

PHYSICS KNOWLEDGE ORGANISERS

Mechanical	Force acts upon an object
Electrical	Electric current flow
Heat	Temperature difference between objects
Radiation	Electromagnetic waves or sound

Energy pathways

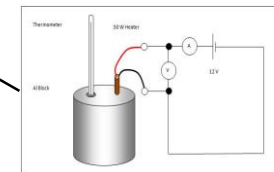
Change in thermal energy = mass **X** specific heat capacity **X** temperature change $\Delta E = m \times c \times \Delta \theta$

Specific Heat Capacity	Energy needed to raise 1kg of substance by 1°C	Depends on: mass of substance, what the substance is and energy put into the system.
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HIGHER: efficiency can be increased using machines.

Efficiency = $\frac{\text{Useful power output}}{\text{Total power input}}$

Efficiency = $\frac{\text{Useful output energy transfer}}{\text{Total input energy transfer}}$



Kinetic energy	Energy stored by a moving object	$\frac{1}{2} \times \text{mass} \times (\text{speed})^2$ $\frac{1}{2} mv^2$
Elastic Potential energy	Energy stored in a stretched spring, elastic band	$\frac{1}{2} \times \text{spring constant} \times (\text{extension})^2$ $\frac{1}{2} ke^2$ (Assuming the limit of proportionality has not been exceeded)
Gravitational Potential energy	Energy gained by an object raised above the ground	Mass \times gravitational field strength \times height mgh

System	An object or group of objects that interact together	EG: Kettle boiling water.
Energy stores	Kinetic, chemical, internal (thermal), gravitational potential, elastic potential, magnetic, electrostatic, nuclear	Energy is gained or lost from the object or device.
Ways to transfer energy	Light, sound, electricity, thermal, kinetic are ways to transfer from one store to another store of energy.	EG: electrical energy transfers chemical energy into thermal energy to heat water up.
Unit	Joules (J)	

Work	Doing work transfers energy from one store to another	By applying a force to move an object the energy store is changed.	Work done = Force \times distance moved $W = Fs$
Power	The rate of energy transfer	1 Joule of energy per second = 1 watt of power	Power = energy transfer \div time $P = E \div t$ Power = work done \div time, $P = W \div t$

	Units
Specific Heat Capacity	Joules per Kilogram degree Celsius (J/Kg°C)
Temperature change	Degrees Celsius (°C)
Work done	Joules (J)
Force	Newton (N)
Distance moved	Metre (m)
Power	Watts (W)
Time	Seconds (s)

Useful energy	Energy transferred and used
Wasted energy	Dissipated energy, stored less usefully

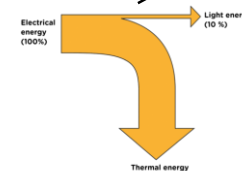
Prefix	Multiple	Standard form
Kilo	1000	10^3
Mega	1000 000	10^6
Giga	100 000 000	10^9

Energy stores and changes

AQA ENERGY – part 1

Energy Conservation and Dissipation

Closed system	No change in total energy in system
Open system	Energy can dissipate



HIGHER: When an object is moved, energy is transferred by doing work.

Work done = Force \times distance moved

Frictional forces cause energy to be transferred as thermal energy. This is wasted.

Dissipate	To scatter in all directions or to use wastefully	When energy is 'wasted', it dissipates into the surroundings as internal (thermal) energy.
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Ways to reduce 'wasted' energy	Energy transferred usefully	Insulation, streamline design, lubrication of moving parts.
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Principle of conservation of energy	The amount of energy always stays the same.	Energy cannot be created or destroyed, only changed from one store to another.
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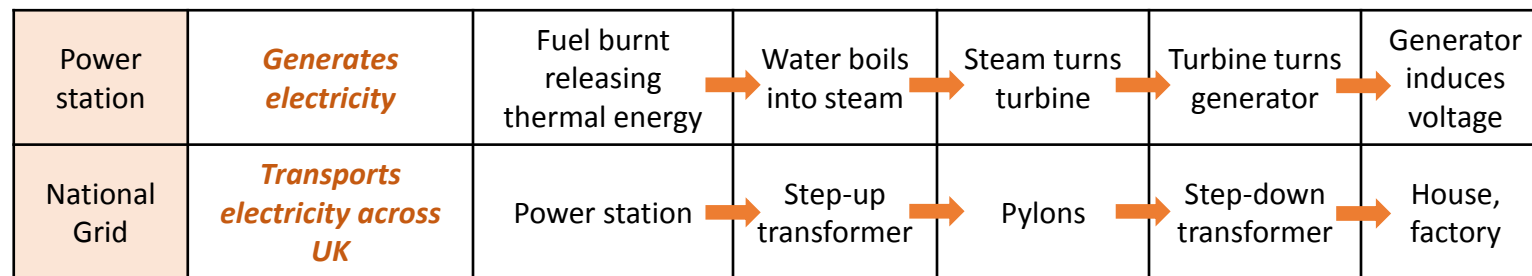
	Units
Energy (KE, EPE, GPE, thermal)	Joules (J)
Velocity	Metres per second (m/s)
Spring constant	Newton per metre (N/m)
Extension	Metres (m)
Mass	Kilogram (Kg)
Gravitational field strength	Newton per kilogram (N/Kg)
Height	Metres (m)

Reducing friction - using wheels, applying lubrication. Reducing air resistance – travelling slowly, streamlining.

Using renewable energy will need to increase to meet demand.

Transport	<i>Petrol, diesel, kerosene produced from oil</i>	Used in cars, trains and planes.
Heating	<i>Gas and electricity</i>	Used in buildings.
Electricity	<i>Most generated by fossil fuels</i>	Used to power most devices.

Power station – NB: You need to understand the principle behind generating electricity. An energy resource is burnt to make steam to drive a turbine which drives the generator.



Renewable energy makes up about 20% of energy consumption.

Fossil fuel reserves are running out.

Energy demand is increasing as population increases.

Non-renewable energy resource	<i>These will run out. It is a finite reserve. It cannot be replenished.</i>	e.g. Fossil fuels (coal, oil and gas) and nuclear fuels.
Renewable energy resource	<i>These will never run out. It is an infinite reserve. It can be replenished.</i>	e.g. Solar, Tides, Waves, Wind, Geothermal, Biomass, Hydroelectric

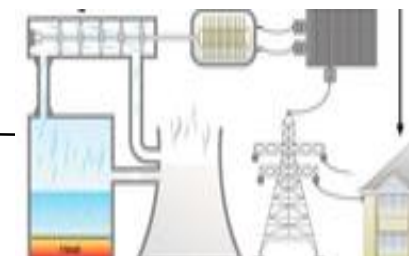
Using fuels

Energy resources

Global Energy Resources

AQA ENERGY part 2

National Grid



Energy resource	<i>How it works</i>	Uses	Positive	Negative
Fossil Fuels (coal, oil and gas)	<i>Burnt to release thermal energy used to turn water into steam to turn turbines</i>	Generating electricity, heating and transport	Provides most of the UK energy. Large reserves. Cheap to extract. Used in transport, heating and making electricity. Easy to transport.	Non-renewable. Burning coal and oil releases sulfur dioxide. When mixed with rain makes acid rain. Acid rain damages building and kills plants. Burning fossil fuels releases carbon dioxide which contributes to global warming. Serious environmental damage if oil spilt.
Nuclear	<i>Nuclear fission process</i>	Generating electricity	No greenhouse gases produced. Lots of energy produced from small amounts of fuel.	Non-renewable. Dangers of radioactive materials being released into air or water. Nuclear sites need high levels of security. Start up costs and decommission costs very expensive. Toxic waste needs careful storing.
Biofuel	<i>Plant matter burnt to release thermal energy</i>	Transport and generating electricity	Renewable. As plants grow, they remove carbon dioxide. They are 'carbon neutral'.	Large areas of land needed to grow fuel crops. Habitats destroyed and food not grown. Emits carbon dioxide when burnt thus adding to greenhouse gases and global warming.
Tides	<i>Every day tides rise and fall, so generation of electricity can be predicted</i>	Generating electricity	Renewable. Predictable due to consistency of tides. No greenhouse gases produced.	Expensive to set up. A dam like structure is built across an estuary, altering habitats and causing problems for ships and boats.
Waves	<i>Up and down motion turns turbines</i>	Generating electricity	Renewable. No waste products.	Can be unreliable depends on wave output as large waves can stop the pistons working.
Hydroelectric	<i>Falling water spins a turbine</i>	Generating electricity	Renewable. No waste products.	Habitats destroyed when dam is built.
Wind	<i>Movement causes turbine to spin which turns a generator</i>	Generating electricity	Renewable. No waste products.	Unreliable – wind varies. Visual and noise pollution. Dangerous to migrating birds.
Solar	<i>Directly heats objects in solar panels or sunlight captured in photovoltaic cells</i>	Generating electricity and some heating	Renewable. No waste products.	Making and installing solar panels expensive. Unreliable due to light intensity.
Geothermal	<i>Hot rocks under the ground heats water to produce steam to turn turbine</i>	Generating electricity and heating	Renewable. Clean. No greenhouse gases produced.	Limited to a small number of countries. Geothermal power stations can cause earthquake tremors.

Radius of an atom
 $1 \times 10^{-10}\text{m}$



Electrons gained
Negative ion

Electrons lost
Positive ion

Atom	Same number of protons and electrons
Ion	Unequal number of electrons to protons
Mass number	Number of protons and neutrons
Atomic number	Number of protons

Particle	Charge	Size	Found
Neutron	None	1	In the nucleus
Proton	+	1	
Electron	-	Tiny	Orbits the nucleus

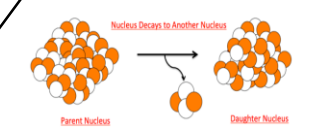
Isotope

${}^6_3\text{Li}$
 ${}^7_3\text{Li}$

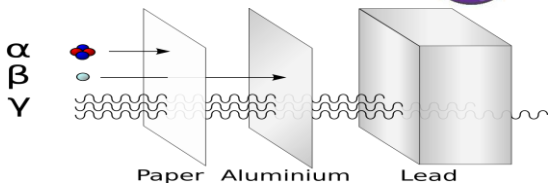
Different forms of an element with the same number of protons but different number of neutrons

Discovery of the nucleus

Democritus	Suggested idea of atoms as small spheres that cannot be cut.
J J Thomson (1897)	Discovered electrons– emitted from surface of hot metal. Showed electrons are negatively charged and that they are much less massive than atoms.
Thomson (1904)	Proposed ‘ <i>plum pudding</i> ’ model – atoms are a ball of positive charge with negative electrons embedded in it.
Geiger and Marsden (1909)	Directed beam of alpha particles (He^{2+}) at a thin sheet of gold foil. Found some travelled through, some were deflected, some bounced back.
Rutherford (1911)	Used above evidence to suggest alpha particles deflected due to electrostatic interaction between the very small charged nucleus, nucleus was massive. Proposed mass and positive charge contained in nucleus while electrons found outside the nucleus which cancel the positive charge exactly.
Bohr (1913)	Suggested modern model of atom – electrons in circular orbits around nucleus, electrons can change orbits by emitting or absorbing electromagnetic radiation. His research led to the idea of some particles within the nucleus having positive charge; these were named protons.
Chadwick (1932)	Discovered neutrons in nucleus – enabling other scientists to account for mass of atom.

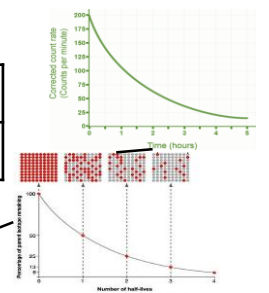
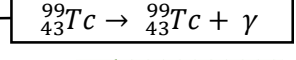
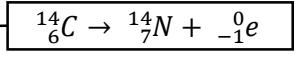
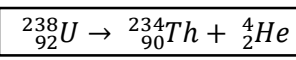


Decay	Range in air	Ionising power	Penetration power
Alpha	Few cm	Very strong	Stopped by paper
Beta	Few m	Medium	Stopped by Aluminium
Gamma	Great distances	Weak	Stopped by thick lead



Radioactive decay	Unstable atoms randomly emit radiation to become stable
Detecting	Use Geiger Muller tube
Unit	Becquerel
Ionisation	All radiation ionises

Decay	Emitted from nucleus	Changes in mass number and atomic number	
Alpha (α)	Helium nuclei (${}^4_2\text{He}$)	-4	-2
Beta (β)	Electron (${}^0_{-1}\text{e}$)	0	+1
Gamma (γ)	Electromagnetic wave	0	0
Neutron	Neutron	-1	0



Atoms and Isotopes

Atoms and Nuclear Radiation

Contamination	Unwanted presence of radioactive atoms
Irradiation	Person is in exposed to radioactive source

AQA ATOMIC STRUCTURE

PHYSICS ONLY: Hazards and uses of Radioactive emissions and of background radiation

Half life	The time taken to lose half of its initial radioactivity
Sievert	Unit measuring dose of radiation
Background	Constant low level environmental radiation, e.g. from nuclear testing, nuclear power, waste

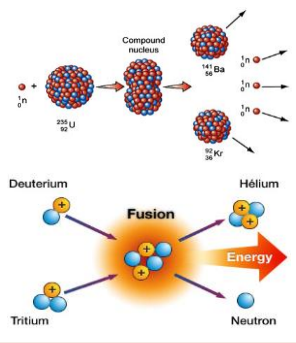
Nuclear fission and fusion

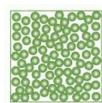
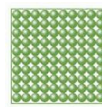
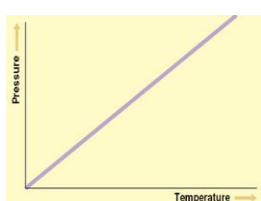
Uses	Different isotopes have different half lives	Short half-lives used in high doses, long half lives used in low doses.
Tracers	Used within body	Isotope with short half life injected, allowed to circulate and collect in damaged areas. PET scanner used to detect emitting radiation. Must be beta or gamma as alpha does not penetrate the body.
Radiation therapy	Used to treat illnesses e.g. cancer	Cancer cells killed by gamma rays. High dose used to kill cells. Damage to healthy cells prevented by focussed gamma ray gun.

Fuel rods	Made of U-238, ‘enriched’ with U-235 (3%). Long and thin to allow neutrons to escape, hitting nuclei.
Control rods	Made of Boron. Controls the rate of reaction. Boron absorbs excess neutrons.
Concrete	Neutrons hazardous to humans – thick concrete shield protects workers.

PHYSICS ONLY: Nuclear energy

Nuclear fission	One large unstable nucleus splits to make two smaller nuclei	Neutron hits U-235 nucleus, nucleus absorbs neutron, splits emitting two or three neutrons and two smaller nuclei. Process also releases energy.	Process repeats, chain reaction formed
Nuclear fusion	Two small nuclei join to make one larger nucleus	Difficult to do on Earth – huge amounts of pressure and temperature needed.	Occurs in stars





State	Particle arrangement	Properties
Solid	Packed in a regular structure. Strong forces hold in place so cannot move.	Difficult to change shape.
Liquid	Close together, forces keep contact but can move about.	Can change shape but difficult to compress.
Gas	Separated by large distances. Weak forces so constantly randomly moving.	Can expand to fill a space, easy to compress.

	Units
Density	Kilograms per metre cubed (kg/m³)
Mass	Kilograms (kg)
Volume	Metres cubed (m³)
Energy needed	Joules (J)
Specific latent heat	Joule per kilogram (J/kg)
Change in thermal energy	Joules (J)
Specific heat capacity	Joule per kilogram degrees Celsius (J/kg°C)
Temperature change	Degrees Celsius (°C)
Pressure	Pascals (Pa)

Pressure of a fixed volume of gas increases as temperature increases (temperature increases, speed increases, collisions occur more frequently and with more force so pressure increases).

Temperature of gas is linked to the average kinetic energy of the particles.

If kinetic energy increases so does the temperature of gas.

No kinetic energy is lost when gas particles collide with each other or the container.

Gas particles are in a constant state of random motion.

$$P = m \div V$$

Density = mass \div volume.



Density	Mass of a substance in a given volume
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Freezing	Liquid turns to a solid. Internal energy decreases.
Melting	Solid turns to a liquid. Internal energy increases.
Boiling / Evaporating	Liquid turns to a gas. Internal energy increases.
Condensation	Gas turns to a liquid. Internal energy decreases.
Sublimation	Solid turns directly into a gas. Internal energy increases.
Conservation of mass	When substances change state, mass is conserved.
Physical change	No new substance is made, process can be reversed.

Kinetic theory of gases

Particle model

Pressure

AQA PARTICLE MODEL OF MATTER

PHYSICS ONLY: when you do work the temperature increases e.g. pump air quickly into a ball, the air gets hot because as the piston in the pump moves the particles bounce off increasing kinetic energy, which causes a temperature rise.

Reducing the volume of a fixed mass of gas increases the pressure.

Halving the volume doubles the pressure.

$$PV = \text{constant.}$$

$$P_1 V_1 = P_2 V_2$$

Specific Heat Capacity

Energy needed to raise 1kg of substance by 1°C

Depends on:

- Mass of substance
- What the substance is
- Energy put into the system.

Change in thermal energy = mass **X** specific heat capacity **X** temperature change.

$$\Delta E = m \times c \times \Delta \theta$$

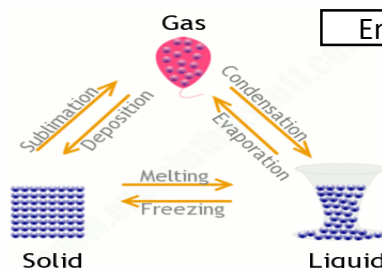
Internal energy and energy transfers

Change of state

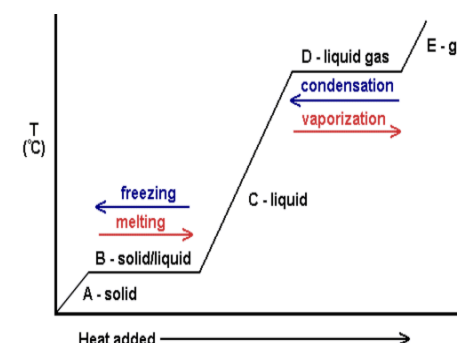
Specific Latent Heat	Energy needed to change 1kg of a substance's state
Specific Latent Heat of Fusion	Energy needed to change 1kg of solid into 1 kg of liquid at the same temperature
Specific Latent Heat of Vaporisation	Energy needed to change 1kg of liquid into 1 kg of gas at the same temperature

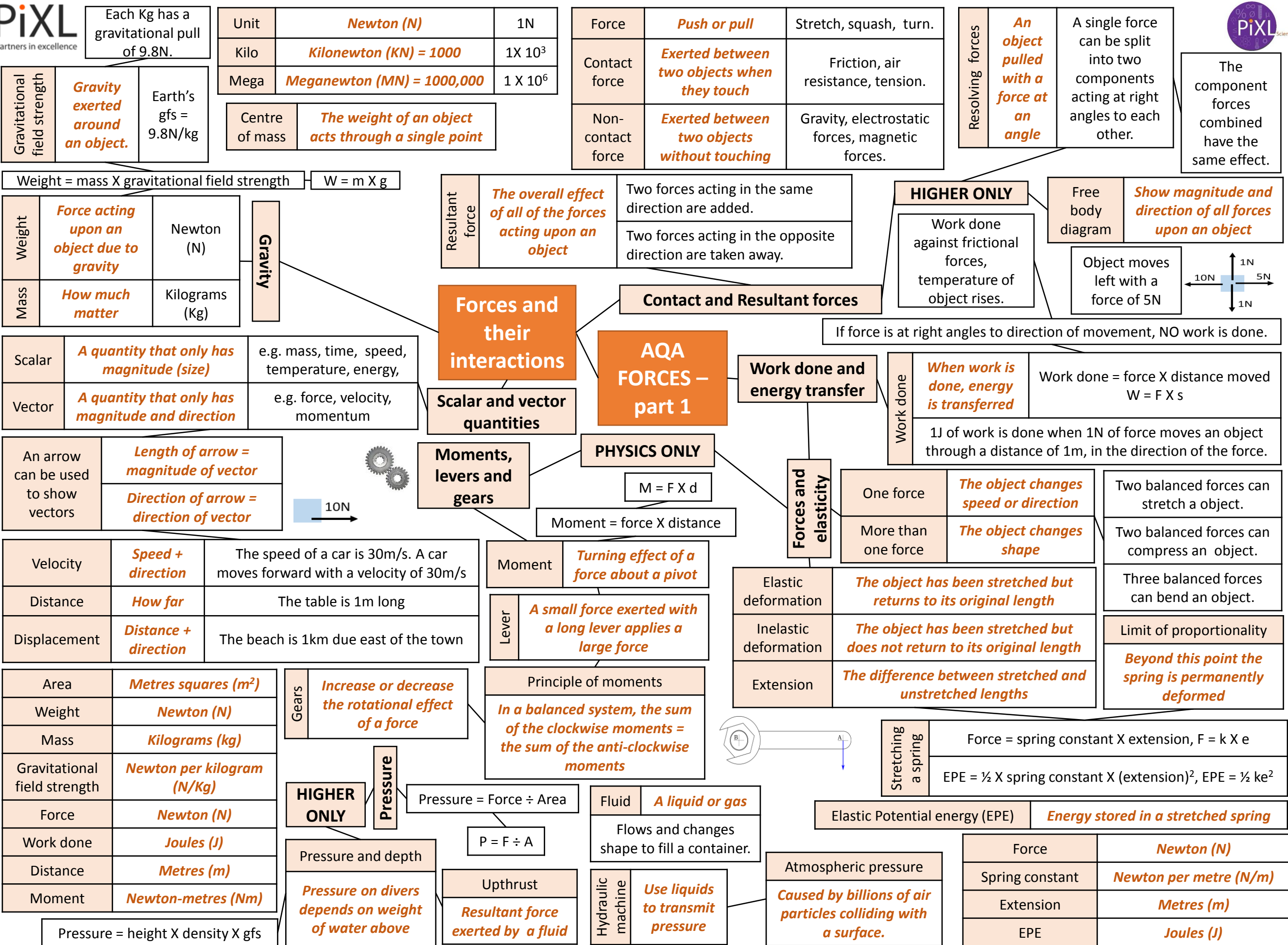
Energy needed = mass **X** specific latent heat.

$$\Delta E = m \times L$$



Internal energy	Energy stored inside a system by particles	Internal energy is the total kinetic and potential energy of all the particles (atoms and molecules) in a system.
	Heating changes the energy stored within a system	Heating causes a change in state. As particles separate, potential energy stored increases. Heating increases the temperature of a system. Particles move faster so kinetic energy of particles increases.





Aeroplane banks to change direction	Velocity changes.
Car travelling around a bend	Constant speed, direction changes.
Satellite orbiting the Earth	Constant speed, direction changes.

Distance travelled	Area under the graph shape
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Constant acceleration
$(\text{final velocity})^2 - (\text{initial velocity})^2 = 2 \times \text{acceleration} \times \text{distance}$ $V^2 - u^2 = 2 \times a \times s$

Gradient = vertical ÷ horizontal

HIGHER ONLY

Accelerating objects	It takes time for objects to reach top speed	Draw a tangent to the curve, work out gradient.
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Velocity-time graph	Shows speed of an object
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Accelerating	Object getting faster
Decelerating	Object slowing down

Falling objects

Falling objects accelerate due to gravity.	In no air resistance, objects accelerate at 9.8m/s^2	Air resistance slows falling objects down.
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Terminal velocity	Weight of an object is balanced by resistive forces	Object moves at a constant velocity. Resultant force = 0.
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PHYSICS ONLY	Parachuting	Size of air resistance depends on area of object and speed
	Larger the area, the larger the air resistance.	
	Larger the speed, the larger the air resistance.	

Inertia	When objects continue in the same state of motion	Speed or direction only changes if a resultant force acts on the object
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HIGHER ONLY

Newton's first Law	Balanced forces	When the resultant force on an still object = 0, the object is stationary.
Newton's second Law	Unbalanced forces	When the resultant force on a moving object = 0, the object is at a constant speed.
Newton's third Law	Equal and opposite forces	When the resultant force is greater than 0, the object accelerates. It could speed up, slow down or change direction.
		When two objects interact the forces exerted are equal and in an opposite direction.

Acceleration is proportional to resultant force.
Acceleration is inversely proportional to mass.

Momentum	HIGHER ONLY
Is a vector	$p = m \times v$

Momentum = mass X velocity

Changes in momentum
Force is applied to stop momentum
If momentum changes slowly, the force applied is small so less damage.

HIGHER ONLY

PHYSICS HIGHER ONLY

Crumple zones

Speed / velocity	Metres per second (m/s)
Distance	Metres (m)
Time	Seconds (s)
Acceleration	Metres per second squared (m/s²)
Force	Newton (N)
Mass	Kilogram (Kg)
Momentum	Kilograms metres per second (Kgm/s)

Conservation of momentum
When two objects collide, the momentum they have before the collision = the momentum they have after the collision
Closed system = no external forces acting on it.

Inertial mass	How difficult it is to change the velocity of an object
Inertial mass = force ÷ acceleration	
If the mass is large, to change velocity a big force is needed.	

HIGHER ONLY

$$F = m \times a$$

AQA FORCES – part 2

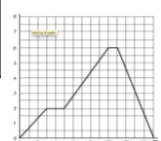
Observing and recording motion

Forces and braking

Speed is rarely constant.

Distance-time graph	Shows how far an object moves along a straight line
Speed of object	Use the gradient of graph

Describing motion



Forces, acceleration and Newton's Laws of motion

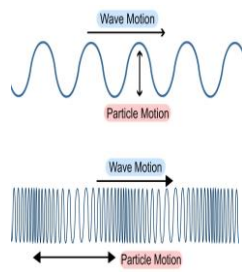
Speed	How fast an object moves	Scalar
Displacement	Includes the distance and direction an object moves	vector
Distance	How far an object moves	scalar

Car on motorway	30m/s	Walking	1.5m/s
Train	60m/s	Running	3m/s
Jet plane	200m/s	Cycling	6m/s

Speed affects both thinking and braking distances.	Frictional forces decelerate a moving object and bring it to rest.
Thinking distance	Distance travelled whilst the driver reacts
Braking distance	Distance travelled whilst the car is stopped by the brakes
Stopping distance	Total thinking and braking distances

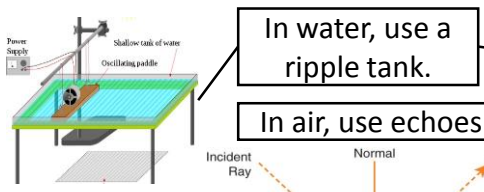
Factors affecting stopping distances	Drivers reaction times	Drinking alcohol, taking drugs, tired.
	Braking distances	Weather conditions, worn brakes or tyres, road surface, size of braking force.
Braking and kinetic energy	Work done by braking force, reduces kinetic energy	Kinetic energy decreases, temperature of brakes increases due to frictional forces.

Wave speed	Wave speed = frequency X wavelength	$V = f \times \lambda$
Wave period	Wave period = $1 \div \text{frequency}$	$T = 1 \div f$
Speed	Speed = distance \div time	$v = d \div t$



Transverse wave	Vibration causing the wave is at right angles to the direction of energy transfer	Energy is carried outwards by the wave.	Water and light waves, S waves.
Longitudinal wave	Vibration causing the wave is parallel to the direction of energy transfer	Energy is carried along the wave.	Sound waves, P waves.

Wavelength	Distance from one point on a wave to the same point of the next wave
Amplitude	The maximum disturbance from its rest position
Frequency	Number of waves per second
Period	Time taken to produce 1 complete wave



Measuring speed

In water, use a ripple tank.

In air, use echoes.

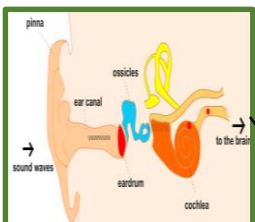
Properties

Air Water

Sound waves travelling through different mediums, the frequency stay constant.

Angle of incidence = angle of reflection
(i) = (r)

Reflection	Wave bounces off the surface.
Refraction	Waves changes direction at boundary.
Transmitted	Passes through the object.
Absorbed	Passes into but not out of, transfers energy and heats up the object.

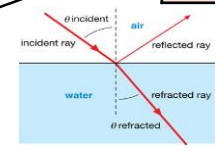


PHYSICS HIGHER ONLY

Hearing

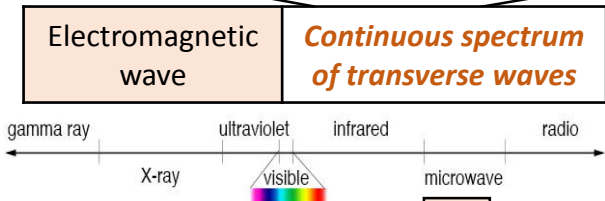
Frequencies between 20 – 20,000 Hz

Longitudinal waves cause ear drum to vibrate, amplified by three ossicles which creates pressure in the cochlea.



Light refracts as it slows down in a denser substance

Electromagnetic waves



Absorbed light changes into thermal energy store.

Transverse and Longitudinal waves

Waves in air, fluids and solids

AQA Waves

Black body radiation

e.g. Gamma

Short wavelengths have high frequency and high energy.

PHYSICS ONLY

Earth and Global warming

Ultraviolet, visible light, infra-red radiation penetrate atmosphere and heat up Earth's surface.

Longer wavelengths are radiated back, trapped by atmosphere.

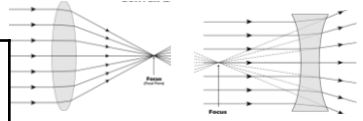
Energy lost is not at the same rate as energy being absorbed so Earth heats up.

Black body radiation	All objects absorb or reflect infrared radiation	Hotter objects emit more infrared radiation.
Constant temperature	Rate of absorption = rate of radiation	Intensity and wavelength of energy affects temperature.

PHYSICS ONLY

Magnification = image size \div object size

HIGHER: Lenses



Convex	Real or virtual images.
Concave	Only virtual images.

2F	Image same size, upside down, real.
2F - F	Image larger, upside down, real.
< F	Image bigger, right way, virtual.

Specular	Flat surface reflection.
Diffuse	Rough surface reflection.

HIGHER: Properties

EM wave	Danger	Use
Radio	Safe.	Communications, TV, radio.
Microwave	Burning if concentrated.	Mobile phones, cooking, satellites.
Infrared		Heating, remote controls, cooking.
Visible	Damage to eyes.	Illumination, photography, fibre optics.
Ultra violet	Sunburn, cancer.	Security marking, disinfecting water.
X-ray	Cell destruction, mutation, cancer.	Broken bones, airport security.
Gamma		Sterilising, detecting and killing cancer.

Low frequency, long wavelength.

High frequency, short wavelength

White Wave lengths reflected

Black Wave lengths absorbed

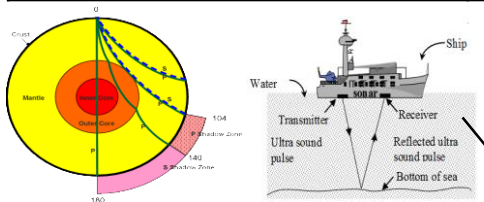
Seismic waves

P wave	S wave	Seismograph
Longitudinal	Transverse	Shows P and S waves arriving at different times.
Fast	Slow	
Travel through solids and liquids	Travels through solids	By using the times the waves arrive at the monitoring centres, the epicentre of earthquake can be found. ($v = x \div t$).
Produced by earthquakes.		

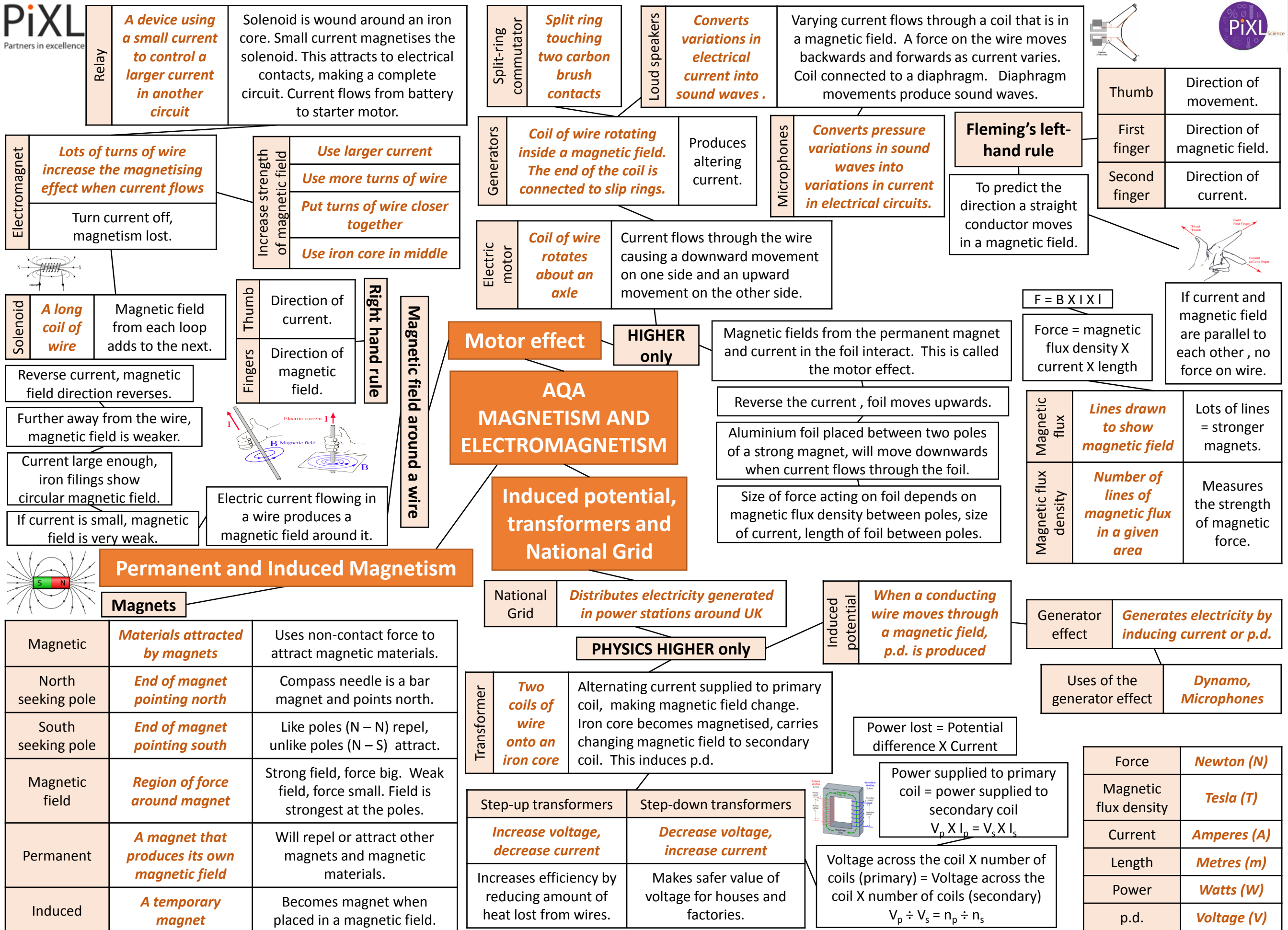
Black surfaces	Good emitters, good absorbers
White surfaces	Poor emitters, poor absorbers
Shiny surfaces	Good reflectors



EM waves refract



Ultra sound	Partially reflected off boundary	Used for medical and foetal scans.
Sonar	Reflected off objects	Used to determine depth of objects under the sea.





Milky Way
our galaxy.

Planet	<i>A large body orbiting the Sun</i>
Moon	<i>A natural satellite orbiting a planet</i>
Dwarf planet	<i>A body large enough to have its own gravity which caused a spherical shape</i>
Solar system	<i>Any object orbiting the Sun due to gravity</i>
Galaxy	<i>Collection of billions of stars</i>
Universe	<i>Collection of galaxies</i>



Comets, asteroids, satellites.

Other objects.

Solar system

Effect of gravity.

Gravity causes moons to orbit planets, planets to orbit the Sun, stars to orbit galaxy centres.

Force of gravity changes the moon's direction not its speed.

Gravity pulls objects towards the ground.

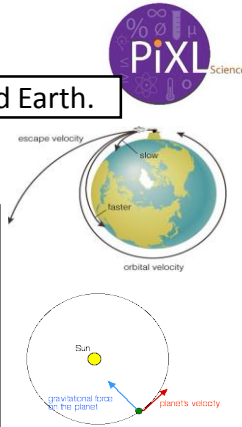
Speed of Orbit.

Too fast = disappears into Space.

Correct speed = steady orbit around Earth.

Too slow = falls to Earth.

To calculate speed of Orbit: distance object moves in 1 orbit, Distance = $2\pi r$, then average speed = distance \div time.



Orbital motions

HIGHER: Circular orbits.

Velocity = a vector.

A planet's velocity changes but speed remains constant.

Due to the Sun's gravity, planets accelerate towards the Sun and so changes direction.

When ambulances go past the sound changes from a high pitch to a low pitch.

Planets close to the Sun, gravity pull is strong. Planets move quickly.

Planets further away from the Sun, gravity pull is weaker. So speed of planet is slower.

Frequency of sound wave decreases, wavelength increases.

The life cycle of a star.

Nebula	<i>A cloud of cold hydrogen gas and dust</i>	Cloud collapses due to gravity, particles move very fast colliding with each other, kinetic energy transfers into internal energy and the temperature increases.
Protostar	<i>The large ball of gas contracts to form a star</i>	High temperature causes Hydrogen nuclei to collide and nuclear fusion begins. A star is 'born'.
Main sequence	<i>Stable period of star</i>	Gravity tries to collapse the star but enormous pressure of fusion energy expands and balances the inward force.



Stars the same size as our Sun.

Red giant	<i>A large star that fuses Helium into heavier elements</i>	Hydrogen runs out, star becomes unstable, pressure inside drops causing star to collapse. Atoms now closer together results in atoms fusing and temperature increases. This increase in temperature causes the core to swell.
White dwarf	<i>Star collapses</i>	Nuclear fuel runs out, fusion stops, dense very hot core.
Black dwarf	<i>Cold dark star</i>	White dwarf cools down.

Stars larger than our Sun.

Red super giant	<i>Star swells greatly</i>	Nuclear fuel begins to run out and star swells (more matter = bigger size).
Supernova	<i>Gigantic explosion due to run away fusion reactions</i>	Rapid collapse, heats to very high temperatures causing run away nuclear reactions, star explodes, flinging remnants out into space. Large gravitational forces collapse the core into a tiny space. Remains of supernova form heavier elements (Iron and above)
Neutron star	<i>Very dense star</i>	Made out of neutrons.

OR if collapse is into a really tiny space.

Black hole

No light escapes

Gravitational forces so strong everything is pulled in.

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Red shift

Understanding models.

Red-shift	<i>The observed increase in wavelength of light from most distant galaxies. Light moves towards the red end of the spectrum.</i>
Hubble (1929)	<i>He studied light from distant galaxies; found as frequency decreases, wavelength increases.</i>
	Light from star in our galaxy.
	Light from star in nearby galaxy.
	Light from star in distant galaxy.
The Big Bang	<i>Universe began 13.8 billion years ago</i>
All matter and space expanded violently from a single point.	
Red—shift provides evidence for expansion.	

Galaxies are moving away from us in all directions.

Light from distant galaxies is red-shifted, so galaxy is moving away from us.

Galaxies further away have bigger red-shift so are moving faster away.

Aristotle (ancient Greek)	<i>Earth at the centre, other heavenly bodies move around the Earth.</i>
Copernicus (1473 - 1543)	<i>Sun at the centre, other heavenly bodies move around the Sun.</i>
Galileo (1610)	<i>Made a telescope, looked at Jupiter, found four moons rotating around planet.</i>

Planets and moons moved at different speeds to stars = reason for different positions.