

since when did calculus
 PROB LEC 4: have so many markers?

19 JAN '17

• HW LATE - MY BAD

• next wk: more phys; this wk - calculus

• BOOKS & RES

→ Wentz / Zee / D'Inverno / --

for this wk

TODAY

• TRANSFORMATION OF Γ HANDS

• COVARIANT DERIV

• $\Gamma(g)$ (metric connection)

• NEWTONIAN LIMIT & TIME DILATION IN GRAV FIELD

KEY POINTS FROM LAST TIME

• EQUIVALENCE PRINCIPLE:

GRAVITY = ACCELERATED FRAME

→ in free falling frame, "no gravity"

↑ LOCAL, INERTIAL FRAME

WE DENOTED COORDS AS y^μ

METRIC IN FREE FALL IS SPECIAL REL: $g_{\mu\nu} = \eta_{\mu\nu}$

WE STARTED TO FLESH OUT THE PHYSICS IN
ANY OTHER FRAME, x^μ

↑
 COULD BE SAME FRAME, CURVY COORDS

OR NON-INERTIAL OBSERVER

eg RIDE OPERATOR OF TOWER OF TERROR

CHANGE IN COORDS

$$\begin{array}{ccc}
 y^a & \longrightarrow & x^m \\
 \eta_{\alpha\beta} & \longrightarrow & g_{\mu\nu}(x) = \eta_{\alpha\beta} \frac{\partial y^\alpha}{\partial x^\mu} \frac{\partial y^\beta}{\partial x^\nu}
 \end{array}$$

$$\frac{d^2 y^\alpha}{d\tau^2} = 0 \longrightarrow \frac{d^2 x^\mu}{d\tau^2} + \boxed{\Gamma_{\mu\nu}^\rho} \frac{dx^\mu}{d\tau} \frac{dx^\nu}{d\tau} = 0$$

free fall/LIF

also:

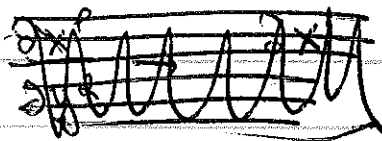
$$\boxed{\frac{\partial y^\alpha}{\partial x^\mu} \Gamma_{\mu\nu}^\rho = \frac{\partial^2 y^\alpha}{\partial x^\mu \partial x^\nu}} \longleftarrow \boxed{\frac{\partial x^\rho}{\partial y^\alpha} \frac{\partial^2 y^\alpha}{\partial x^\mu \partial x^\nu}}$$

CHRISTOFFEL SYMBOLS
AFFINE CONNECTION

NOTE: $\Gamma_{\mu\nu}^\rho$ are not TENSORS
why? the 2nd derivative - not linear

$$\text{TENSOR: } T_{\mu\nu}^\rho \rightarrow \left(\frac{\partial x'^\rho}{\partial x^\sigma} \right) \left(\frac{\partial x^\delta}{\partial x'^\mu} \right) \left(\frac{\partial x^\gamma}{\partial x'^\nu} \right) T_{\delta\gamma}^\sigma$$

$\uparrow \quad \quad \uparrow \quad \quad \uparrow$
 "R" "R⁻¹" "R⁻¹"



$$(\Gamma')^{\rho}_{\mu\nu} \equiv \frac{\partial x'^{\rho}}{\partial y^{\alpha}} \frac{\partial}{\partial x'^{\mu}} \frac{\partial y^{\alpha}}{\partial x'^{\nu}}$$

\swarrow \searrow

$\frac{\partial x'^{\rho}}{\partial x^{\sigma}} \frac{\partial x^{\sigma}}{\partial y^{\alpha}}$

TRANSFORM.
"R"

\uparrow
PIECES
OF Γ

$\frac{\partial x^{\delta}}{\partial x'^{\mu}} \frac{\partial}{\partial x^{\delta}} \left(\frac{\partial x^{\gamma}}{\partial x'^{\nu}} \frac{\partial y^{\alpha}}{\partial x^{\gamma}} \right)$

"R"

CHAIN RULE

$$\frac{\partial x^{\delta}}{\partial x'^{\mu}} \left[\frac{\partial^2 x^{\gamma}}{\partial x^{\delta} \partial x'^{\nu}} \frac{\partial y^{\alpha}}{\partial x^{\gamma}} + \frac{\partial x^{\gamma}}{\partial x'^{\nu}} \frac{\partial^2 y^{\alpha}}{\partial x^{\delta} \partial x^{\gamma}} \right]$$

nothing
OTHER PIECE OF Γ

$$(\Gamma')^{\rho}_{\mu\nu} = \frac{\partial x'^{\rho}}{\partial x^{\sigma}} \underbrace{\frac{\partial x^{\sigma}}{\partial y^{\alpha}} \frac{\partial^2 y^{\alpha}}{\partial x^{\delta} \partial x^{\gamma}}}_{\Gamma^{\sigma}_{\delta\gamma}} \frac{\partial x^{\delta}}{\partial x'^{\mu}} \frac{\partial x^{\gamma}}{\partial x'^{\nu}} \quad \left. \vphantom{\frac{\partial x^{\delta}}{\partial x'^{\mu}} \frac{\partial x^{\gamma}}{\partial x'^{\nu}}} \right\} \text{tensorial transp.}$$

\uparrow \uparrow \uparrow \uparrow
 R $\Gamma^{\sigma}_{\delta\gamma}$ R^{-1} R^{-1}

junk transp.

$$+ \left[\frac{\partial x'^{\rho}}{\partial x^{\sigma}} \frac{\partial x^{\sigma}}{\partial y^{\alpha}} \frac{\partial x^{\delta}}{\partial x'^{\mu}} \frac{\partial^2 x^{\gamma}}{\partial x^{\delta} \partial x'^{\nu}} \frac{\partial y^{\alpha}}{\partial x^{\gamma}} \right]$$

\downarrow
 $= \delta^{\sigma}_{\gamma}$

$$\rightarrow \left[\frac{\partial x'^{\rho}}{\partial x^{\gamma}} \frac{\partial^2 x^{\gamma}}{\partial x'^{\mu} \partial x'^{\nu}} \right]$$

JUNK TERM: $\left[\frac{\partial x'^P}{\partial x^\gamma} \frac{\partial^2 x^\gamma}{\partial x'^\mu \partial x'^\nu} \right]$

USE: $\frac{\partial x'^P}{\partial x^\gamma} \frac{\partial x^\gamma}{\partial x'^\nu} = \delta_\nu^P$

DIFFERENTIATE: $\partial/\partial x'^\mu$ BOTH SIDES

$\frac{\partial x^\gamma}{\partial x'^\mu} \frac{\partial}{\partial x'^\mu} \rightarrow \frac{\partial^2 x'^P}{\partial x'^\mu \partial x^\gamma} \frac{\partial x^\gamma}{\partial x'^\nu} + \left[\frac{\partial x'^P}{\partial x^\gamma} \frac{\partial^2 x^\gamma}{\partial x'^\mu \partial x'^\nu} \right] = 0$

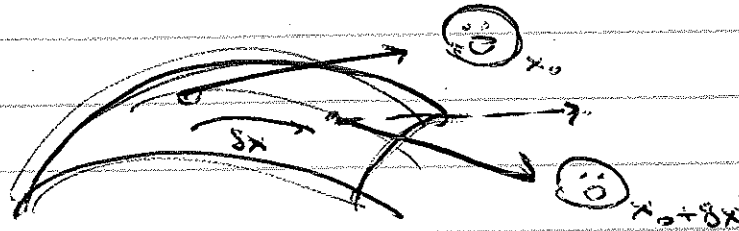
S.t. $\left(\Gamma_1 \right)^\rho_{\mu\nu} = \frac{\partial x'^P}{\partial x^\sigma} \frac{\partial x^\sigma}{\partial x'^\mu} \frac{\partial x^\gamma}{\partial x'^\nu} \Gamma_{\gamma\sigma}^\rho$

$- \frac{\partial x^\gamma}{\partial x'^\mu} \frac{\partial x^\sigma}{\partial x'^\nu} \frac{\partial^2 x'^P}{\partial x'^\mu \partial x'^\nu}$

ONE USEFUL INTERP/OBS: THE JUNK APPEARS
TO BE RELATED TO THE FACT THAT
WE'RE DIFFERENTIATING SOMETHING W/ AN
INDEX:

$$\frac{\partial}{\partial x^\mu} \text{ (smiley face) }^\nu (x)$$

BOTH OF THESE TRANSFORM



THIS IS A HINT THAT Γ WILL HAVE
SOMETHING TO DO W/ DIFFERENTIATING
TENSORS.

TOTAL: PIRS
CALCULUS WAY.

Wemb. 4-6 IN FACT, WE MIGHT AS WELL SEE THIS
CONSIDER A VECTOR V^μ

transf law: AS $x \rightarrow x'$, the vector V^μ
transforms to V'^μ

$$V'^\mu = \frac{\partial x'^\mu}{\partial x^\nu} V^\nu$$

BOTH FUNCTIONS OF x'

$$\frac{\partial V'^\mu}{\partial x'^\lambda} = \underbrace{\frac{\partial x'^\mu}{\partial x^\nu} \frac{\partial x^\rho}{\partial x'^\lambda} \frac{\partial V^\nu}{\partial x^\rho}}_{\text{"GOOD" TRANSFORMATION}} + \left[\frac{\partial^2 x'^\mu}{\partial x^\nu \partial x^\rho} \frac{\partial x^\rho}{\partial x'^\lambda} \right] V^\nu$$

is this an AMERIGANISM

looks familiar!

OBSERVE:

$$\Gamma_{\lambda\kappa}^{\mu} V'^\kappa = \frac{\partial x'^\mu}{\partial x^\nu} \frac{\partial x^\rho}{\partial x'^\lambda} \left[\frac{\partial x^\sigma}{\partial x'^\kappa} \right] \Gamma_{\rho\sigma}^\nu \left[\frac{\partial x'^\kappa}{\partial x'^\eta} \right] V'^\eta$$

δ^σ_η

$$= \frac{\partial x^\nu}{\partial x'^\lambda} \left[\frac{\partial x^\rho}{\partial x'^\kappa} \frac{\partial^2 x'^\mu}{\partial x^\nu \partial x^\rho} \frac{\partial x'^\kappa}{\partial x'^\eta} \right] V'^\eta$$

δ^ρ_η

$$= \left[- \frac{\partial x^\nu}{\partial x'^\lambda} \frac{\partial^2 x'^\mu}{\partial x^\nu \partial x^\rho} V^\rho \right]$$

(cancels δ^ρ_η)! (note $\nu \leftrightarrow \rho$ sym.)

SO AS A RESULT: (dropping primes)

$$D_\mu V^\nu \equiv \left(\frac{\partial V^\nu}{\partial x^\mu} + \Gamma_{\mu\sigma}^\nu V^\sigma \right)$$

is COVARIANT (is a tensor)

$$D_\mu V^\nu \rightarrow \left[\frac{\partial x^\rho}{\partial x'^\mu} \frac{\partial x'^\nu}{\partial x^\sigma} D'_\rho V^\sigma \equiv D'_\mu V'^\nu \right]$$

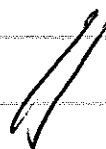
transforms the way its indices "want to"

$$D_\mu = \left(\frac{\partial}{\partial x^\mu} + \Gamma_{\mu\sigma}^\sigma \right)$$

↑ will generalize to more complicated tensors

Q: D_μ on ϕ ?

↑ no indices



NOW LET'S STEP BACK A BIT & GO BACK
TO OUR 'TOWER OF TERROR' EXAMPLE

nice coords : $y^a \leftarrow$ LOCAL, INERTIAL FRAME

gen coords : $x^\mu \leftarrow$ e.g. RIDE OPERATOR

WE ARGUED THAT @ A GIVEN POINT IN
SPACETIME (= an event), CAN COOK UP
 y^a COORDS AS A FUNCTION OF x^μ ,
given $\Gamma :: \dot{\tau} \leftarrow$ proper time

"curvature"

evidently gravity lives in these
objects

THESE OBJECTS CAN BE RELATED

$$g_{\mu\nu} = \frac{\partial y^a}{\partial x^\mu} \frac{\partial y^b}{\partial x^\nu} \eta_{ab}$$

DIFFERENTIATE $g_{\mu\nu}$ (KNOW THIS BRINGS OUT Γ 's!)

$$\frac{\partial g_{\mu\nu}}{\partial x^\rho} = \left(\underbrace{\frac{\partial^2 y^a}{\partial x^\rho \partial x^\mu} \frac{\partial y^b}{\partial x^\nu}}_{\frac{\partial y^a}{\partial x^\sigma} \Gamma^\sigma_{\rho\mu}} + \frac{\partial y^a}{\partial x^\mu} \underbrace{\frac{\partial^2 y^b}{\partial x^\rho \partial x^\nu}}_{\frac{\partial y^b}{\partial x^\sigma} \Gamma^\sigma_{\rho\nu}} \right) \eta_{ab}$$

* warning - I'm not obviously allowed to do this!

SO WE HAVE FOUND:

$$\frac{\partial g_{\mu\nu}}{\partial x^\rho} = \underbrace{\Gamma_{\mu\rho}^\sigma}_{\text{sym}} \underbrace{g_{\sigma\nu}}_{\text{sym}} + \Gamma_{\rho\nu}^\sigma g_{\sigma\mu}$$

great! now we want to massage into $\Gamma = \Gamma(g)$

BUT THIS IS HARD - CANNOT
SIMPLY GROUP the Γ 's —
they have different indices!

HAVE TO BE MORE CLEVER!

WRITING $\partial_\mu \leftrightarrow \partial/\partial x^\mu$

some
tensor
struct!

$$\begin{aligned} \partial_\lambda g_{\mu\nu} &= g_{\sigma\nu} \Gamma_{\lambda\mu}^\sigma + [g_{\sigma\mu} \Gamma_{\lambda\nu}^\sigma] \\ + \partial_\mu g_{\lambda\nu} &= g_{\sigma\nu} \Gamma_{\mu\lambda}^\sigma + [g_{\sigma\lambda} \Gamma_{\mu\nu}^\sigma] \\ - \partial_\nu g_{\mu\lambda} &= [g_{\sigma\lambda} \Gamma_{\mu\nu}^\sigma] - [g_{\sigma\mu} \Gamma_{\lambda\nu}^\sigma] \\ &= \boxed{2g_{\sigma\nu} \Gamma_{\mu\lambda}^\sigma} \end{aligned}$$

but w/ $\frac{1}{2} g^{\nu\sigma}$ on both sides:

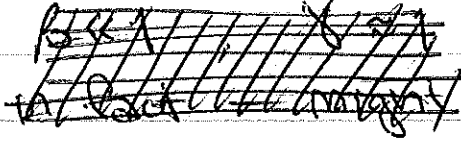
$$\boxed{\Gamma_{\mu\lambda}^\sigma = \frac{1}{2} g^{\sigma\nu} (\partial_\lambda g_{\mu\nu} + \partial_\mu g_{\lambda\nu} - \partial_\nu g_{\mu\lambda})}$$

(metric connection)

Wk 3-4

Newtonian Limit

$$\text{EOM: } \frac{d^2 x^\mu}{d\tau^2} + \Gamma_{\rho\sigma}^\mu \frac{\partial x^\rho}{\partial \tau} \frac{\partial x^\sigma}{\partial \tau} = 0$$

NEWTON / NON-REL LIMIT \rightarrow low "velocity"
 $\frac{\partial x^i}{\partial \tau} \ll 1$
 \rightarrow WEAK, STATIONARY FIELD

$$\text{so: } \left[\frac{\partial x}{\partial \tau} \ll \frac{\partial g}{\partial \tau} \right]$$

$$\text{EOM} \rightarrow \left[\frac{d^2 x^\mu}{d\tau^2} + \Gamma_{00}^\mu \left(\frac{\partial x^0}{\partial \tau} \right)^2 \right] \quad i=1,2,3$$

$$+ 2\Gamma_{0i}^\mu \frac{\partial x^0}{\partial \tau} \frac{\partial x^i}{\partial \tau} + \Gamma_{ij}^\mu \frac{\partial x^i}{\partial \tau} \frac{\partial x^j}{\partial \tau}$$

} SUBDOM.

$$\Gamma_{00}^\mu = \frac{1}{2} g^{\mu\nu} (\partial_0 g_{0\nu} + \partial_0 g_{\nu 0} - \partial_\nu g_{00})$$

NEWTONIAN GRAVITY LIVES IN HERE.

STATIC, so $\partial_0 = \partial/\partial x^0 = \partial/\partial t$
 PARTS SHOULD VANISH

$$\Gamma_{00}^{\mu} = -\frac{1}{2} g^{\mu\nu} \partial_{\nu} [g_{00}]$$

WEAK FIELD LIMIT

SPACETIME IS ALMOST MINKOWSKI

$$g_{\mu\nu} = \underbrace{\eta_{\mu\nu}}_{\text{CONST.}} + h_{\mu\nu}$$

τ may have x-dep.

LEADING ORDER:

$$\Gamma_{00}^{\mu} = -\frac{1}{2} \eta^{\mu\nu} \partial_{\nu} h_{00}$$

PLUG BACK INTO FORM:

$$\frac{\partial^2 x^{\mu}}{\partial \tau^2} = \frac{1}{2} \eta^{\mu\nu} \partial_{\nu} h_{00} \cdot \left(\frac{\partial t}{\partial \tau} \right)^2$$

mostly minus metric

$$\frac{\partial^2 x^i}{\partial \tau^2} = -\frac{1}{2} \nabla_i h_{00} \left(\frac{\partial t}{\partial \tau} \right)^2$$

ORDINARY GRADIENT

$$\frac{\partial^2 t}{\partial \tau^2} = 0$$

$$\text{b/c } \frac{1}{2} \eta^{00} \partial_0 h_{00} = 0$$

$$\Rightarrow \left[\frac{\partial t}{\partial \tau} = \text{CONST} \right]$$

$$\frac{\partial}{\partial t} \left(\frac{\partial x}{\partial t} \right) = \frac{\partial^2 x}{\partial t^2}$$

but: $\frac{\partial}{\partial t} \left(\frac{\partial t}{\partial t} \right) = 0$

by EOM

$$= \left(\frac{\partial t}{\partial \tau} \right)^2 \frac{\partial^2 x}{\partial t^2}$$

$$= -\frac{1}{2} \nabla^2 h_{00} \left(\frac{\partial t}{\partial \tau} \right)^2$$

SPATIAL
EOM

$$\Rightarrow \frac{\partial^2 x}{\partial t^2} = -\frac{1}{2} \nabla^2 h_{00} \stackrel{?}{=} -\nabla^2 \Phi$$

↑
acceleration

$$\Rightarrow \boxed{h_{00} = 2\Phi + \text{const}}$$

const shift in Φ
unphysical

$$\boxed{\begin{aligned} g_{00} &= (1 + 2\Phi) \\ g_{ij} &= -\delta_{ij} \end{aligned}}$$

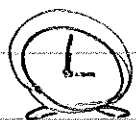
EXERCISE:

WHAT HAPPENS
TO SIGNS
IN FAST
COAST METRIC?

vs. from SR James!

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TIME DILATION IN A GRW FIELD not nec in free fall
eg on a table



CLOCK IN SAME FRAME
IN A CURVED SPACETIME
in a grav. field

IN NICE FRAME (free fall / loc. inert.)

$$\Delta T^2 = \eta_{\alpha\beta} dy^\alpha dy^\beta = g_{\mu\nu} dx^\mu dx^\nu$$

TIME BTWN TICKS
IN ABS. OF GRW.

$$\left(\eta_{\alpha\beta} \frac{\partial y^\alpha}{\partial x^\mu} \frac{\partial y^\beta}{\partial x^\nu} \right)$$

AS OBS BY NICE FRAME (manufacturer spec.)

$$\left(\frac{\Delta T}{dt} \right)^2 = g_{\mu\nu} \frac{dx^\mu}{dt} \frac{dx^\nu}{dt}$$

\uparrow velocity of clock

then time btwn ticks in x frame
is

$$\frac{dt}{\Delta T} \cdot (\Delta T) = \frac{1}{\sqrt{g_{\mu\nu} \frac{dx^\mu}{dt} \frac{dx^\nu}{dt}}}$$

CLOCK
@ REST
IN X'
FRAME

$$\frac{1}{\sqrt{g_{00}}}$$

$$\frac{1}{\sqrt{1 + 2\Phi}}$$

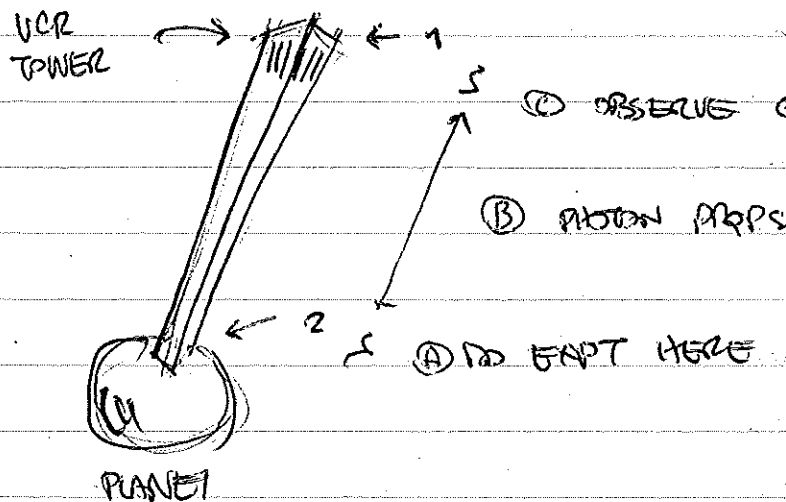
newtonian limit

SO DO EXPT. MEASURE CLOCKS NEAR
BH & FAR AWAY. WHAT DO YOU
FIND?

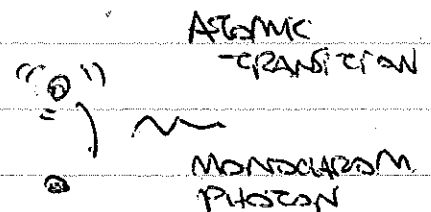
→ no measured discrepancy
from manufacturer's
description, ΔT !

How do you measure a clock?
w/ another clock ... which is
also dilated.

EXPERIMENT



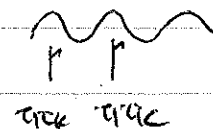
EXPERIMENT:



① OBSERVE CRESTS HERE

② PHOTON PROPS

③ TO EXP. HERE



Time for photon to go from ground to top
is something we "calculated"

SWE : $0 = g_{\mu\nu} dx^\mu dx^\nu$

got: $dt = \frac{1}{g_{00}} \left[-g_{ij} dx^i + \sqrt{(g_{00}g_{ij} - g_{i0}g_{j0})} \right] dx^i dx^j$

↑
integrate this
to get time for
a photon to traverse
height of tower.

(ASSUME
SPH. SYM GRV FIELD, eg. NEWTONIAN,
S₀ - TRAJECTORY IS STRAIGHT RADIAL)

BUT : $\int_{\text{time}} dt$ IS JUST SOME CONSTANT
FOR EACH PHOTON / CREST

SO TIME SEPARATION BWN
CRESTS IS UNAFFECTED BY
PROPAGATION FROM 2 → 1.

@ the floor: $dt_2 = \Delta T / \sqrt{g_{00}(x_2)}$

@ the top: $dt_1 = \Delta T / \sqrt{g_{00}(x_1)}$

then the ratio of obs. frequencies are:

$$\frac{\nu_2}{\nu_1} = \sqrt{\frac{g_{00}(x_2)}{g_{00}(x_1)}}$$

emitted @ 2 ↑
emitted @ 1
Both obs @ 1

Newtonian limit: weak field, $g_{00} = 1 + \phi$, $\phi \ll 1$

$$\frac{\nu_2 - \nu_1}{\frac{1}{2}(\nu_2 + \nu_1)} = \sqrt{\frac{2(\phi_2 - \phi_1)}{1 + \phi_1 + \phi_2}}$$

$$\approx \boxed{\phi_2 - \phi_1}$$

$\phi_i = \phi(x_i)$

shift in weak field limit.