xleftarrow[T_G]{} \end{array}$$

Whitening filter - example



S_V directly transfers through the out.

SI sees the capacitor as load

(no resistance), so spectrum will be

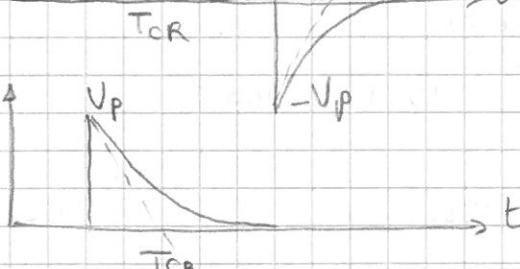
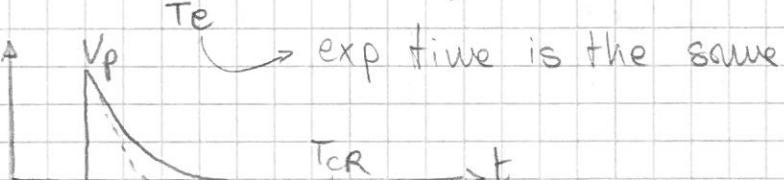
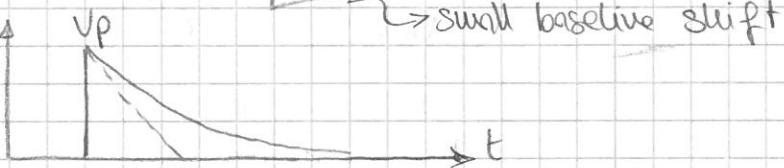
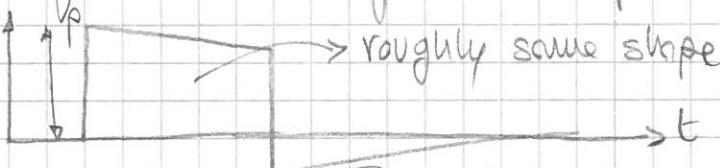
SI gets integrated through $\left(\frac{1}{sc}\right)^2$

Spectrum is no more white.

There's the need to place a HPF

With $f_{WHT} = \frac{1}{T_{pole}}$ so that spectrum returns white.

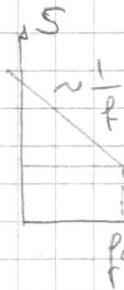
Note: HPF filtering means that signal can be affected by the low frequency cut. If I_S has a baseline, that will be removed. If I_S is slow enough, it will be changed. Examples:



→ HPF pole is faster (dominant)

exp. time is therefore changed

Flicker noise



It's defined through the slope ($\frac{dS}{df}$) and the corner frequency with white noise.

If there is no upper/lower bound, integrated noise will be ∞ . Consider:

- RC-LPF higher limit sets $f_{HI} = \frac{1}{2\pi T_{fRC}}$

- CR-HPF lower limit sets $f_{LO} = \frac{1}{2\pi T_{fCR}}$

If CR-HPF freq is low enough, then $\sigma_w = \sqrt{\sum_i \frac{\pi}{2} f_{HI}}$
While integrated $1/f$ noise will be $\sigma_{1/f} = \sqrt{\sum_i f_c \cdot \ln(f_{HI}/f_{LO})}$

Total noise will be $\sigma_{TOT} = \sqrt{\sigma_w^2 + \sigma_{1/f}^2}$

Goal: reduce $1/f$ noise by setting f_{HI}/f_{LO} in order to have negligible flicker noise when summed quadratically with σ_w

CDS $\delta(t)$ We could do a zero setting by double sampling:
 $\xrightarrow{T} \xrightarrow{\Delta T} t$ ① uncorrelated fast samples will be sampled two times
 $\downarrow \delta(t + \Delta t)$ (delta sign does not matter in terms of noise)

② corrected (very slow) samples will be "added and then subtracted" by the two deltas \rightarrow High pass action of $1/f$ and low frequency signals.

By doing computations, HPF cut is set to $f_{LO} = \frac{1}{2\pi \Delta T}$

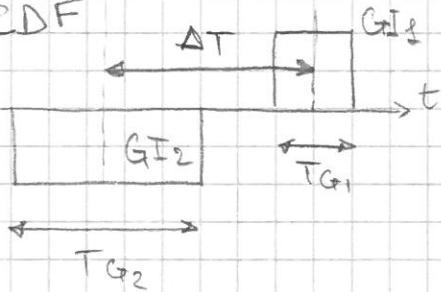
While no high limit is imposed because of deltas \rightarrow only the preamp limits noise on the higher side.

Because of double acquisition of all frequencies (δ) at ①, noise in CDS is doubled:

$\sigma_w = \sqrt{2 \sum_i \frac{\pi}{2} f_{PA}} \sim f_{PA}$ should be considered as low limit but it is
 \hookrightarrow noise doubling usually negligible

$$\sigma_{1/f} = \sqrt{2 \sum_i f_c \ln(f_{PA}/f_{LO})}$$

CDF



To solve noise doubling caused by deltas, it is possible to use a double GI system. ΔT is defined as the difference between middle positions of each GI.

(Be careful when selecting GI, see the example later).

Noise will still be integrated two times, but now there's freedom to reduce noise introduced by zero setting.

GI_1 : filter that integrates useful signal \rightarrow set this like there is no CDF action going on \rightarrow specs set by SNR and signal

GI_2 : filter used for HP action on $1/f$ and low freq signals:

Characteristic frequencies:

$$f_{GI_1} = \frac{1}{2TG_1}$$

$$f_{GI_2} = \frac{1}{2TG_2}$$

$$f_{lo} = \frac{1}{2\pi\Delta T}$$

$$\bar{\sigma}_w = \sqrt{\sum_i [f_{GI_1} - f_{lo}] + [f_{GI_2} - f_{lo}]} \approx \sqrt{\sum_i f_{GI_1}}$$

$f_{lo} \ll f_{GI_1}, f_{GI_2}$ if $f_{GI_1} \gg f_{GI_2}$

Always check for simpler calculations

$$\bar{\sigma}_{1/f} = \sqrt{\frac{\sum_i}{2} \left[\text{fc} \left[\text{lu} \left(\frac{f_{GI_1}}{f_{lo}} \right) + \text{lu} \left(\frac{f_{GI_2}}{f_{lo}} \right) \right] \right]} \approx \sqrt{\frac{\sum_i}{2} \text{fc} \text{lu} \left(\frac{f_{GI_1}}{f_{lo}} \right)}$$

Note: ideally $TG_2 \rightarrow \infty$ to have negligible noise. On the other hand if $TG_2 \rightarrow \infty$, $\Delta T \propto \frac{TG_1}{2} + \frac{TG_2}{2} \rightarrow \infty \Rightarrow f_{lo} \ll 1$

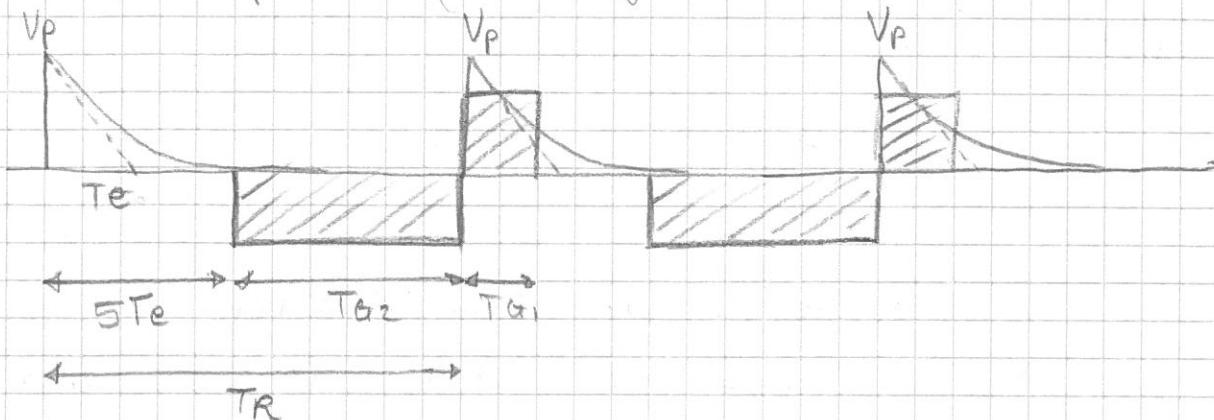
If f_{lo} is too low \rightarrow high pass action on $1/f$ would be too scarce. This would lead to good σ_w but $\sigma_{1/f}$ would increase too much. Remember Hint:

$$f_{hi} = 100\text{MHz} \quad f_{lo} = 1\text{Hz} \quad \text{lu} \left(\frac{100\text{MHz}}{1\text{Hz}} \right) = 18,4$$

- If f_{hi} halves: $100\text{MHz} \rightarrow 50\text{MHz}$ (50MHz loss) $\text{lu} \left(\frac{50\text{MHz}}{1\text{Hz}} \right) = 17,7$
- If f_{lo} doubles: $1\text{Hz} \rightarrow 2\text{Hz} \rightarrow 1\text{Hz}$ more leads to same result

Conclusion: controlling lower freq limit is way more important than lowering higher limit \rightarrow ratio is the thing that matters (14)

Example of a CDF setting



T_{G1} : set by single pulse integration to max SNR out.

for example, for an exp signal $T_{G1} = 1,25 T_e$

T_R : repetition period of a periodic signal

T_e : exponential time constant

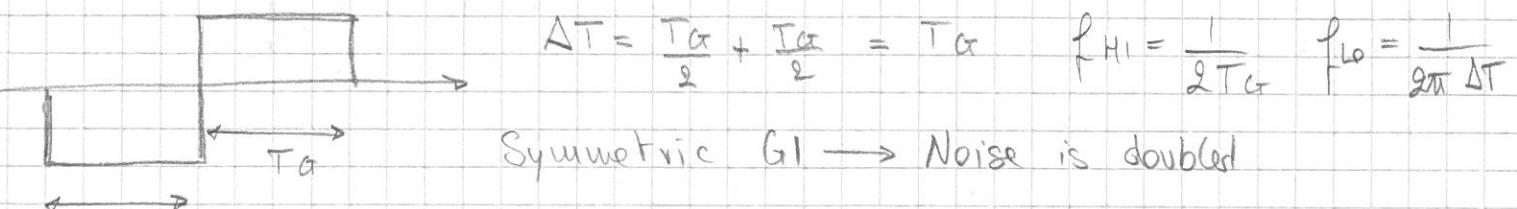
$5T_e$: exponential transient has passed

To have good zero settling signal has to be constant, so

$$T_{G2} = T_R - 5T_e$$

T_{G1} is already set by other specs to fulfill SNR requirements

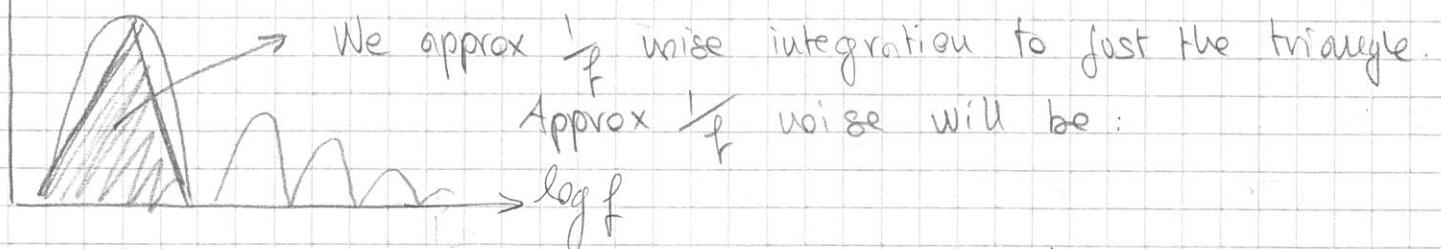
Particular case: symmetric CDF



Symmetric G1 \rightarrow Noise is doubled

$$\sigma_W = \sqrt{S_{\text{sum}} [(f_{H1} - f_{lo}) + (f_{H1} - f_{lo})]} = \sqrt{2 S_{\text{sum}} (f_{H1} - f_{lo})}$$

$$|W_{DF}|^2$$



$$\frac{\sigma_1}{f} \approx \sqrt{2,11 S_{\text{sum}} f_c \ln \left(\frac{\frac{1}{2\pi\Delta T}}{\frac{1}{2T_G}} \right)} = \sqrt{--- \ln \left(\frac{\frac{1}{2\pi T_G}}{\frac{1}{2T_G}} \right)} = \sqrt{2,11 S_{\text{sum}} f_c \ln(\pi)}$$

This symm. CDF is useful when there is not much room left for G12

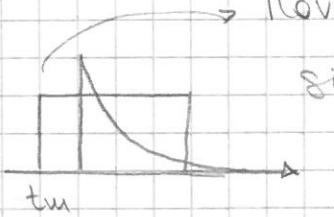
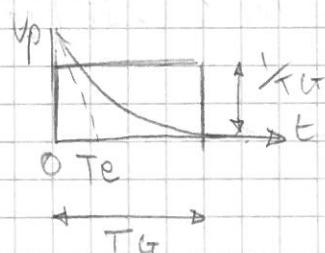
BLR - Base Line Restorer

It is very specific, it will not be discussed here

(check theory as always ☺)

FILTERING SIGNALS

- Exp × GI filter



→ Moving tm before does not integrate signal → It does not make sense since it just integrates noise

$$\text{Sig} = \int_{-\infty}^{+\infty} s(\tau) w(\tau) d\tau = \int_0^{T_g} \frac{V_p}{T_g} \cdot e^{-\frac{\tau}{T_g}} d\tau = \frac{V_p T_g}{T_g} \left[e^{-\frac{\tau}{T_g}} \right]_0^{T_g} = \frac{V_p T_g}{T_g} \left(1 - e^{-\frac{T_g}{T_g}} \right) = \frac{V_p T_g}{T_g} (1 - e^{-1}) = \frac{V_p T_g}{T_g} \cdot 0,638$$

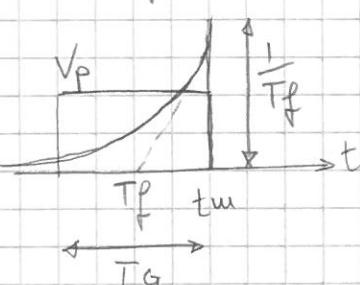
$$\text{noise} = \int_{-\infty}^{+\infty} R_{WW}(\tau) K_{WW}(\tau) d\tau = \frac{S_{WW}}{2} \cdot \frac{1}{T_g}$$

$$\text{SNR} = \frac{\frac{V_p T_g}{T_g} \cdot (1 - e^{-1})}{\sqrt{\frac{S_{WW}}{2} \cdot \frac{1}{T_g}}} \cdot \sqrt{\frac{V_p}{T_g}} = \frac{\frac{V_p}{T_g} \cdot \sqrt{T_g}}{\sqrt{\frac{S_{WW}}{2}}} \cdot \sqrt{1 - e^{-1}} \quad T_g = 1,25 T_e$$

$$\text{SNR} \approx \frac{V_p}{\sqrt{\frac{S_{WW}}{2}}} \cdot \sqrt{T_g} \cdot 0,638 \quad T_g = 1,25 T_e$$



- Rect pulse × RC-LPF

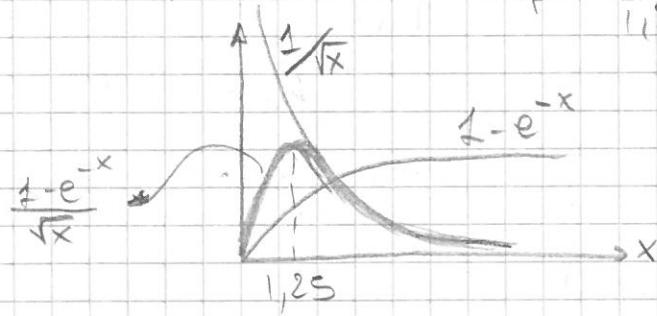


$$\text{sig} = \int_0^{T_f} \frac{V_p}{T_f} \cdot e^{-\frac{\tau}{T_f}} d\tau = V_p \left(1 - e^{-\frac{T_f}{T_f}} \right) = V_p (1 - e^{-1}) = V_p \cdot 0,638$$

$$\text{noise} = \sigma_n^2 = \frac{S_{WW}}{2} \cdot \frac{1}{2 T_f}$$

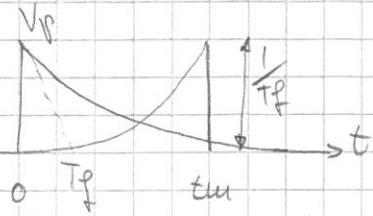
$$\text{SNR} = \frac{V_p}{\sqrt{\frac{S_{WW}}{2}}} \cdot \frac{1 - e^{-1}}{\sqrt{\frac{1}{T_f}}} \cdot \sqrt{\frac{V_p}{T_f}} = \frac{V_p}{\sqrt{\frac{S_{WW}}{2}}} \cdot \sqrt{T_g} \cdot \frac{1 - e^{-1}}{\sqrt{x}} \quad x = 1,25 \text{ so } T_f = \frac{T_g}{1,25} = 0,8 T_g$$

$$\text{SNR}_{\text{max}} \approx \frac{V_p}{\sqrt{\frac{S_{WW}}{2}}} \cdot \sqrt{T_g} \cdot 0,638$$



• Exp pulse \times RC-LPF

For simplicity, both exp time constant are the same:



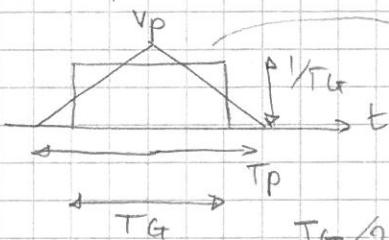
$$\text{Signal} = \int_0^{t_m} V_p e^{-\frac{\tau}{T_f}} dt = \frac{V_p}{T_f} \cdot t_m e^{-\frac{t_m}{T_f}}$$

$$\text{Noise} = \sigma_n^2 = \frac{S_{\text{uni}}}{2} \cdot \frac{1}{2T_f} = \sqrt{S_{\text{uni}} \frac{\pi}{2} f_{\text{pole}}}$$

$$\text{SNR} = \frac{(V_p/T_f) \cdot t_m e^{-\frac{t_m}{T_f}}}{\sqrt{S_{\text{uni}} \frac{\pi}{2} f_{\text{pole}}}} \quad \text{SNR is max for } t_m = T_f$$

• Tri pulse \times GI

It makes sense to choose t_m so that GI is centered on the peak (to max SNR). Now we need to find T_G that maxes SNR.



Take half of the triangle from 0 to T_G

$$\text{Signal} = \frac{1}{2} \int_0^{T_G} \frac{1}{T_p} V_p \left(1 - \frac{2\tau}{T_p}\right) d\tau = \frac{2V_p}{T_p} \left[1 - \frac{2\tau^2}{T_p^2}\right]_0^{T_G/2} = V_p \left(1 - \frac{T_G}{2T_p}\right)$$

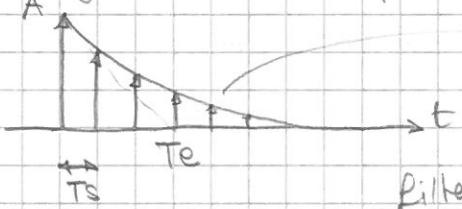
$$\text{noise} = \sigma_n^2 = \sqrt{\frac{S_{\text{uni}}}{2} \cdot \frac{1}{T_G}}$$

Max for $T_G = \frac{2}{3} T_p$

$$\text{SNR} = \frac{V_p \left(1 - \frac{T_G}{2T_p}\right)}{\sqrt{\frac{S_{\text{uni}}}{2} \cdot \frac{1}{T_G}}} = \frac{V_p}{\sqrt{\frac{S_{\text{uni}}}{2}}} \sqrt{T_G \left(1 - \frac{T_G}{2T_p}\right)}$$

$$\text{SNR}_{\text{MAX}} = \frac{V_p}{\sqrt{\frac{S_{\text{uni}}}{2}}} \cdot \sqrt{\frac{2}{3} T_p} \cdot \frac{2}{3}$$

• Digital OF of exponential signal



→ K sample has r^K weight where $r = e^{-Ts/Te}$

filter has same shape of the signal → autocorr of digital filter

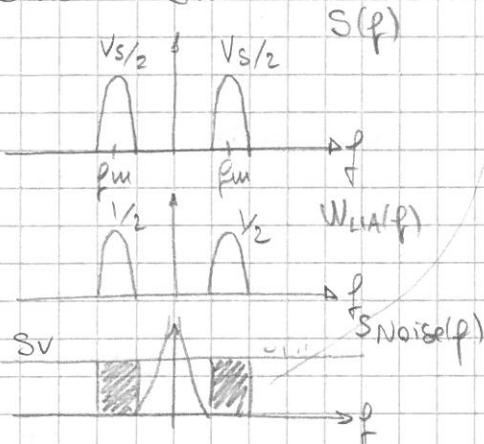
$$\text{Signal} = A \sum_{n=0}^{\infty} r^n = \frac{A}{1-r^2} \quad \text{noise} = \sigma_n^2 = \sigma_{PA}^2 \sum_{n=0}^{2\infty} r^{2n} = \frac{\sigma_{PA}^2}{1-r^2}$$

$$|\text{SNR}|_{\text{out}} = \frac{A}{\sigma_{PA}} \cdot \frac{\frac{1}{1-r^2}}{\sqrt{\frac{1}{1-r^2}}} = \frac{A}{\sigma_{PA}} \cdot \frac{1}{\sqrt{1-r^2}} \quad \text{if } 2Ts \ll Te \text{ then } 1-r^2 = 1-e^{-\frac{2Ts}{Te}} \approx \frac{2Ts}{Te}$$

$$|\text{SNR}|_{\text{out}} \approx \frac{A}{\sigma_{PA}} \sqrt{\frac{Te}{2Ts}} = |\text{SNR}|_{\text{input}} \sqrt{\frac{Te}{2Ts}}$$

LIA

- Sine \times Sine



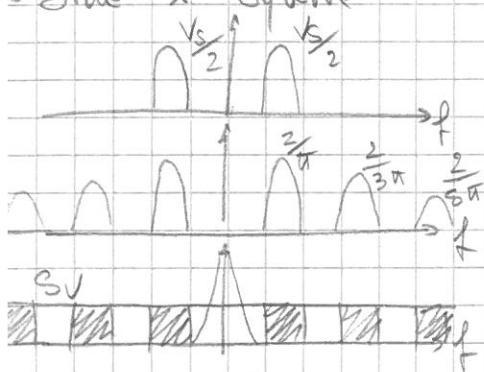
$$\text{Signal} = \frac{V_s}{2} \cdot 2 = V_s \quad A_{f_u} = 2 \cdot \frac{\pi}{2} f_{LPF}$$

$$\sigma_n = \sqrt{2 \cdot \frac{S_{TOT}}{2} \cdot A_{f_u}} = \sqrt{2} \sigma_{LPF}$$

$$\text{Where } \sigma_{LPF} = \sqrt{\frac{S_V}{2} \cdot 2 \cdot \frac{\pi}{2} f_{LPF}} = \sqrt{\frac{S_V \pi}{2} f_{LPF}}$$

$$\text{SNR} = \frac{V_s}{\sqrt{2} \sigma_{LPF}}$$

- Sine \times Square

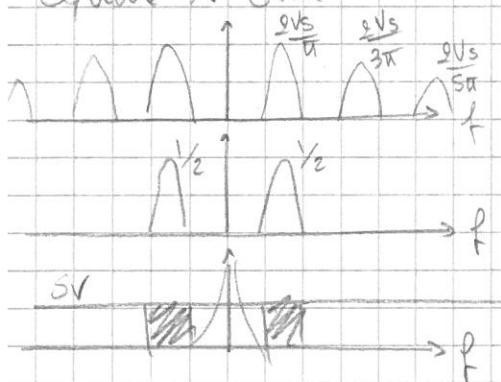


$$\text{Signal} = 2 \cdot \frac{V_s}{2} \cdot \frac{2}{\pi} = \frac{2}{\pi} V_s$$

$$\begin{aligned} \sigma_n &= \sqrt{2 \cdot \frac{S_V}{2} \cdot A_{f_u} \left(\frac{2}{\pi} \cdot \frac{2}{\pi} \cdot \frac{1}{2k+1} \right)^2} \\ &= \sqrt{2 \frac{S_V}{2} \cdot 2 \cdot \frac{\pi}{2} f_{LPF} \cdot \frac{4}{\pi^2} \cdot \frac{1}{8}} = \sqrt{S_V \frac{\pi}{2} f_{LPF}} \end{aligned}$$

$$\text{SNR} = \frac{2 V_s}{\pi \sigma_{LPF}}$$

- Square \times Sine

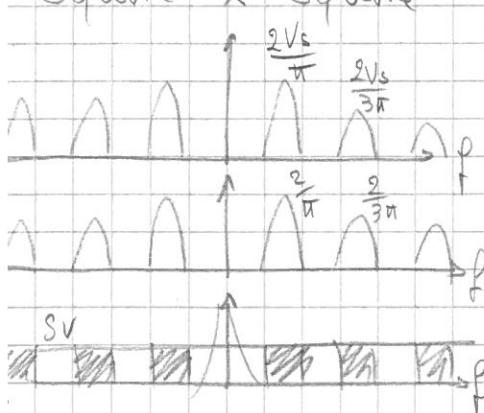


$$\text{Signal} = 2 \cdot \frac{2}{\pi} V_s \cdot \frac{1}{2} = \frac{2}{\pi} V_s$$

$$\sigma_n = \sqrt{2} \sigma_{LPF}$$

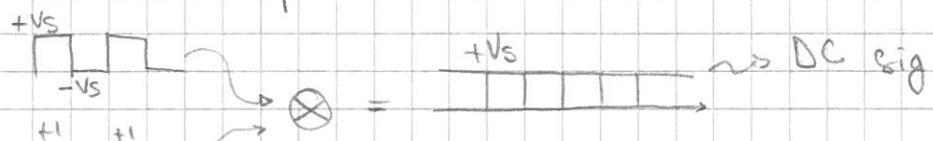
$$\text{SNR} = \frac{2}{\pi \sqrt{2}} \frac{V_s}{\sigma_{LPF}}$$

- Square \times Square



$$\text{Signal} = 2 V_s \left(\sum_{n=0}^{\infty} \frac{2}{\pi} \cdot \frac{1}{2k+1} \right)^2 = 2 V_s \cdot \frac{4}{\pi^2 / 8} = V_s$$

Intuitive explanation:



$$\sigma_n = \sigma_{LPF} \quad \text{SNR} = V_s / \sigma_{LPF}$$

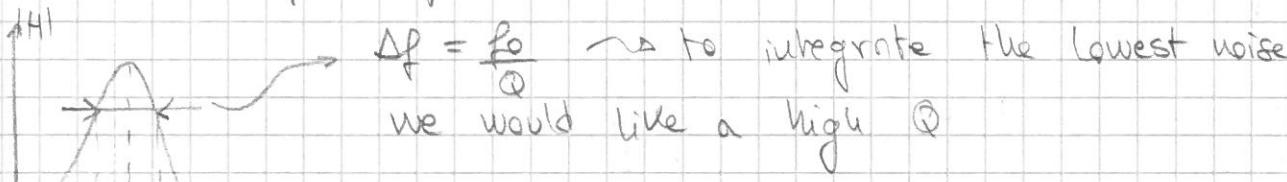
Note:

if signal

$$\begin{array}{c|cc|c} & V_s & V_s \\ \hline & 0 & 0 \\ +1 & +1 & +1 \\ -1 & -1 & -1 \end{array} \rightarrow \otimes = \begin{array}{c|cc|c} & +V_s & +V_s \\ \hline & 0 & 0 \end{array} \rightarrow \text{DC signal} = \frac{V_s}{2}$$

BPF - RLC filter

In case we don't want a LIA implementation we could use a RLC bandpass filter



$$\text{SNR} = \sqrt{\frac{S}{N}} \cdot \frac{\pi}{2} \frac{\Delta f}{f_0} = \sqrt{\frac{S}{N} \frac{\pi}{2} \frac{f_0}{Q}}$$

For $f_0 \gg 100\text{MHz} \rightarrow Q > 10$ (best technology $Q \approx 100$)

For $1\text{MHz} < f_0 < 100\text{MHz} \rightarrow Q \approx 10$

For $f_0 < 1\text{MHz} \rightarrow Q \approx 5$ max

Note: for a given Q , $\Delta f_{\text{noise}} = \frac{\pi}{2} \frac{f_0}{Q}$ is reduced as f_0 decreases

SECOND PART - Photo detectors

$P_p = n_p h\nu \rightarrow$ optical power n_p = photon generation rate $\frac{\text{photons}}{\text{second}}$

$I_D = n_e \cdot q \rightarrow$ photogenerated current n_e = electron gen. rate

$$\eta_D = \frac{n_e}{n_p} = \frac{N_e}{N_p} = \text{quantum detection efficiency} \quad N_e = \text{electrons generated} \\ N_p = \text{photons generated}$$

$$= \frac{\text{photogenerated } e^- \text{ or } e^-/\text{hole pair}}{\text{photons arrived to PD}}$$

$$S_D = \frac{I_D}{P_p} \text{ or } \frac{V_D}{P_L} = \frac{\text{output current/voltage}}{\text{input power on PD area}} = \text{Radiant sensitivity } \left[\frac{A}{W} \right], \left[\frac{V}{W} \right] \\ \rightarrow \text{wavelength expressed}$$

$$S_D = \frac{I_D}{P_p} = \frac{n_e q}{n_p h\nu} = \frac{n_e}{N_p} \cdot \frac{\lambda}{hc/q} = \eta_D \cdot \frac{\lambda [\mu\text{m}]}{1,24} \text{ in micrometers}$$

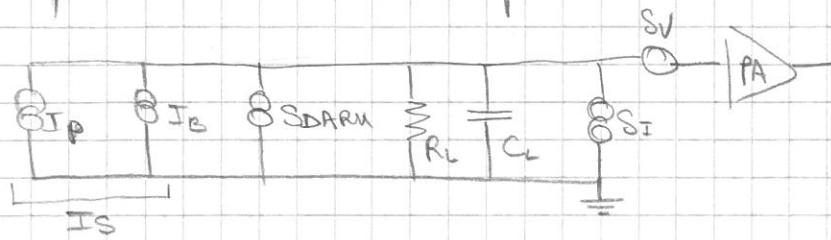
NEP = noise equivalent power (detector noise referred to input).

It takes into account just detector intrinsic noise

Note: to compute NEP SNR = 1 and noise band = 1Hz dark current

$$\text{SNR}_{\text{NEP}} = \frac{I_S}{\sqrt{2qI_B \Delta f}} = \frac{S_D P_p}{\sqrt{2qI_B}} = 1 \rightarrow \text{NEP} = P_p \Big|_{\min} = \frac{\sqrt{2qI_B}}{S_D}$$

Acquisition chain for Photo detectors



I_p = pulse current I_b = baseline current (ambient light, etc...)

$R_L C_L$ = sensor load where $C_L = C_{PA} + C_{sensor}$

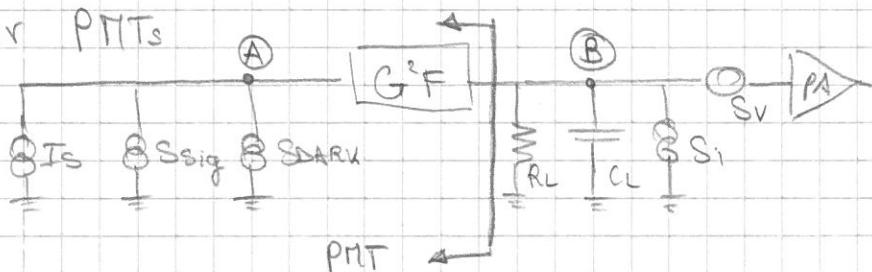
$S_{DARK} = \sqrt{2qI_p} =$ shot noise induced by dark current

$S_I, S_V =$ preamplifier noise sources ($\sqrt{S_I}/\text{voltage} = \sqrt{\frac{S_V}{R^2}}$)

$\sqrt{\frac{kT}{R_L}} =$ resistor voltage noise

$\sqrt{2qI_S} =$ shot noise induced by signal current itself

For PMTs



PA input referred noises (current) :

$$\sqrt{2qI_S G^2 F} = \text{signal noise}$$

$$\sqrt{2qI_D G^2 F} = \sqrt{2q^2 n_b G^2 F} \quad \text{where } n_b = \text{dark count rate}$$

$$\sqrt{\frac{kT}{R}} = \text{resistor noise}$$

$$\sqrt{S_I} = \rightarrow \text{PA input current noises}$$

$$\sqrt{S_V/R_L^2} = \boxed{ }$$

It is possible to refer all the noises in ① instead of ② by correctly multiplying by $G^2 F$

PMTs

$G \sim 10^3 \div 10^6$ easily

$F \approx 2$ or less for most PMTs ($F=1$ best tech)

η depends a lot from PMT type

S20: range from 300 to ~ 800 nm

$\eta = 20\%$ at 350nm (highest), 0,5% at 800nm

S11: range from 300 to ~ 600 nm

$\eta = 15\%$ at 450nm (highest), 1% at 800nm

Silicon PDs

Absorption lengths:

$$\text{at } 400\text{ nm} \rightarrow L_a = 100\mu\text{m}$$

$$500\text{ nm} \rightarrow L_a = 1\mu\text{m}$$

$$600\text{ nm} \rightarrow L_a = 2\mu\text{m}$$

$$700\text{ nm} \rightarrow L_a = 5\mu\text{m}$$

$$800\text{ nm} \rightarrow L_a = 10\mu\text{m}$$

Neutral region $W_n = 200\text{ nm} \div 2\mu\text{m}$

(100nm best tech available)

Depleted region has to be $W_d \geq L_a$

$W_d \sim 5L_a$ is preferable

Reflectivity $R = 0,2 \div 0,4$

(0,1 best tech available)

$$\eta_D = (1-R) e^{-\frac{W_n}{L_a}} (1-e^{-\frac{W_d}{L_a}})$$

Ⓐ Ⓡ Ⓢ

$$C|_{PD} = E_0 \epsilon_s \cdot \frac{A}{W_d} \xrightarrow[W_d]{\substack{\nearrow \text{area of} \\ \nearrow \text{detector}}} \xrightarrow[\text{depleted layer}]{\substack{\nearrow \\ \nearrow}}$$

Ⓐ Reflect the sunlest amount of photons as possible

Ⓑ low number of photons must be absorbed here (W_n has to be very thin)

Ⓒ Absorb the highest amount of photons as possible
(W_d has to be thick, typically $\approx 5 L_a$)

Single Electron Response (SER): few tens of picoseconds

APDs

Same acquisition scheme for PMTs, but:

$G = \text{up to } 500 \text{ max} \rightarrow \text{strongly varies with temperature}$

$G = 1000$ can not be feasible

$F = 2$ lowest possible

$F \approx 2,5$ for $G = 100$

$F \approx 5$ for $G = 500$

Low noise preamplifier

$$\sqrt{S_{V,U}} = 2 \div 5 \text{ nV}/\sqrt{\text{Hz}} \quad \boxed{\text{Reasonable values}}$$

$$\sqrt{S_{I,U}} = 0,5 \text{ pA}/\sqrt{\text{Hz}}$$

Best LNA values available $\sqrt{S_{V,U}} \approx 1 \text{ nV}/\sqrt{\text{Hz}}$ $\sqrt{S_{I,U}} = 0,01 \text{ pA}/\sqrt{\text{Hz}}$

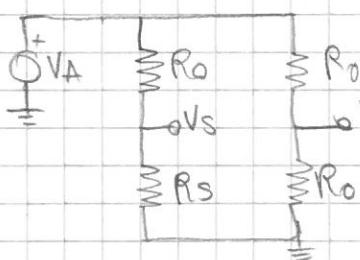
Temperature Sensors

Thermistors: non linear response with temperature \rightarrow can be compensated using look-up tables in microprocessors

PT100: very linear response through a larger temperature range ($\sim 13,8K$ to $303K$)

$$\Delta T = T - T_0$$

Wheatstone bridge consider $R_s = R_0 (1 + \alpha \Delta T)$



$$R_s = R_0 + \Delta R_s \quad \left\{ \begin{array}{l} V_o = V_A \frac{R_0}{2R_0} = \frac{V_A}{2} \\ V_s = V_A \cdot \frac{R_s}{R_0 + R_s} \end{array} \right.$$

\rightarrow Large signal

$$\Delta V_s = V_s - V_A = V_A \left(\frac{R_0 + \Delta R_s}{2R_0 + \Delta R_s} - \frac{1}{2} \right) = \frac{V_A}{2} \frac{\Delta R_s}{2R_0 + \Delta R_s} \text{ expression}$$

$$\text{For small variations } \Delta R_s \ll 2R_0, \text{ so } \Delta V_s = \frac{V_A}{4} \cdot \frac{\Delta R_s}{R_0} = \frac{V_A \alpha \Delta T R_0}{4 R_0}$$

$$\Delta V_s = \frac{V_A}{4} \cdot \alpha \Delta T \rightsquigarrow \text{small signal expression}$$

Resulting noise from the wheatstone bridge will be

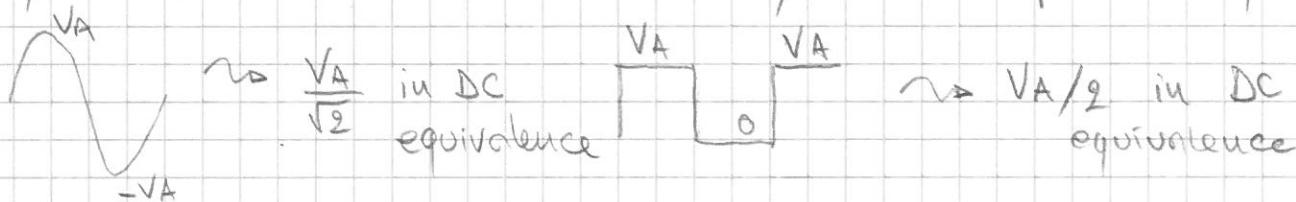
$$\sqrt{S_V}, \sqrt{S_I R_s^2}, \sqrt{4 k T R_s}, 1/f \text{ components}$$

Usually, a power requirement is set not to cause self heating of the system, for example $\left(\frac{V_A/2}{R_0}\right)^2 \leq P_{MAX}$

We find V_A value when solved

For sinusoidal V_A , voltage bias can be $\sqrt{2}$ higher to its DC equivalent

Note: When using LIA's, sine/square V_A can change SNR by a lot \rightarrow remember to always check power requirement:



This can be misleading in the type of wave to use in the LIA, but the end result (considering P_{MAX} requirement) will be the same regardless of the approach used. Example:

$$\circ \text{Sine } V_A \rightarrow \frac{\left(\frac{V_A}{2\sqrt{2}}\right)^2}{R_0} \leq P_{MAX} \rightarrow |V_A|_{MAX} = \sqrt{8R_0P_{MAX}} \quad SNR = \frac{\sqrt{8R_0P_{MAX}} \times \Delta T}{\sqrt{2} \sigma_{LPF}}$$

$$\circ \text{Square } V_A \rightarrow \frac{\left(\frac{V_A}{2}\right)^2}{R_0} \leq P_{MAX} \rightarrow |V_A|_{MAX} = 2\sqrt{R_0P_{MAX}} \quad SNR = \frac{2\sqrt{R_0P_{MAX}} \times \Delta T}{\sigma_{LPF}}$$

$$\text{We can see } SNR|_{\text{Sine LIA}} = SNR|_{\text{square LIA}}$$

STRAIN GAUGES

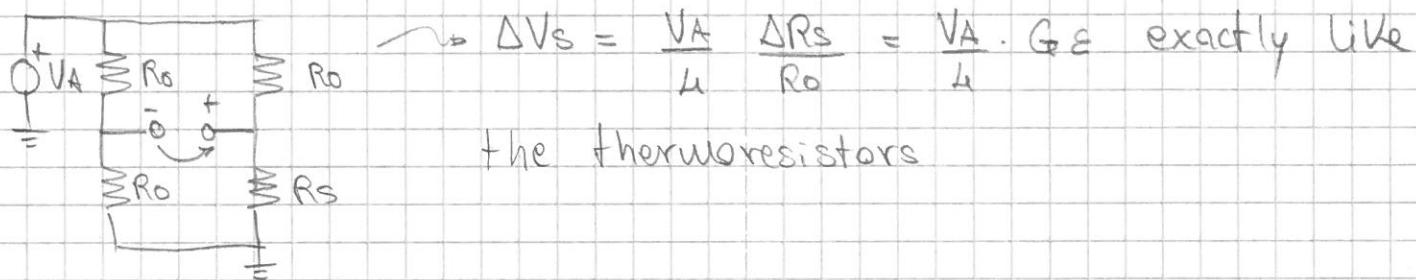
Strain $\epsilon = \frac{N}{E}$ \rightarrow Force applied
 \hookrightarrow Young Modulus

$$R = \rho \frac{L}{A} \rightarrow \frac{\Delta R}{R_0} = \frac{\Delta L}{L_0} - \frac{\Delta A}{A_0} + \frac{\Delta \rho}{\rho_0} = \epsilon + \alpha \nu \epsilon + \beta E \epsilon = G \epsilon$$

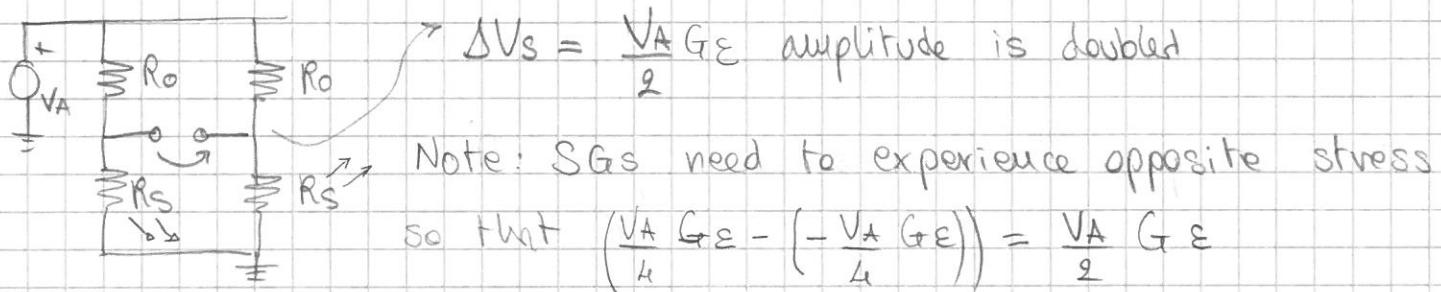
Where G = gauge factor

When an extension of the resistor is applied \rightarrow length increases, area decreases and resistivity changes because of the piezoresistive effect $\rho = \rho_0 (1 + \beta N)$

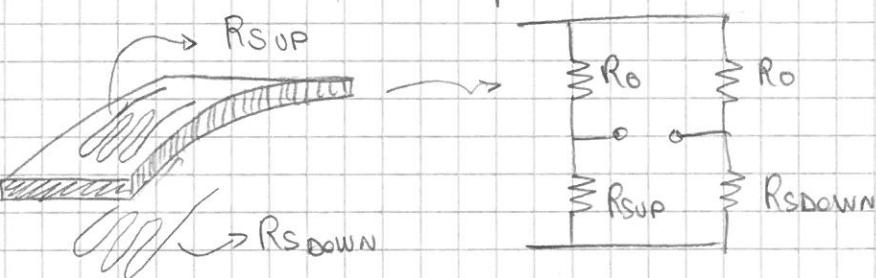
Wheatstone bridge



If two SG are used :



How can they be positioned?



Using this configuration it is possible to measure compression and extension

Temperature dependency

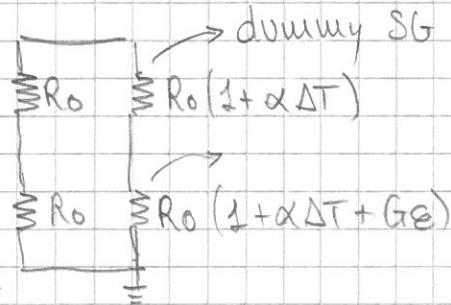
SGs are often placed in hot places or on pieces which temperature changes rapidly

$$\Delta V_s \Big|_{\text{temperature}} = \frac{V_A}{4} \propto \Delta T \rightarrow \text{same expression}$$

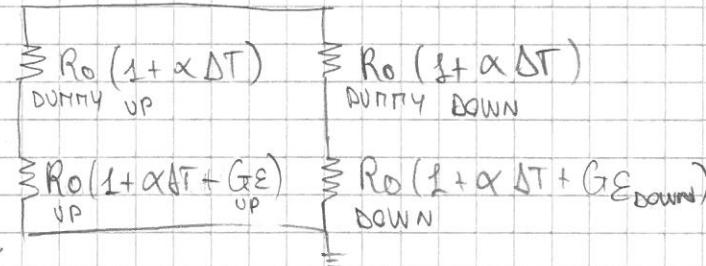
for the PT100



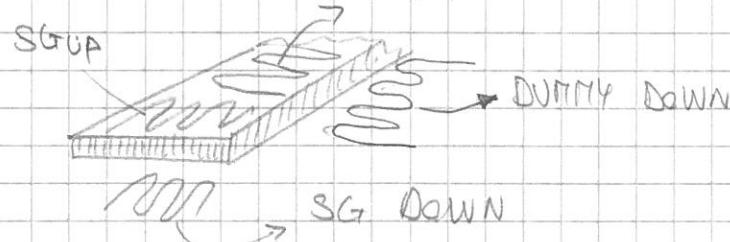
To solve this, dummy SGs are used:



↳ Single SG



↳ Comp/extension double SG configuration



Dummy orientation is perpendicular to SG so that it does not experience the wanted stress

$$\text{If } \Delta T \Big|_{\text{dummy}} = \Delta T \Big|_{\text{SG}} \rightarrow \Delta V_s = 0 \text{ but } \Delta T_D \neq \Delta T_{SG}$$

(They usually differ by $\pm 0,1^\circ\text{C}$ or something similar) So

$$\Delta V_s \Big|_{\text{temperature}} = \frac{V_A}{4} \propto (\Delta T_{SG} - \Delta T_D) \text{ or } \frac{V_A}{2} \propto (\Delta T_{SG} - \Delta T_D)$$

example:

$$V_A = 50\text{mV}, \alpha = 4\text{m} \cdot \frac{1}{^\circ\text{C}}, \Delta T = 100^\circ\text{C}, (\Delta T_{SG} - \Delta T_D) = 0,1^\circ\text{C}$$

$$\Delta V_s \Big|_{\text{w/o dummy}} = \frac{V_A}{4} \propto \Delta T = 5 \cdot 10^{-3} \text{ strain} \rightarrow \text{Changes by}$$

$$\Delta V_s \Big|_{\text{with dummy}} = \frac{V_A}{4} \propto (\Delta T_{SG} - \Delta T_D) = 5 \cdot 10^{-6} \text{ strain}$$

3 orders of magnitude

