Two Coupled Swarmalators with Chirality

July 12, 2024

1 Reference

1.1 Farrell F D C, Marchetti M C, Marenduzzo D, et al. Pattern formation in self-propelled particles with density-dependent motility[J]. Physical review letters, 2012, 108(24): 248101.

Microscopic dynamics:

$$\dot{\mathbf{r}}_{i} = v\mathbf{e}_{\theta_{i}},$$

$$\dot{\theta}_{i} = \gamma \sum_{j=1}^{N} F(\theta_{j} - \theta_{i}, \mathbf{r}_{j} - \mathbf{r}_{i}) + \sqrt{2\epsilon}\tilde{\eta}_{i}(t).$$
(1)

The microscopic density of particles at position \mathbf{r} with angle θ is given by

$$f(\mathbf{r}, \theta) = \sum_{i=1}^{N} \delta(\mathbf{r} - \mathbf{r}_i) \, \delta(\theta - \theta_i) \,. \tag{2}$$

Using Itô's formula, a stochastic dynamical equation for the density Eq. (2) can be derived:

$$\partial_t f(\mathbf{r}, \theta) + \mathbf{e}_{\theta} \cdot \nabla [vf]$$

$$= \epsilon \frac{\partial^2 f}{\partial \theta^2} - \frac{\partial}{\partial \theta} \sqrt{2\epsilon f} \eta - \gamma \frac{\partial}{\partial \theta} \int d\theta' d\mathbf{r}' f(\mathbf{r}', \theta') \times f(\mathbf{r}, \theta) F(\theta' - \theta, \mathbf{r} - \mathbf{r}') .$$
(3)

Drop the noise term, and Fourier transform Eq. (3) to get equations of motion for

$$f_k \equiv \int f(\mathbf{r}, \theta) e^{ik\theta} d\theta$$
 (4)

2 Our Work

We replace the finite range alignment interaction by a pseudopotential (δ -interaction) in the model:

$$\dot{\mathbf{r}}_i = v\mathbf{p}_i$$

$$\dot{\theta}_i = \omega_i + \lambda \sum_{j \neq i} \delta\left(\mathbf{r}_j - \mathbf{r}_i\right) \sin\left(\theta_j - \theta_i\right)$$
(5)

where $\mathbf{p}_i = (\cos \theta_i, \sin \theta_i)$. Assuming that we have M species, consisting of N_1, \dots, N_M particles with identical frequencies $\tilde{\Omega}_1, \dots, \tilde{\Omega}_M$ respectively, and that N_1, \dots, N_M are all macroscopic in an area element over which macroscopic quantities (density, polarization) vary significantly, allows us to derive a

continuum theory for the particle dynamics. The combined probability density to find a particle of given species j at position \mathbf{r} with angle θ at time t is given by

$$f^{(j)}(\mathbf{r}, \theta, t) = \sum_{i=1}^{N} \delta(\mathbf{r} - \mathbf{r}_{i}(t)) \delta(\theta - \theta_{i}(t)) \delta_{\Omega_{i}, \tilde{\Omega}_{j}}.$$
 (6)

A Boltzmann-like equation for the combined density $f^{(j)}$ can be derived:

$$\dot{f}^{(j)}(\mathbf{r},\theta,t) = -\operatorname{Pe}\mathbf{p} \cdot \left[\nabla_{\mathbf{r}} f^{(j)}(\mathbf{r},\theta,t)\right] - \Omega_{j} \partial_{\theta} f^{(j)}(\mathbf{r},\theta,t) + \partial_{\theta}^{2} f^{(j)}(\mathbf{r},\theta,t)
- g \partial_{\theta} \left[f^{(j)}(\mathbf{r},\theta,t) \int d\theta' \sin(\theta'-\theta) \sum_{i=1}^{M} f^{(i)}(\mathbf{r},\theta',t)\right] - \partial_{\theta} \sqrt{2f^{(j)}(\mathbf{r},\theta,t)} \eta_{j}(\mathbf{r},\theta,t)$$
(7)

where η_j represents spatiotemporal white noise with zero mean and unit-variance (the subscript j denotes that the noise-realization of a given ensemble is individual for each species).

In the following, we focus on a mean-field description and neglect the (multiplicative) noise term in Eq. (7). Now transforming Eq. (7) to Fourier space yields a hierarchy of dynamical equations for the Fourier modes $f_k^{(j)}(\mathbf{r},t) = \int \mathrm{d}\theta f^{(j)}(\mathbf{r},\theta,t) \mathrm{e}^{\mathrm{i}k\theta}$ and $2\pi f^{(j)}(\mathbf{r},\theta,t) = \sum_{k=-\infty}^{\infty} f_k^{(j)}(\mathbf{r},t) \mathrm{e}^{-\mathrm{i}k\theta}$, reading

$$\dot{f}_{k}^{(j)}(\mathbf{r},t) = -\frac{\text{Pe}}{2} \left[\partial_{x} \left(f_{k+1}^{(j)} + f_{k-1}^{(j)} \right) - i \partial_{y} \left(f_{k+1}^{(j)} - f_{k-1}^{(j)} \right) \right]
+ \left(ik\Omega_{j} - k^{2} \right) f_{k}^{(j)} + \frac{igk}{2\pi} \sum_{m=-\infty}^{\infty} f_{k-m}^{(j)} F_{-m} \sum_{i=1}^{M} f_{m}^{(i)}$$
(8)

After a long calculation, we find the following equations

$$\dot{\rho} = -v\nabla \cdot \mathbf{w}$$

$$\dot{\mathbf{w}} = (\lambda \rho - 2)\frac{\mathbf{w}}{2} - \frac{v}{2}\nabla \rho + \frac{v^2}{2b}\nabla^2 \mathbf{w} - \frac{\lambda^2}{b}|\mathbf{w}|^2 \mathbf{w}$$

$$+ \frac{\lambda v}{4b} \left[5\nabla \mathbf{w}^2 - 10\mathbf{w}(\nabla \cdot \mathbf{w}) - 6(\mathbf{w} \cdot \nabla)\mathbf{w} \right]$$

$$+ \omega \mathbf{w}_{\perp} + \frac{v^2 \omega}{4b}\nabla^2 \mathbf{w}_{\perp} - \frac{\lambda^2 \omega}{2b}|\mathbf{w}|^2 \mathbf{w}_{\perp}$$

$$- \frac{\lambda v \omega}{8b} \left[3\nabla_{\perp} \mathbf{w}^2 - 6\mathbf{w} \left(\nabla_{\perp} \cdot \mathbf{w} \right) - 10 \left(\mathbf{w} \cdot \nabla_{\perp} \right) \mathbf{w} \right]$$
(9)

where

$$\omega = \langle \omega_i \rangle + \omega (\mathbf{x}, t) + \sqrt{\frac{\Delta_\omega}{f}} \eta$$

$$b = 2 (4 + \omega^2)$$

$$\mathbf{w}_\perp = (-w_y, w_x)$$

$$\nabla_\perp = (-\partial_y, \partial_x)$$
(10)