

# Hamiltonian Systems

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- 1 Brief overview of Classical Mechanics
- 2 General Hamiltonian Systems
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# Very classical mechanics

For a particle in 3-space, we have the position  $\begin{bmatrix} x(t) \\ y(t) \\ z(t) \end{bmatrix} = r(t) \in \mathbb{R}^3$

$$\vec{v} = \frac{d}{dt}r, \quad \vec{a} = \frac{d}{dt}\vec{v}, \quad \vec{p} = m\vec{v}.$$

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The particle may be subject to a force  $\vec{F}(r, \vec{v}, t)$ , obeying Newton's laws:

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Several particles in one "vector":

$$R = \begin{bmatrix} x_1 \\ y_1 \\ \vdots \\ y_l \\ z_l \end{bmatrix}, \quad \begin{bmatrix} m_1 & & & \\ & m_1 & & \\ & & \ddots & \\ & & & m_l \\ & & & & m_l \end{bmatrix} \begin{bmatrix} \ddot{x}_1 \\ \ddot{y}_1 \\ \vdots \\ \ddot{y}_l \\ \ddot{z}_l \end{bmatrix} \stackrel{!}{=} \begin{bmatrix} F_{1,x} \\ F_{1,y} \\ \vdots \\ F_{l,y} \\ F_{l,z} \end{bmatrix}$$

# Motivation

Consider  $l$  particles in 3-space on holonomic constraints:

$$\underline{q} = (q^i) \mapsto R = [R^\lambda] \in N = f^{-1}[\{0\}] \cap \text{some open set} \stackrel{\text{SM}}{\subseteq} \mathbb{R}^{3l}.$$

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The masses  $m_\lambda$  give the (extrinsic) momentum  $P$  on a trajectory as a Riemannian metric  $m$  as on the last slide. This descends to  $N$ :

$$\bar{P} = m(\bar{\dot{R}}) \Rightarrow \tilde{p} = m(\vec{\dot{R}}) \doteq (p_i) = (m_{ij}\dot{q}^j)$$

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Newton's equation (extrinsic):

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Newton's equation (extrinsic):

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Kinetic energy:

$$2T = \sum_{\lambda} m_{\lambda} (v^{\lambda})^2 = \|\dot{R}\|_m^2 = \|P\|_{m^{-1}}^2 = m_{ij} v^i v^j = m^{ij} p_i p_j$$

# Generalized forces

The parallel force  $\tilde{Q} = Q_i \varepsilon^i = g(\vec{F}_{\parallel})$  may later depend on velocity, but here we consider conservative forces with the potential  $V \in C^\infty(N, \mathbb{R})$

$$\vec{F} = -\vec{\nabla}V = -g^{-1}\left(\frac{\partial V}{\partial r}\right), \quad \text{i.e.} \quad \tilde{Q} = -d\tilde{V} = -\frac{\partial V}{\partial q^i} dq^i$$

Newton's equation becomes:

$$\tilde{Q} = m\left(\frac{D}{dt}\vec{v}\right) = (\dot{v}^j)\vec{e}_j + v^i v^j \tilde{\Gamma}_{ij}$$

The potential and kinetic energy constitute the **Hamiltonian** or total energy function

$$H(q^i, p_i, t) = T(q^i, p_i, t) + V(q^i, p_i, t)$$

# Example 1 - $l$ -body problem

Unrestricted motion of  $l$  celestial bodies with gravitation, i.e. the forces

$$\vec{F}_i = \sum_{j \neq i} -\frac{m_i m_j}{4\pi |r_i - r_j|^2} \vec{e}_{r_i - r_j}$$

generated by the shared potential

$$4\pi V(r_1, \dots, r_l) = -\frac{m_1 \cdot m_2}{|r_1 - r_2|} - \dots - \frac{m_1 \cdot m_l}{|r_1 - r_l|} - \frac{m_2 \cdot m_3}{|r_2 - r_3|} - \dots$$

This has the Hamiltonian

$$H(r_1, \dots, r_l, \vec{p}_1, \dots, \vec{p}_l) = \sum_i \frac{|\vec{p}_i|^2}{2m_i} - \sum_{i < j} \frac{m_i m_j}{4\pi |r_i - r_j|}$$

We have to remove the collision set  $\{R \mid V = \infty\}$ , so  $N = \text{dom } V$ .

## Lagrangian mechanics

- $M = TN \ni \vec{v} @ R \doteq \underline{v} @ \underline{q} \stackrel{!}{=} \dot{\underline{q}} @ \underline{q}$
- Lagrangian:  $L = T - V \in \mathcal{F}(M)$
- We have the canonical momenta  $p_i = m_{ij} v^j = \frac{\partial}{\partial \dot{q}^i} T = \frac{\partial}{\partial \dot{q}^i} L$ .
- Lagrange equation:  $\dot{q}^j = v^j, \quad \dot{p}_j = Q_j + \frac{\partial T}{\partial q^j} = \frac{\partial L}{\partial q^j}$

# Abstraction to the phase space

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## Hamiltonian mechanics

- $M = T^*N \ni \hat{p} @ R \doteq \underline{p} @ \underline{q}$
- Hamiltonian:  $H = \hat{p} \hat{v} - L = T + V \in \mathcal{F}(M)$
- $v^j = m^{ij} p_i = \frac{\partial}{\partial p_i} T = \frac{\partial}{\partial p_i} H$
- Hamilton equations:  $\dot{q}^j = \frac{\partial H}{\partial p_j}, \quad \dot{p}_j = -\frac{\partial H}{\partial q^j}$

They are duals via a Legendre transformation.

## Example 2 - Central potential

If  $V(R) = V\left(\begin{pmatrix} r \cos \varphi \\ r \sin \varphi \end{pmatrix}\right) = V(r, \varphi)$  only depends on  $r$ , we have

$$F = \frac{\partial V}{\partial r} \cdot (-\vec{e}_r) \quad \text{or} \quad (Q_r, Q_\varphi) = \underline{g} \begin{pmatrix} -\partial_r V \\ 0 \end{pmatrix} = (-\partial_r V, 0)$$

The canonical momenta are

$$\begin{pmatrix} p_r \\ p_\varphi \end{pmatrix} = \begin{pmatrix} m & 0 \\ 0 & mr^2 \end{pmatrix} \begin{pmatrix} \dot{r} \\ \dot{\varphi} \end{pmatrix}$$

We can identify  $L = p_\varphi = mr^2\dot{\varphi}$  as the angular momentum.

$$H = \frac{1}{2} m^{ij} p_i p_j + V(\underline{q}) = \frac{p_r^2}{2m} + \frac{p_\varphi^2}{2mr^2} + V(r)$$

## Example 2 - Central potential

$$H = \frac{p_r^2}{2m} + \frac{p_\varphi^2}{2mr^2} + V(r)$$

Hamiltonian equations:

$$\begin{pmatrix} \dot{r} \\ \dot{\varphi} \end{pmatrix} = \frac{\partial H}{\partial p} = \begin{pmatrix} \frac{p_r}{m} \\ \frac{p_\varphi}{mr^2} \end{pmatrix}$$

$$(\dot{p}_r, \dot{p}_\varphi) = -\frac{\partial H}{\partial q} = \left( -\frac{\partial V}{\partial r} + \frac{p_\varphi^2}{mr^3}, 0 \right)$$

Notice the “fictional” centrifugal force  $m\dot{\varphi}^2 r$ .  $p_\varphi$  is conserved, because  $\varphi$  is a **cyclic coordinate**.

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# Generalization

We consider a system with  $n$  degrees of freedom, described by a smooth  $n$ -dimensional manifold  $N$ , called the **configuration space** and its cotangent bundle  $M = T^*N$ , the **phase space** with a function  $H$ .

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With canonical coordinates  $\underline{u} = \begin{pmatrix} \underline{q} \\ \underline{p} \end{pmatrix} \mapsto U \in M$  the Hamiltonian equations turn into:

$$\dot{\underline{U}} \doteq \underline{\dot{U}} = \begin{pmatrix} \dot{\underline{q}} \\ \dot{\underline{p}} \end{pmatrix} \stackrel{!}{=} \begin{pmatrix} \frac{\partial H}{\partial p_j} \\ -\frac{\partial H}{\partial q^i} \end{pmatrix} = \underbrace{\begin{pmatrix} 0 & 1_n \\ -1_n & 0 \end{pmatrix}}_{(J^{\mu\nu})} \frac{\partial H^\top}{\partial \underline{u}} = \underline{J}(\underline{DH}) \doteq -\Omega(dH) =: \underline{X}_H$$

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Here  $\Omega = \omega^{-1}$ , where  $\omega = d\lambda = \sum_i dp^i \wedge dx^i$  with the Liouville 1-form

$$\lambda(z) = \underbrace{\pi_{TT^*N \rightarrow T^*N}(z)}_{\text{Basepoint } \tilde{p}} \underbrace{(T\pi_{T^*N \rightarrow N}(z))}_{\text{spatial part of } z} = \tilde{p}(\vec{v}) \Leftrightarrow \lambda_{\tilde{p}} = \sum_i p^i dx^i$$

which has the matrix representation  $\omega_{\mu\nu} = J^{\mu\nu} = -\Omega^{\mu\nu}$

# General Definition of Hamiltonian Systems

Let  $(M, \omega)$  be a symplectic manifold, with tangent-cotangent isomorphism

$$\omega : TM \xrightarrow{\cong} T^*M, z \mapsto \omega(z, \cdot) \quad \text{and} \quad \Omega = \omega^{-1}.$$

The triple  $(M, \omega, H)$  is called a **Hamiltonian system** where  $H \in \mathcal{F}(M)$  is the **Hamiltonian** or energy function.

**Canonical coordinates**  $\underline{u} = (u^\mu) = (x^1, \dots, x^n, p^1, \dots, p^n)^\top$  are such that

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## Important Definition

The **Hamiltonian vector field** associated to  $f \in \mathcal{F}(M)$  is

$$X_f := -\Omega(df) \quad \text{i.e.} \quad \omega(X_f, z) = -df(z) = -\partial_z f.$$

The Hamiltonian equations now are:

$$\frac{d}{dt}\gamma = -\Omega(dH)\circ\gamma = X_H\circ\gamma$$

The **trajectories** of the Hamiltonian system (solutions to the Hamiltonian equations) are exactly the integral curves of  $X_H$ .

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These exist at least locally (first order ODE):

$$U_0 \mapsto \gamma_{U_0} : t \mapsto \gamma_{U_0}(t) \leftarrow U_0 : \psi_{X_H}^t \leftarrow t$$

This is the **Hamiltonian flow** (flow of  $X_H$ ).

# Special vector fields

A vector field  $z \in \mathcal{X}(M)$  is ...

- **symplectic** iff  $\omega$  is preserved along the integral curves of  $z$ :  $\psi_z^{t*} \omega = \omega$ ,  
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That follows from the fact that the Lie-derivative wrt  $z$  commutes with the flow of  $z$ : (at point  $P' = \psi_z^t(P) = \gamma_P(t)$ )

$$\frac{d}{dt} \psi_z^{t*}(\omega) = \frac{d}{ds} \Big|_{s=0} (\psi_z^{t+s})^*(\omega) = \frac{d}{ds} \Big|_{s=0} \psi_z^{t*}(\psi_z^{s*} \omega) = \psi_z^{t*}(\mathcal{L}_z \omega)$$

So if  $\mathcal{L}_z \omega = 0$ ,  $\psi_z^{t*}(\omega(\gamma_P(t)))$  must be constant wrt  $t$ ;  $\psi_z^{0*}(\omega(\gamma_P(0))) = \omega(P)$ .  
If not, then there is some place where  $\psi_z^{t*}(\omega(\gamma_P(t))) \neq \psi_z^{0*}(\omega(P))$

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## Relation between these definitions

$z$  symplectic  $\Leftrightarrow \omega(z)$  closed  $\Leftrightarrow \omega(z)$  is locally exact  $\Leftrightarrow z$  locally Hamiltonian.

$$\mathcal{L}_z \omega \stackrel{\text{Cartan}}{=} d(\iota_z \omega) + \iota_z d\omega = d(\omega(z))$$

One also defines the **Poisson bracket** of  $g, f \in \mathcal{F}(M)$ :

$$\{g, f\} := \omega(X_g, X_f) = df(X_g) = \partial_{X_g} f = -\Omega(dg, df) = -\Omega^{\mu\nu} \cdot \partial_\mu g \cdot \partial_\nu f$$

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A quantity  $f$  is conserved iff it Poisson commutes with  $H$ , because on a trajectory  $\gamma$ :

$$\frac{d}{dt}(f \circ \gamma) = \partial_{X_H} f \circ \gamma = \{H, f\} \circ \gamma$$

Energy is conserved along trajectories, because  $\{H, H\} = 0$ .

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Canonical commutation relations:  $\{u^\mu, u^\nu\} = -\Omega^{\kappa\iota} \partial_\kappa u^\mu \partial_\iota u^\nu = -\Omega^{\mu\nu} = J^{\mu\nu}$

Hamiltonian equations:  $\frac{d}{dt} u^\mu \circ \gamma = \{H, u^\mu\} \circ \gamma = -\Omega^{\kappa\mu} \frac{\partial H}{\partial u^\kappa} = X_H^\mu$

$$X_{\{f,g\}} = [X_f, X_g]$$

The Poisson bracket generalizes to the commutator (divided by  $i\hbar$ ) in quantum mechanics.

$$\left[ (M, \omega) \xrightarrow{\text{Symp}} (M', \omega') \right] = \left\{ \psi : M \xrightarrow{\text{SM}} M' \mid \psi^*(\omega') = \omega' \circ T\psi = \omega \right\}$$

$$\text{Symp}(M, \omega) := \text{Aut}_{\text{Symp}}(M, \omega) = \left\{ \psi : M \longrightarrow M \mid \omega_{\mu\nu} = \omega'_{\mu\nu} \right\}$$



## Canonical transformations are symplectomorphisms

The canonical transformations are changes of coordinates  $\begin{pmatrix} q \\ p \end{pmatrix} \mapsto \begin{pmatrix} Q \\ P \end{pmatrix}$  that leave the form of Hamilton's equations the same, i.e. such that the new coordinates are also canonical, i.e. symplectomorphisms:

$$-\{u^\mu, u^\nu\} = \Omega^{\mu\nu} \stackrel{!}{=} \tilde{\Omega}^{\mu\nu} = -\{\tilde{u}^\mu, \tilde{u}^\nu\} = \frac{\partial \tilde{u}^\mu}{\partial u^\kappa} \frac{\partial \tilde{u}^\nu}{\partial u^\iota} \Omega^{\iota\kappa}$$

New Hamiltonian vector field:  $X_{\psi^*H} = \psi^*X_H$ :

$$\omega(X_{\psi^*H}) = d(\psi^*H) = \psi^*dH = \psi^*\omega(X_H) = \omega(\psi^*X_H)$$

## Fact

If for  $t \in \mathbb{R}$ , the flow function  $\phi_H^t$  is well-defined, it is a symplectomorphism onto its image.

It is a diffeomorphism with smooth inverse  $\phi_H^{-t}$ , which is defined on  $\text{im } \phi_H^t$ . That it is symplectic is the fact that  $X_H$  is symplectic.

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# Electromagnetism

Consider the vector potential of the electromagnetic force  $(\phi, \vec{A})$  generating the Lorentz force

$$\vec{F}(r, \vec{v}, t) = q \cdot (\vec{E} + \vec{v} \times \vec{B}) = q \cdot \left( -\vec{\nabla}\phi - \frac{\partial \vec{A}}{\partial t} + \vec{v} \times \vec{\nabla} \times \vec{A} \right).$$

This can be described by the Lagrangian

$$L(r, \vec{v}) = \frac{m}{2} |\vec{v}|^2 - q \cdot (\phi(r) - \vec{v} \cdot \vec{A}(r)) = \frac{1}{2} m(\vec{v}, \vec{v}) - q\phi + \vec{A}\vec{v}$$

giving the canonically conjugated momentum

$$\vec{p} = \frac{\partial L}{\partial \vec{v}} = m(\vec{v}) + q\vec{A} \Leftrightarrow \vec{v} = m^{-1}(\vec{p} - q\vec{A})$$

and the Hamiltonian  $(\vec{p}\vec{v} - L)$

$$H = \frac{1}{2} m(\vec{v}, \vec{v}) + q\phi = \frac{1}{2} m^{-1}(\vec{p} - q\vec{A}, \vec{p} - q\vec{A}) + q\phi$$

# Twisted cotangent bundle

A **twisted cotangent bundle** is  $T^*N$  as a symplectic manifold, but the symplectic form differs from  $d\lambda$ :

$$\omega = d\lambda + \sigma.$$

This gives rise to a kind of “Lorentz force”  $Y : TN \longrightarrow TN$ ,

$$g(\vec{Y}(\vec{v})) = \sigma(\vec{v})$$

This is an alternative approach to the magnetism problem from the last slide.