

# Local and global models

- We can make models of light-matter interaction in two ways
  - Locally: we treat each object in a scene separately from any other object
  - Globally: we treat all objects together, and model the interactions between objects
- we'll develop a simple local illumination model in detail
- we'll look at global models in COMP37111 next year

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## **Illumination models**

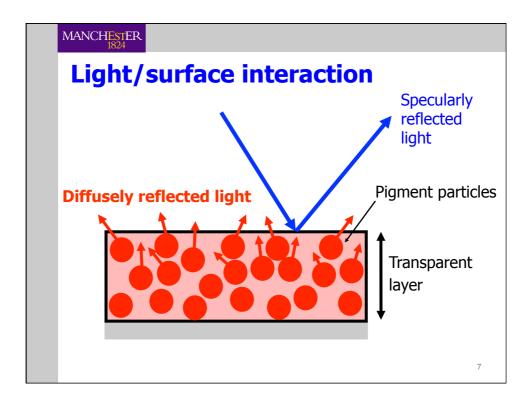
- The interaction of light and matter is an extremely complex process
- In computer graphics we try to model this process. In other words, we approximate it
- Computer pictures are digital, with finite precision. We can only ever approximate.

## **Local illumination: elements**

- We'll develop a model step-by-step, to include the following:
  - Ambient illumination
  - Diffuse reflection
  - Positional light source
  - Specular reflection
  - Colour of lights and surfaces

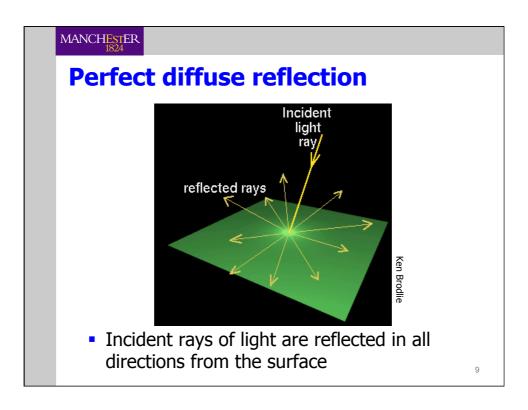
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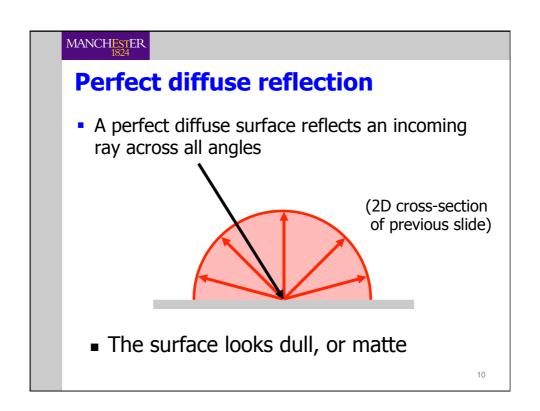
# Reflectivity • There are three kinds of reflection: • Perfect diffuse reflection • Perfect specular reflection • Imperfect specular reflection



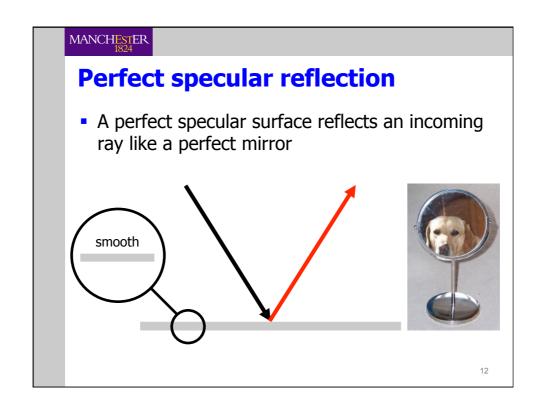
# **Diffuse and specular reflection**

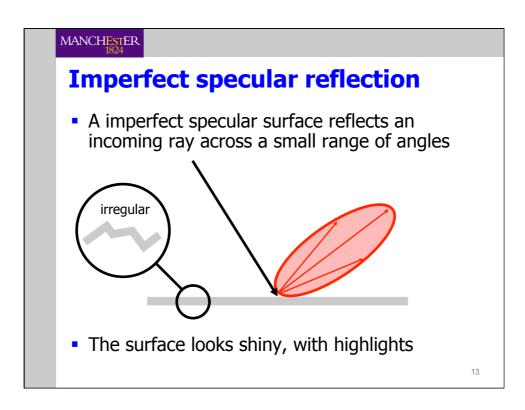
- Diffuse reflection is absorption and uniform reradiation
  - Some wavelengths are absorbed, some are reflected
  - a green object looks green because it only reflects green
- Specular reflection is reflection at the air/ surface interface
  - To a first approximation, the colour of the specular reflection is that of the light source

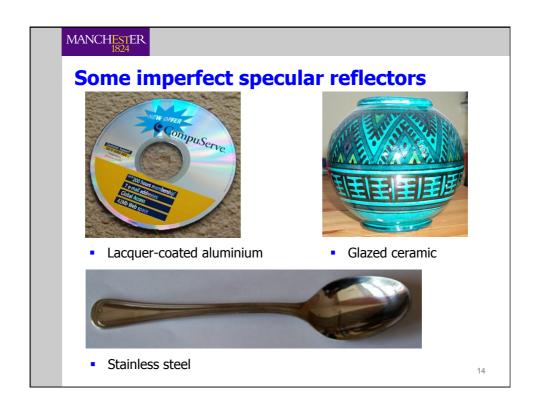






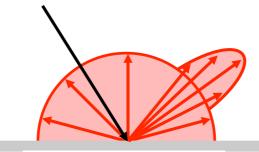






# **Diffuse/specular surfaces**

 Most surfaces exhibit a combination of diffuse and specular reflection



 We can model these effects with varying degrees of realism

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## **Illumination sources**

- We begin our development of the local illumination model by considering diffuse reflection, and sources of illumination:
  - Ambient illumination
  - Point illumination source at infinity (directional illumination)
  - Point illumination source in the scene

# **Ambient illumination**

- Consider an environment with a light source
- Multiple reflections cause a general level of illumination in the scene



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# **Ambient illumination**

- ullet If intensity of ambient light is  $I_a$
- Amount of light diffusely reflected from a surface is

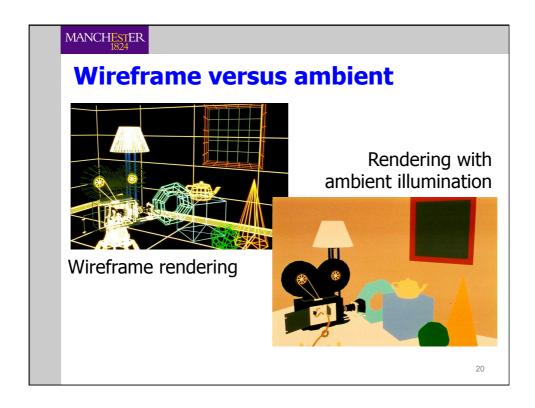
$$I_{\text{ambient}} = k_a I_a$$

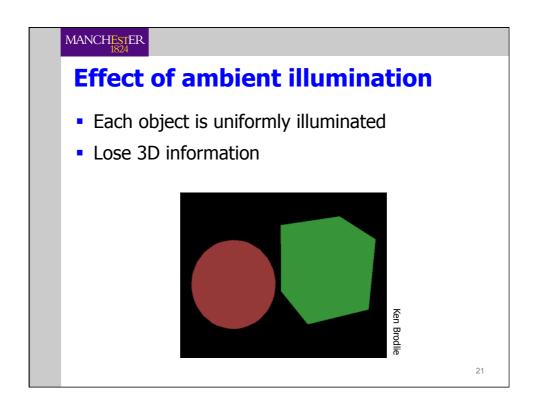
Where k<sub>a</sub> is the ambient reflection coefficient of the surface

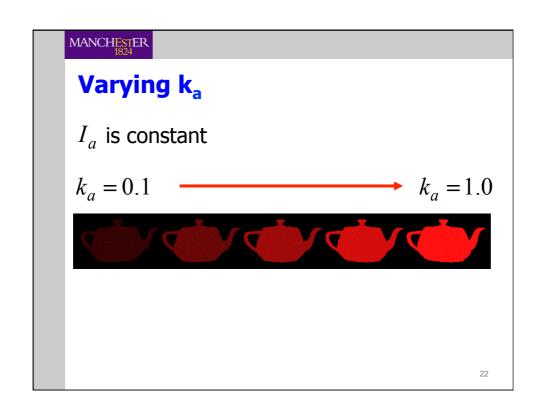
$$0 \le k_a \le 1$$

# Local illumination model v1

- We now have the first term in the model we're developing
- I = ambient
- $I = k_a I_a$





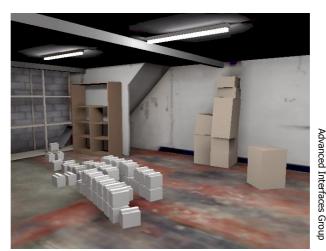


# **True ambient lighting**

Note: the

 $I = k_a I_a$ 

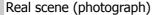
term is a gross simplification of true ambient illumination, which is **not** constant in a scene



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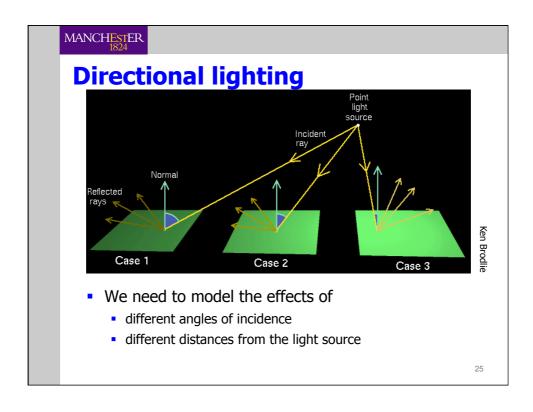
# **True ambient lighting**

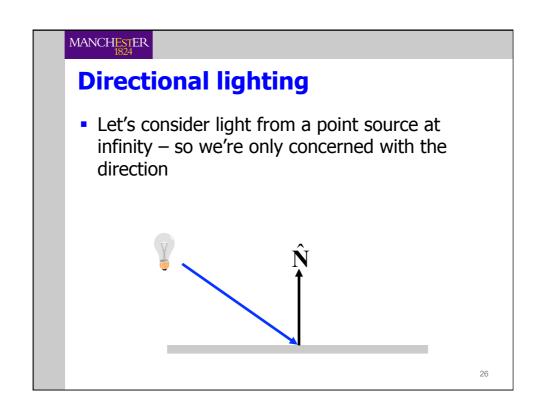




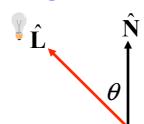


Scene modelled and rendered with accurate global illumination model





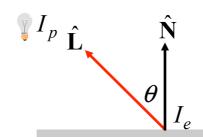
# **Describing surface orientation**



- $\ \ \, \hat{N}$  is surface normal
- $\hat{\mathbf{L}}$  is direction of light source
- ullet heta is angle of incidence
- Note the vectors are normalised

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# **Diffuse reflection: Lambert's Law**



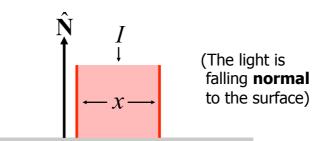
- Light source of intensity  $\boldsymbol{I}_p$  Effective intensity received is  $\boldsymbol{I}_e$
- Lambert's Law:  $I_e = I_p \cos \theta$



Johann Heinrich Lambert (1728-1777)

## Lambert's Law derived

 Consider light of intensity I and cross-sectional width x falling on a surface:



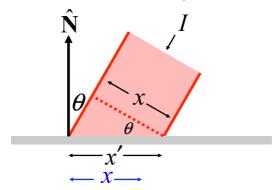
 So width x on surface receives all of intensity I

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# Lambert's Law derived

- Now consider the light inclined at heta
- $x' = x/\cos\theta$ , so  $x = x'\cos\theta$
- So width x receives intensity  $I\cos\theta$



# **Diffuse reflectivity**

- We express how good a diffuse reflector a surface is using k<sub>d</sub>
- $k_d$  is the **diffuse reflection coefficient** of the surface,  $0 \le k_d \le 1$
- Amount of diffusely reflected light is
  - $I_{\text{diffuse}} = I_p k_d \cos \theta$

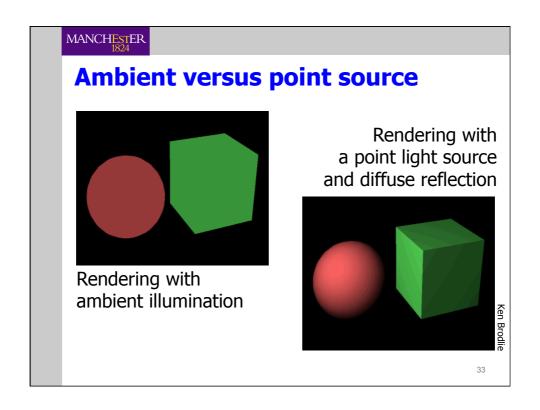
 $I_{\text{diffuse}} = I_p k_d (\hat{\mathbf{N}} \cdot \hat{\mathbf{L}})$ 

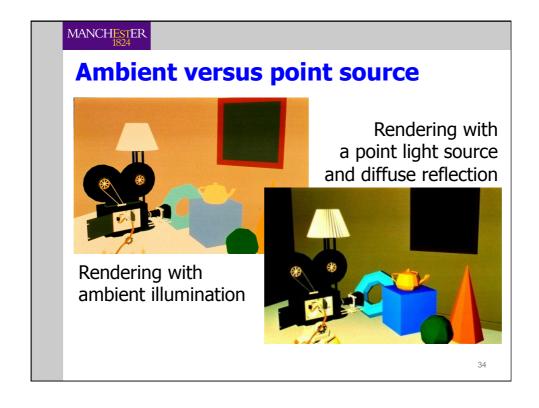
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## Local illumination model v2

- I = ambient + diffuse
- $I = k_a I_a + I_p k_d (\hat{\mathbf{N}} \cdot \hat{\mathbf{L}})$





# Effect of k<sub>d</sub>

Varying k<sub>d</sub> from 0 to 1



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# **Light source distance**

- Physically, light intensity falls off with the square of distance travelled
- After travelling  $\boldsymbol{d}$  , original intensity  $\boldsymbol{I}_p$  is now  $\boldsymbol{I}_e$

$$I_e = \frac{I_p}{4\pi d^2}$$

# Light source distance

- While this is physically correct, it doesn't always work well for computer graphics
- We only have a limited number of pixel intensities, and often the  $d^2$  term changes too rapidly, so instead we use:

$$I_e = \frac{I_p}{\frac{k_c + k_l d + k_q d^2}{k_c + k_l d + k_q d^2}}$$
 We can choose  $\frac{k_c, k_l, k_q}{k_c + k_l d + k_q d^2}$  best results

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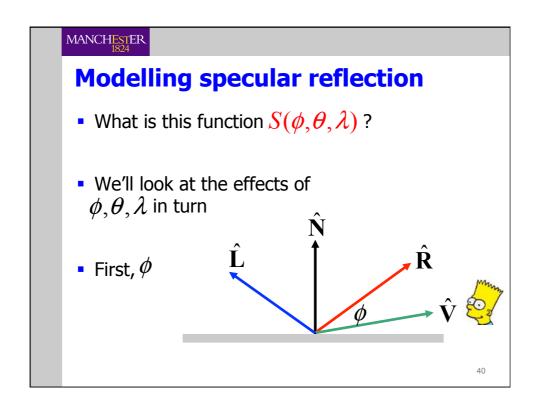
## **Local illumination model v3**

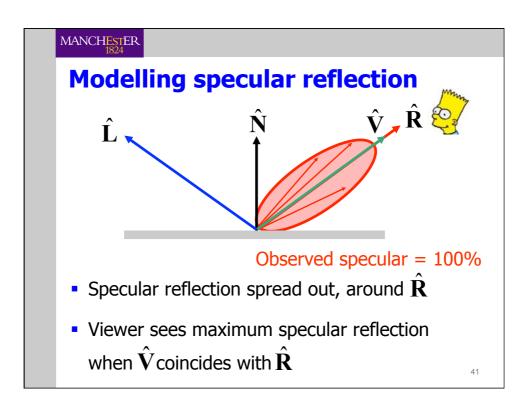
• I = ambient + distance (diffuse)

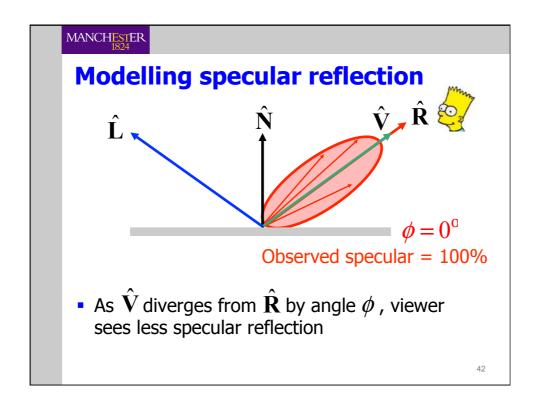
$$I = k_a I_a + \frac{I_p}{d'} k_d (\hat{\mathbf{N}} \cdot \hat{\mathbf{L}})$$

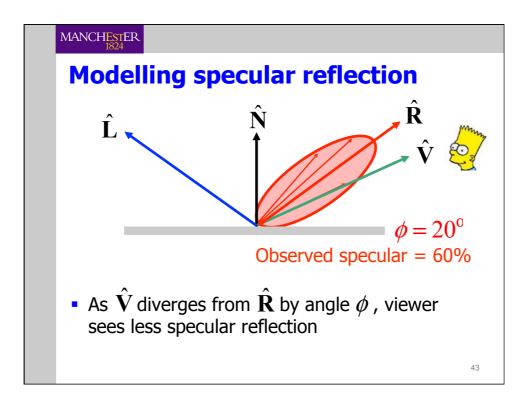
• Where  $d' = k_c + k_l d + k_q d^2$ 

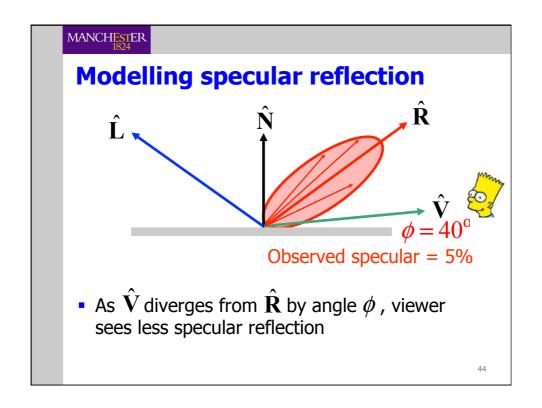
# Modelling specular reflection $\hat{\mathbf{L}} \qquad \hat{\mathbf{N}} \qquad \hat{\mathbf{R}} \qquad \hat{\mathbf{V}} \qquad \hat{\mathbf{$





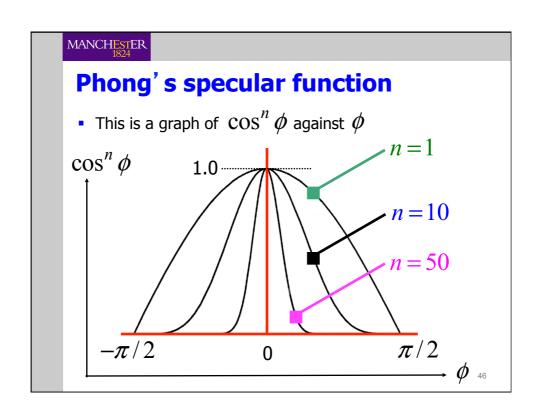






# **Modelling specular reflection**

- Variation of observed specular =  $F(\phi)$
- But what is the function F?
- Bui-Tuong Phong (1942-1975) proposed using the function  $\cos^n \phi$



# Phong's specular function

So we now have

• 
$$I_{\text{specular}} = I_p \cos^n \phi$$

• Which we can rewrite using vectors as

$$I_{\text{specular}} = I_p (\hat{\mathbf{R}} \cdot \hat{\mathbf{V}})^n$$

• Normally we use  $1 \le n \le 200$ 









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$$n = 10$$

n = 20

n = 80

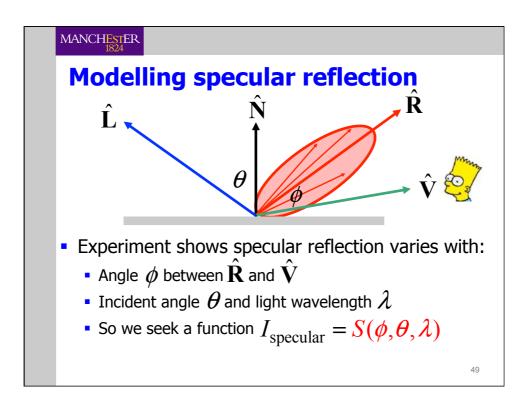
n = 200

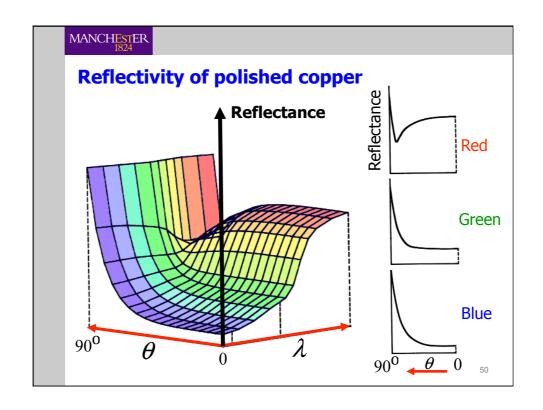
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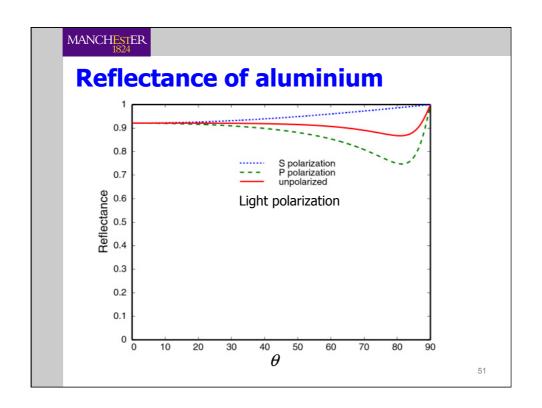
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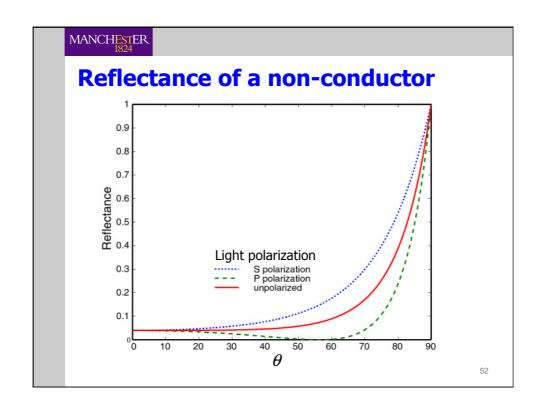
# **Incident angle and wavelength**

- Recall  $I_{\text{specular}} = S(\phi, \theta, \lambda)$
- We've accounted for  $\phi$
- Now we look at the effects of  $\, heta\,$  and  $\,\lambda\,$









# **Incident angle and wavelength**

- This complex variation is expressed by the Fresnel equation (Augustin-Jean Fresnel, 1788-1827)
  - > A founder of the wave theory of light.
  - > Developed theory of diffraction of light.
  - Obtained circularly polarised light
  - Developed the use of compound lenses instead of mirrors for lighthouses.



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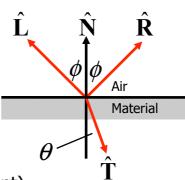
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# **Incident angle and wavelength**

 This complex variation is expressed by the Fresnel equation (Augustin-Jean Fresnel, 1788-1827)

$$F = \frac{1}{2} \left[ \frac{\sin^2(\phi - \theta)}{\sin^2(\phi + \theta)} + \frac{\tan^2(\phi - \theta)}{\tan^2(\phi + \theta)} \right]$$

- *F* is the fraction of light reflected
- $\sin \theta = \sin \phi / \mu$
- $\mu$  is the refractive index of the material (  $\lambda$  dependent)



# **Incident angle and wavelength**

- In practice, we usually approximate F with a single constant  $k_{\scriptscriptstyle S}$
- $k_s$  is the **specular reflection coefficient** of the surface,  $0 \le k_s \le 1$
- $I_{\text{specular}} = I_p \mathbf{k}_s (\mathbf{R} \cdot \mathbf{V})^n$
- But, we sacrifice accuracy for efficiency

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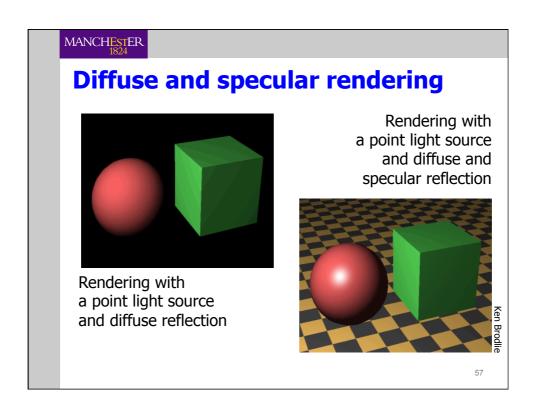
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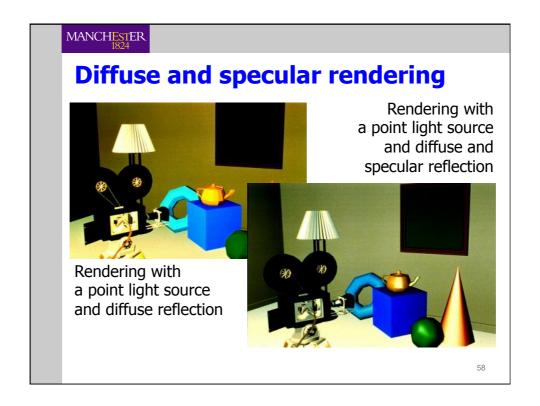
## Local illumination model v4

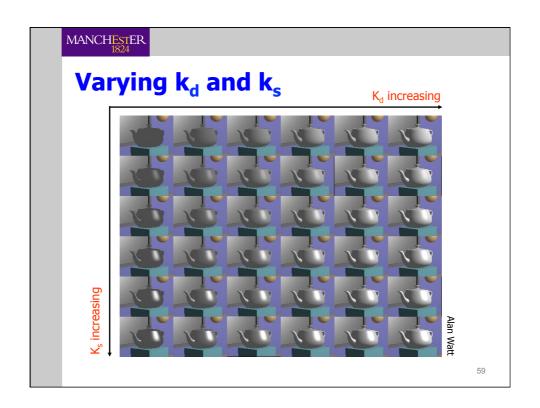
I = ambient + distance (diffuse + specular)

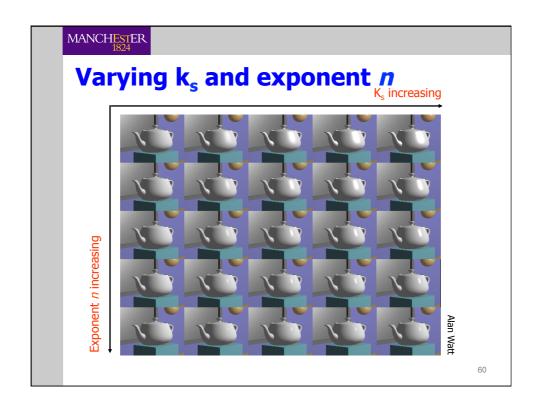
$$I = k_a I_a + \frac{I_p}{d'} \left[ k_d (\hat{\mathbf{N}} \cdot \hat{\mathbf{L}}) + k_s (\hat{\mathbf{R}} \cdot \hat{\mathbf{V}})^n \right]$$

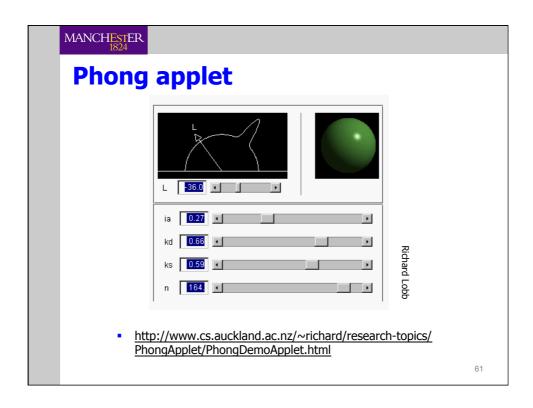
• Where  $d' = k_c + k_l d + k_q d^2$ 

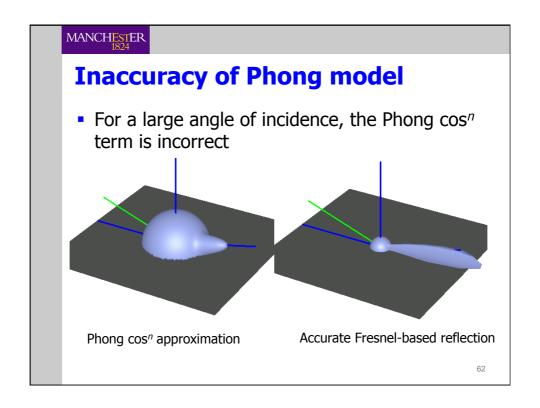












# **Incorporating colour**

- So far, we've only considered light intensity, not colour
- It's easy to incorporate we express light colour as a triple of RGB intensities:
  - $I_{pR}, I_{pG}, I_{pB}$
- And correspondingly we express surface colour using
  - $k_{aR}, k_{aG}, k_{aB}$
  - $k_{dR}, k_{dG}, k_{dB}$

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## **Local illumination model v5**

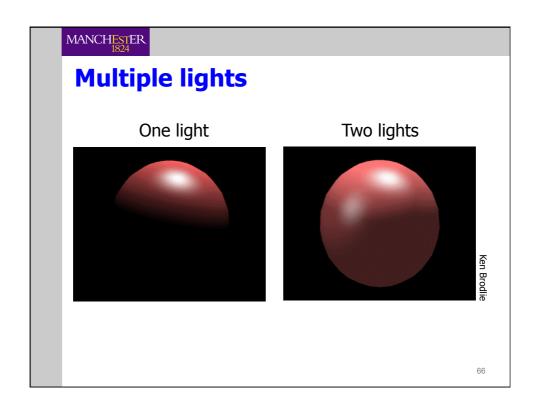
- We now have a separate expression for each colour component. For example, for red:
- $I_R$ = ambient<sub>Red</sub> + distance (diffuse<sub>Red</sub> + specular)

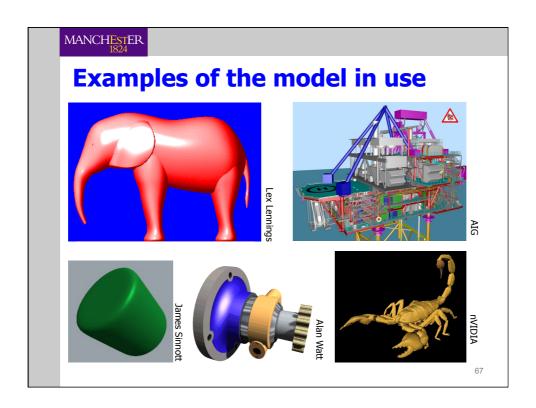
$$I_R = k_{aR} I_{aR} + \frac{I_{pR}}{d'} \left[ k_{dR} (\hat{\mathbf{N}} \cdot \hat{\mathbf{L}}) + k_s (\hat{\mathbf{R}} \cdot \hat{\mathbf{V}})^n \right]$$

# **Multiple lights**

- Finally, what if we have more than one light?
- Easy, compute illumination separately for each and sum. For M lights:

• 
$$I = ambient + \sum_{i=1}^{M} (diffuse_i + specular_i)$$





# The "standard" model

- The local illumination model we've developed is the "standard" model in use today
- Implemented in OpenGL
- Implemented in hardware on consumer 3D graphics cards