Problem Solving

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Sheet of equations useful for the Problem solving paper

"Mo money mo Problems"

- Samson

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1 Mechanics

1.1 Constant acceleration

• For constant acceleration a the equations of motion can be written in three useful forms. The variables here are initial velocity u final velocity v, time t and displacement s:

$$v = u + at$$

$$v^{2} = u^{2} + 2as$$

$$s = ut + \frac{1}{2}at^{2}$$

1.2 Force

 $\bullet\,$ The force due to a potential U is:

$$\mathbf{F} = -\nabla U$$

1.3 Friction

• Friction is given by the following expression, where μ is the co-efficient of friction and N is the normal force acted on the object by the surface it is sliding across.

$$F_{\mathrm{Frict}} = \mu N$$

1.4 Work

• For constant Force work is just defined $W = \mathbf{F} \cdot \mathbf{s}$. (Force times displacement). If the force is not constant:

$$W = \int_{a}^{b} \mathbf{F} \cdot d\mathbf{s}$$

1.4.1 Power

• Power is defined as:

$$P = \frac{dW}{dr} = \mathbf{F} \cdot \mathbf{v}$$

Where \mathbf{v} is velocity.

1.5 Circular motion & Rotation

• In circular motion the force acting as the *centripetal* force (force pulling to wards the center, gravity, tension in rope, ect) is balanced by a *centrifugal* force equal in magnitude but opposite direction, given by:

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$$F_{\text{fugal}} = \frac{mv^2}{r} = mr\omega^2, \quad (\text{as } v = \omega r)$$

Where here v is the tangential velocity, r is the radius, m is the mass of the object in motion and ω is the angular velocity.

1.5.1 Rotational Kinetic energy

• Usually kinetic energy is just $\frac{1}{2}mv^2$, but for rigid bodies rotation it is:

$$K = \frac{1}{2}I\omega^2$$

Where I is the moment of inertia. A list of these can be found in the formula and tables booklet for different geometries.

1.5.2 Parallel axis Theorem

• This is the rule that tells us how to add the self rotation moment of inertia I_{self} to the orbital rotation of a mass M at radius R:

$$I_{\text{total}} = I_{\text{self}} + MR^2$$

1.5.3 Torque

• This is defined as:

$$oldsymbol{ au} = \mathbf{r} imes \mathbf{F}$$

We also have that:

$$|\boldsymbol{ au}|=\alpha I$$

Where $\alpha = \frac{d\omega}{dt}$. We can then find that the work done by a source the torque is:

$$W = \int_{\theta_0}^{\theta_1} \boldsymbol{\tau} d\theta, \quad \Longrightarrow P = \tau \omega$$

Where P is the power.

1.5.4 Angular momentum

• The standard definition is just:

 $\mathbf{L} = \mathbf{r} \times \mathbf{p}$, for a particle $I\boldsymbol{\omega}$, Ridged body motion

We can also write the torque as $\tau = \frac{d\mathbf{L}}{dt}$.

1.6 Center of Mass

• This is given by:

$$\mathbf{R}_{CM} = \frac{\sum_{i} m_{i} \mathbf{r}_{i}}{\sum_{i} m_{i}}$$

1.7 Stress/strain

• Shear modulus G is:

$$G = \frac{\text{shear stress}}{\text{shear strain}}$$

Where shear stress = F/A (Force per Area) and shear strain = $\frac{\Delta s}{l}$. Length displaced Δs over the length of the object l. The same sort of equation holds for the other moduli. Youngs modulus is G with the word "shear" replaced with "Tensile". Same goes with Bulk and Elastic.

1.8 Gravitation

• The Gravitational potential energy of a mass m in a gravitational field produced by a body of mass M is:

$$U = -\frac{GMm}{r} \implies \mathbf{F}_{\text{grav}} = -\frac{GMm}{r^2}\hat{r}$$

1.8.1 Acceleration due to gravity

 \bullet This is just the gravitational force $F_{\rm grav}$ per unit mass:

$$g = \frac{GM}{r^2}$$

1.9 Orbits

• To find the escape velocity we just find the velocity that makes the total energy 0 when r = R, this results in:

$$v_{\rm esc} = \sqrt{\frac{2GM}{R}}$$

The Schwartzchild radius is when this velocity v is the speed of light c.

1.9.1 Kepler's Period

• This is the orbital period for a body in elliptical motion with semi-major axis a:

$$T^2 = \frac{4\pi^2 a^3}{GM}$$

1.10 Periodic motion

• Simple Harmonic motion is defined as motion where the force is:

$$F = -kx$$

- This motion will then have a angular frequency ω , which is related to the period of the motion T, by: $T = 2\pi/\omega$.
- Frequency is defined as one over the period $f = 1/T = \omega/2\pi$.

1.10.1 Pendulum

• For a pendulum of length L we have that $\omega = \sqrt{\frac{g}{L}}$. If this pendulum is a rigid body with moment of inertia I, then this takes the form: $\omega = \sqrt{\frac{mgL}{I}}$.

1.10.2 Damped Harmonic Motion

• The General solution to damped harmonic motion takes the form:

$$x(t) = Ae^{-\gamma t}cos(\omega' t)$$

Where
$$\omega' = \sqrt{\frac{k}{m} - \frac{k^2}{4m^2}}$$
.

1.11 Density and Pressure

• Density is defined as:

$$\rho = \frac{m}{V}$$

• Pressure is in its most general form perpendicular Force F_{\perp} divided by area A:

$$P = \frac{F_{\perp}}{A}$$

1.11.1 Pressure in a fluid

• At a height h below the surface is:

$$P = P_0 + \rho g h$$

1.11.2 Fluid Flow

• For a fluid flowing through a pipe of changing cross sectional area the conserved quantity is the area A times the velocity of the fluid v. $(A_1v_1 = A_2v_2)$.

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1.11.3 Bernoulli's equation

• The relations ship between pressure and velocity in a fluid is given by:

$$P + \frac{1}{2}\rho v^2 = \text{constant}$$

1.11.4 Wave equation

• This takes the form:

$$\frac{\partial^2 y}{\partial x^2} = \frac{1}{v^2} \frac{\partial^2 y}{\partial t^2}$$

Where y - y(x,t) and $v = \lambda * f$ (wavelength by frequency) is the speed of the wave. The general form to the solutions of this equation are:

$$y(x,t) = A\cos(kx - \omega t)$$

1.11.5 Wave Power

• Given the tension T in a string with mass per length $\mu = m/L$ the wave power is:

$$P = \frac{1}{2}\sqrt{\mu T}\omega^2 A^2$$

1.11.6 Beat frequency

• When there is two sources of oscillatory motion the beat frequency is the difference between the two underlying frequencies $f_{\text{beat}} = |f_1 - f_2|$.

1.11.7 Doppler affect

• The observed frequency of a source emitting a frequency f_0 when the receiver is moving at velocity v_r and the source at velocity v_s is:

$$f = \left(\frac{c \pm v_r}{c \mp v_s}\right)$$

Where we have + in the numerator if receiver is moving towards source and a - for away. And we have a + in the denominator if the source is moving towards the medium and a - if away.

2 Thermodynamics

2.1 Heat

2.1.1 Heat Capacity

• Heat capacity is defined as the temperature derivative of the Energy:

$$C_V = \left(\frac{\partial E}{\partial T}\right)_V$$

For C_P just replace V with P.

• Heat energy change due to a temperature change ΔT for a substance with heat specific capacity $c \equiv C/m$ and mass m:

 $\Delta Q = mc\Delta T$

• For a mass m and specific latent heat l:

$$\Delta E = ml$$

2.1.3 Heat Flux

• This is the heat flux radiated by a source of temperature T with emissivity ε :

$$F = \varepsilon \sigma T^4$$

Where σ is the Stephan Boltzmann constant. Flux is also related to power P and area A via:

$$F = P/A$$

2.2 Ideal Gas

• The ideal Gas Law is:

$$PV = NkT$$

N is the number of particles and k the Boltzmann factor.

2.2.1 Energy of a Gas

• The relation for energy of a Gas is:

$$E = \frac{d}{2}NkT$$

Where d is the degrees of freedom of the gas molecules, for ideal Gas this is just d = 3.

 \bullet This can also be expressed as:

$$E = \frac{1}{2}mN\bar{v}^2$$

Where \bar{v} is the average velocity given by $\bar{v} = \sqrt{\frac{dkT}{m}}$

2.3 Mean free Path

• This is just defined as the velocity of the particles times the mean free time t_{mean} . Which can be expressed as:

$$\lambda = \frac{1}{\sigma n}$$

Where σ is the cross section and n the number of particles per unit volume. This is assuming stationary

Problem Solving 2 Thermodynamics