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1 Introduction

[1]

1.1 Conventions

2 Basic environments

2.1 Eigenstate and Euklidean path integral in general ✓

We now introduce the Euclidean path integral because we want to find the eigenstates of the left and right Rindler wedges.

First of all we have a ground state $\langle\Omega|$ of H and a vacuum state $\langle 0|$, so we can write for a very long time T

$$\begin{aligned} e^{-iHT}\langle 0| &= \sum_n e^{-iE_n T} \langle n| \langle n|0\rangle \\ &= e^{-iE_0 T} \langle\Omega| \langle\Omega|0\rangle + \sum_{n\neq 0} e^{-iE_n T} \langle n| \langle n|0\rangle \end{aligned}$$

Now we enclose and get the ground state

$$\langle\Omega| = \lim_{T\rightarrow\infty} \frac{e^{iE_0 T}}{\langle\Omega|0\rangle} e^{-iHT} \langle 0|$$

We can define E_0 with $H_0\langle 0| = \langle 0|$ so

$$\langle\Omega| = \frac{1}{\langle\Omega|0\rangle} \lim_{T\rightarrow\infty} e^{-iHT} \langle 0|$$

(see p.86 in [2]) But this equation does still tell us nothing about the entanglement. So let's continue:

Let a time-independent field ϕ act on this ground state:

$$\langle\phi|\Omega\rangle = \frac{1}{\langle\Omega|0\rangle} \lim_{T\rightarrow\infty} \langle\phi|e^{-iHT}|0\rangle$$

Now use the Euclidean path integral formalism, rotate t about 90° into the complex plane: $t \rightarrow -it_E$ and choose the early boundary condition $\phi = 0$, so that

$$\langle\phi|\Omega\rangle \propto \int_{\hat{\phi}(t_E=-\infty)=0}^{\hat{\phi}(t_E=0)=\phi} D\hat{\phi} e^{-I_E} \quad (2.1)$$

with the Euclidean action for a free massive scalar field

$$I_E[\hat{\phi}] = \frac{1}{2} \int d^3x dt_E \left[(\partial_{t_E} \hat{\phi})^2 + (\vec{\nabla} \hat{\phi})^2 + m^2 \hat{\phi}^2 \right] \quad (2.2)$$

Note that $\hat{\phi}$ compared to ϕ is time-dependent.

References

- [1] D. Harlow. *Jerusalem Lectures on Black Holes and Quantum Information*. URL: <https://arxiv.org/abs/1409.1231>.
- [2] M.E. Peskin and D.V. Schroeder. *An Introduction to Quantum Field Theory*. 1995.