

Trigger Efficiency Analysis in the ICARUS Neutrino Detector

Neutrino Group Final Presentation

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Agenda

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Detector Overview

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Detector Overview

ICARUS Detector

Imaging Cosmic And Rare Underground Signals

- Liquid Argon Time Projection Chamber (LArTPC) Detector
 - Creates digital images of neutrino interactions to better understand their properties
- Part of the Short Baseline Neutrino Program at Fermilab
- Composed of two semi-independent cryostats, each containing two LArTPCs

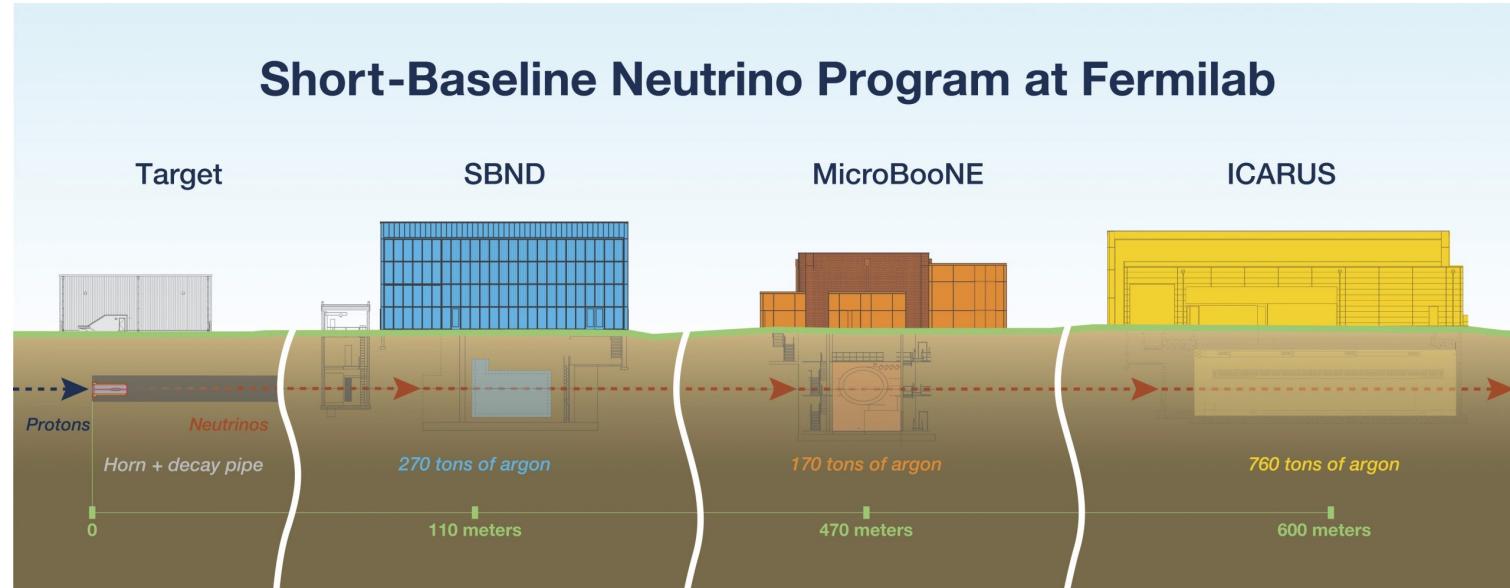


Figure 2. Overview of SBN at Fermilab. Image credits: ICARUS collaboration

ICARUS Detector

Imaging Cosmic And Rare Underground Signals

- LArTPC reads drifting charge → slow! (ms)
- PMTs (photomultipliers) detect light → fast! (ns)
- Within drift time, many cosmic rays pass through detector creating background
- PMTs help us identify cosmic rays

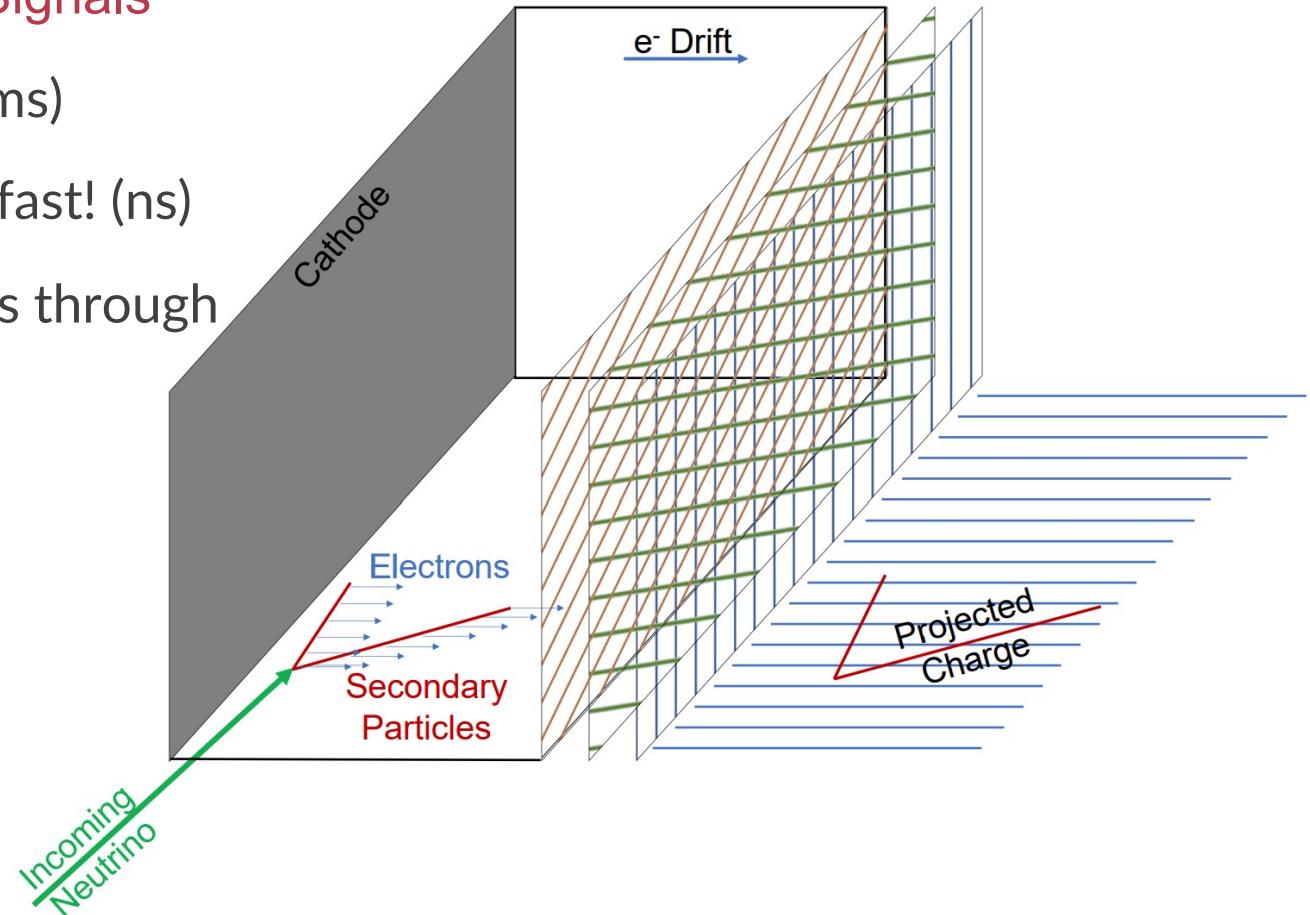


Figure 3. Overview of Liquid Argon Time Projection Chamber (LArTPC).
Image Credits: ANL, FNAL.

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Trigger System Overview

Trigger System

- Hardware system that filters out background in real time
- Test different light requirement levels using software emulation
 - Each pair of PMTs “triggering” means that one detected light above fixed threshold
 - M1: 1 PMT pair triggers within the entire detector
 - S3, S5, S8, S10, S15: # of PMT pairs that trigger within 1 of 3 6m sections of detector
- Select requirement level that maximizes efficiency of recording desired tracks while minimizing the background we accumulate

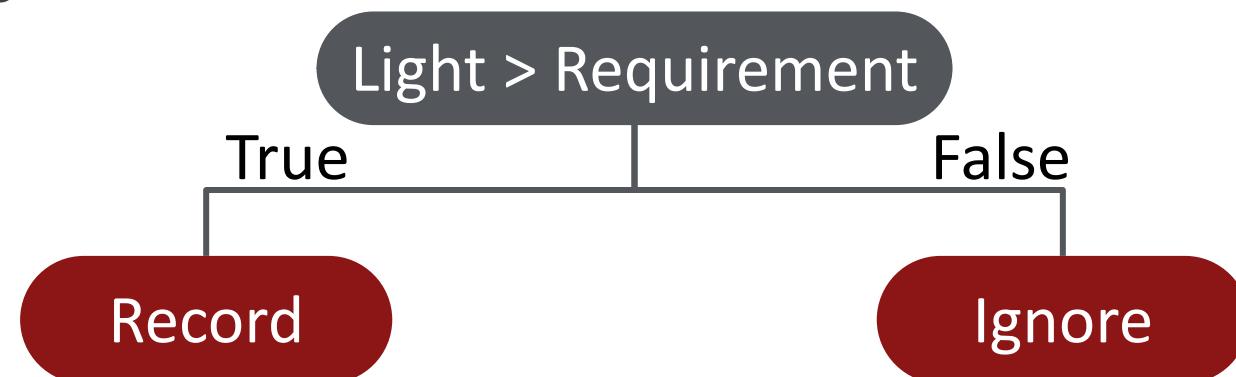


Figure 4. ICARUS Trigger System Logic.

The background of the slide features a complex, abstract pattern of wavy, translucent lines in various colors, primarily green, blue, and yellow, set against a dark, almost black, background. The lines create a sense of depth and motion.

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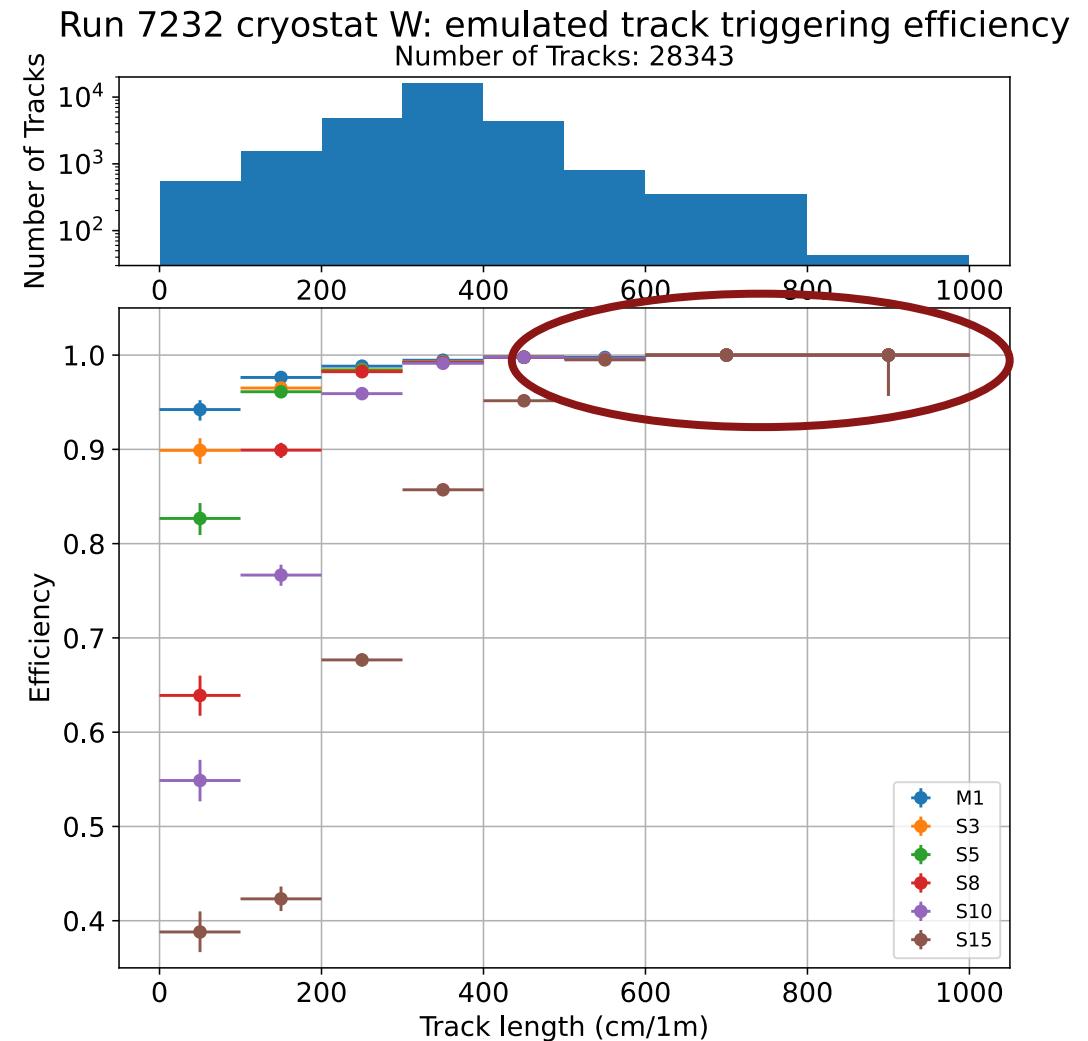
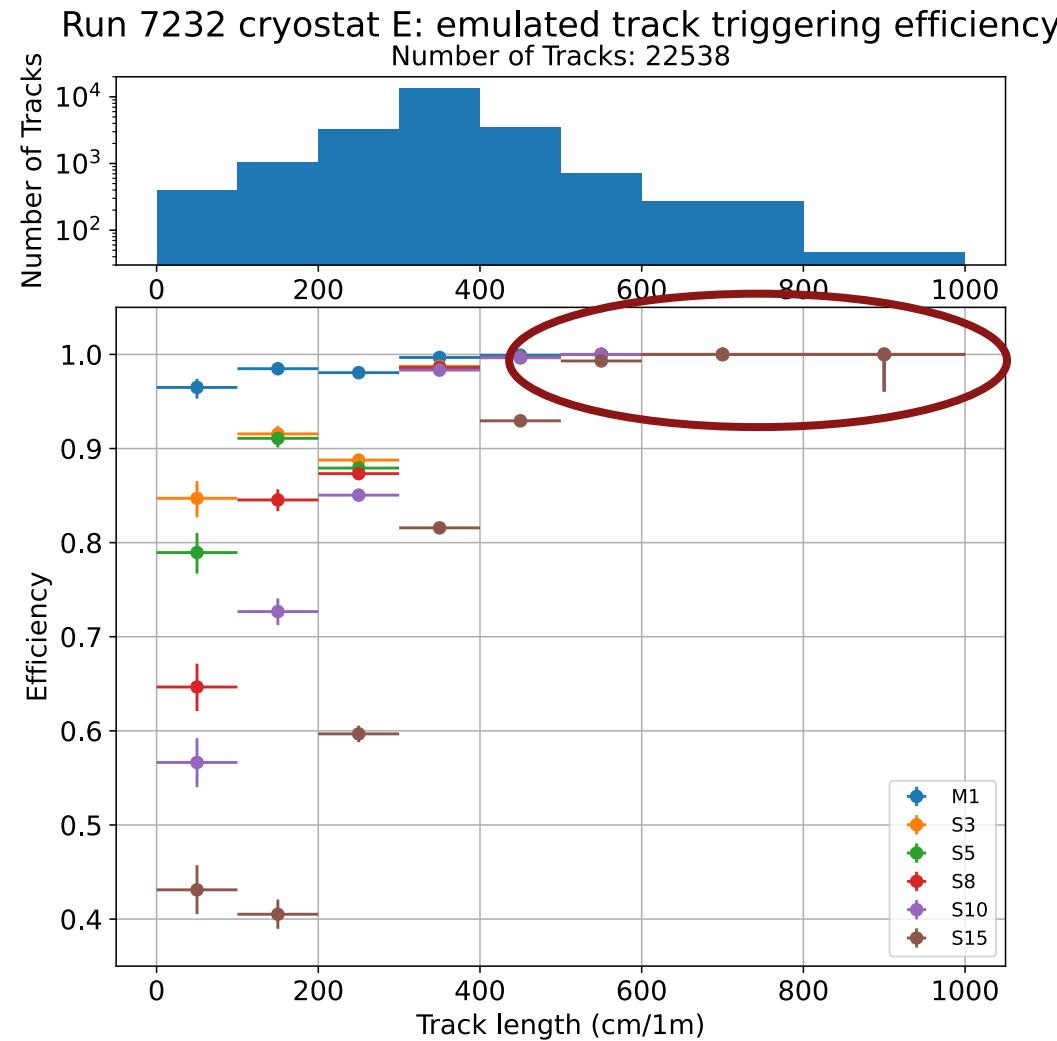
Efficiency Analysis

What does our data look like?

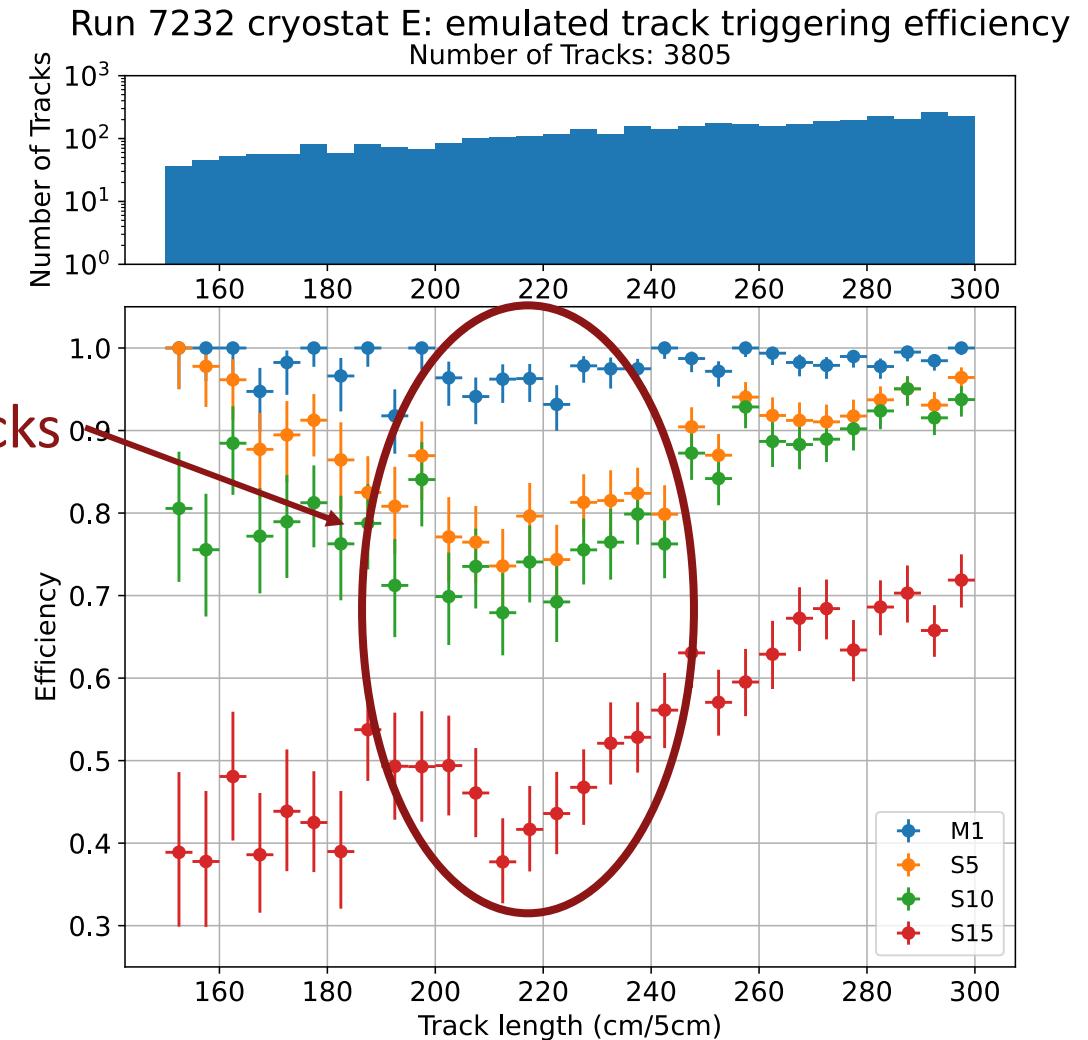
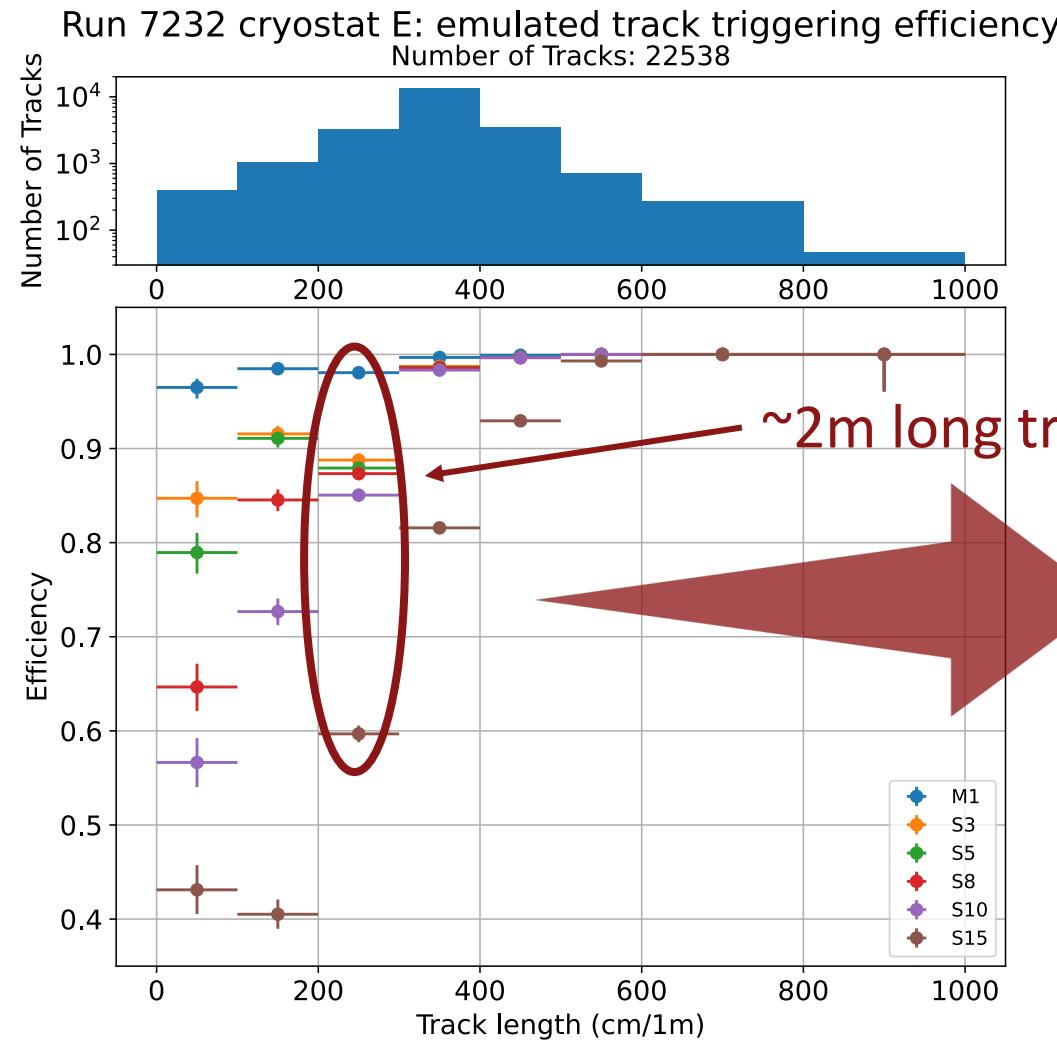
- “Minimum bias” run from November 30, 2021
 - Data collected without hardware trigger constraints
 - Software used to emulate trigger performance under different light requirement levels
- Cathode-crossing tracks only
 - Only tracks for which we can reconstruct the time without biasing trigger efficiency measurement
 - Hope to look at tracks that don’t cross cathode in the future, reducing sample bias

$$\text{Efficiency} = \frac{\text{selected tracks that would trigger}}{\text{selected tracks}}$$

Efficiency of Trigger as a function of Track Length



Efficiency of Trigger as a function of Track Length



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The 2m Track Anomaly

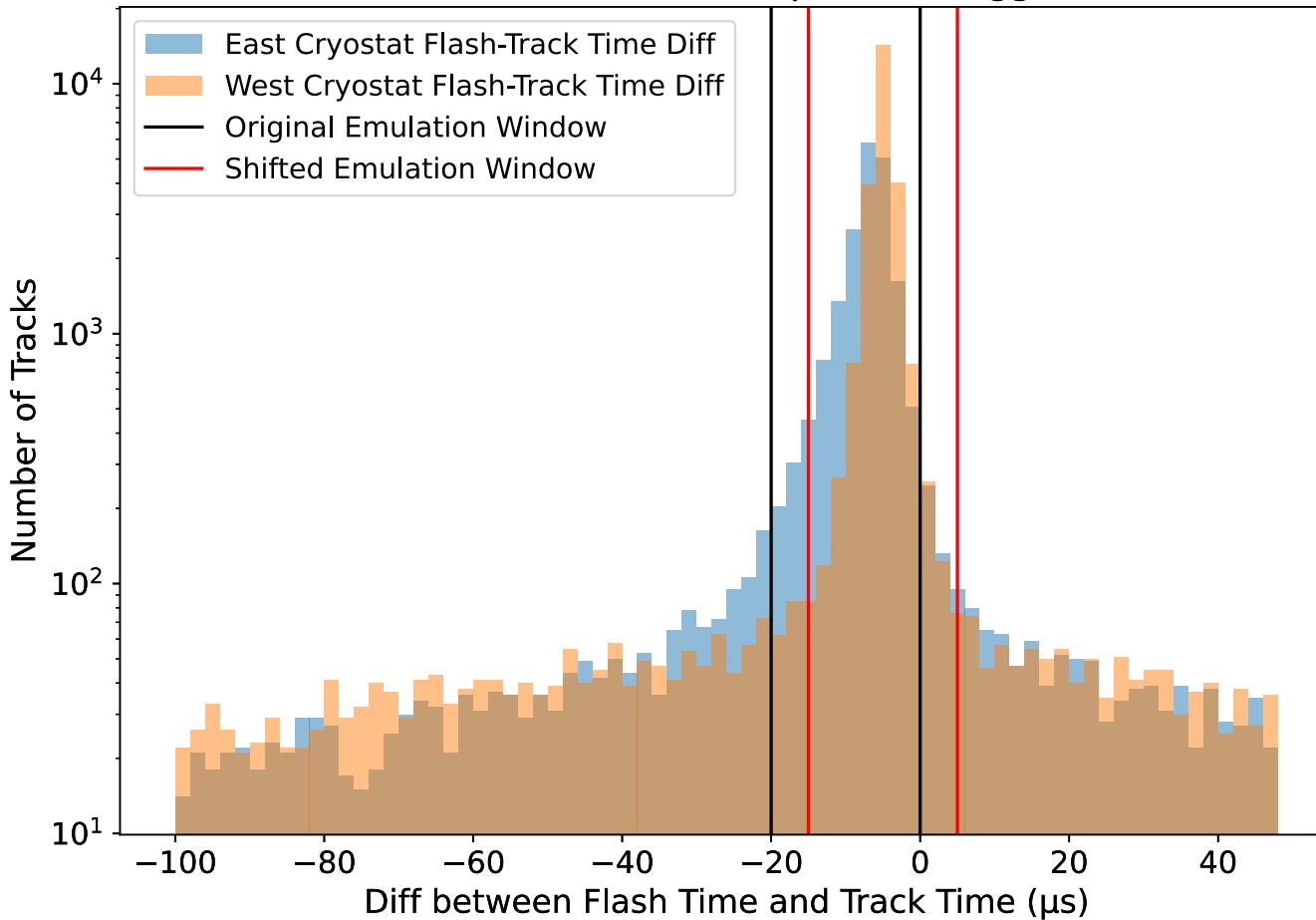
The 2m Track Anomaly

Tracks ~2m in length have a noticeably lower efficiency than slightly shorter or longer tracks

- Especially evident in East cryostat but also present in West cryostat (less severely)
- No noticeable spatial pattern for non-triggering tracks
- Statistics limited for current dataset, hoping a newer run will provide more insight

Shifting the Trigger Emulation Window

Run 7232: Flash-Track Time Difference Compared to Trigger Emulation Windows

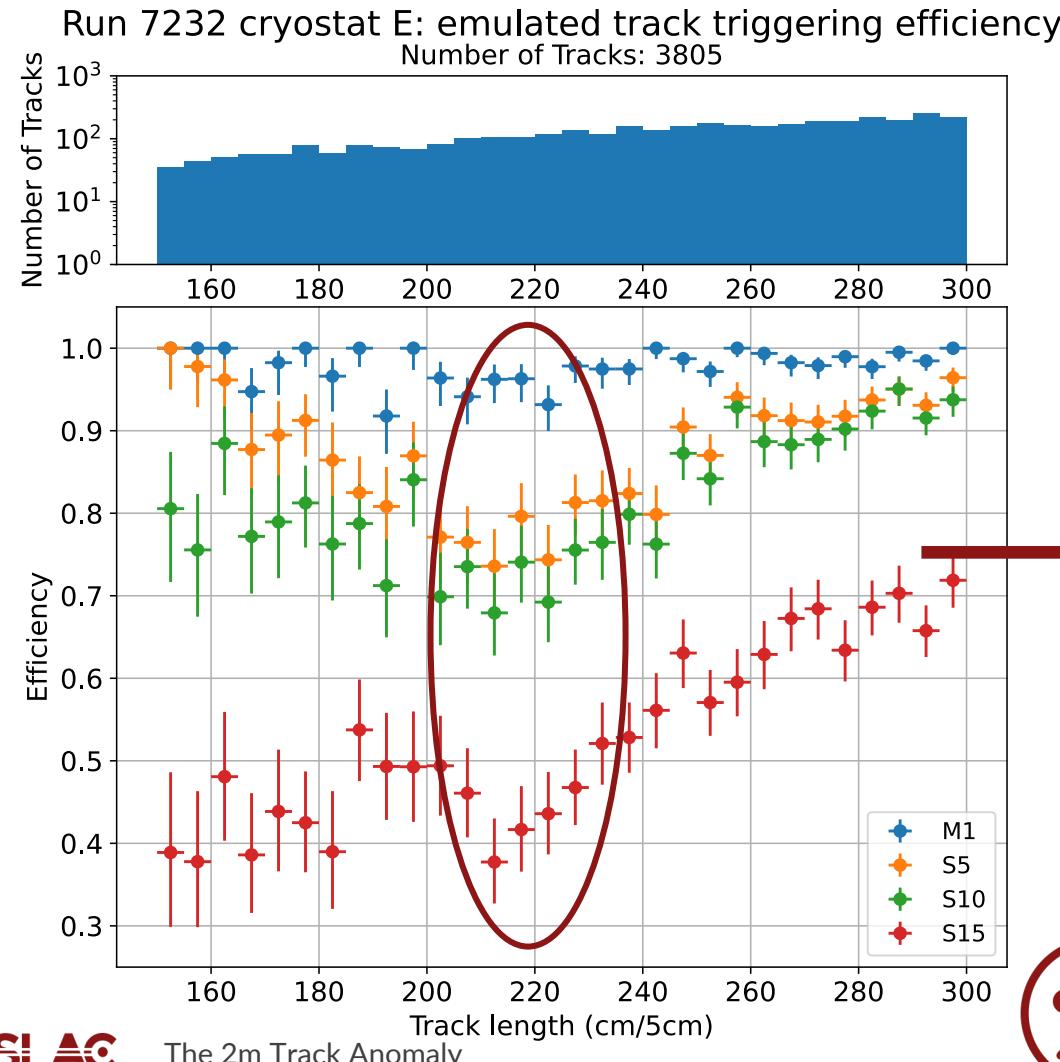


Possible Solution:

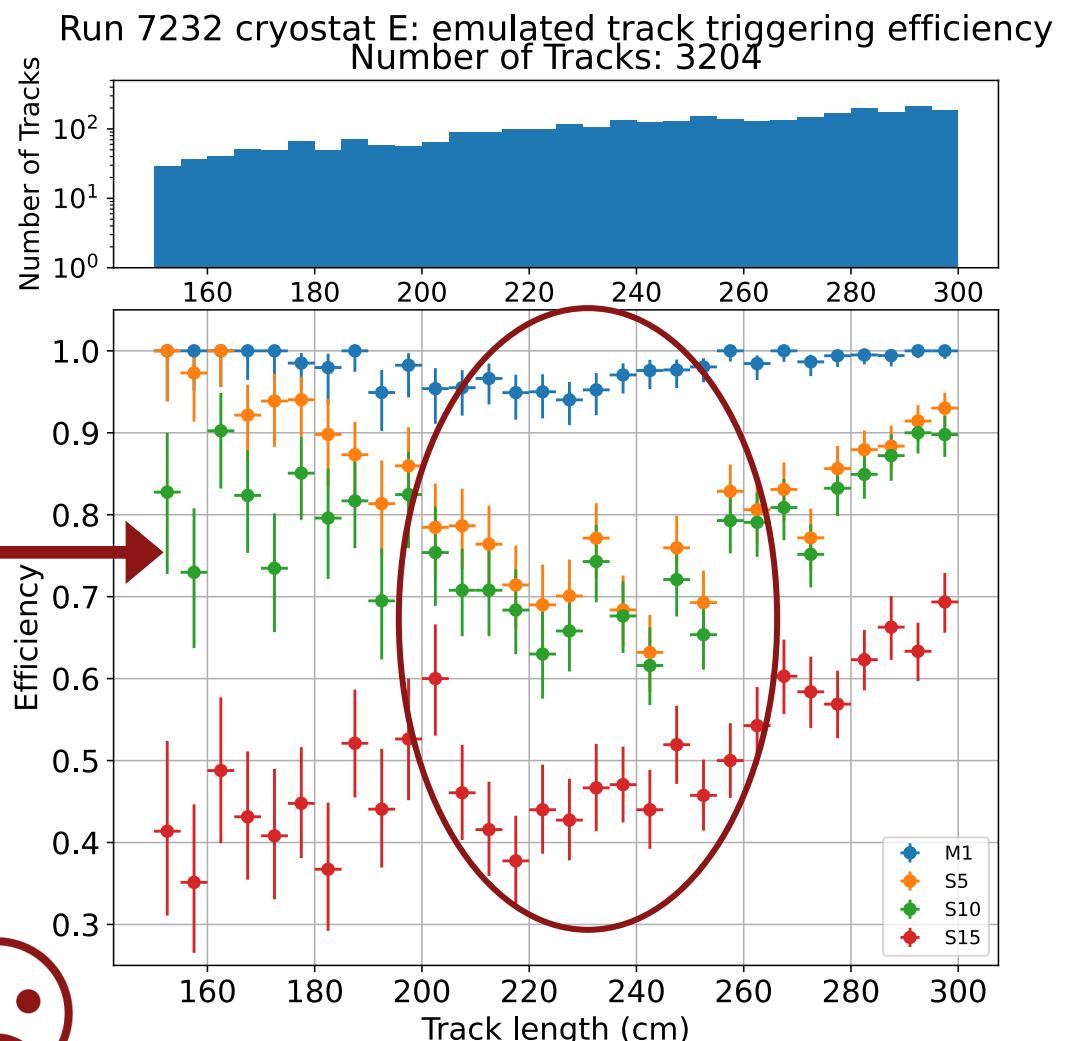
- In Trigger Emulation software, we look for light in a $20 \mu\text{s}$ window before the track time (t_0), and check whether that light exceeds chosen light requirement level
- Most tracks of length $\sim 2\text{m}$ that failed to trigger matched to light occurring $< 5 \mu\text{s}$ after t_0
- Shifted window later by $5 \mu\text{s}$ to see whether overall efficiency improved

Shifting the Trigger Emulation Window: East Cryostat

Original Emulation Window

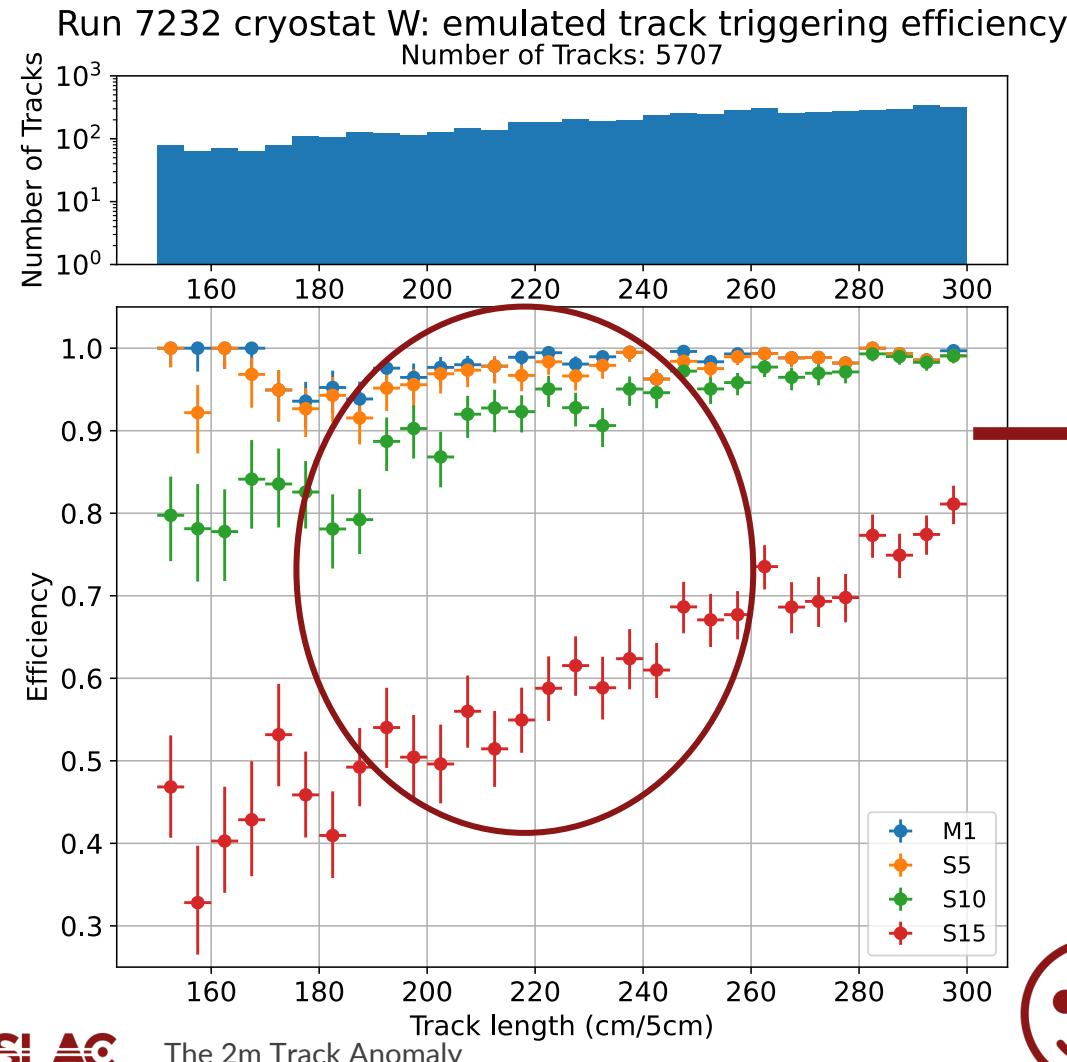


Shifted Emulation Window

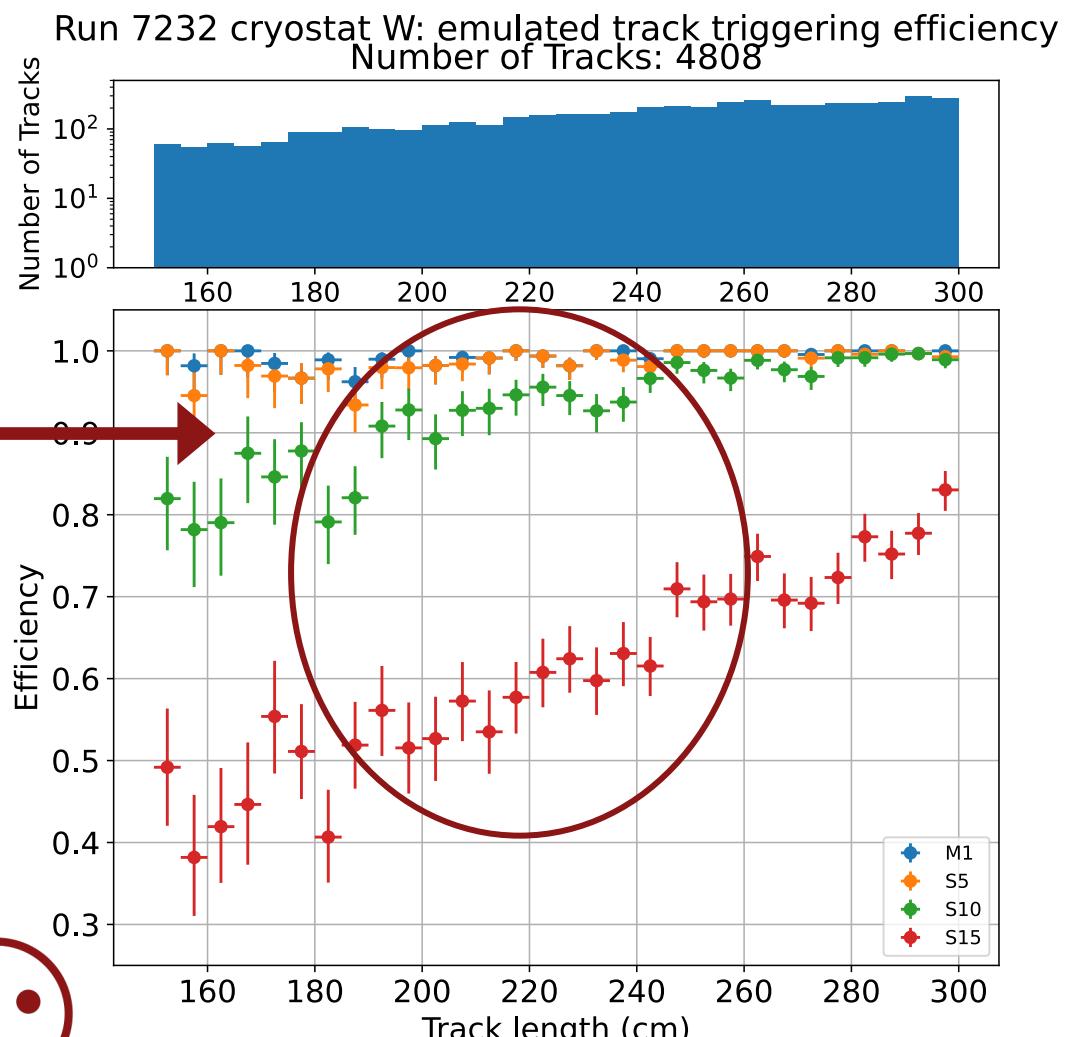


Shifting the Trigger Emulation Window: West Cryostat

Original Emulation Window



Shifted Emulation Window



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Other Possible Considerations

Potential Causes of the 2m track anomaly

Possible connections

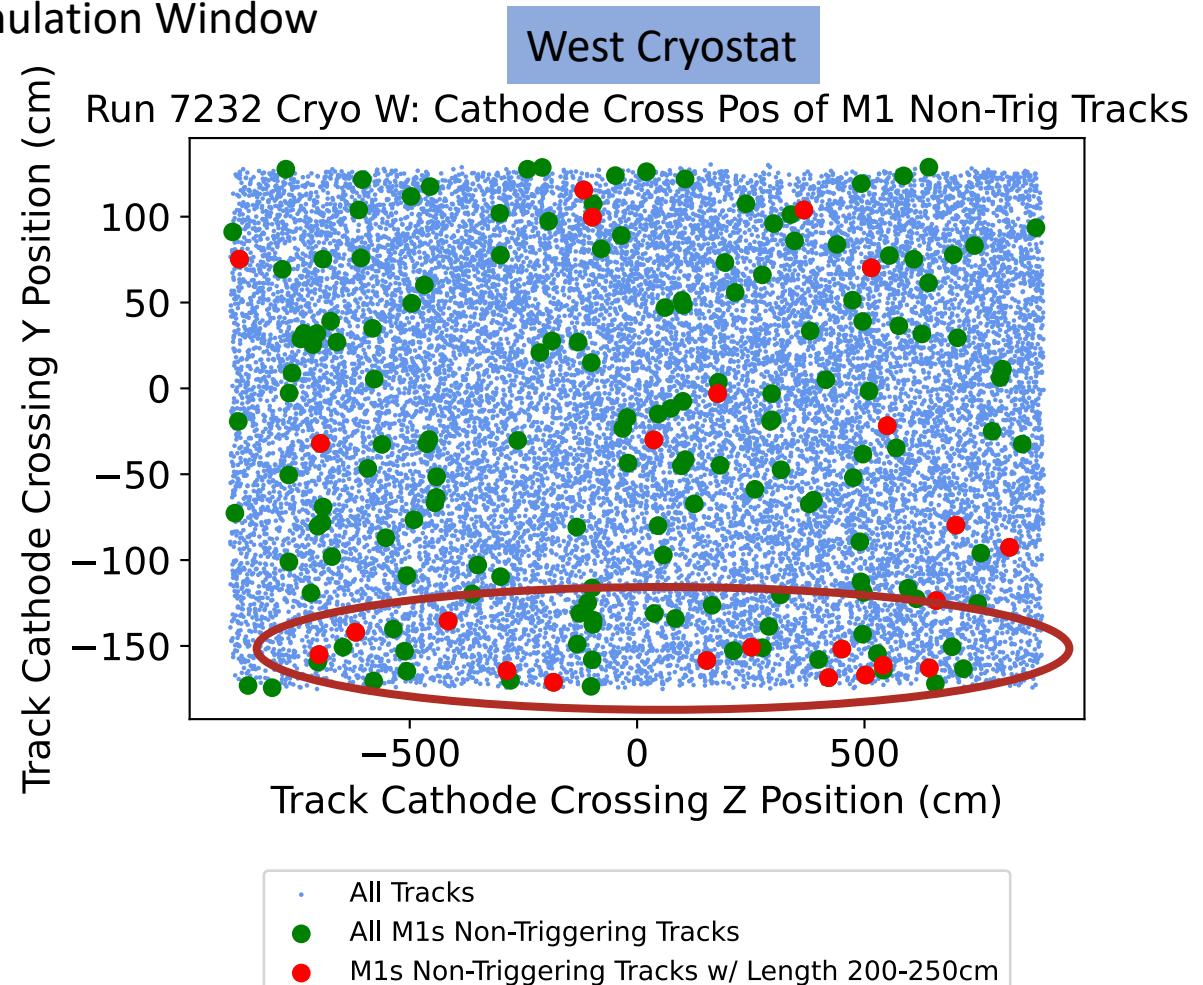
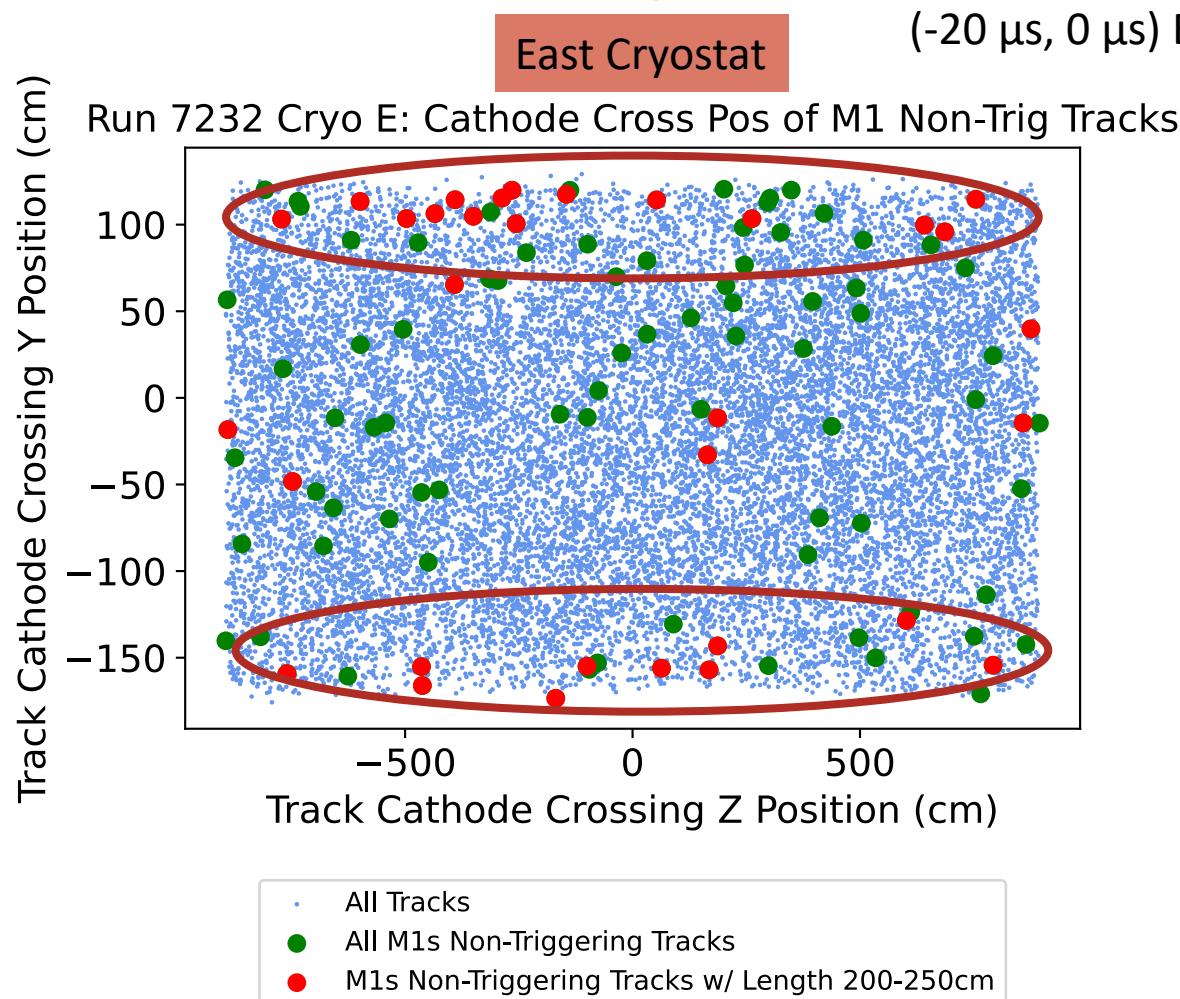
1. Shifting trigger emulation window, t0 reconstruction
 - Somewhat for 2m tracks
2. Length balance ratio of tracks on either side of cathode
3. Smaller track component along the z-axis, large track component along the x-axis
 - Not specific to 2m tracks (geometric bias)
4. Border distortions of electric field
5. Y-axis cathode-crossing position
 - Lower visibility of photons at high y-values
 - Photons lost if they hit the field cage, fewer PMTs at high y-values

Ruled out connections

1. Not time dependent
2. Not from cathode bending
 - Cathode bending shouldn't greatly affect t0 reconstruction
3. Not specific to y-direction of tracks
4. Likely not due to split tracks during track reconstruction
 - From what we saw of the event display (not a statistically significant sample)
 - Also, most tracks appear to go from one edge of cryostat to another (>60% have at least 1 endpoint at an edge of the cryostat)

Possible connection:

Y-axis cathode-crossing position

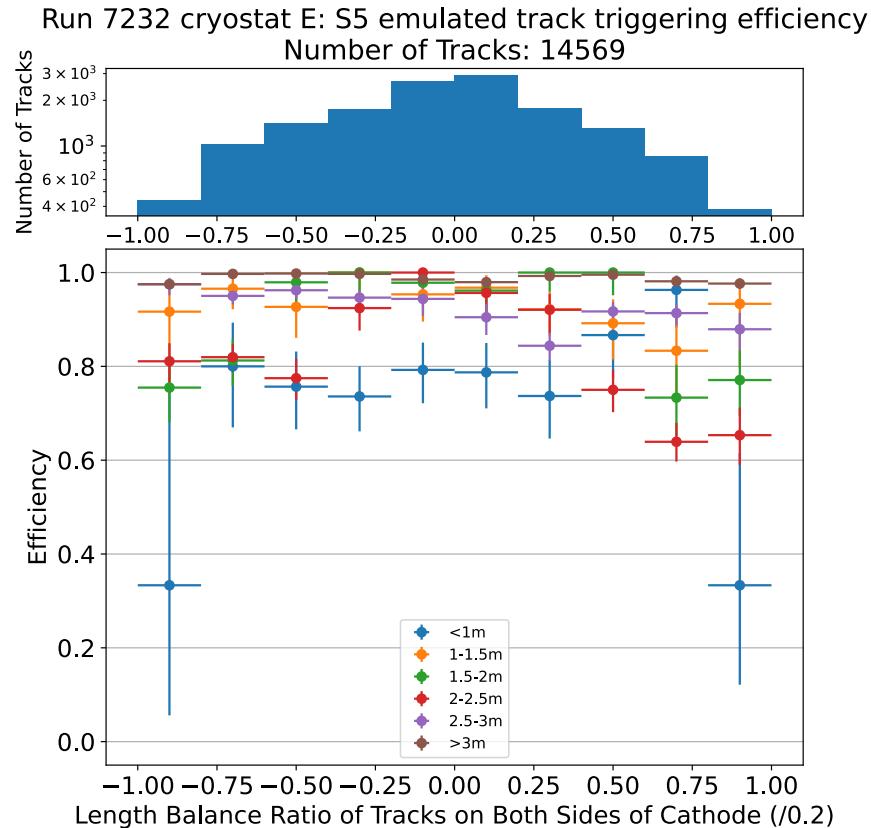


Possible connection:

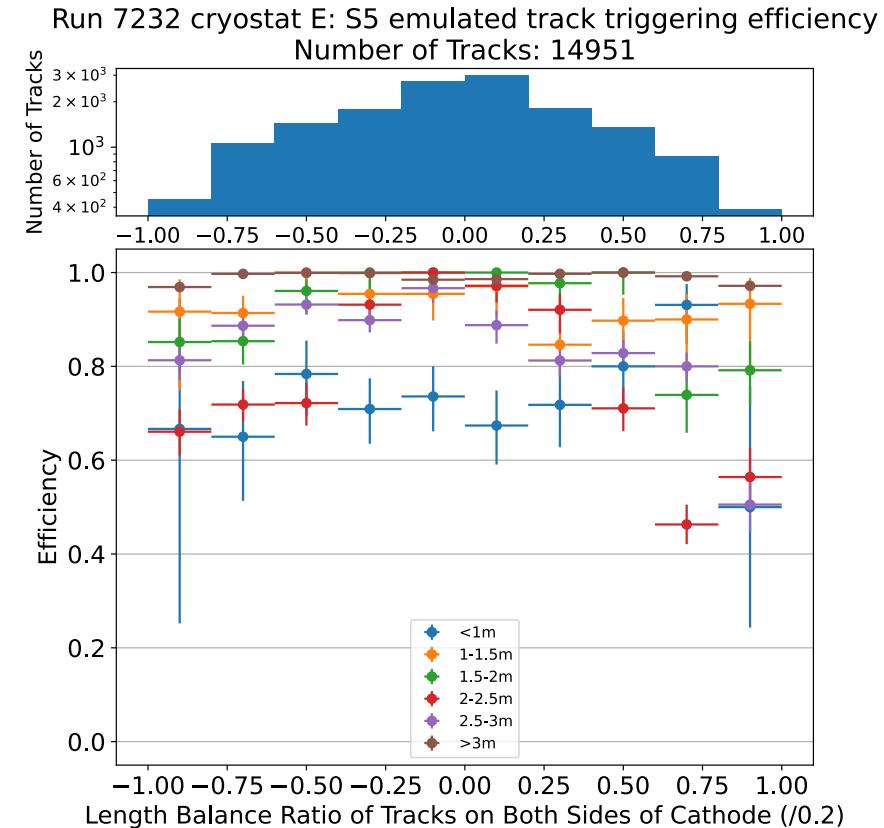
Length balance ratio of tracks on either side of cathode

East Cryostat

(-20 μ s, 0 μ s) Emulation Window



(-15 μ s, 5 μ s) Emulation Window



Balance Ratio	≈ -1	≈ 0	≈ 1
Primary TPC of Track	East	Both	West

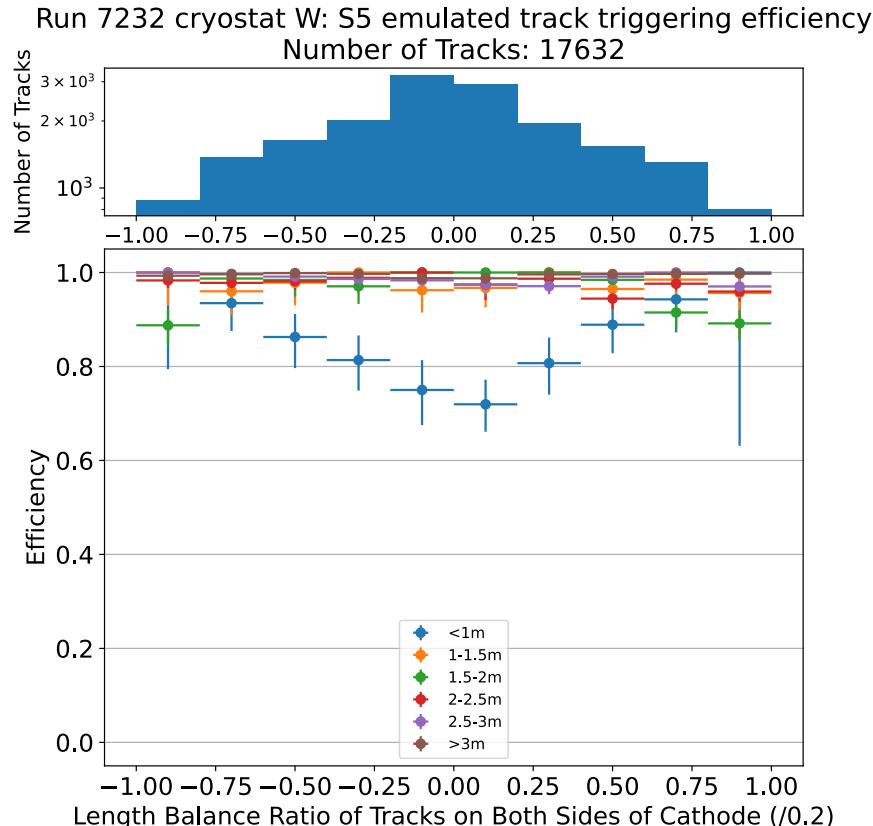
Possible connection:

Length balance ratio of tracks on either side of cathode

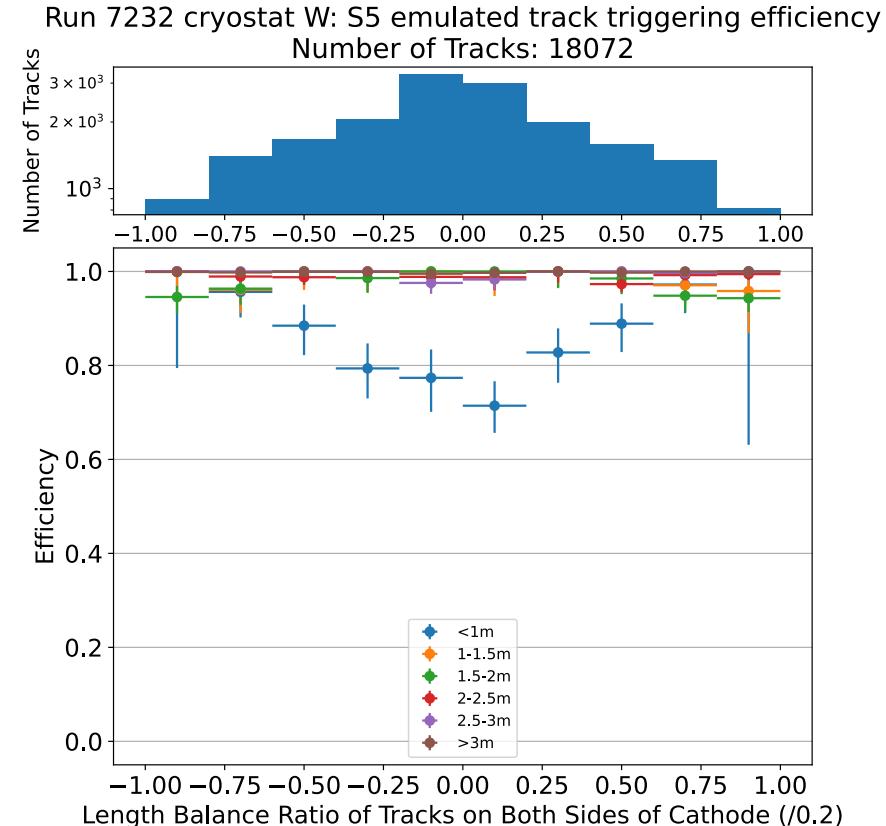
Balance Ratio	≈ -1	≈ 0	≈ 1
Primary TPC of Track	East	Both	West

West Cryostat

(-20 μ s, 0 μ s) Emulation Window

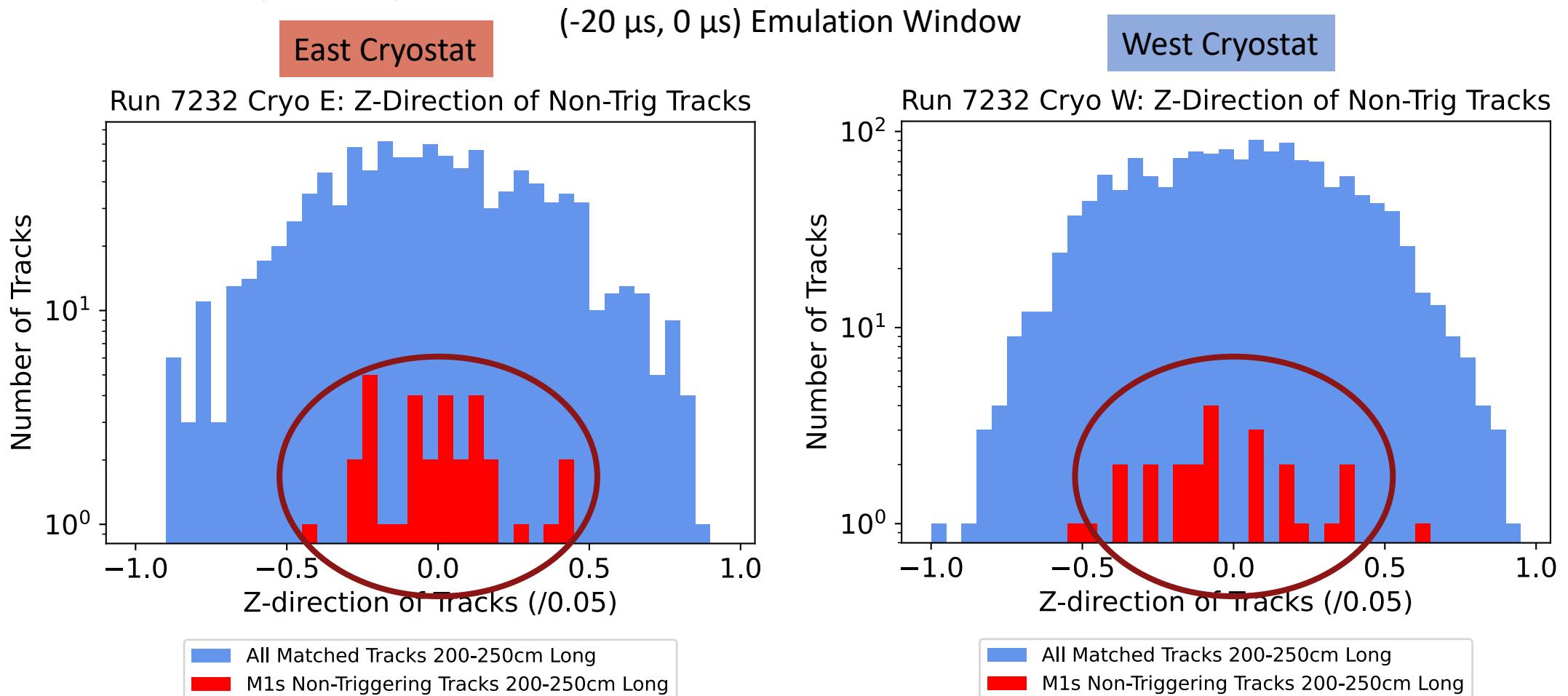


(-15 μ s, 5 μ s) Emulation Window



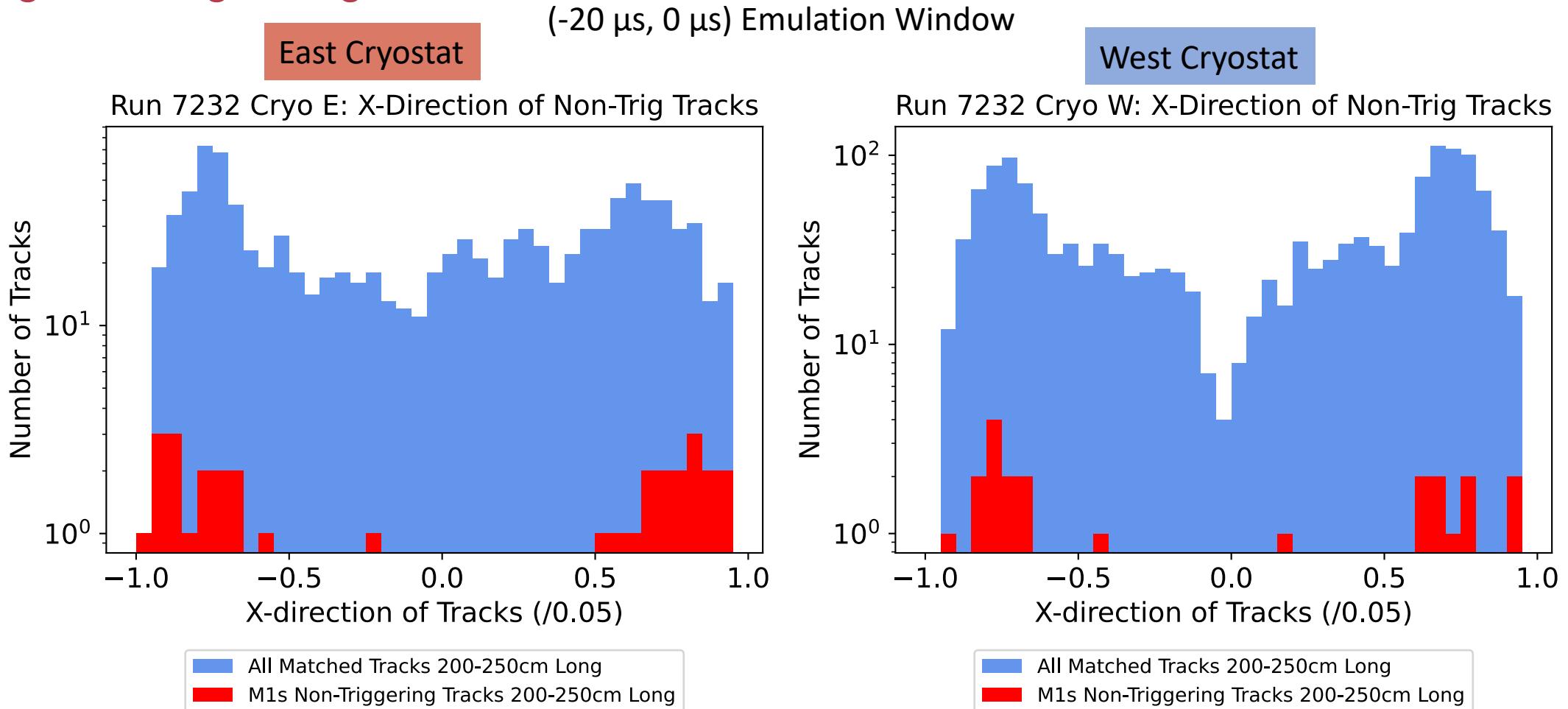
Possible Connection:

Smaller track angle along the z-axis



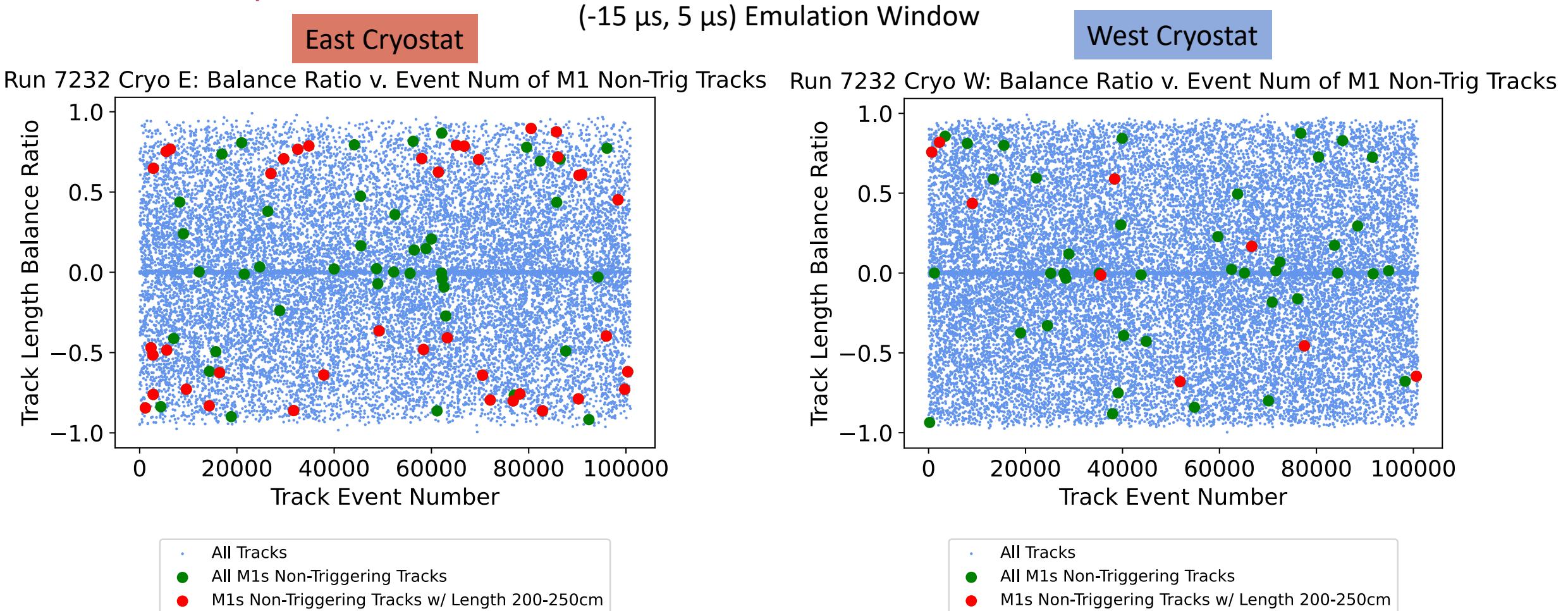
Possible Connection:

Large track angle along the x-axis



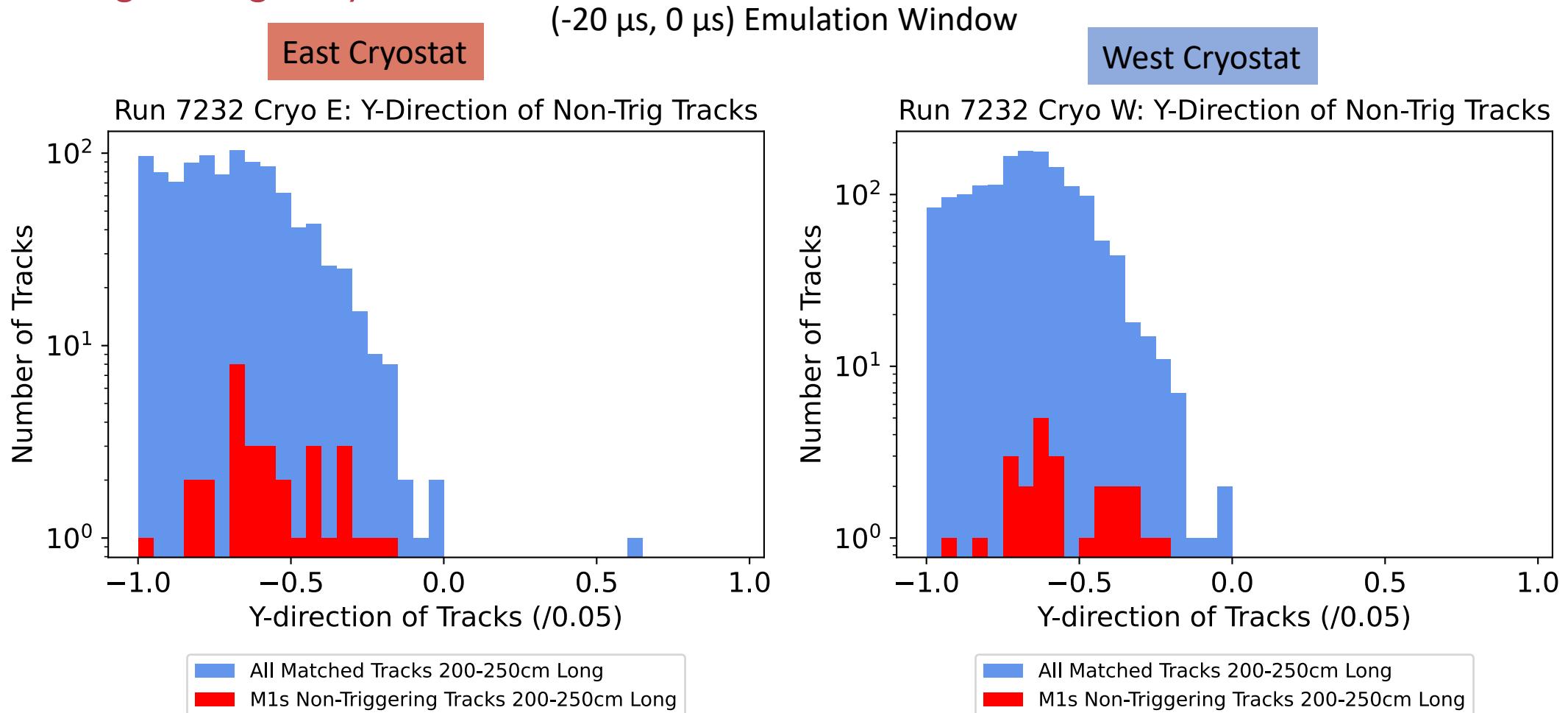
Ruled Out Connection:

Not time dependent



Ruled Out Connection:

Track angle along the y-axis



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Conclusions and Next Steps

Conclusions and Next Steps

- Still investigating the 2m track anomaly
- New minimum bias run (similar to this data) taken last week
 - Repeat these analyses to understand changes
 - Look for efficiency improvements or any new features to study
- Later analyze CRT-matched tracks
 - CRT (Cosmic Ray Tagger) is a set of sensors covering outside of detector
 - Detect cosmic particles entering or leaving the cryostat
 - CRT hits can be matched to tracks within the TPCs and can provide us with time and position information for tracks that don't necessarily cross the cathode

Recap

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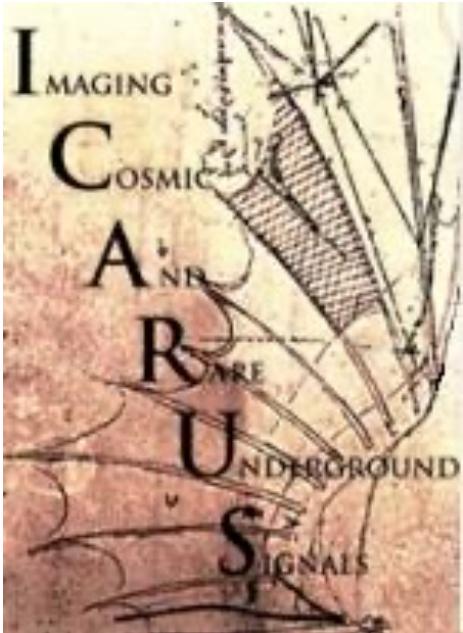
Other Possible Considerations

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Conclusions and Next Steps

Acknowledgements

I would like to acknowledge Dr. Gianluca Petrillo, the ICARUS Trigger Working Group, the SLAC Neutrino Group, Erin and Hillary, my fellow interns, and the DOE

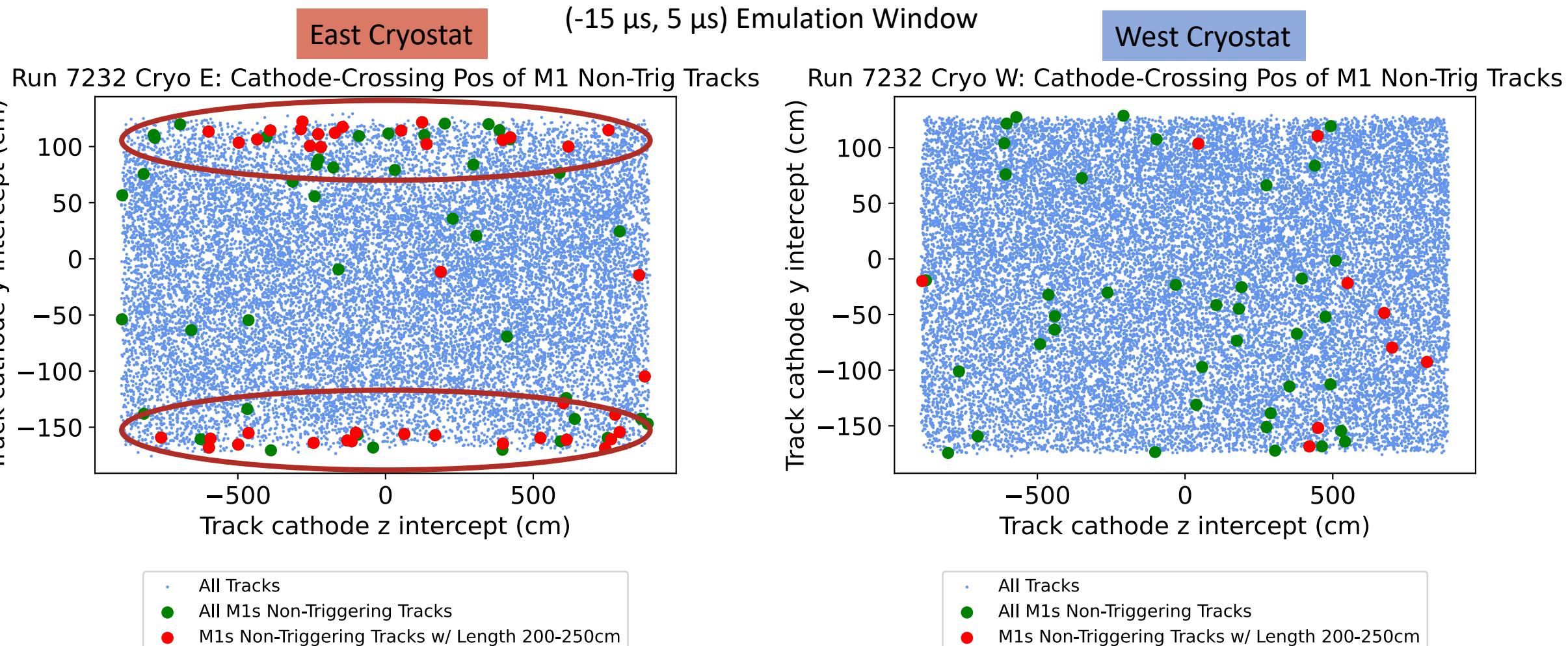


Thank you for listening! Any questions?

Additional Slides

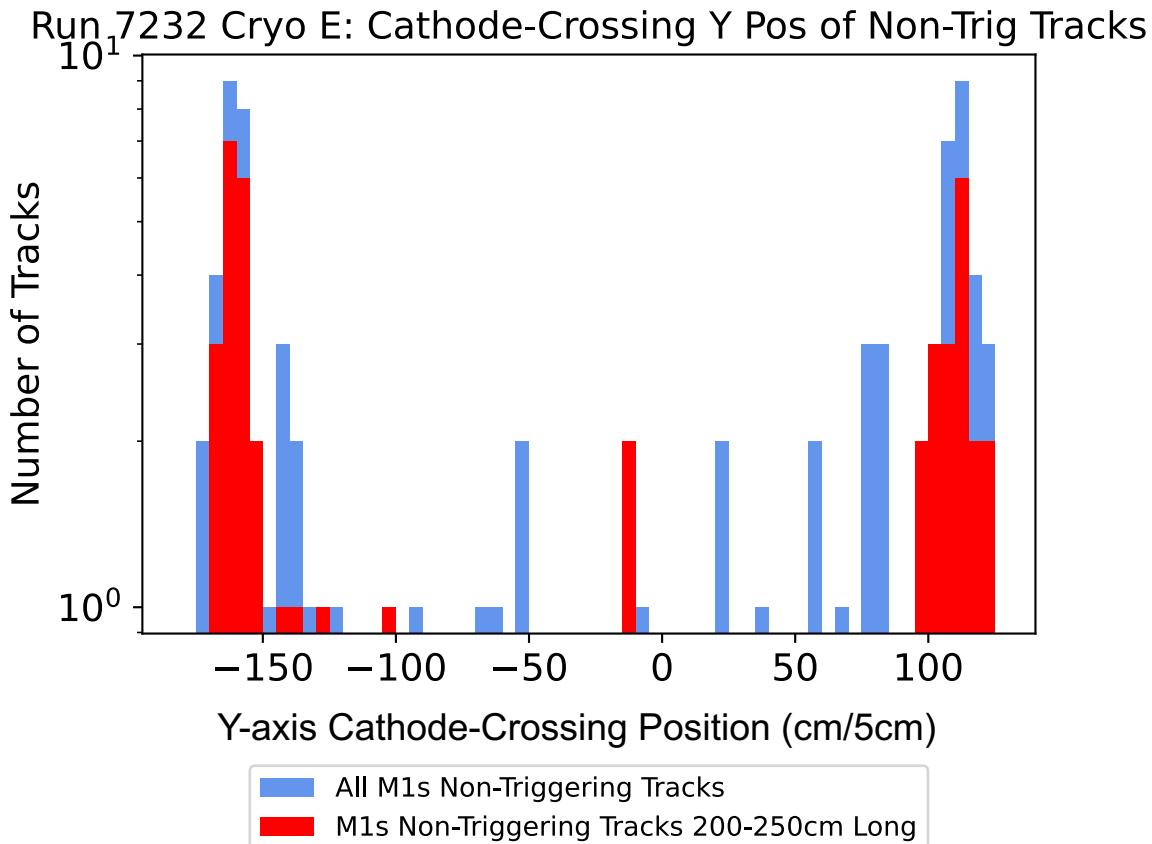
Possible connection:

Y-axis cathode-crossing position



Possible connection:

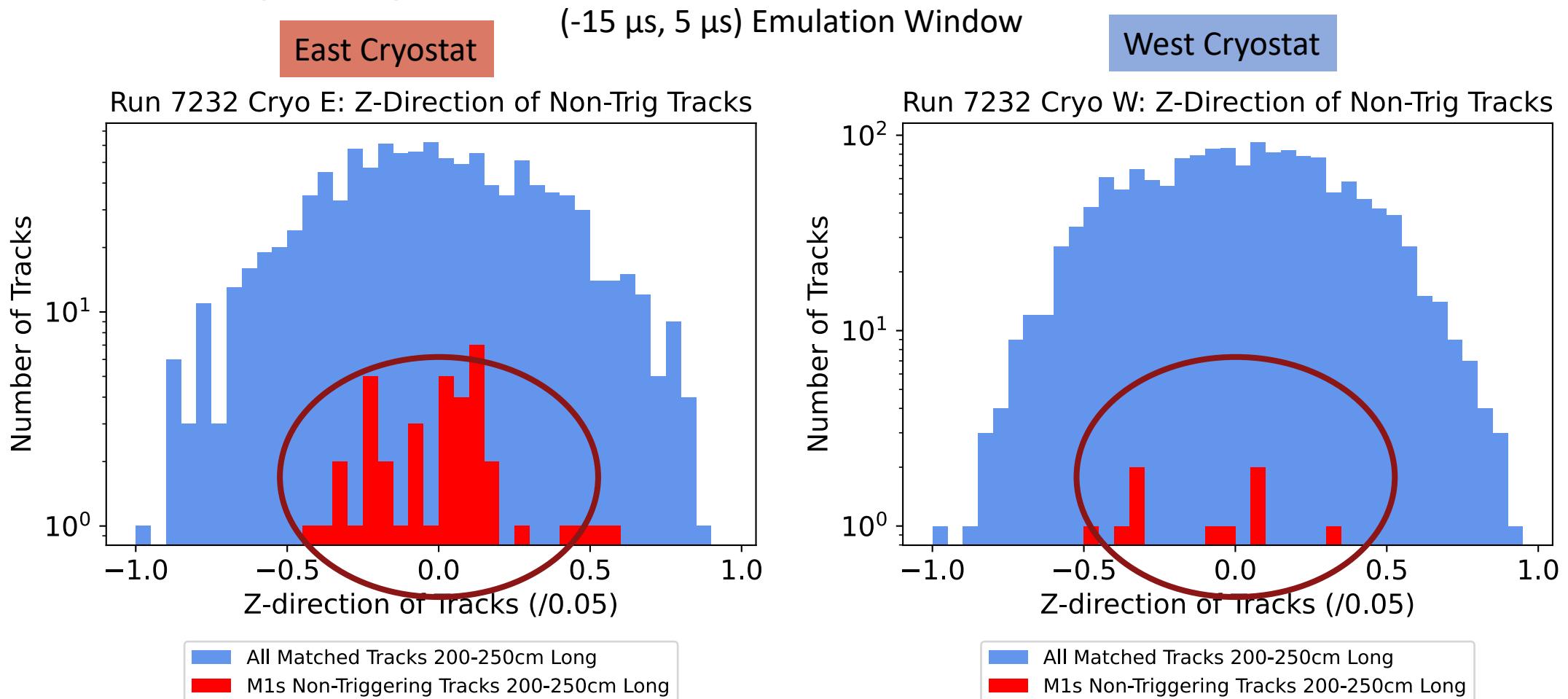
Y-axis cathode-crossing position



- Proportion of M1 non-triggering tracks with y-axis cathode-crossing position between -150 cm and 90 cm = **0.362**
- Number of M1s non-triggering tracks with lengths 200-250 cm = **42**
- Estimated Number of M1 non-triggering tracks with lengths 200-250 cm with y-axis cathode-crossing position between -150 cm and 90 cm = $0.36 \times 42 = 15.2$
- Actual Number of M1 non-triggering tracks with lengths 200-250 cm with y-axis cathode-crossing position between -150 and 90 cm = **6**

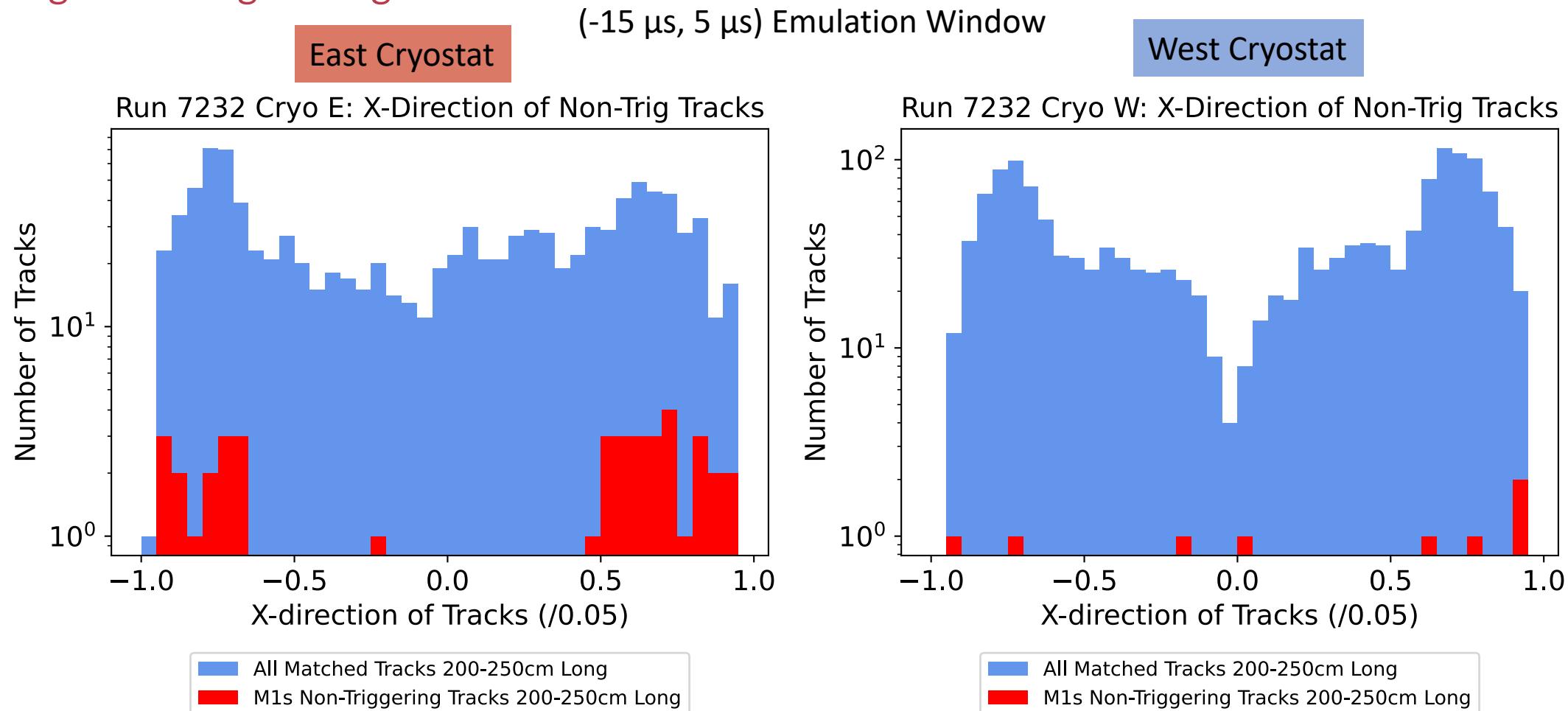
Possible Connection:

Smaller track angle along the z-axis



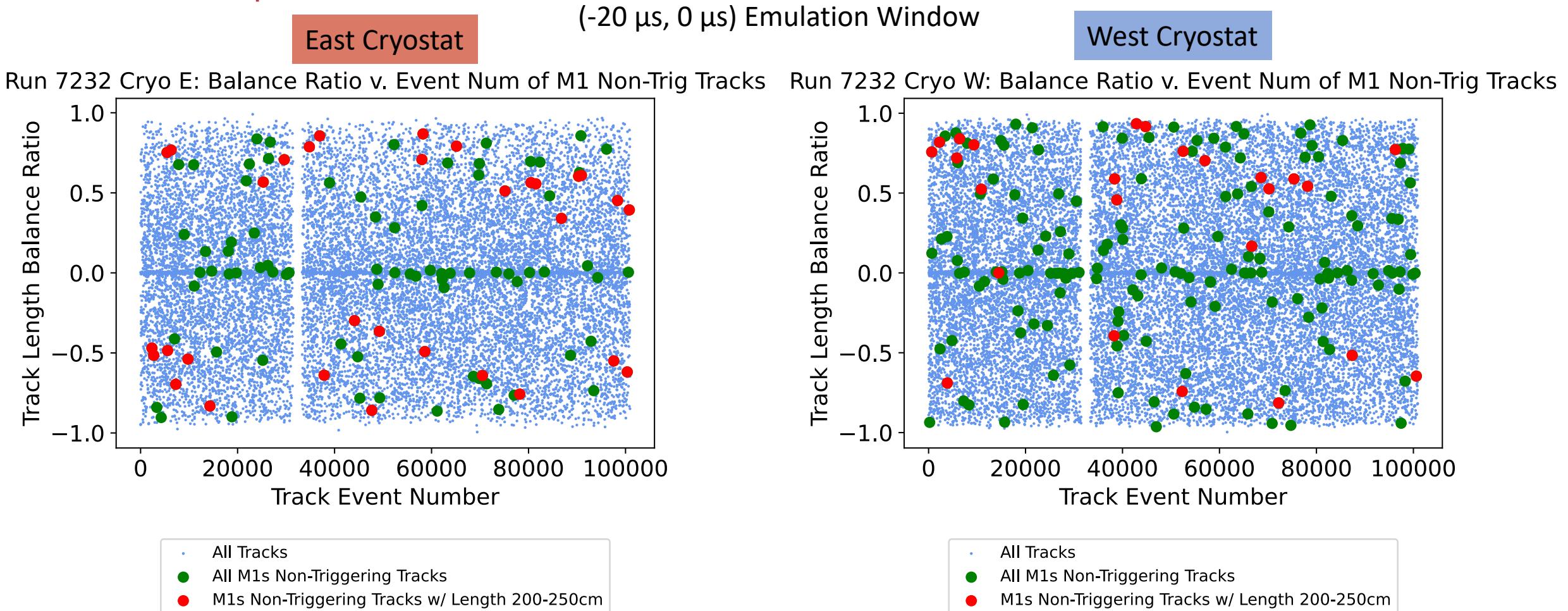
Possible Connection:

Large track angle along the x-axis



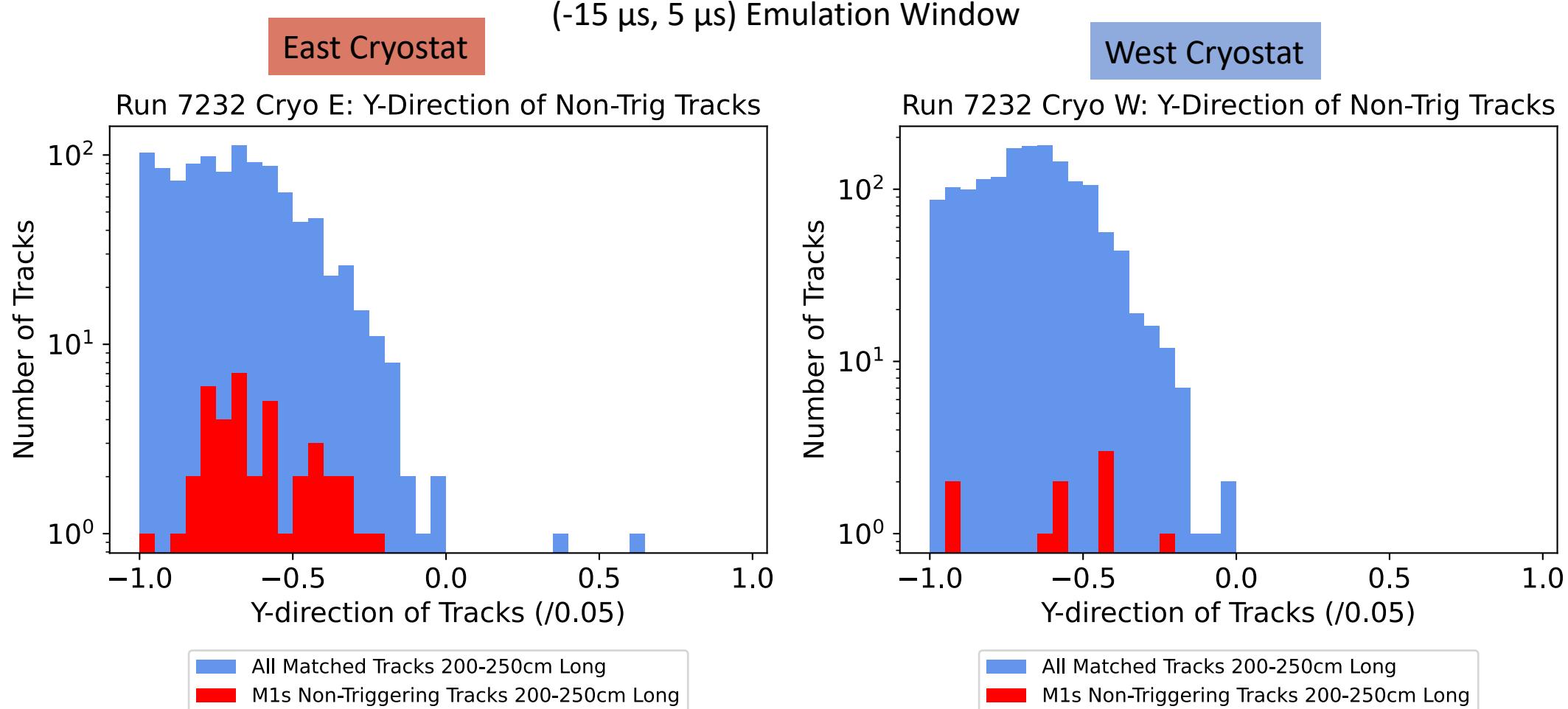
Ruled Out Connection:

Not time dependent



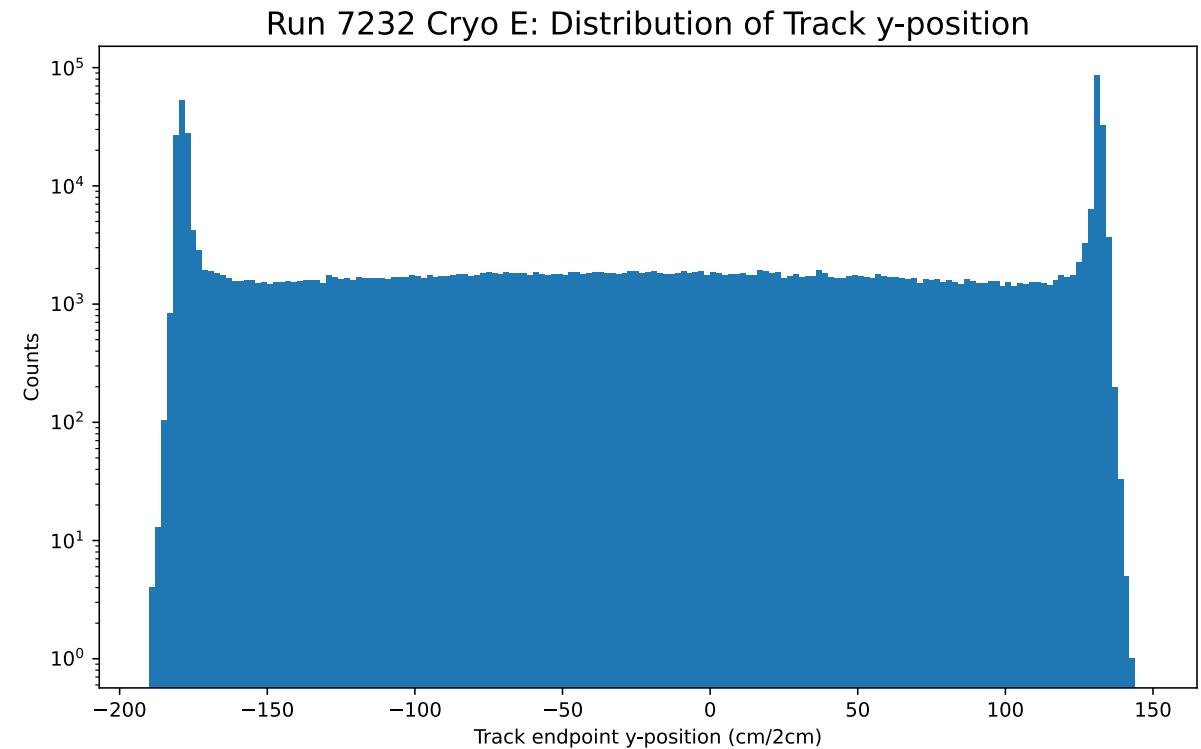
Ruled Out Connection:

Track angle along the y-axis

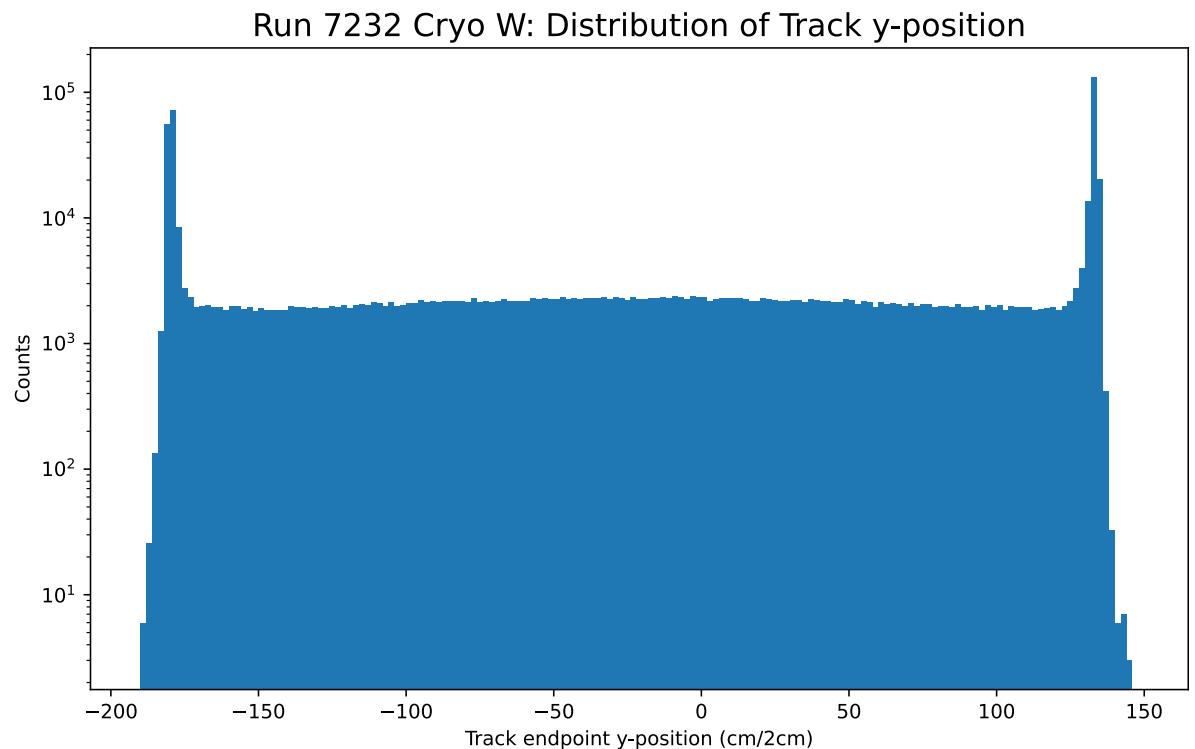


Y-axis Track Endpoint Distribution

East Cryostat



West Cryostat



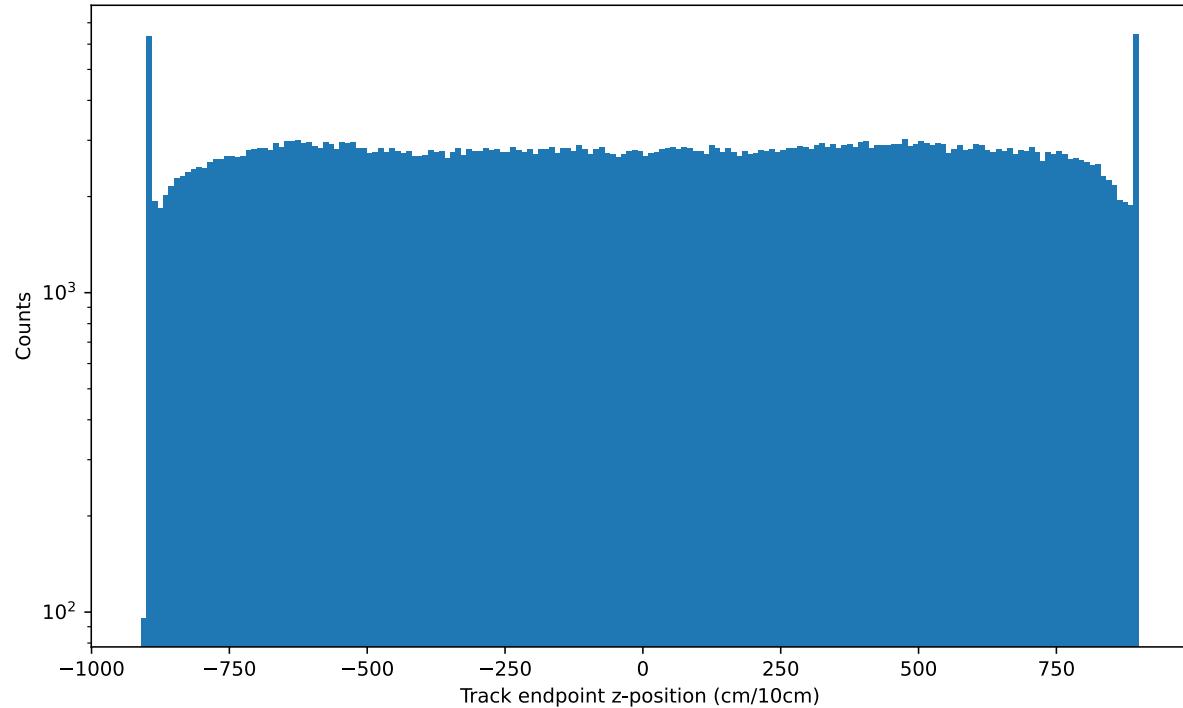
Z-axis Track Endpoint Distribution

East Cryostat

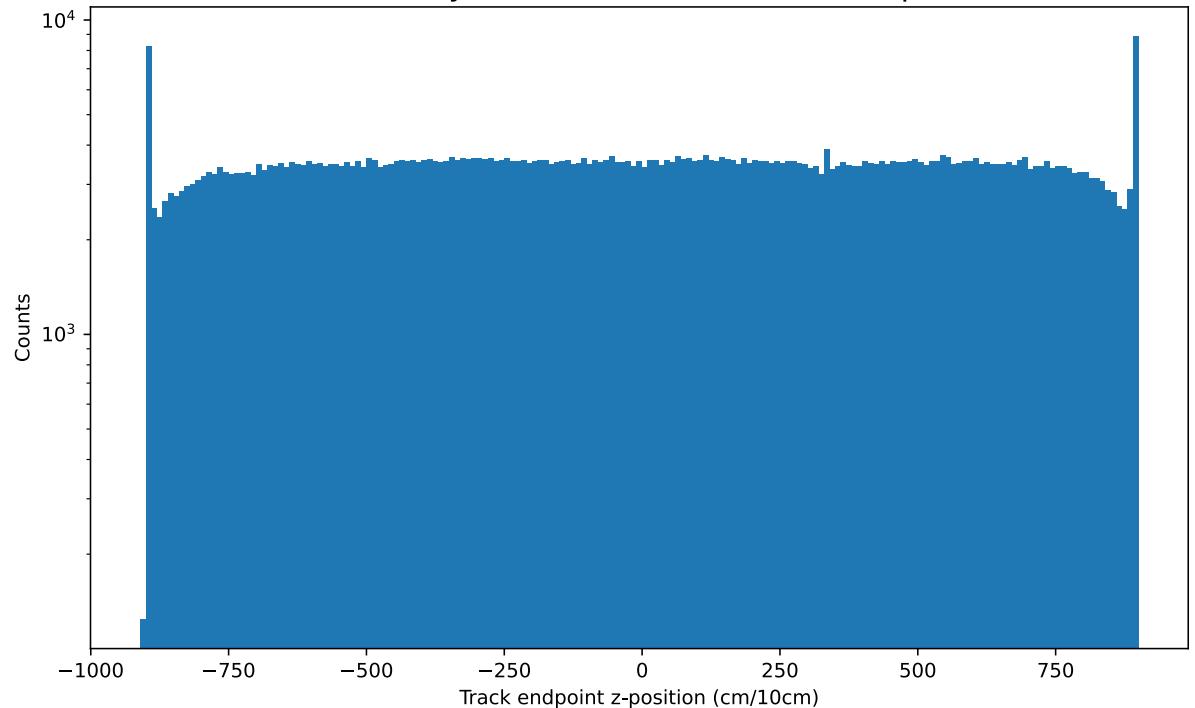
(-15 μ s, 5 μ s) Emulation Window

West Cryostat

Run 7232 Cryo E: Distribution of Track z-position



Run 7232 Cryo W: Distribution of Track z-position



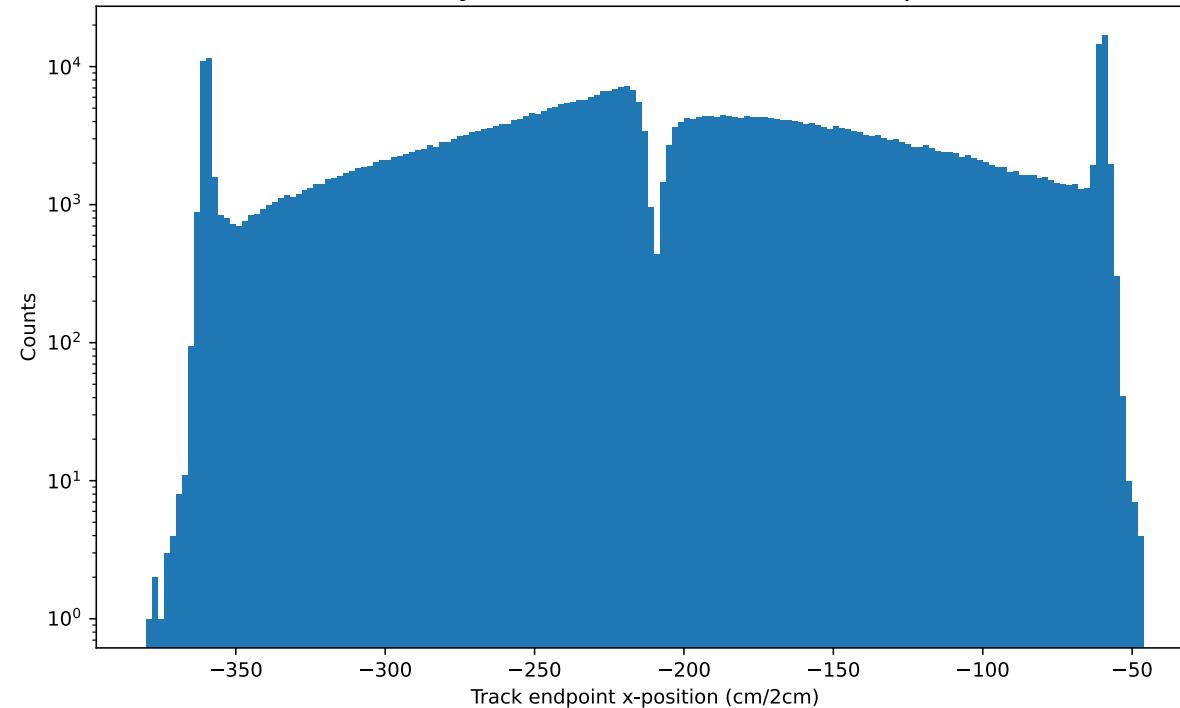
X-axis Track Endpoint Distribution

East Cryostat

(-15 μ s, 5 μ s) Emulation Window

West Cryostat

Run 7232 Cryo E: Distribution of Track x-position



Run 7232 Cryo W: Distribution of Track x-position

