

Preliminary Results

**A constraint of [NII] 122 μm and
a new dust continuum detection of
a $z = 7.15$ Lyman Break Galaxy with ALMA**

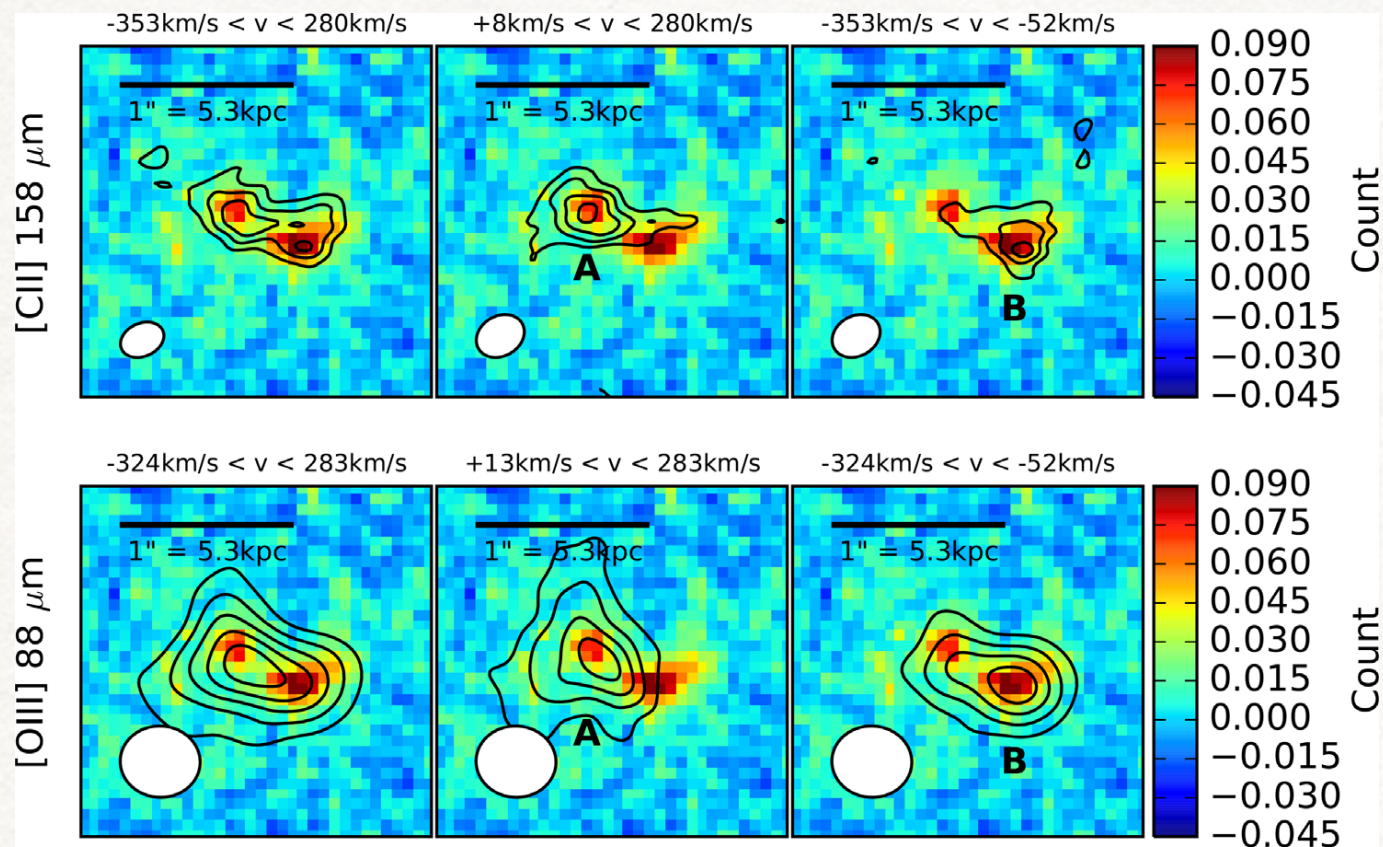
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Y. Tamura, H. Matsuo, E. Zackrisson, C. Binggeli

B14-65666 @ $z = 7.15$

○ [OIII] 88 μ m, [CII] 158 μ m, dust continuum detected

□ Big Three Dragon (大三元), Hashimoto+19

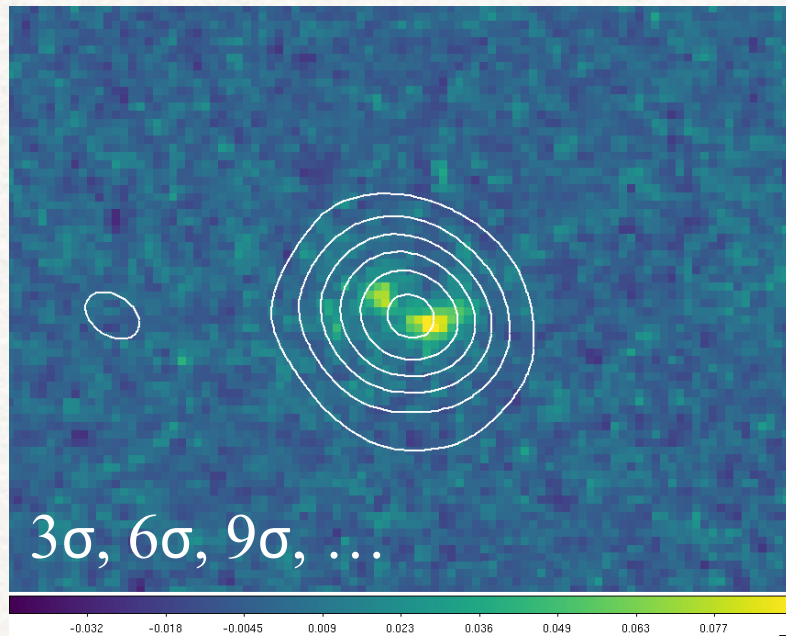


ALMA new observation

- Cycle 7, Band 7 observation (PI: A. K. Inoue)
 - targets: [NII] 122um, dust continuum
 - exp time: 2.4 h

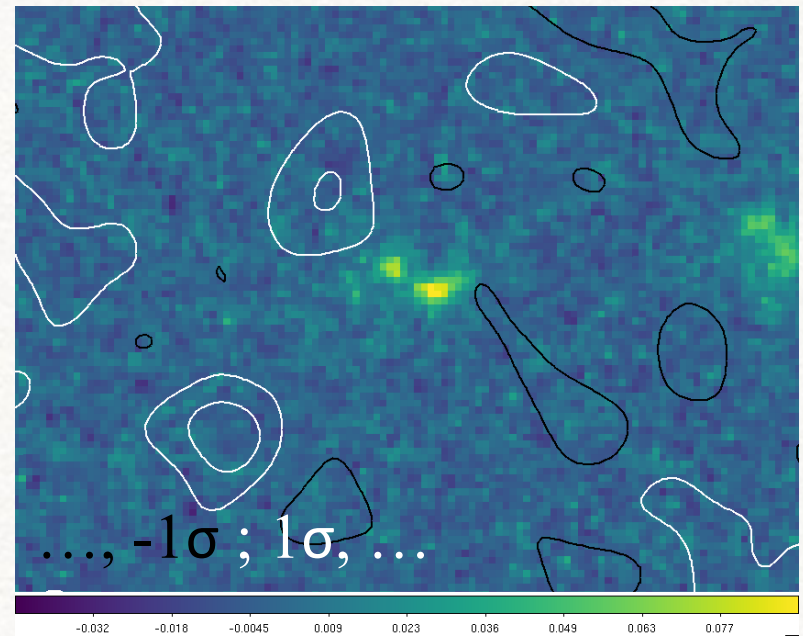
Image

○ Dust (120 μm)



- $\sim 19\sigma$ detect
- 218 ± 19 uJy

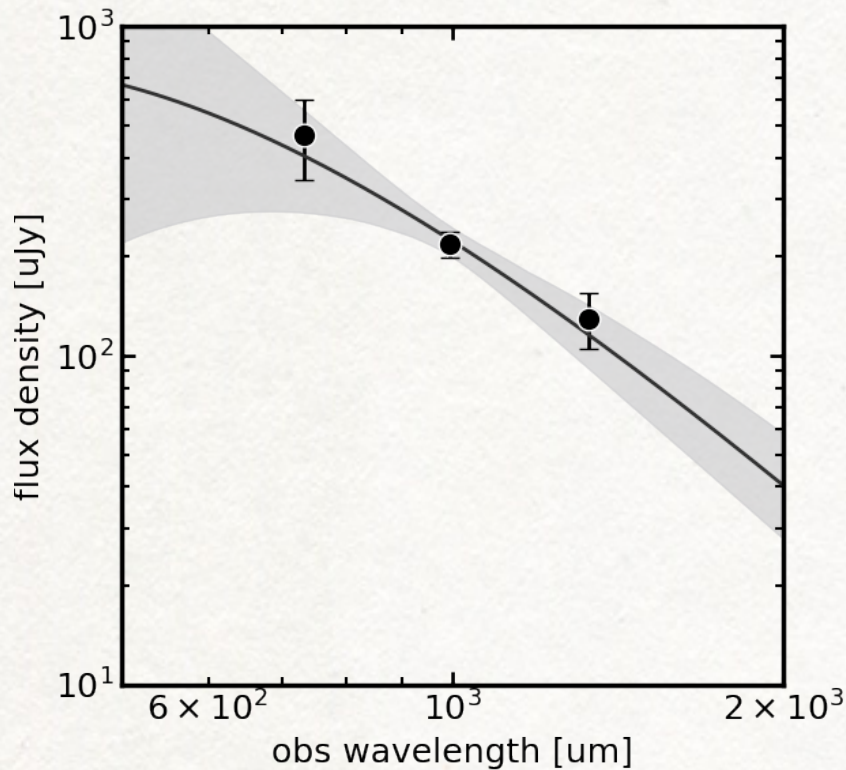
○ [NII] 122 μm



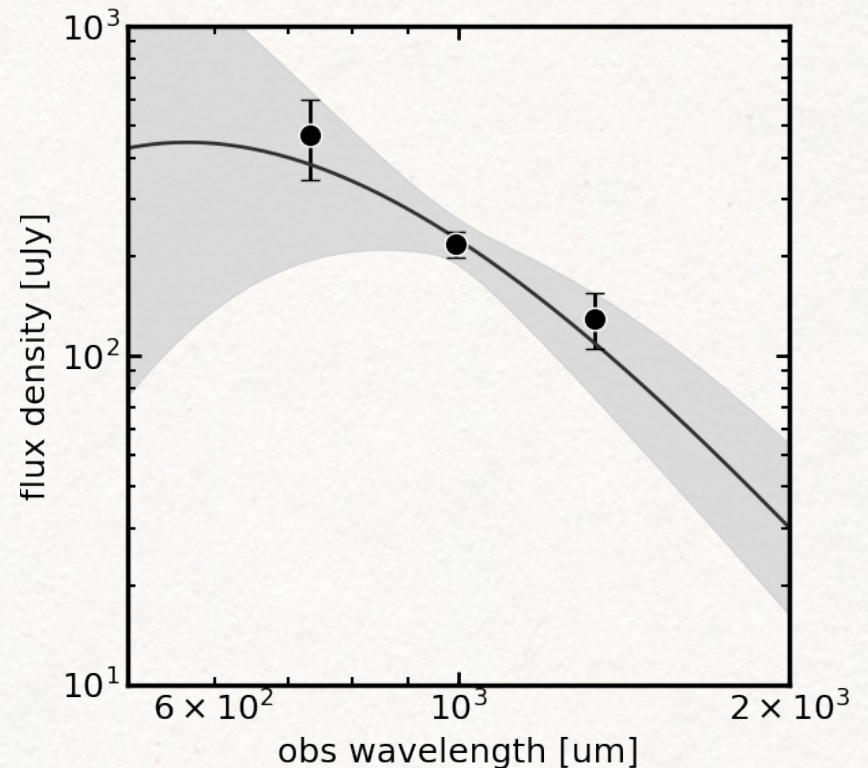
- 3σ upper limit
- < 0.025 Jy/beam km/s

FIR SED Fitting

○ Beta = -1.0 fix

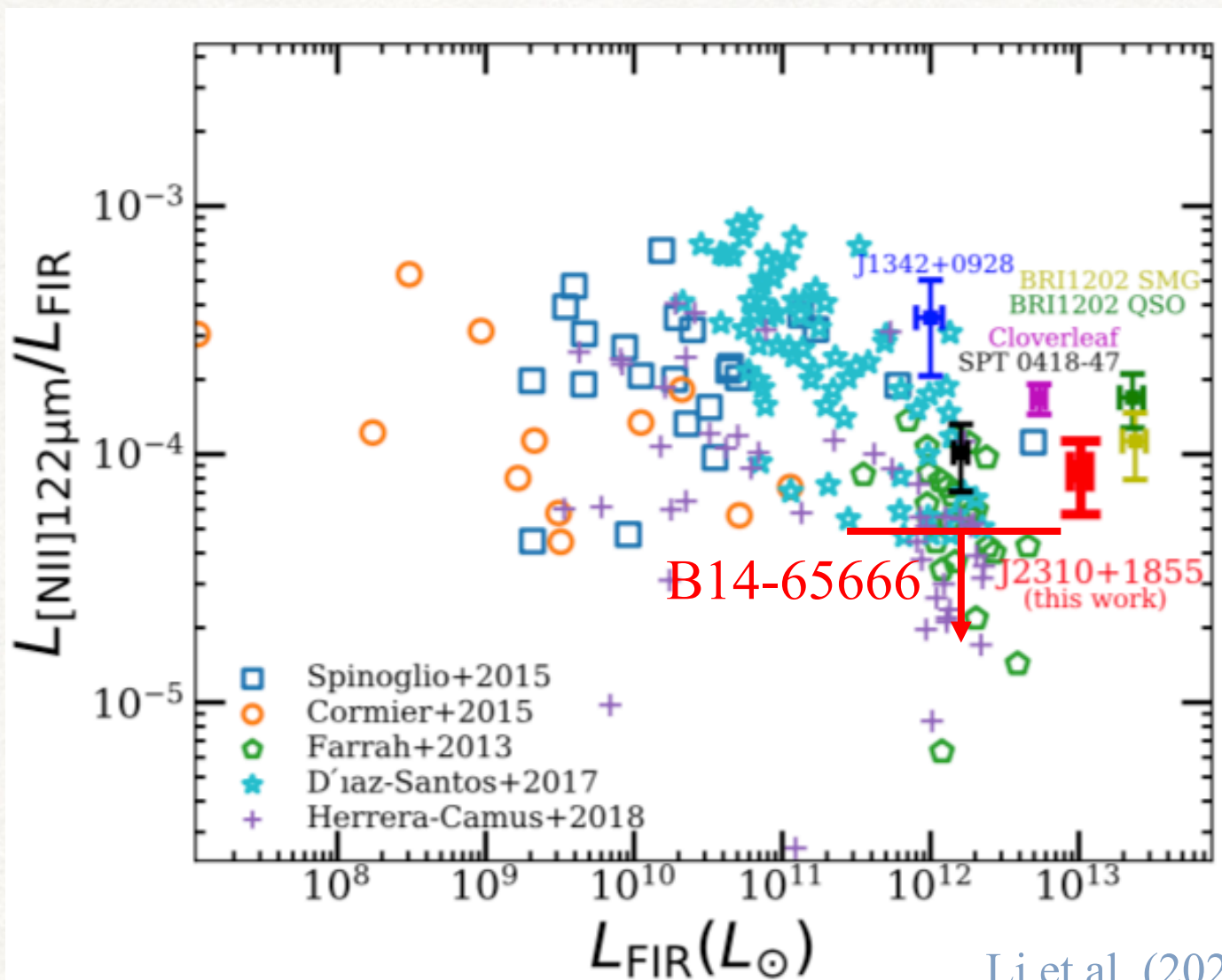


○ Beta = -2.0 fix



$T > 30$ K, but difficult to set upper limit

$[NII]/L_{FIR}$



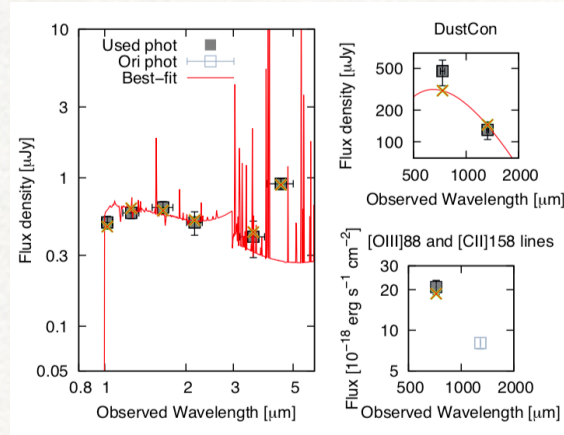
Li et al. (2020)

Discussion

- What can we know from [NII] 122 μ m observation?
 - N/O abundance at high redshift

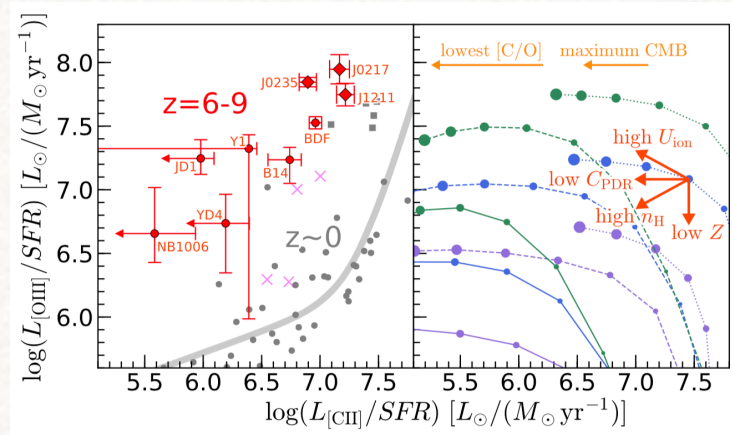
Strategy

○ SED fitting (Hashimoto+19)



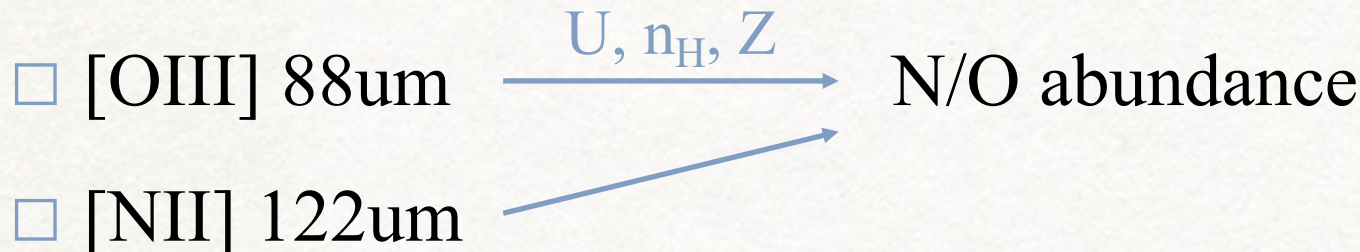
SFR, metallicity Z

○ $[\text{OIII}]/[\text{CII}]$ (Harikane+20)

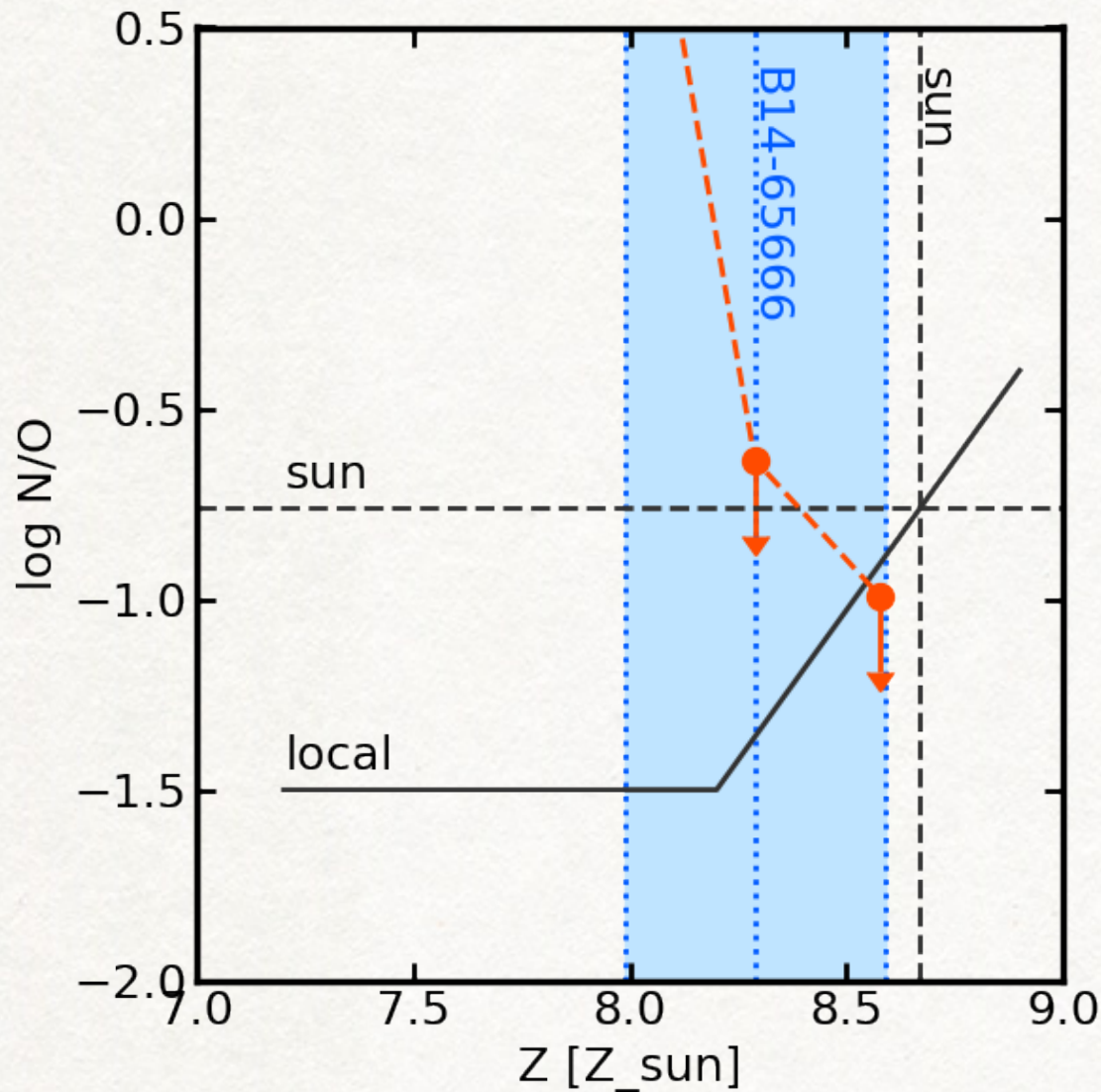


Ionization parameter U
Hydrogen density n_{H}

○ Cloudy calculation



N/O abundance



At high- Z :
lower than solar N/O

At low- Z :
no strong limit

Discussion

○ What can we know from [NII] 122 μ m observation?

□ N/O abundance at high redshift

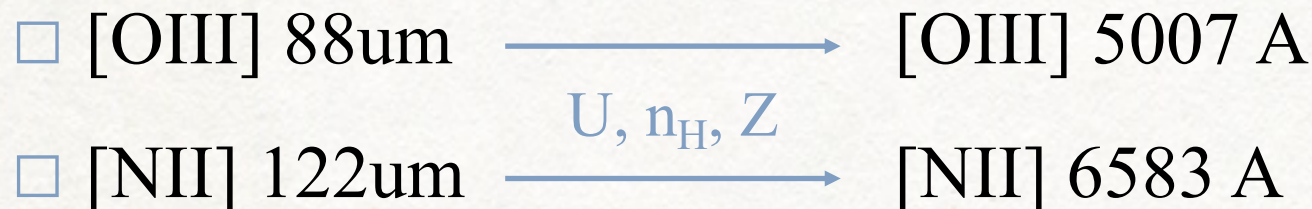
□ Redshift evolution of BPT diagram

BPT diagram

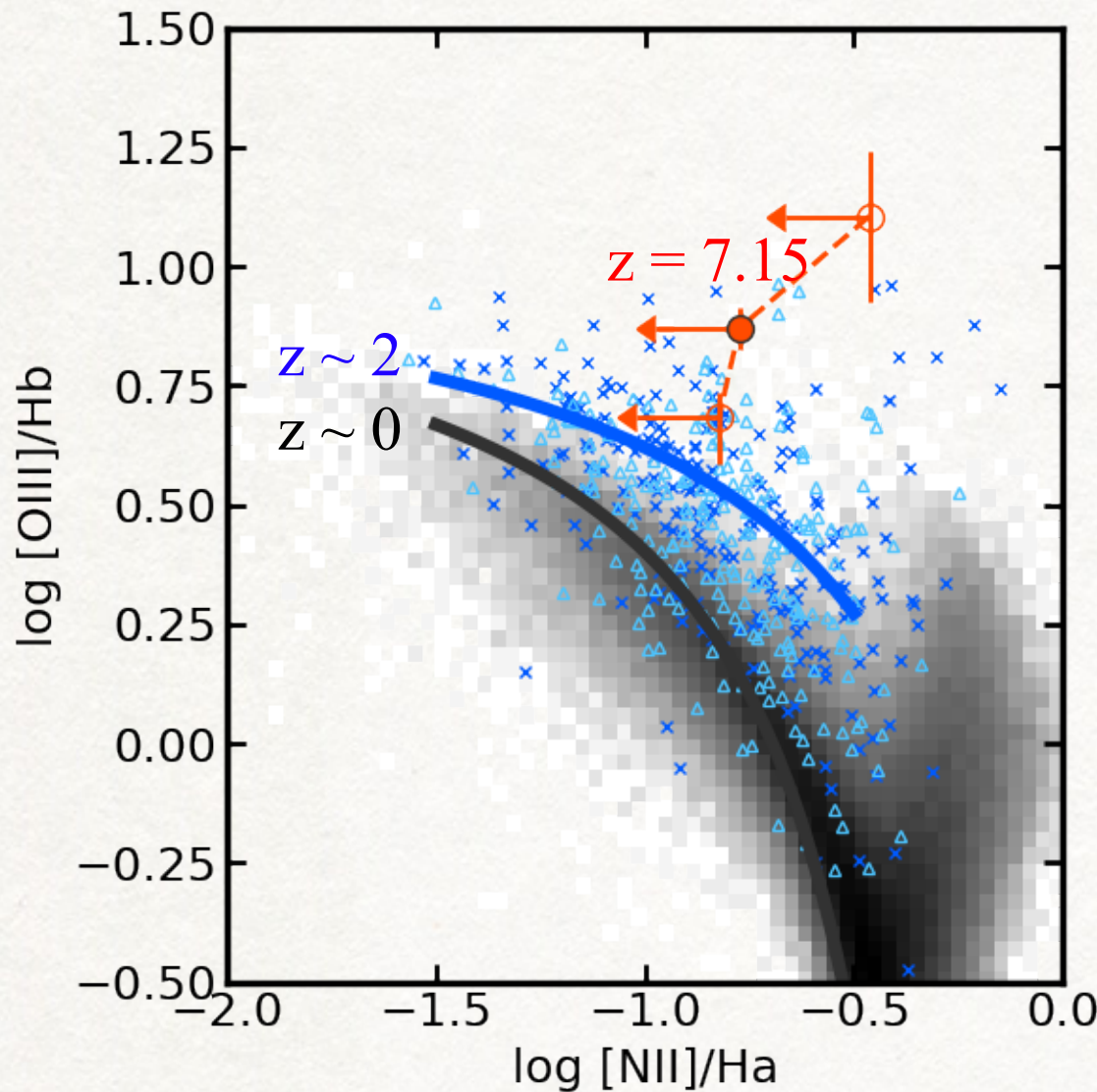
○ From multi-wavelength SED fitting (Hashimoto+19)



○ Cloudy calculation



BPT diagram



Ionization states in galaxies show evolutionary trend at $z \sim 0 \rightarrow 7$.

$z \sim 0$ SDSS

$z \sim 2$ Strom+18

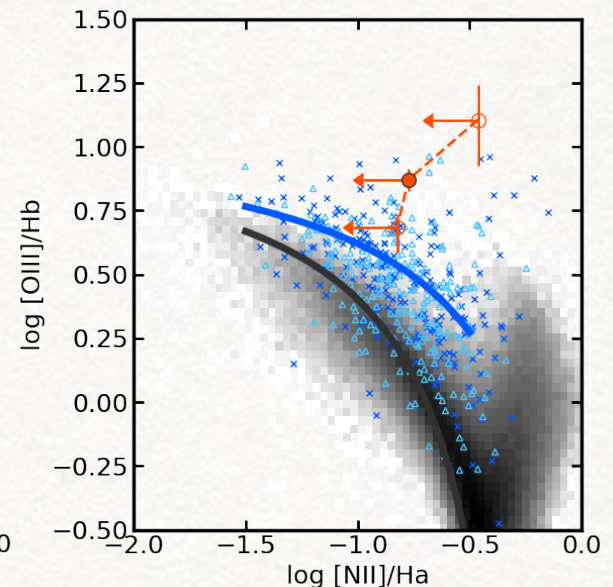
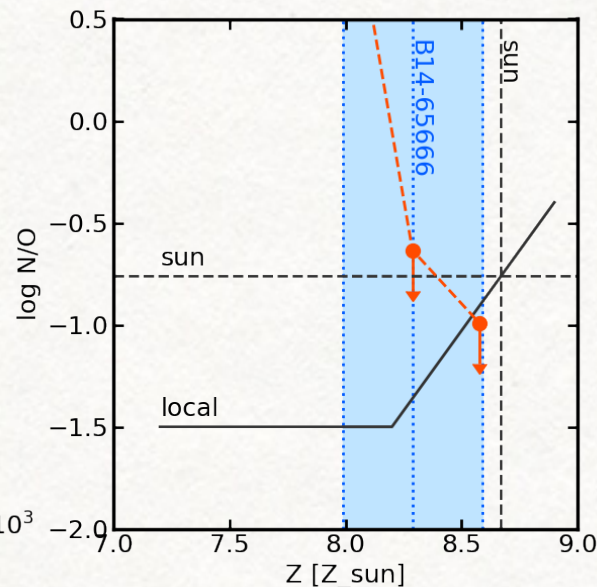
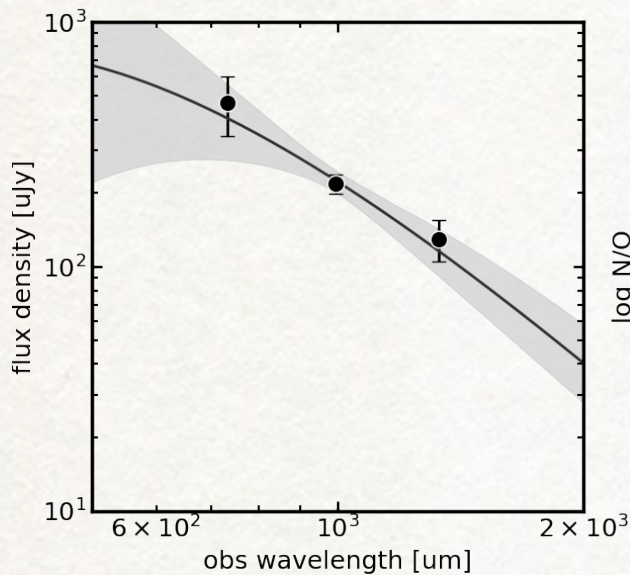
$z \sim 2$ Shivaei+19

Summary

Preliminary Results, so Stay Tuned!

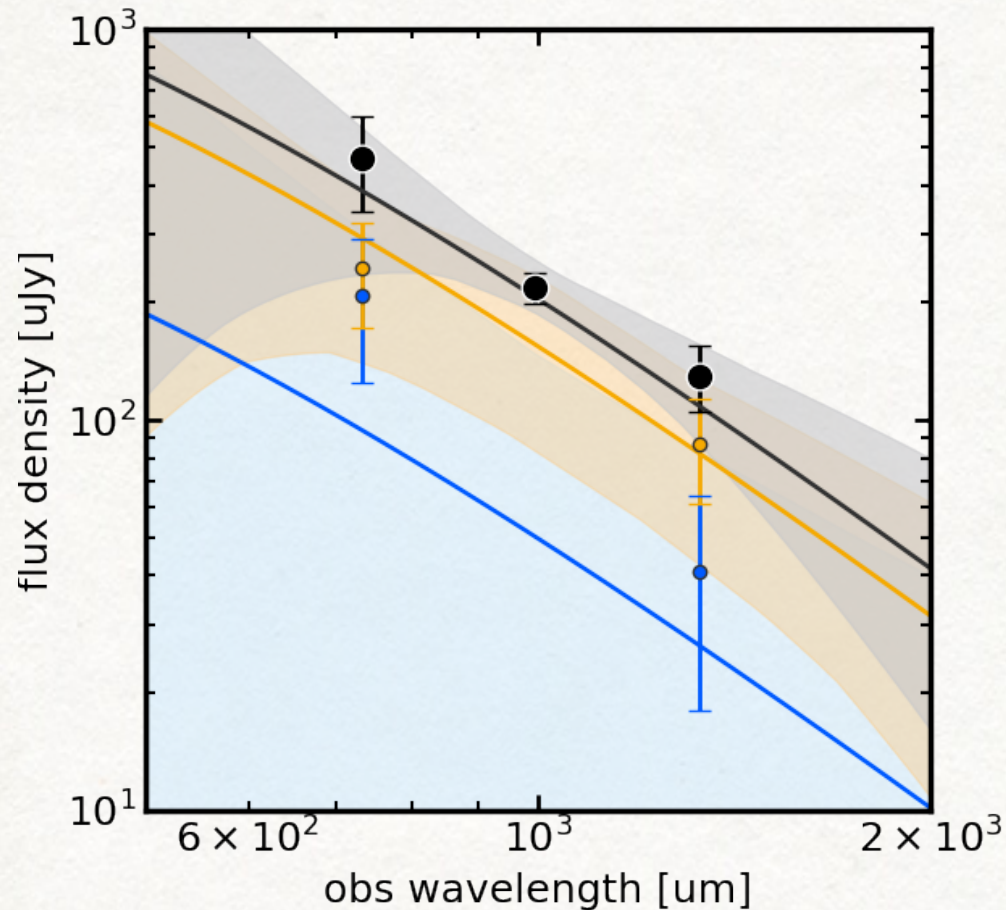
○ 18σ dust detection & 3σ [NII] upper limit on B14-65666

- Dust temperature $T > 30\text{K}$
- Sub solar N/O if high-Z
- Evolutional trend at $z \sim 0 \rightarrow 7$ on BPT diagram



FIR SED Fitting

○ Beta, T free



Strategy

