

Lecture 15. Quasisymmetry

• Symmetry and conservation

$A = \oint \sqrt{2m(H - \mu B)} d\ell \rightarrow$ Quasi-symmetry in $B \equiv |\vec{B}|$ allows magnetic confinement even without symmetry in actual \vec{B}

• Radial drift and Isodynamics

$$H = \frac{1}{2} m v_{||}^2 + \mu B + e\phi = \text{conserved} \Rightarrow \vec{V}_1 \equiv \vec{X} = \frac{v_{||} (\vec{B} + \vec{\nabla} \times (\rho_{||} \vec{B}))}{B + \hat{b} \cdot \vec{\nabla} \times (\rho_{||} \vec{B})} \simeq v_{||} \hat{b} + \underbrace{\frac{v_{||}}{B} \vec{\nabla} \times (\rho_{||} \vec{B})}_{\vec{v}_d}$$

(drift across field lines $\equiv \vec{v}_d$)

$$\vec{v}_d \equiv \vec{V}_d = \frac{v_{||}}{B} \vec{\nabla} \times (\rho_{||} \vec{B}) = \rho_{||} \vec{\nabla} \times (v_{||} \hat{b}) = \rho_{||} (v_{||} \vec{\nabla} \times \hat{b} + \vec{\nabla} v_{||} \times \hat{b})$$

$$= \rho_{||} \left[v_{||} \vec{\nabla} \times \hat{b} - \frac{1}{v_{||}} \left(\frac{\mu}{m} \vec{\nabla} B - \frac{e}{m} \vec{\nabla} \phi \right) \times \hat{b} \right]$$

$$\left(\text{where } \vec{\nabla} v_{||} = \pm \vec{\nabla} \sqrt{\frac{2}{m} (H - \mu B - e\phi)} = \pm \frac{1}{2 v_{||}} \vec{\nabla} \left(\frac{\mu}{m} (H - \mu B - e\phi) \right) = \frac{1}{v_{||}} \left(\frac{\mu}{m} \vec{\nabla} B - \frac{e}{m} \vec{\nabla} \phi \right) \right)$$

(radial drift)

$$\vec{v}_d \cdot \vec{\nabla} \psi = \rho_{||} v_{||} (\vec{\nabla} \times \hat{b}) \cdot \vec{\nabla} \psi + \frac{\rho_{||}}{v_{||}} \left[\frac{\mu}{m} (\hat{b} \times \vec{\nabla} B) \cdot \vec{\nabla} \psi \right]$$

$$\left(\begin{aligned} \text{note, } \vec{\nabla} \times \hat{b} &= \vec{\nabla} \times \left(\frac{\vec{B}}{B} \right) = \frac{\vec{\nabla} \times \vec{B}}{B} + \vec{\nabla} \left(\frac{1}{B} \right) \times \vec{B} = \frac{1}{B} \vec{\nabla} \times \vec{B} - \frac{\vec{\nabla} B \times \vec{B}}{B^2} \\ \hat{b} \times \vec{\nabla} \ln B &= \frac{\vec{B}}{B} \times \frac{\vec{\nabla} B}{B} = \frac{-\vec{\nabla} B \times \vec{B}}{B^2} \rightarrow \vec{\nabla} \times \hat{b} = \frac{1}{B} \vec{\nabla} \times \vec{B} + \hat{b} \times \vec{\nabla} \ln B \\ &= \frac{1}{B} (\hat{b} \times \vec{\nabla} B) \end{aligned} \right)$$

$$\begin{aligned} \therefore \vec{v}_d \cdot \vec{\nabla} \psi &= \frac{m v_{||}^2}{e B} \left[\frac{1}{B} \vec{\nabla} \times \vec{B} + \hat{b} \times \vec{\nabla} \ln B \right] \cdot \vec{\nabla} \psi + \frac{m}{e B} \frac{v_{||}^2}{2 B} B (\hat{b} \times \vec{\nabla} \ln B) \cdot \vec{\nabla} \psi \\ &= \frac{m v_{||}^2}{e B} (\hat{b} \times \vec{\nabla} \ln B) \cdot \vec{\nabla} \psi + \frac{m v_{||}^2}{2 e B} (\hat{b} \times \vec{\nabla} \ln B) \cdot \vec{\nabla} \psi = \frac{m (v_{||}^2 + \frac{1}{2} v_{||}^2)}{e B} (\hat{b} \times \vec{\nabla} \ln B) \cdot \vec{\nabla} \psi \end{aligned}$$

\therefore isodynamic condition of equilibrium (no radial drift at all, $\vec{v}_d \cdot \vec{\nabla} \psi = 0$)

$$\boxed{\vec{\nabla} \psi \times (\hat{b} \cdot \vec{\nabla} \ln B) = 0}$$

in low ρ , isodynamic condition \Leftrightarrow zero geodesic curvature $\vec{k}_g = (\hat{n} \times \hat{b}) \cdot \vec{K} = 0$
 $\left(\hat{n} = \frac{\vec{\nabla} \psi}{|\vec{\nabla} \psi|}, \vec{K} \simeq \vec{\nabla} \ln B \right)$

• Omnigenity

Isodynamic : Ideal but impossible to meet everywhere

Omnigenity : $\langle \vec{v}_d \cdot \vec{\nabla} \psi \rangle_b = 0$

where bounce average : $\langle A \rangle_b = \frac{\int A dt}{\int dt} = \frac{\int A \frac{dl}{v_{||}}}{\int \frac{dl}{v_{||}}} = w_b \int A \frac{dl}{v_{||}} \quad (w_b = \tau^{-1} = \frac{1}{\int \frac{dl}{v_{||}}})$

• Action variation

$$\vec{B} = \vec{\nabla} \psi \times \vec{\nabla} (\vartheta - \frac{N}{M} \varphi) = \vec{\nabla} \psi \times \vec{\nabla} \alpha \quad (\alpha = \vartheta - \frac{N}{M} \varphi)$$

one can make non-flux-coordinate (ψ, α, θ) where $\theta = \vartheta - \frac{N}{M} \varphi$.

Then, $B^\theta = \vec{B} \cdot \vec{\nabla} \theta = (\vec{\nabla} \psi \times \vec{\nabla} \alpha) \cdot \vec{\nabla} \theta = \frac{1}{\sqrt{g}}$.

$$B_\theta = \vec{B} \cdot \vec{e}_\theta = (\vec{\nabla} \psi \times \vec{\nabla} \alpha) \cdot \sqrt{g} (\vec{\nabla} \psi \times \vec{\nabla} \alpha) = \sqrt{g} B^2$$

Expand radial drift to this (ψ, α, θ) coordinate.

$$\vec{v}_d = \frac{v_{||}}{B} \vec{\nabla} \times (p_{||} \vec{B}) = \frac{v_{||}}{\sqrt{g} B} \left(\frac{\partial}{\partial \alpha} (p_{||} B_\theta) - \frac{\partial}{\partial \theta} (p_{||} B_\alpha) \right) \vec{e}_\psi + (\dots) \vec{e}_\alpha + (\dots) \vec{e}_\theta$$

$$\vec{v}_d \cdot \vec{\nabla} \psi = \frac{v_{||}}{\sqrt{g} B} \left(\frac{\partial}{\partial \alpha} (p_{||} \sqrt{g} B^2) - \frac{\partial}{\partial \theta} (p_{||} B_\alpha) \right)$$

note, $\langle A \rangle_b = w_b \int A \frac{dl}{v_{||}} = w_b \int A \frac{B}{v_{||}} \frac{d\theta}{B\theta} = w_b \int d\theta \frac{\sqrt{g} B}{v_{||}} A \quad (dl = \sqrt{g} B d\theta)$

$$\therefore \langle \vec{v}_d \cdot \vec{\nabla} \psi \rangle_b = w_b \int d\theta \left(\frac{\partial}{\partial \alpha} (p_{||} \sqrt{g} B^2) - \frac{\partial}{\partial \theta} (p_{||} B_\alpha) \right)$$

$$= w_b \frac{\partial}{\partial \alpha} \int d\theta \left(\frac{mv_{||}}{eB} \sqrt{g} B^2 \right) = \frac{w_b}{e} \frac{\partial}{\partial \alpha} \int mv_{||} dl = \frac{w_b}{e} \frac{\partial J}{\partial \alpha} \quad (J = \int mv_{||} dl)$$

\Rightarrow average radial drift is proportional to the variation in the action $J = \int mv_{||} dl$

$$\therefore \langle \vec{v}_d \cdot \vec{\nabla} \psi \rangle_b = \frac{w_b}{e} \frac{\partial J}{\partial \alpha}, \quad \text{similarly} \quad \langle \vec{v}_d \cdot \vec{\nabla} \alpha \rangle_b = -\frac{w_b}{e} \frac{\partial J}{\partial \psi}$$

$$\langle \vec{v}_d \cdot \vec{\nabla} \rangle_b J = \langle \vec{v}_d \cdot \vec{\nabla} \psi \rangle_b \frac{\partial J}{\partial \psi} + \langle \vec{v}_d \cdot \vec{\nabla} \alpha \rangle_b \frac{\partial J}{\partial \alpha} = 0 \quad \neq \text{2nd adiabatic invariance}$$

Omnigenity and Quasisymmetry

omnigenity $\left\{ \begin{array}{l} \text{radial orbit drift vanishes} \quad \langle \vec{v}_d \cdot \vec{\nabla} \psi \rangle_b = 0 \\ \text{no action variation on flux surfaces} \quad dJ/d\alpha = 0 \end{array} \right\}$

$$\langle \vec{v}_d \cdot \vec{\nabla} \psi \rangle_b = w_b \int d\theta \frac{d}{d\alpha} (p_{||} \sqrt{g} B^2)$$

If \sqrt{g} depends only on B , the entire \int will depend only on B .

\Rightarrow Boozer coordinate !!

That is; If B is independent of α on Boozer coordinates,

$$B = B(\psi, \theta) = B(\psi, M\theta_B - N\psi_B),$$

the omnigenity will be automatically achieved (Quasi-symmetry)

$$(M, N) = (1, 0) \quad (QA)$$

$$= (1, \pm N \cdot mfp) \quad (QH) \quad \Leftarrow \text{Type of quasi-symmetry}$$

$$= (0, 1) \quad (QP)$$

Basic principle: Particle will hang on a surface if there is a way to conserve its action

$$\langle \vec{v}_d \cdot \vec{\nabla} \rangle_b J = \langle \vec{v}_d \cdot \vec{\nabla} \psi \rangle_b \frac{dJ}{d\psi} + \langle \vec{v}_d \cdot \vec{\nabla} \alpha \rangle_b \frac{dJ}{d\alpha} = 0$$

If $\frac{dJ}{d\alpha} = 0 \rightarrow$ then, to conserve action, $\frac{dJ}{d\psi}$ must be zero.

Isomorphism

Tokamaks and stellarators are indistinguishable to the leading order for particles.

$$L_{gc} = (m v_{||} \hat{b} + e \vec{A}) \cdot \dot{\vec{X}} + \frac{m}{e} \dot{\varphi} - H_{gc} = e (\rho_{||} \vec{B} + \vec{A}) \cdot \dot{\vec{X}} + \frac{m}{e} \dot{\varphi} - H_{gc}$$

Take Boozer coordinates, $\theta = \theta_B - \frac{N}{M} \varphi_B$, and assume quasi-symmetry.

Then, $B = B(\psi, \theta)$, $H_{gc} = H_{gc}(\psi, \theta)$.

$$\left(\begin{aligned} \vec{B} &= \vec{\nabla} \psi \times \vec{\nabla} \alpha = \vec{\nabla} \times (\psi \vec{\nabla} \alpha), \Rightarrow \vec{A} = \psi \vec{\nabla} \alpha = \psi \vec{\nabla} \theta_B - \chi \vec{\nabla} \varphi_B \\ \vec{B} &= I \vec{\nabla} \theta_B + G \vec{\nabla} \varphi_B + \psi \vec{\nabla} \psi, \quad \vec{X} = \dot{\theta}_B \vec{e}_{\theta_B} + \dot{\varphi}_B \vec{e}_{\varphi_B} + \dot{\psi} \vec{e}_{\psi} \end{aligned} \right)$$

$$\begin{aligned} \Rightarrow L_{gc} &= e (\rho_{||} (I \vec{\nabla} \theta_B + G \vec{\nabla} \varphi_B) + \psi \vec{\nabla} \theta_B - \chi \vec{\nabla} \varphi_B) \cdot (\dot{\theta}_B \vec{e}_{\theta_B} + \dot{\varphi}_B \vec{e}_{\varphi_B} + \dot{\psi} \vec{e}_{\psi}) + \frac{m}{e} \dot{\varphi} - H_{gc} \\ &= e (\psi + \rho_{||} I) \dot{\theta}_B + e (\rho_{||} G - \chi) \dot{\varphi}_B + \frac{m}{e} \dot{\varphi} - H_{gc} \end{aligned}$$

note that this is a canonical form $L_{gc} = \sum_i p_i \dot{q}_i - H_{gc}$.

$\theta_B \rightarrow \theta$, with $\theta = \theta_B - (\frac{N}{M}) \varphi_B$, then.

$$L_{gc} = e (\psi + \rho_{||} I) \dot{\theta} + e \left(\rho_{||} \left(G + \frac{N}{M} I \right) - \left(\chi - \frac{N}{M} \psi \right) \right) \dot{\varphi}_B + \frac{m}{e} \dot{\varphi} - H_{gc}(\psi, \theta)$$

$$p_{\varphi_B} = \frac{\partial L_{gc}}{\partial \dot{\varphi}_B} = e \left(\rho_{||} \left[G + \frac{N}{M} I \right] - \left[\chi - \frac{N}{M} \psi \right] \right) = \text{const}$$

($\because G, I, \chi, \psi$ are flux function)

$$\left(\begin{aligned} \theta_B &\leftrightarrow \theta_B - \left(\frac{N}{M} \right) \varphi_B \\ G &\leftrightarrow G + \left(\frac{N}{M} \right) I \\ \chi &\leftrightarrow \chi - \left(\frac{N}{M} \right) \psi \end{aligned} \right)$$

Isomorphism

One can simply transform coordinates to make particle trajectory under quasisymmetry same as ones for tokamak. (only for guiding center particle)