



# Self Supervised Learning Methods for Imaging

## Part 1: Introduction

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# Tutorial Schedule

## PART I: Introduction to Imaging Inverse Problems

Inverse problem framework; ill-posed problems; dimensionality; noise; deep learning solutions; supervised versus unsupervised; learning vs inductive bias.

## PART II: Unsupervised methods for Invertible Forward Operator

Blind Denoising; SURE; Noise2X; denoising autoencoders; measurement splitting; GSURE.

## PART III: Learning from multiple operators

The impossibility of learning from an incomplete measurement operator; learning from multiple operators; Noise2Noise revisited; measurement splitting revisited; multi-operator consistency; handling noise.

## PART IV: Unsupervised methods for ill-conditioned inverse Problems with a single operator

Exploiting Invariance and symmetries; equivariant and non-equivariant operators; enforcing equivariance; handling noise

## PART V: Identifiability Theory

Identifiability and dimension; learning with noise; learning from incomplete measurements; Cramer-Wold theorem; generic identifiability

## PART VI: Summary and Future Perspectives

# The Inverse problem

**Goal:** estimate signal  $x$  from  $y$


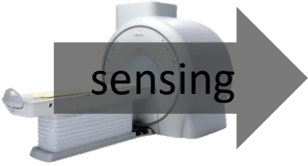
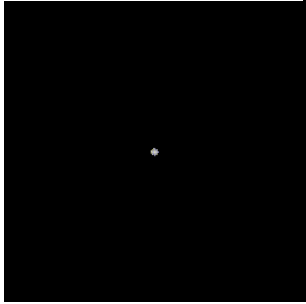

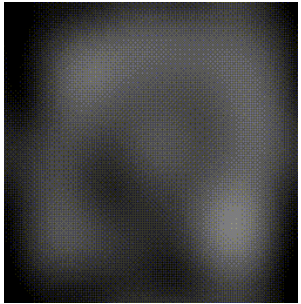

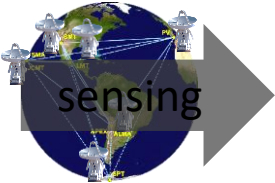
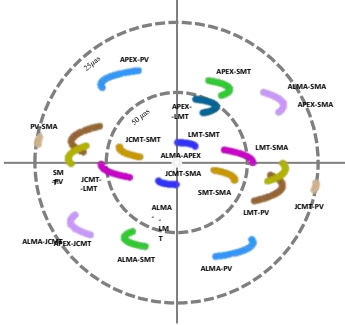

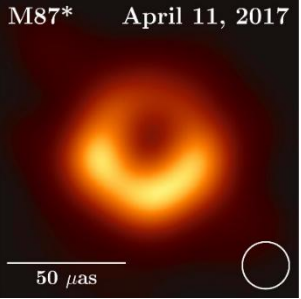





The diagram shows the equation  $y = A(x) + \epsilon$  with several handwritten red annotations. An arrow points from the text "measurements  $\in \mathbb{R}^m$ " to the variable  $y$ . Another arrow points from the text "signal  $\in \mathbb{R}^n$ " to the variable  $x$ . A third arrow points from the text "noise/error" to the variable  $\epsilon$ . A fourth arrow points from the word "Physics" to the operator  $A$ .

$$\text{measurements } \in \mathbb{R}^m \rightarrow y = A(x) + \epsilon \leftarrow \text{noise/error}$$

$\uparrow$   
Physics

We will focus on linear problems where the forward operator  $A$  is a matrix

# Examples

	$x$	$A$	$y$	reconstruction	
<b>Magnetic Resonance Imaging (MRI)</b> $A$ : undersampled Fourier models					 <p>Source: Brian Hargreaves</p>
<b>Black Hole Imaging</b> $A$ : spatial-frequency e.g. Event Horizon Telescope (EHT)					 <p>The Astrophysical Journal Letters, vol. 875, no. L1, 2019.</p>
<b>Cryogenic electron microscopy (Cryo-EM)</b> $A$ : 2D projections of protein particles					 <p>Covid-19 virus' structure            D. Wrapp et al. <i>Science</i>,            vol. 367, no. 6483, 2020.</p>

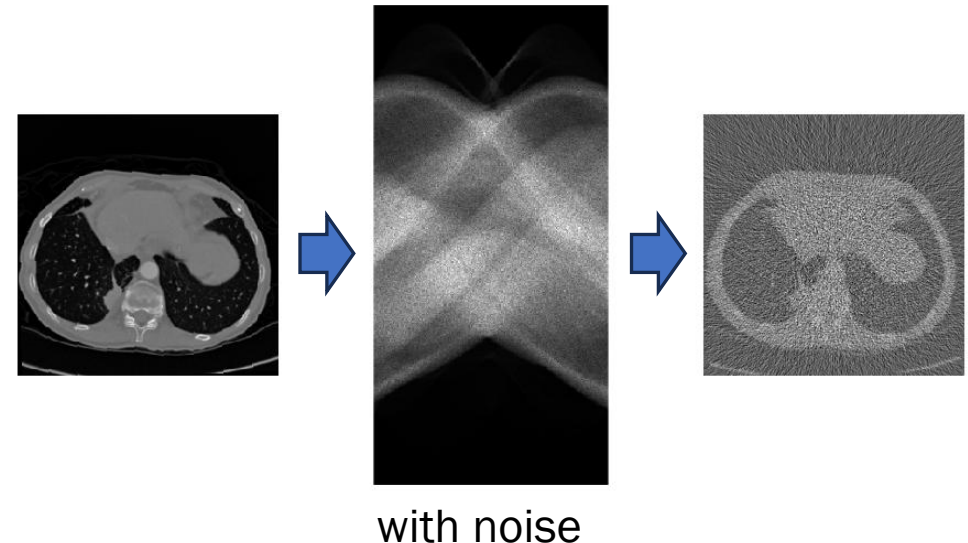
# Why it is hard to invert?

Measurements are usually corrupted by noise, e.g.

$$y = Ax + \epsilon$$

Can be additive, as above, or more complex, e.g. Poisson.

- Often, we do not know the exact noise distribution
- The forward operator may be poorly conditioned



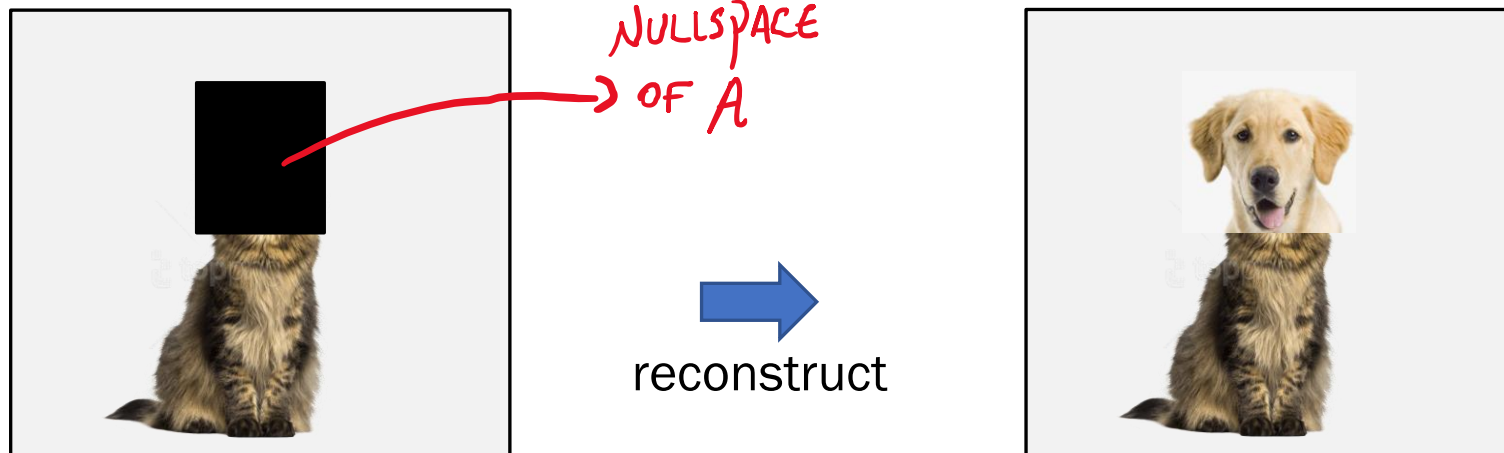
# Why it is hard to invert?

Even in the absence of noise,  $A$  may not be invertible, giving infinitely many  $\hat{x}$  consistent with  $y$ :

$$\hat{x} = A^\dagger y + v$$

where  $A^\dagger$  is the pseudo-inverse of  $A$  and  $v$  is any vector in nullspace of  $A$

*Unique solution only possible if set of signals  $x$  is low-dimensional*



# Low Dimensional Signal Models

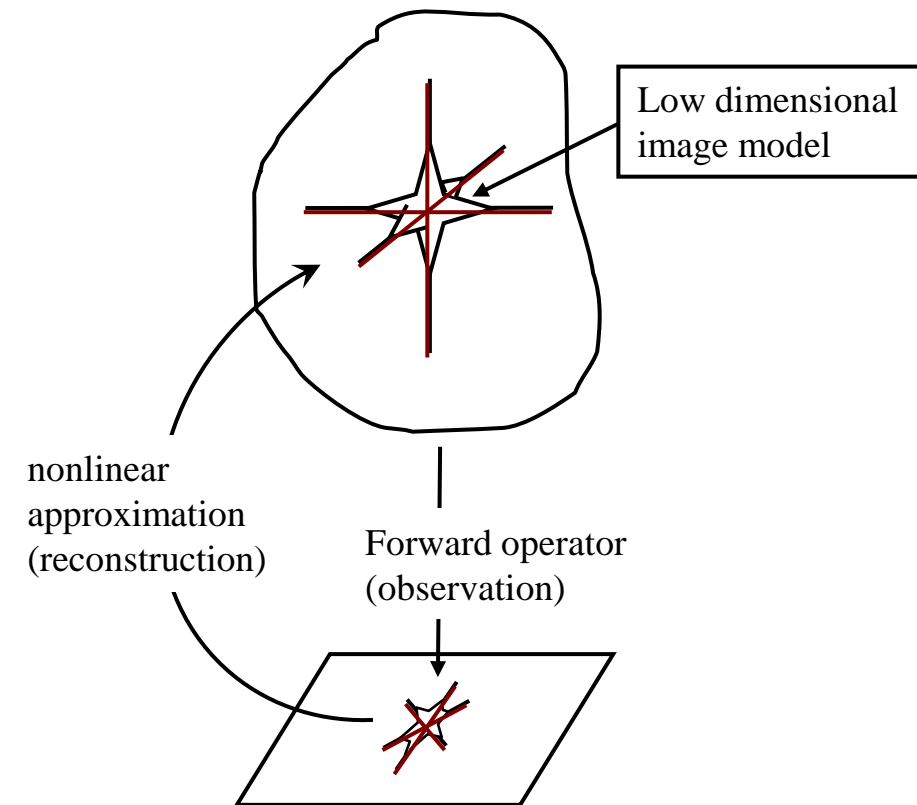
**Idea:** assume approximate low dimensional image model:

$$\dim \mathcal{X} = k \ll n$$

**Examples:** sparsity, low-rank, manifolds

**Signal Embedding:** if  $m \geq \mathcal{O}(k)$  then the problem is approximately one-to-one and (nonlinearly) invertible

This is the principle behind compressed sensing, but is implicit in most inverse imaging problems



# Regularised reconstruction

**Idea:** define a loss  $\rho(\mathbf{x})$  that promotes plausible reconstructions

$$\hat{\mathbf{x}} = \operatorname{argmin}_{\mathbf{x}} \left\| \mathbf{y} - A\mathbf{x} \right\|^2 + \rho(\mathbf{x})$$

**Examples:** total-variation, sparsity, etc.

**Disadvantages:** hard to define a good  $\rho(\mathbf{x})$  in real world problems, loose with respect to the true signal distribution





# Learning approach

## Advantages:

- State-of-the-art reconstructions
- Once trained,  $f_\theta$  is easy to evaluate

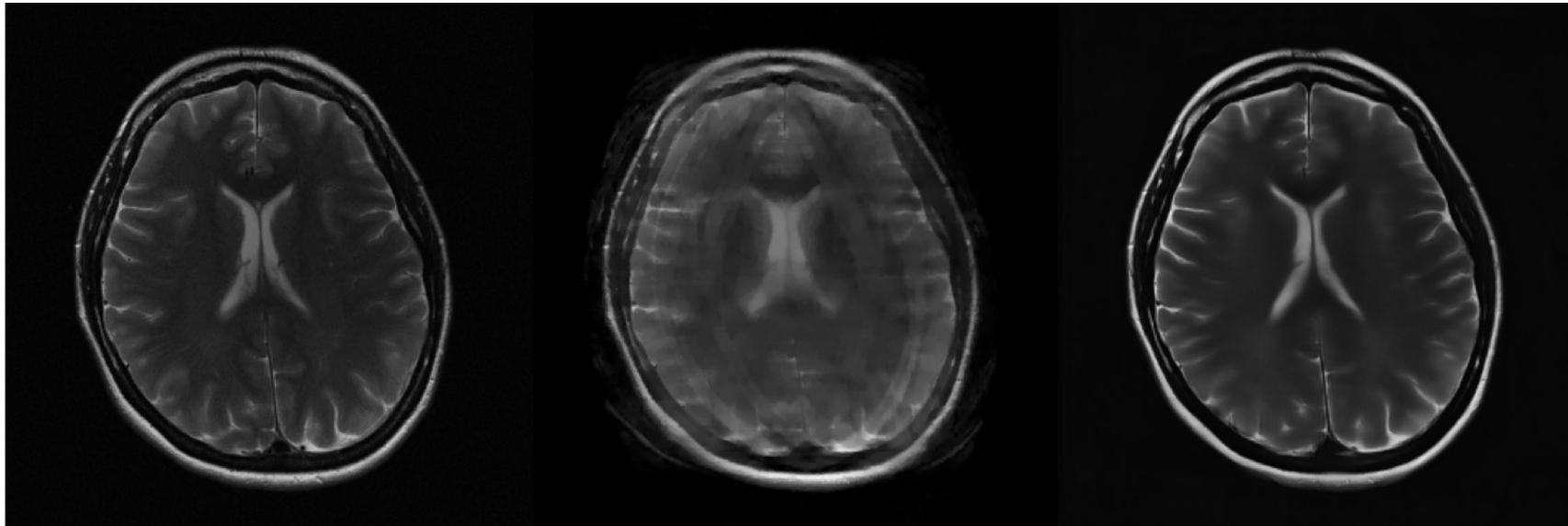
**fastMRI**

Accelerating MR Imaging with AI

Ground-truth

Total variation  
(28.2 dB)

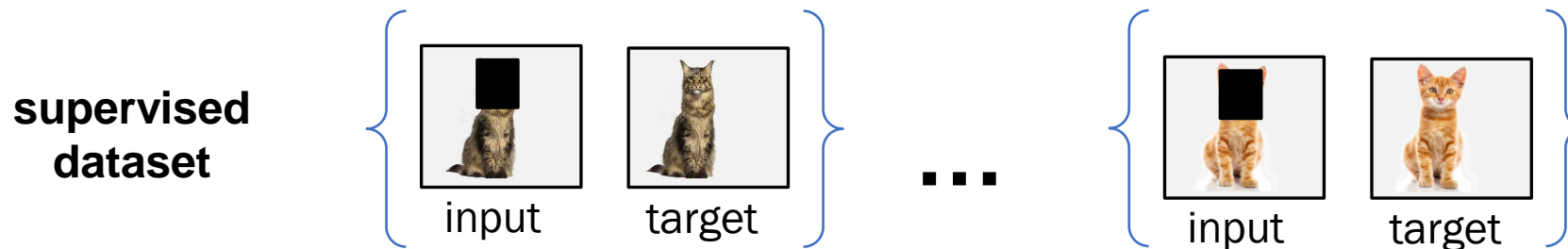
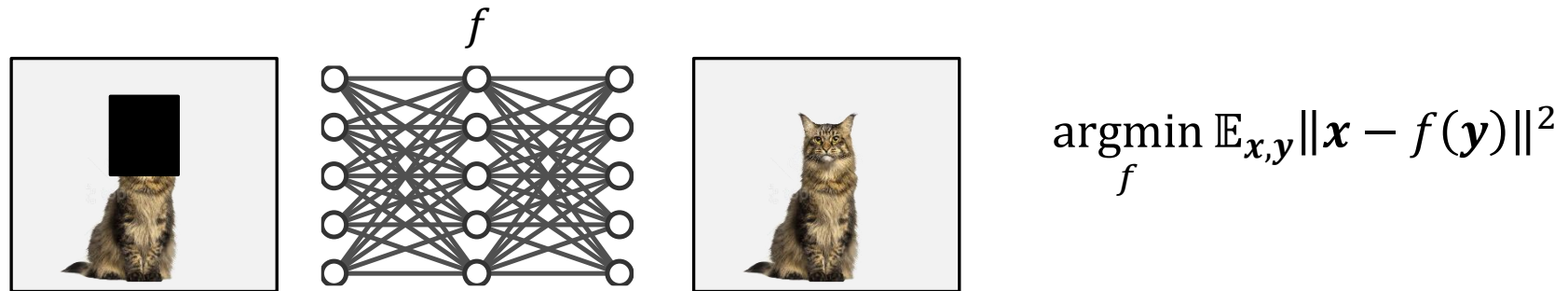
Deep network  
(**34.5 dB**)



x8 accelerated MRI [Zbontar et al., 2019]

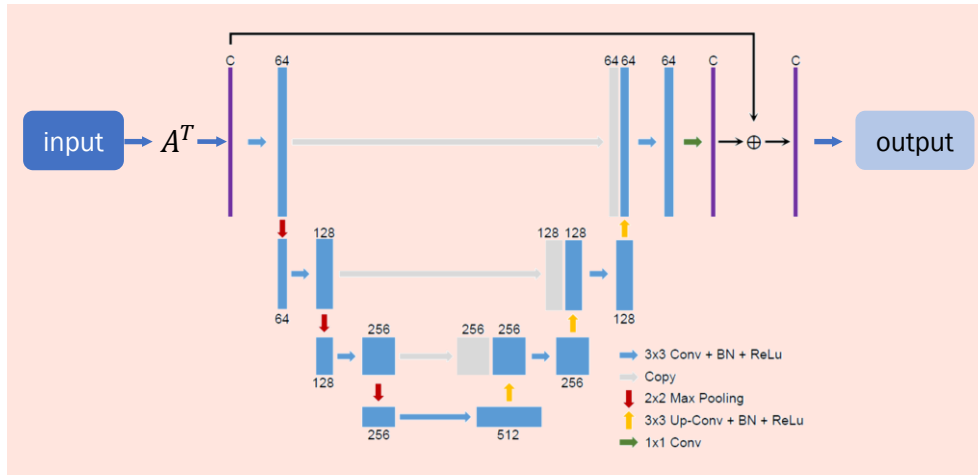
# Learning approach

**Idea:** use training pairs of signals and measurements to directly learn the inversion function

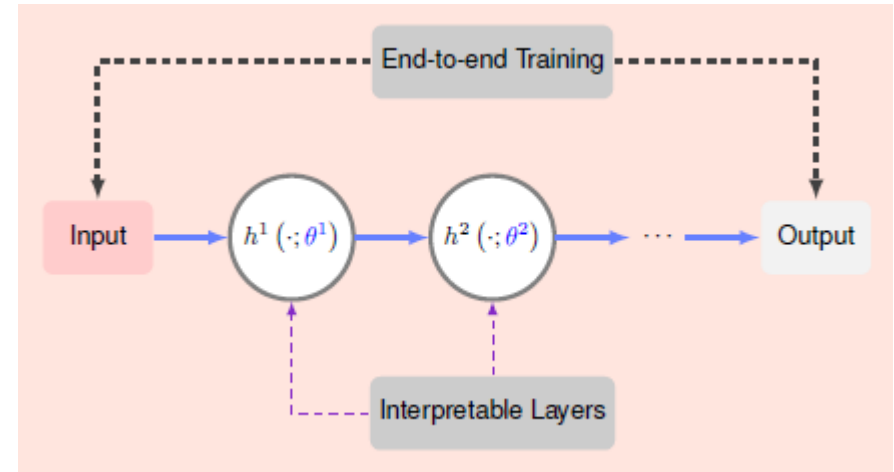


# Learning approach

Many DNN architecture choices, e.g.



Back projected U-Net:  $\hat{x} = f(A^T y)$ , e.g. [Jin, 2017]



Unrolled networks:  $\hat{x} = f(y, A)$ , e.g. [Monga, 2020]

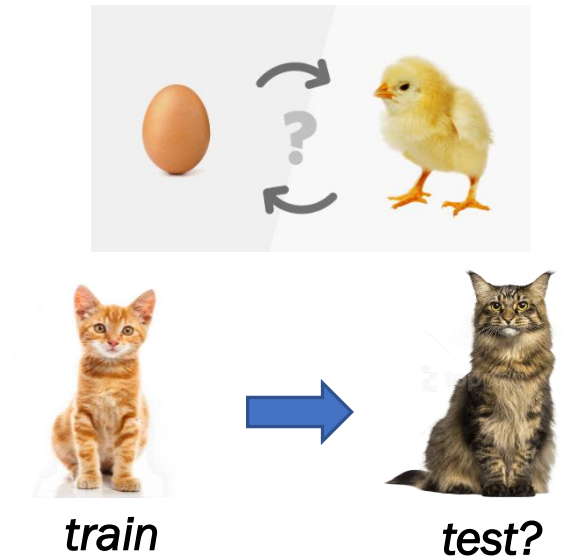
But also DnCNNs, DRUNet, SCUNet, DEQ, restormer, SwinIR, DiffPIR...

Here our focus will be on learning that is typically **architecture agnostic**

# Learning approach

**Main disadvantage:** reference data can be expensive or impossible to get.

- Medical and scientific imaging
- Problems which we already 'solved'
- Distribution shift

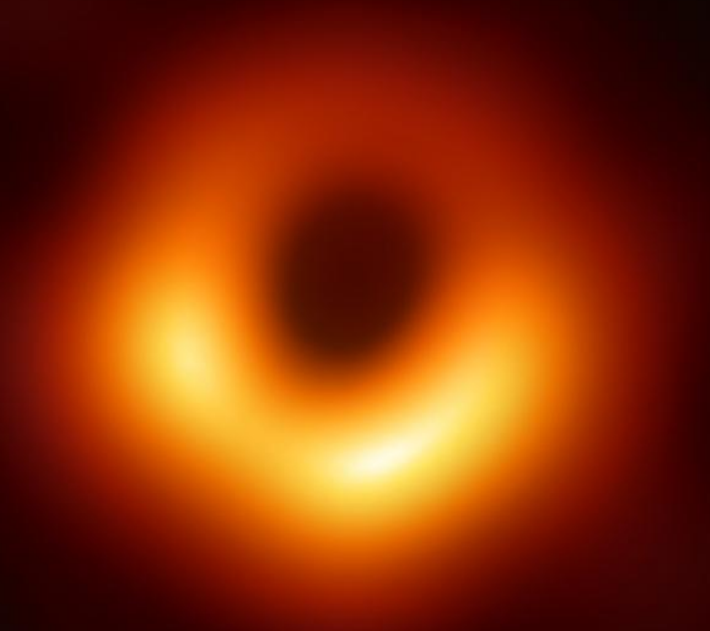


- **Raises the question:**

Can AI be used for data-driven knowledge discovery in imaging?



# AI for Knowledge Discovery?



**The  
Guardian**

Black hole picture captured for first time in space breakthrough



**The  
Guardian**

DeepMind uncovers structure of 200m proteins in scientific leap forward



# Learning vs Inductive Bias

**Inductive bias:** all networks carry inductive bias

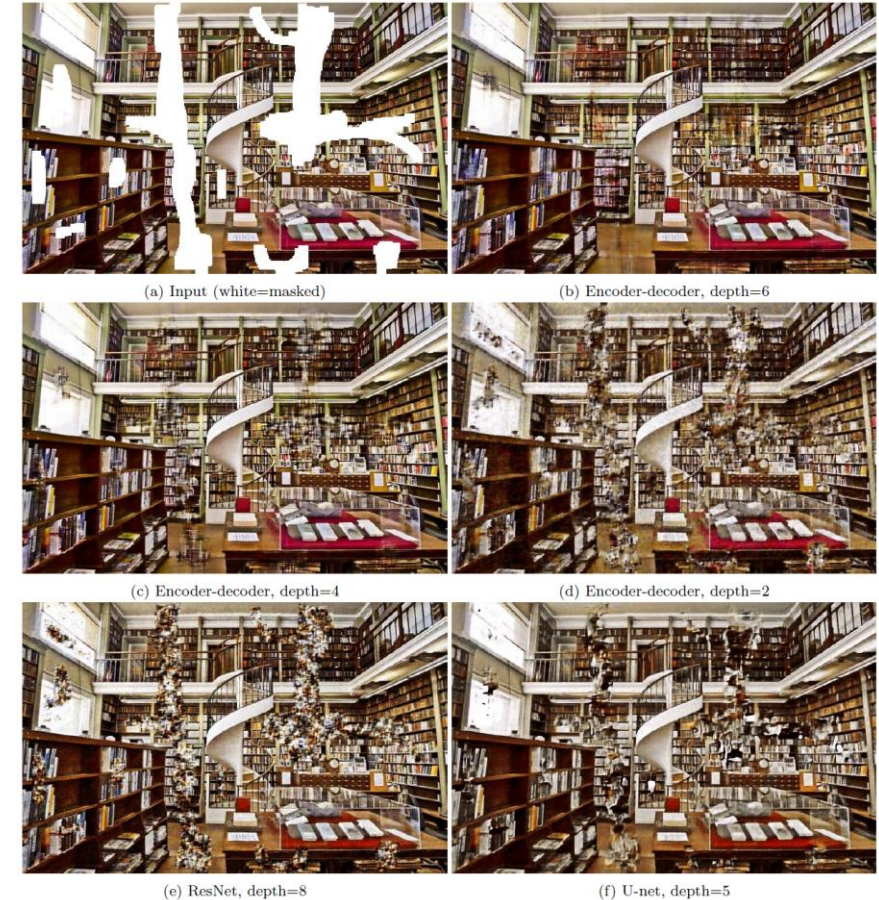
## Deep Image Prior [Ulyanov 2018]

- Architecture *implicitly* encodes “preferred” images
- Minimise data consistency with early stopping

$$\operatorname{argmin}_f \| \mathbf{y} - Af(\mathbf{y}) \|^2$$

## Hard to categorise:

- Inductive bias?
- Learning non-local structure [Tachella 2021]?



Deep Image Prior [Ulyanov 2018]

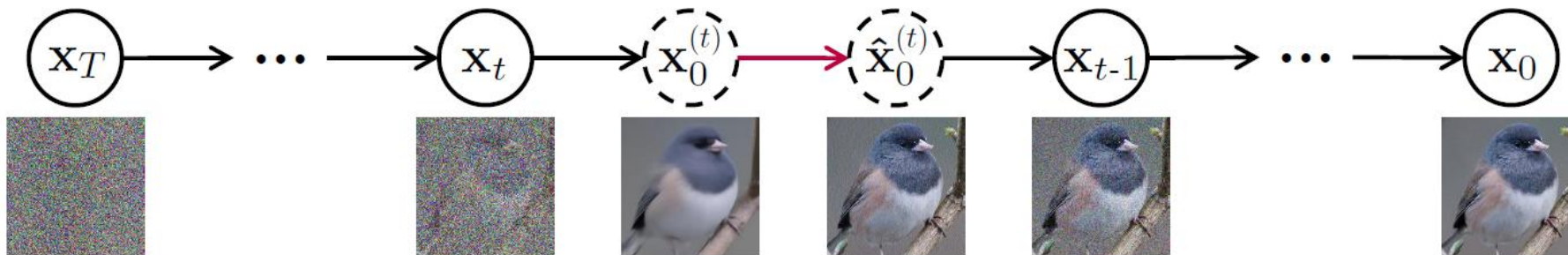
# Learning vs Pre-Trained

**Exploiting Deep Pre-Trained Denoisers:** Exploit powerful pretrained DNN denoisers,  $D(\mathbf{u}, \sigma)$ , e.g., [Zhu, 2023]

**Plug and Play methods:** e.g. PnP proximal gradient descent

$$\mathbf{u}_{k+1} = D(\mathbf{x}_k - \gamma A^T(A\mathbf{x}_k - \mathbf{y}), \sigma)$$

**Conditional Diffusion models:** use pre-trained denoiser in reverse SDE to attempt to sample conditional distribution



Generally pre-trained but can leverage self-supervised denoisers (see part III)

# References

The full reference list for this tutorial can be found here:

<https://tachella.github.io/projects/selfsuptutorial/>

