

PV Concept and a Limit on Collection Efficiency

- When a **photon is absorbed** in a block of a semiconducting material such as silicon, it can **excite an electron** from the **valence band** to the **conduction band**.
- Once this has occurred, the **electron can move freely** in the conduction band. Likewise, the hole created in the valence band can **propagate like a positively charged electron**.
- The aim of the *photovoltaic cell* is to use these **electron and hole excitations** to drive an external circuit.

Collection efficiency in the context of photovoltaic (PV) cells refers to the **fraction of photo-generated charge carriers (electrons and holes)** that are successfully **collected at the electrodes** and contribute to the **external current**, rather than being lost to recombination.

$$\text{Collection Efficiency} = \frac{\text{Number of carriers collected (electrons and holes)}}{\text{Number of carriers generated by absorbed photons}}$$

Why collection efficiency matters: Even if a material absorbs light well (high absorption coefficient), **poor collection efficiency** means much of that energy is **wasted**. So, both **optical absorption** and **carrier collection** must be optimized for high-efficiency solar cells.

- When a photon of energy $E = \hbar\omega$ is incident on a material with a band gap E_{gap} between its (filled) valence band and (empty) conduction band, it can only lead to a conduction electron if $E > E_{gap}$.
- Thus, for example, only photons with an energy $\hbar\omega > 1.1 \text{ eV}$ can excite an electron from the valence band to the conduction band of a silicon crystal.
- Knowing the energy spectrum of the photons in solar radiation – we can immediately determine an upper bound on the collection efficiency of a solar cell with gap E_{gap} . The **limit on collection efficiency** using the Planck radiation spectrum relation

$$\frac{dP}{dAd\omega} = \frac{\hbar}{4\pi^2 c^2} \frac{\omega^3}{e^{\frac{\hbar\omega}{k_B T}} - 1}.$$

- Then the total incident power per unit area from sun:

$$P_{total} = CA \int_0^\infty dE \frac{E^3}{e^{\frac{E}{k_B T}} - 1}$$

- where C is an overall proportionality constant not relevant here, $E = \hbar\omega$ is the energy of a single photon of angular frequency ω , and A is the area exposed to the solar radiation.

- A solar cell with band gap E_{gap} can only absorb photons with energy $E > E_{gap}$. However, **each absorbed photon contributes only E_{gap}** , to electrical power (the rest is lost as heat).
- So, the maximum power the cell can extract is: $P_{max}(E_{gap}) = n. (E_{gap})E_{gap}$
- Where, $n. (E_{gap})$ is the number of photons with $E > E_{gap}$, given by:

$$n. (E_{gap}) = CA \int_{E_{gap}}^{\infty} dE \frac{E^2}{e^{\frac{E}{k_B T}} - 1}$$

- So,

$$P_{max}(E_{gap}) = n. (E_{gap})E_{gap} = CAE_{gap} \int_{E_{gap}}^{\infty} dE \frac{E^2}{e^{\frac{E}{k_B T}} - 1}$$

- Let $x = \frac{E}{k_B T}$, so $dE = k_B T dx$. Then $E = x. k_B T$ and $E_{gap} = x_{gap}. k_B T$

So, the integrals become

$$P_{total} = CA \int_0^{\infty} dE \frac{E^3}{e^{\frac{E}{k_B T}} - 1} = CA(k_B T)^4 \int_0^{\infty} \frac{x^3}{e^x - 1} dx$$

$$P_{max} = CAE_{gap} \int_{E_{gap}}^{\infty} dE \frac{E^2}{e^{\frac{E}{k_B T}} - 1} = CA(k_B T)^4 x_{gap} \int_{x_{gap}}^{\infty} \frac{x^2}{e^x - 1} dx$$

- **Collection efficiency, $\eta_{collection}^{max}$** = $\frac{P_{max}}{P_{total}} = \frac{CA(k_B T)^4 x_{gap} \int_{x_{gap}}^{\infty} \frac{x^2}{e^x - 1} dx}{CA(k_B T)^4 \int_0^{\infty} \frac{x^3}{e^x - 1} dx} = x_{gap} \cdot \frac{\int_{x_{gap}}^{\infty} \frac{x^2}{e^x - 1} dx}{\int_0^{\infty} \frac{x^3}{e^x - 1} dx}$

- **Collection efficiency, $\eta_{collection}^{max}$** = $x_{gap} \cdot \frac{\int_{x_{gap}}^{\infty} \frac{x^2}{e^x - 1} dx}{\int_0^{\infty} \frac{x^3}{e^x - 1} dx}$

- Then, **for the case of Silicon**, where $E_{gap} \cong 1.1 \text{ eV}$. The usable fraction of energy can then be computed as the ratio of the two integrals. Changing variables to $x = \frac{E}{k_B T}$ with $T=6000 \text{ K}$ and $k_B = 8.617 \times 10^{-5} \text{ eV/K}$ gives $x_{gap} = \frac{E_{gap}}{k_B T} = \frac{(1.124 \text{ eV})}{(8.617 \times 10^{-5} \frac{\text{eV}}{\text{K}}) \cdot (6000 \text{ K})} \cong 2.1739$

$$\eta_{collection}^{max} = \frac{P_{max}^{(Si)}}{P_{total}} = 2.1739 \frac{\int_{2.17}^{\infty} \frac{x^2}{e^x - 1} dx}{\int_0^{\infty} \frac{x^3}{e^x - 1} dx} \cong 0.439$$

- Thus, a silicon photovoltaic solar cell can extract at most **43.9%** of the energy in 6000 K blackbody radiation.

- An **upper bound on the collection efficiency of a solar cell** can be derived in this way for any material for which E_{gap} is known; the 1.124 eV gap of silicon is very close to optimal for collecting energy from a 6000 K blackbody spectrum.

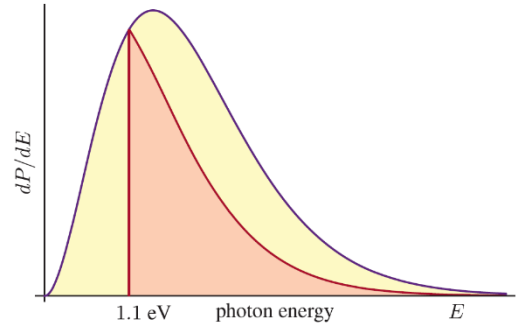


Figure 25.12 The solar power that can be collected by a silicon solar cell with $E_{gap} = 1.1 \text{ V}$ is shown in orange. The total (yellow + orange) area indicates the total power available in a thermal spectrum at 6000 K. The collection efficiency is the ratio of the orange and total areas.

- Photons with energy **less than 1.1 eV do not have enough energy** to excite electrons across the bandgap. Therefore, **cannot generate electricity** in a silicon solar cell.