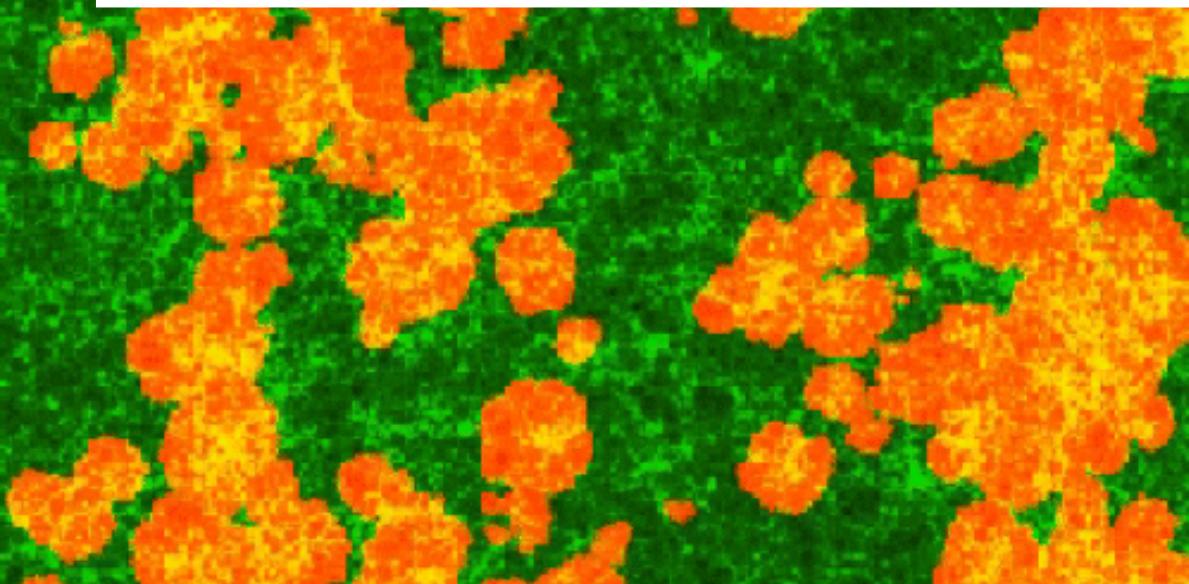




Mapping the Spatially Inhomogeneous Cosmic Reionization with Subaru HSC



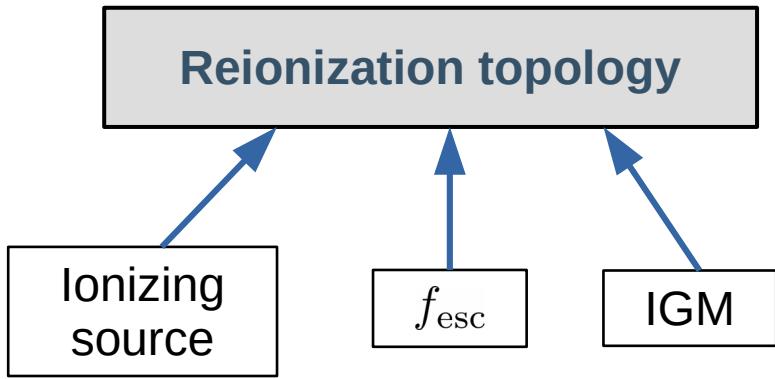
Takehiro Yoshioka (U. of Tokyo),

Nobunari Kashikawa (U. of Tokyo), Akio Inoue (Waseda U.),
Satoshi Yamanaka (Toba College), and CHORUS members

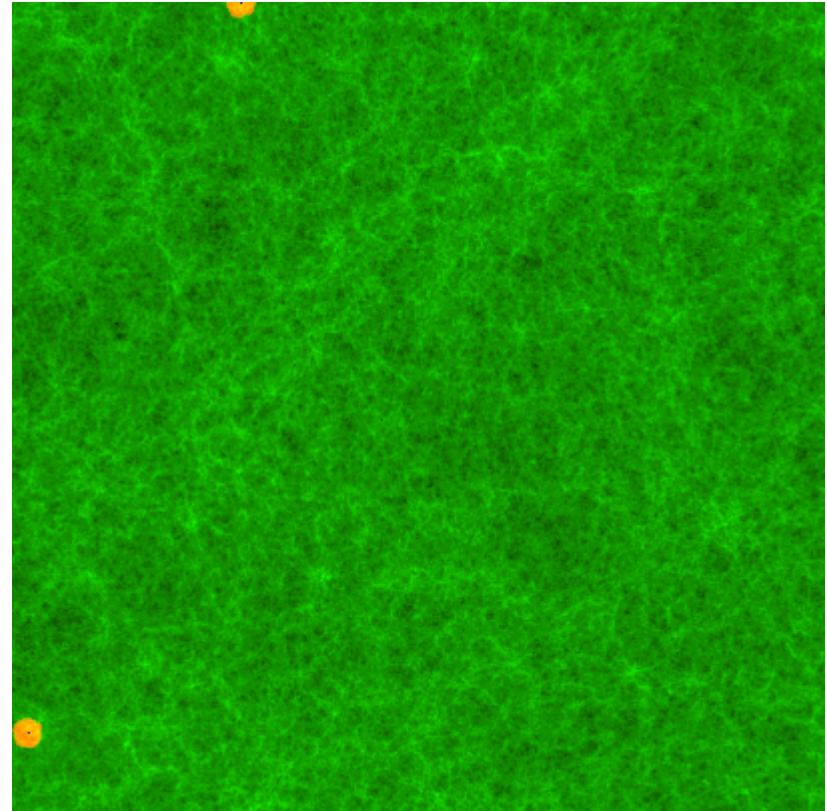


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Spatially inhomogeneous reionization



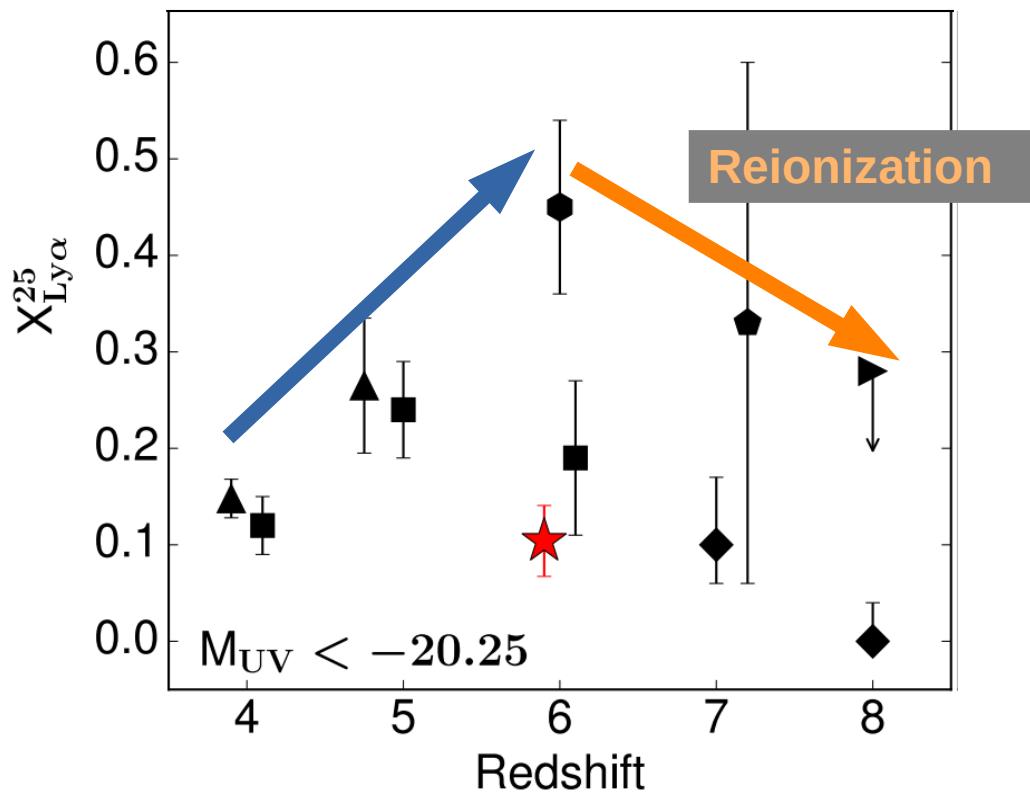
**Investigation of reionization
topology is important for
studying these parameters.**



Iliev et al. (2006)

Ly α fraction as a probe of reionization

- Ly α fraction
 - the fraction of LBGs that show Ly α emission lines.
- Ly α fraction measurements require spectroscopic observations.
 - Difficult to obtain large sample.
 - **Unable to investigate spatially inhomogeneous reionization.**



De Barros et al. (2017)

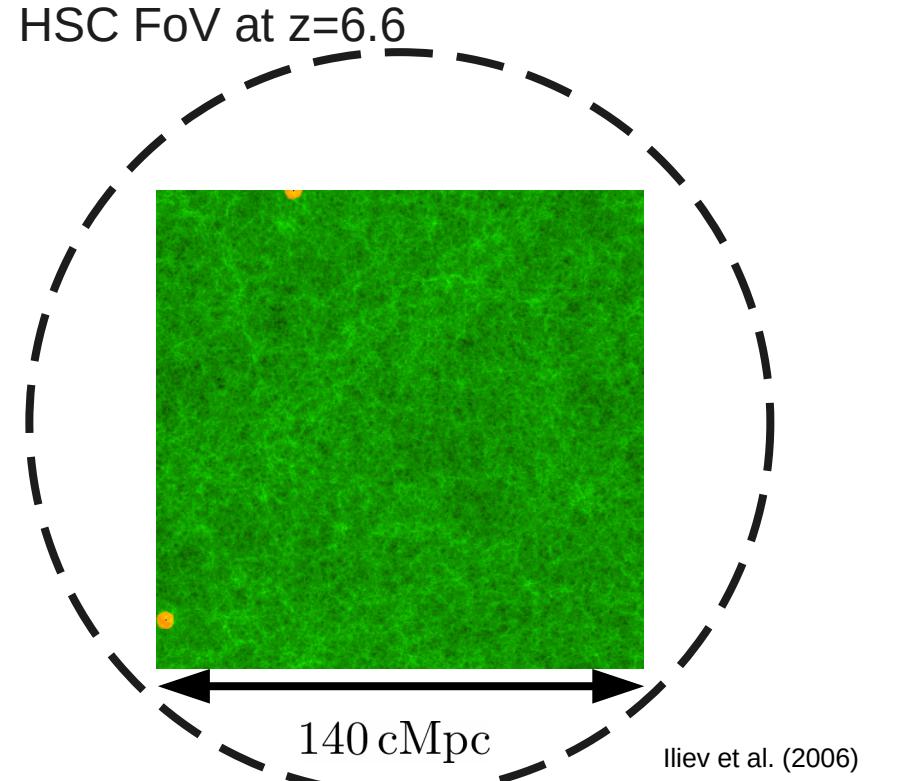
Data points are taken from Stark et al (2011); Schenker et al. (2014); Tilvi et al. (2014); Ono et al. (2012); Curtis-Lake et al. (2012); Cassata et al.(2015).

Motivation

Subaru/Hyper Suprime-Cam (HSC)

- Wide field of view (FoV) $\sim 1.5 \text{ deg}^2$.
 - Efficiently detect LAEs and LBGs.

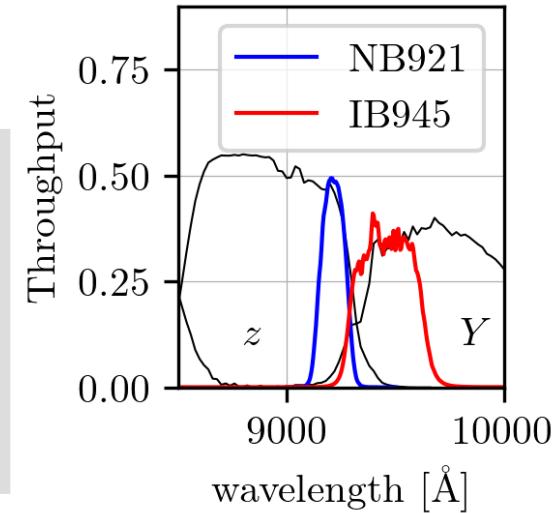
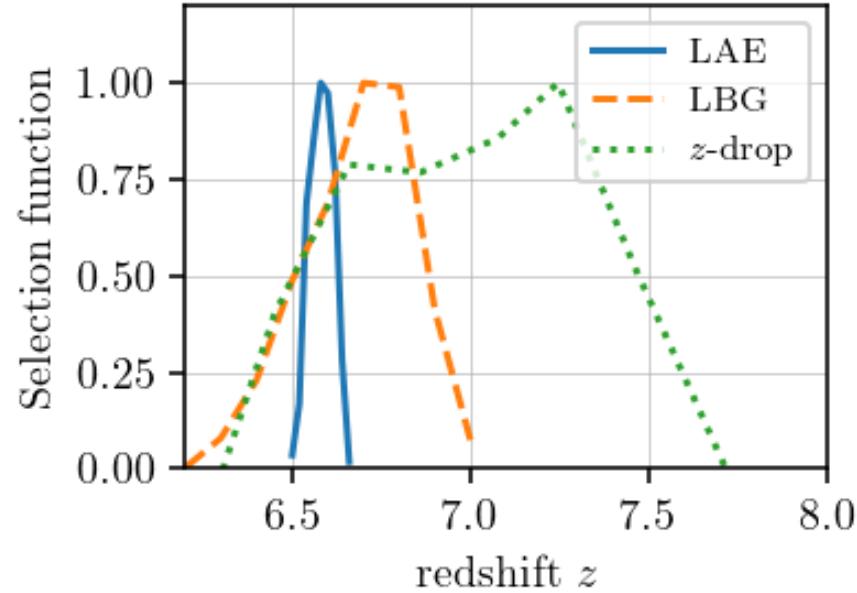
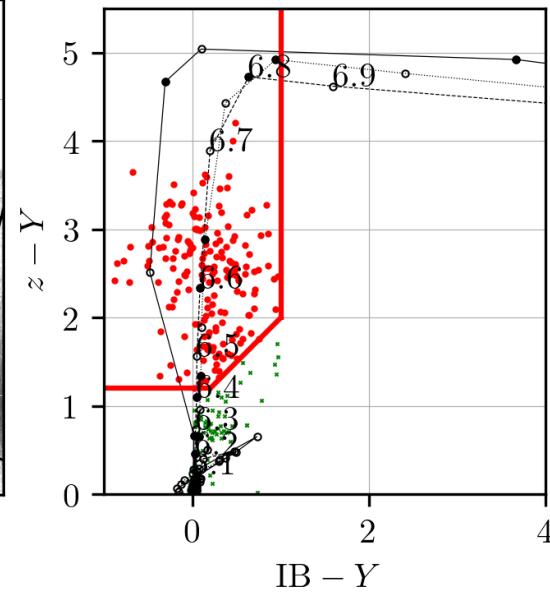
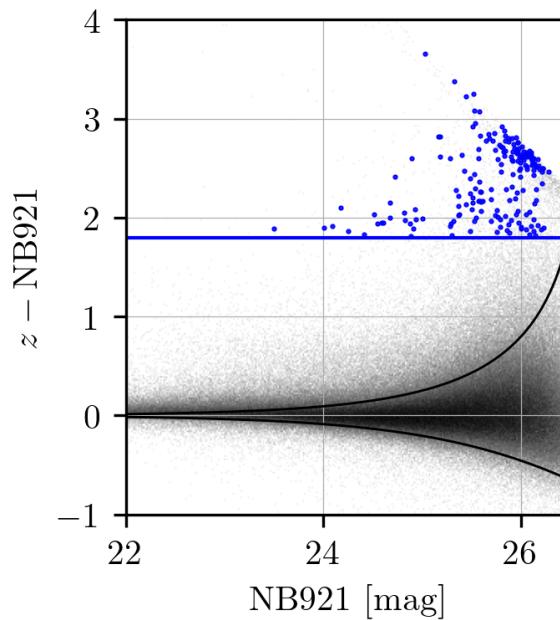
Observe spatially inhomogeneous
reionization by efficient survey of LAEs
and LBGs at $z = 6.6$.



LAE, LBG selections at $z = 6.6$

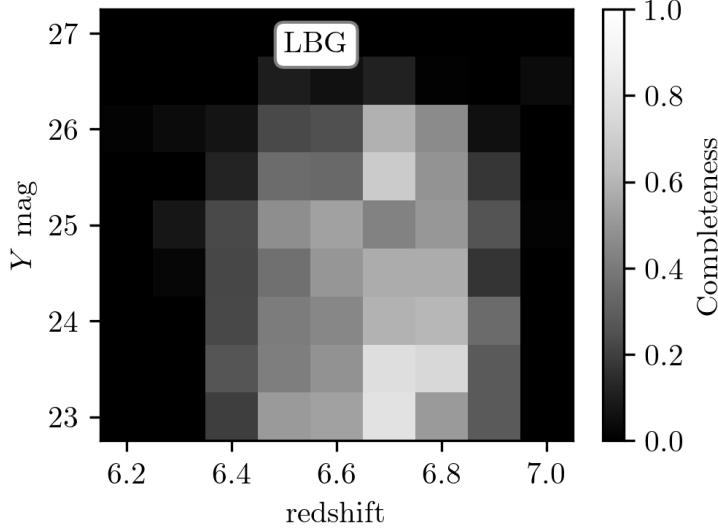
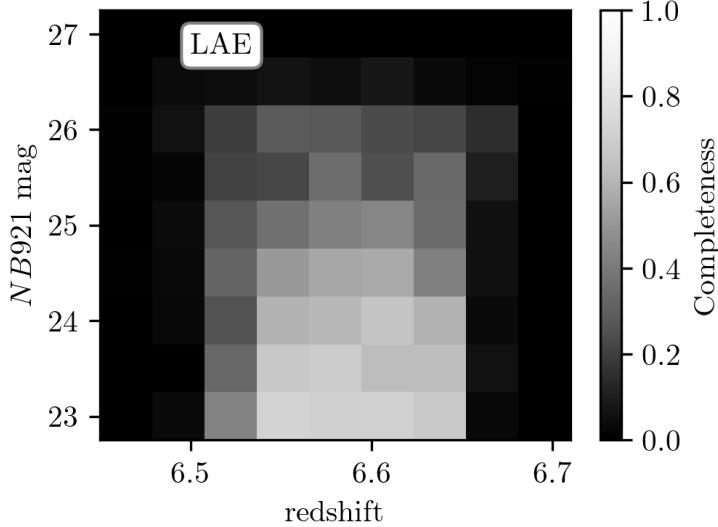
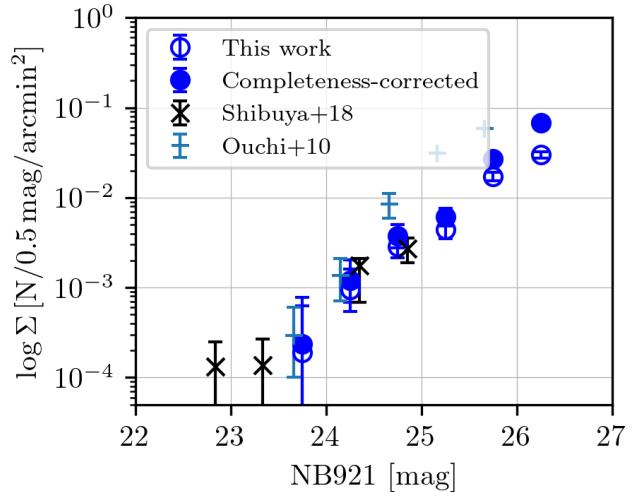
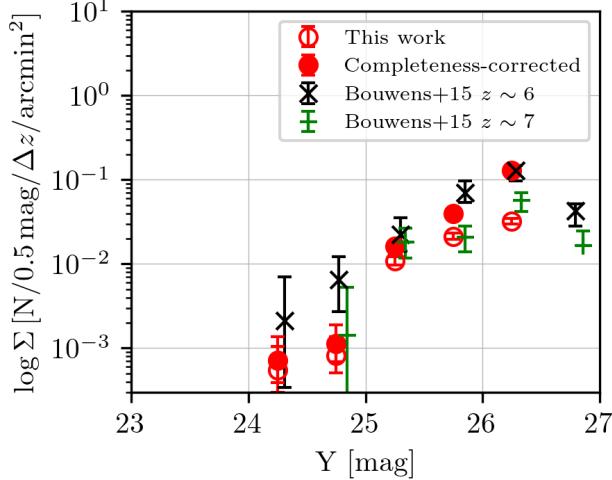
- HSC-SSP UD-COSMOS
~1.5 deg² field.
- NB excess — 189 LAEs
- Lyman break — 179 LBGs

- Brand new filter, **IB945**.
 - Detect LAEs and LBGs **simultaneously at the same redshift $z=6.6$.**



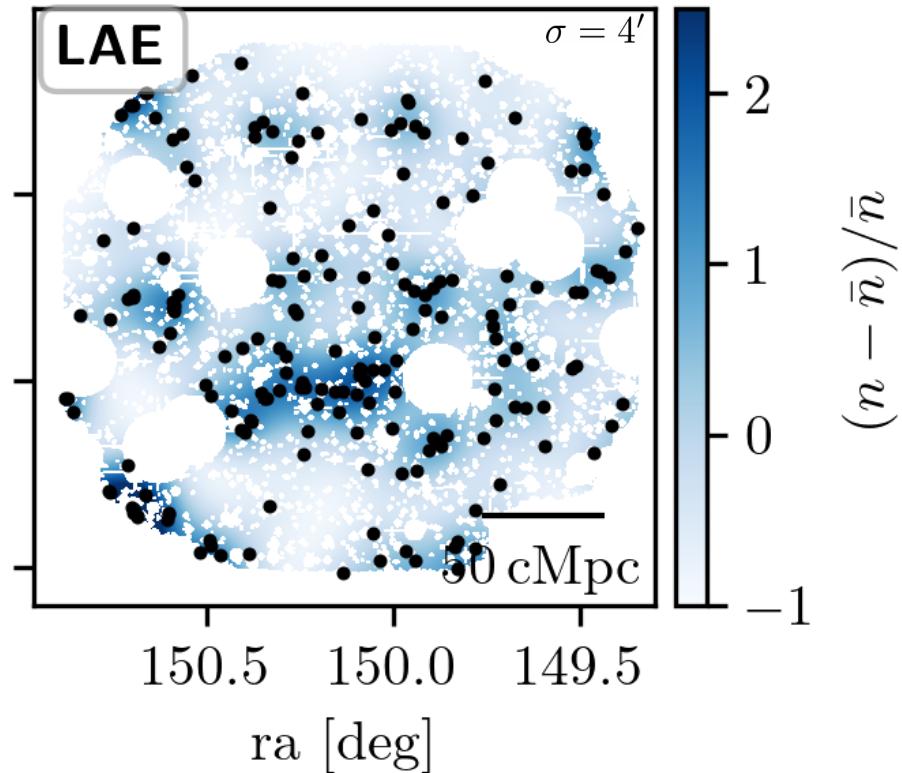
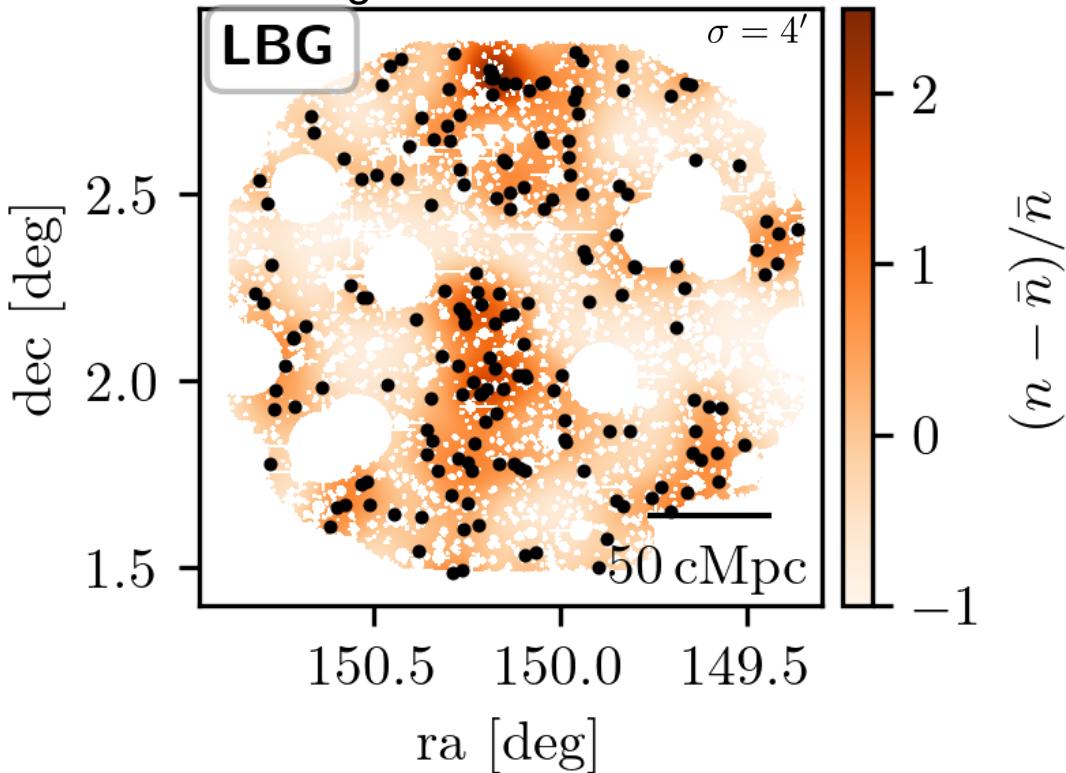
Completeness

- Calculate selection completeness by mock galaxy simulations applying the selection criteria.
- Surface densities of LAEs and LBGs are consistent with previous studies.



Distribution

- Color maps show the density distribution of the objects.
- Their values indicate the overdensity $(n - \text{mean}(n)) / \text{mean}(n)$.
- Mask regions are shown in white.



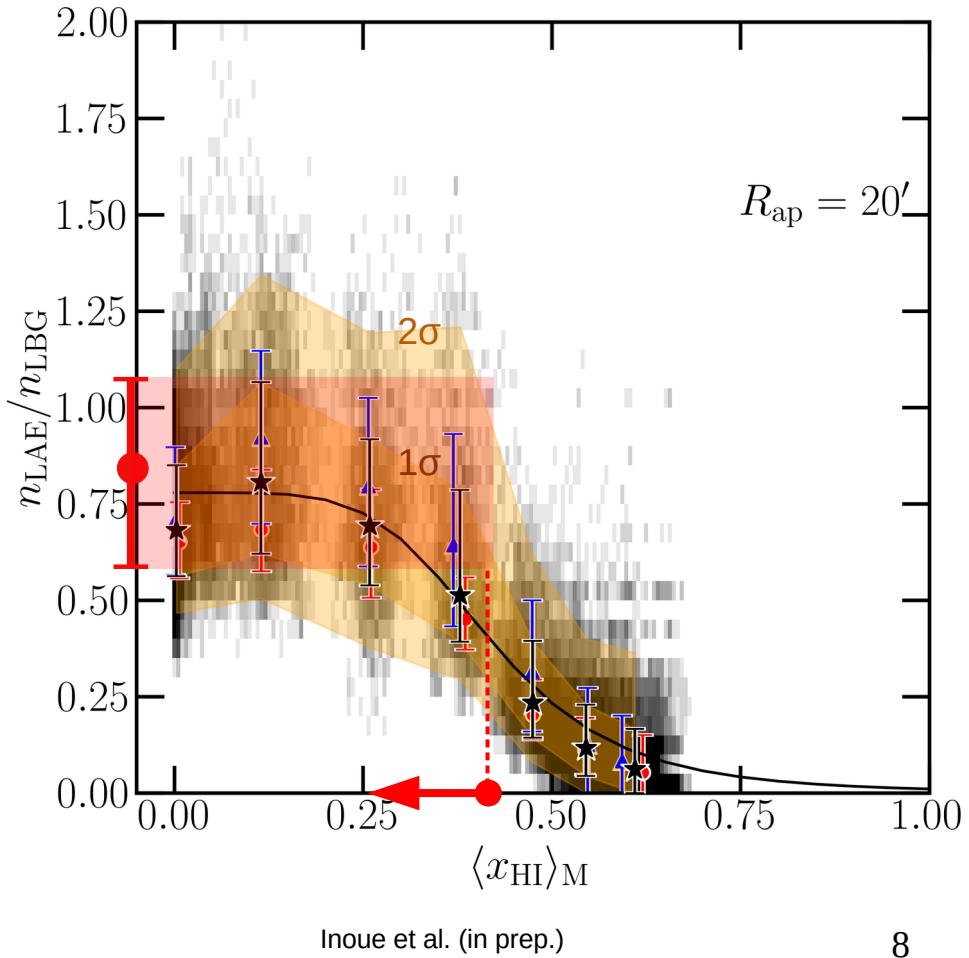
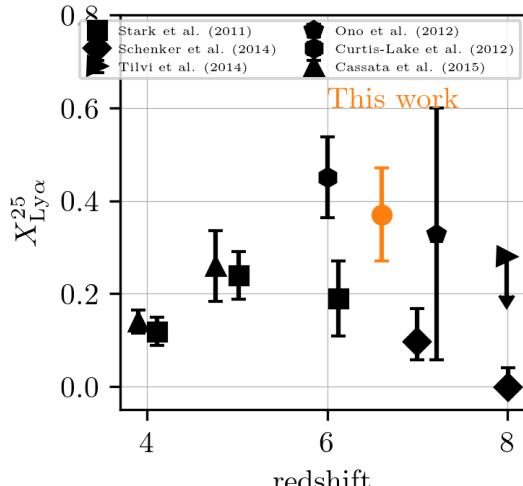
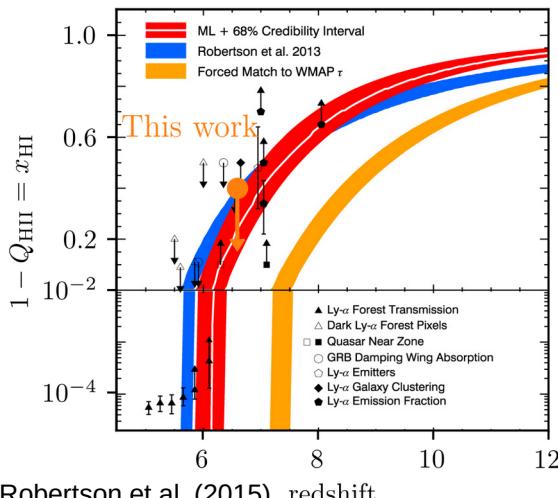
Neutral fraction

- LAE/LBG number ratio averaged in the entire FoV is

$$n(\text{LAE})/n(\text{LBG}) = 0.84^{+0.23}_{-0.27}.$$

- Averaged neutral fraction in the entire FoV is

$$\langle x_{\text{HI}} \rangle < 0.4.$$

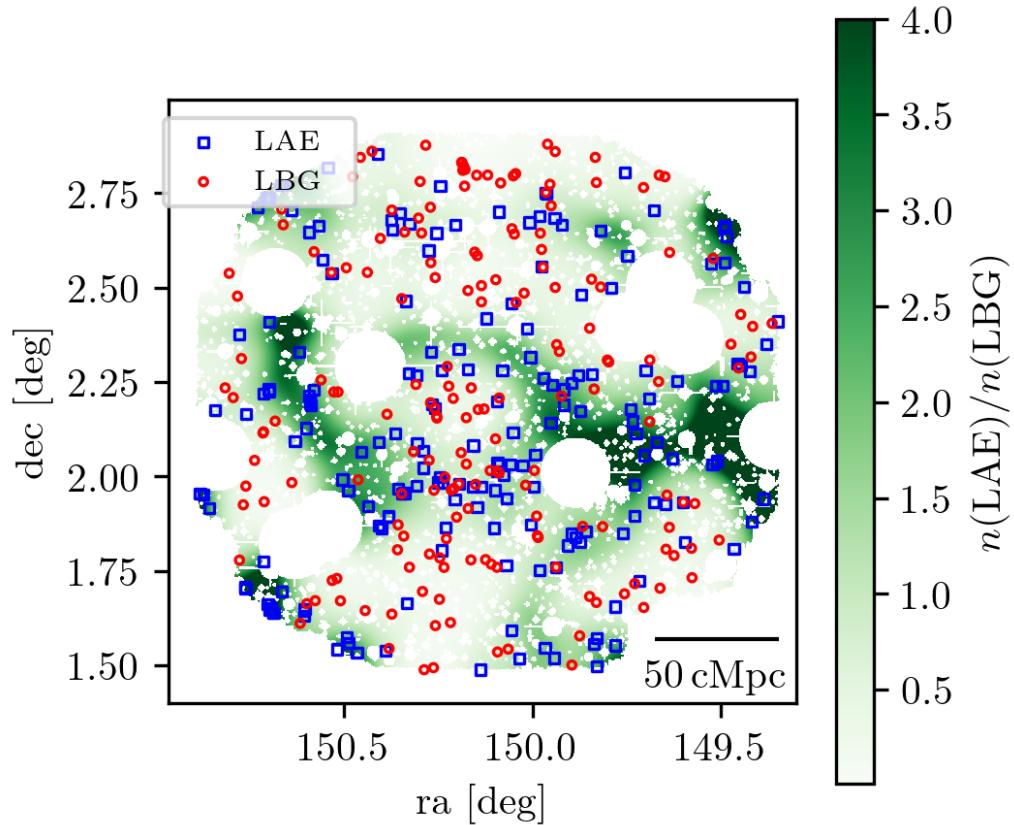


LAE/LBG ratio map

- Spatial variation of LAE/LBG is a factor of three.

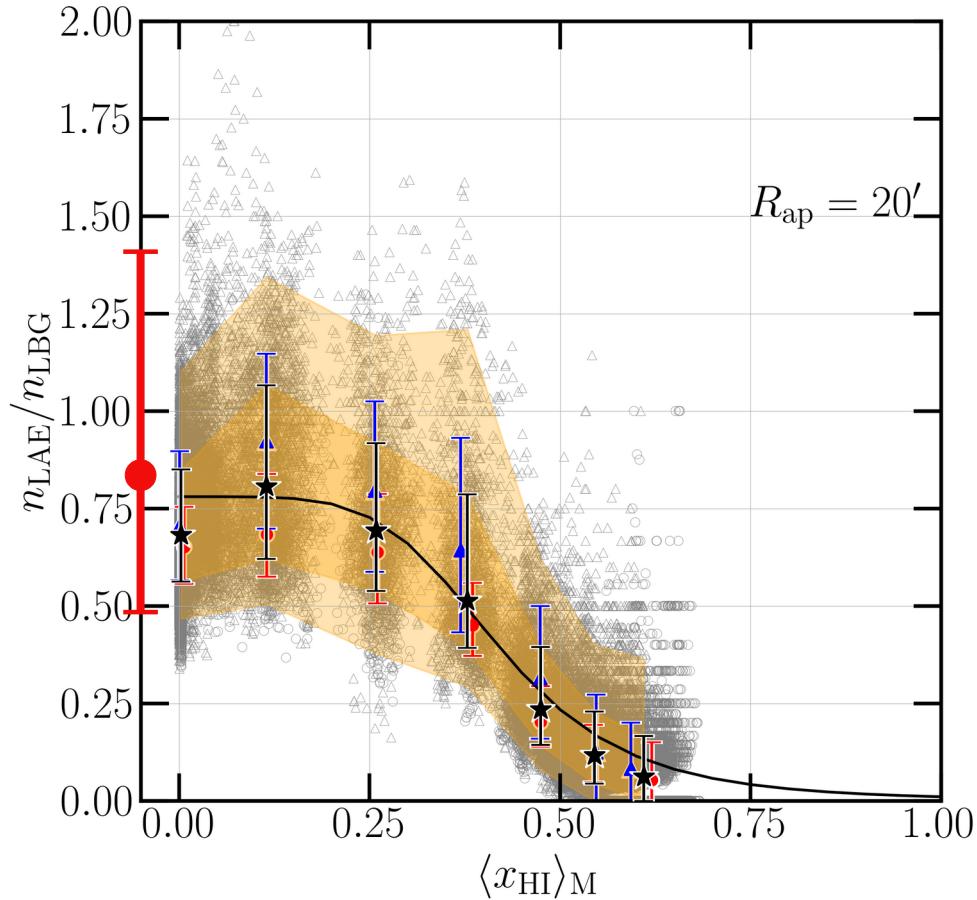
Imply inhomogeneous reionization

- High LAE density region:
 - Reionization is in progress.
- Low LAE density region:
 - Reionization is delayed.



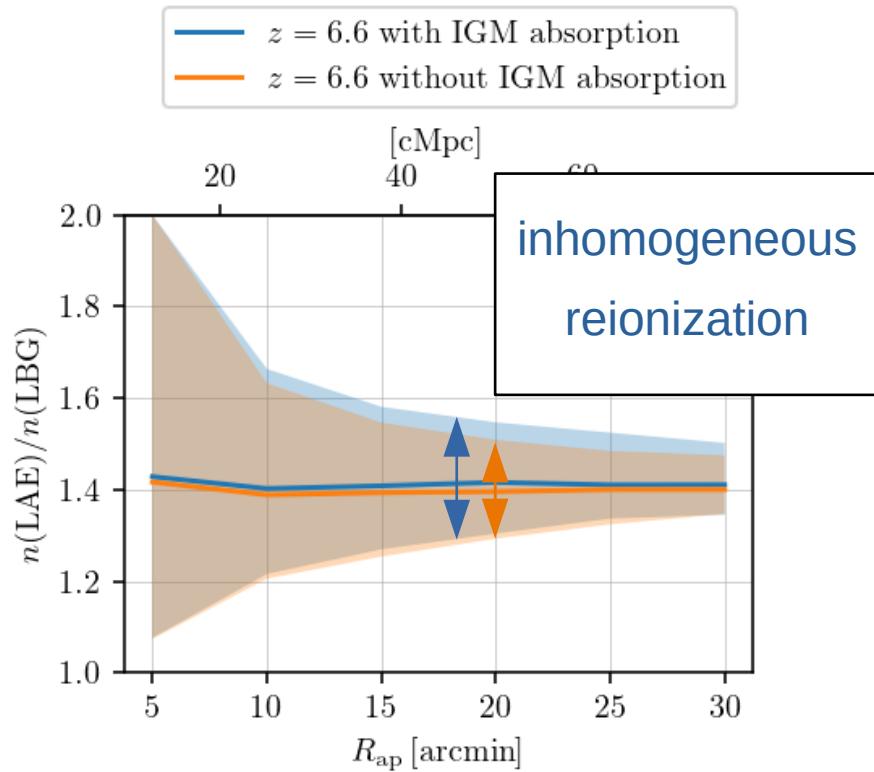
Neutral fraction

- Spatial variation of LAE/LBG is a factor of three.
- Difficult to quantify the spatial variation of neutral fraction.



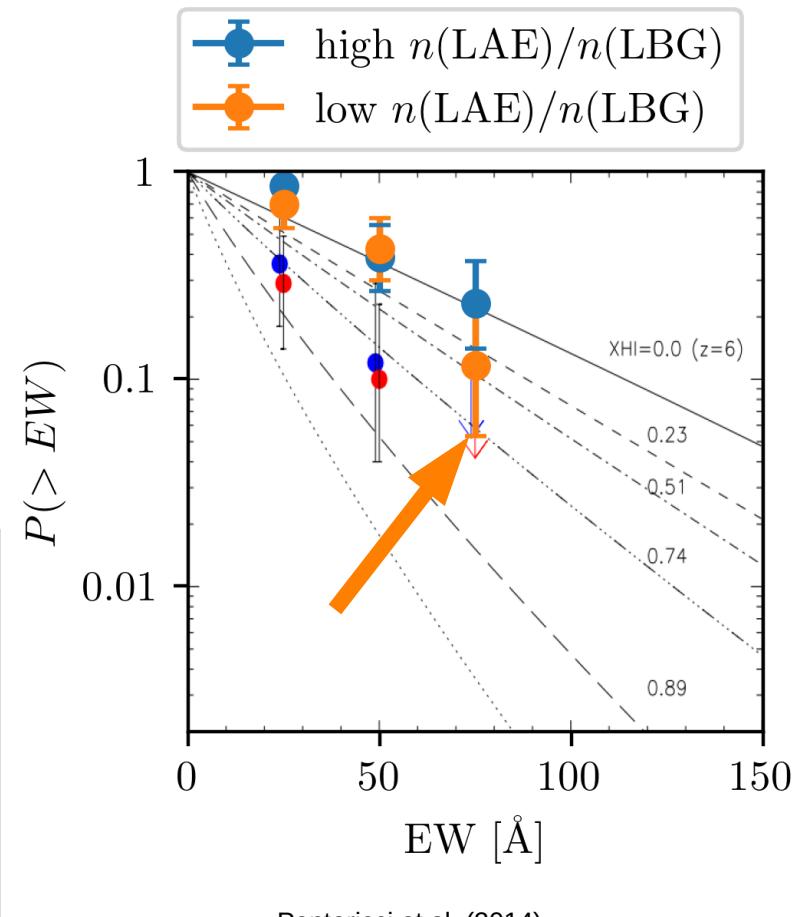
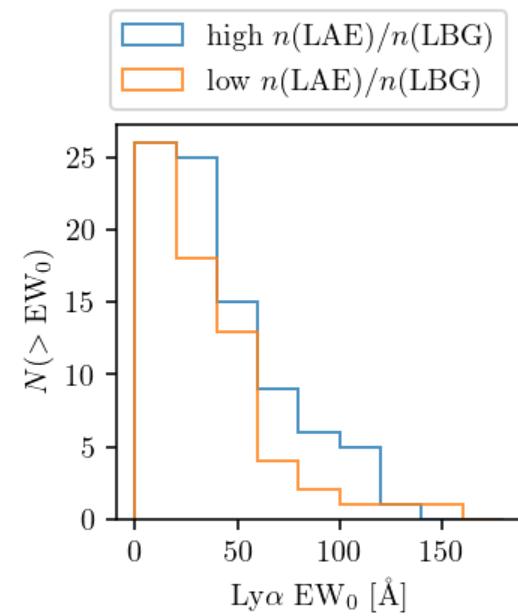
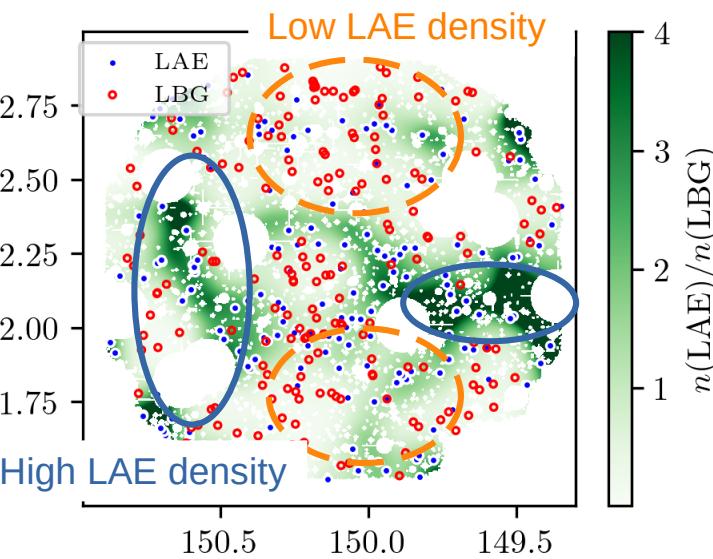
Distinguish the reionization topology from the intrinsic large scale structure

- Large aperture is necessary for this aim.
- Observed large variation in LAE/LBG may be due to intrinsic LSS without IGM absorption.
- Compare the scatter of LAE/LBG ratio between with and without IGM absorption in the simulation data.
- Hard to distinguish them even with ~ 2 mag deeper observations.



Ly α EW distribution

- Compare Ly α EW distribution between high and low LAE density regions.
- **Although the statistical significance is low, Low LAE density region implies higher neutral fraction.**



Pentericci et al. (2014)

Summary

- We **simoultaneously detect LAEs and LBGs at $z \sim 6.6$** in the large FoV of the HSC observations with the new filter IB945.
- From the efficient survey, we attempt to **investigate the spatially inhomogeneous reionization for the first time**.
- From observed LAE/LBG ratio and reionization simulations, we derived **the averaged neutral fraction in the FoV** $x_{\text{HI}} < 0.4$.
- Mapping the spatial inhomogeneity of reionization requires much deeper imagings.
- It is **difficult to distinguish the reionization topology from the LSS** even with ~ 2 deeper observations.