



Lepton Number And Neutrino Masses

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I review the arguments that led theoretical physicists in the 1970's to suspect that lepton number is not an exact symmetry of nature, and to guess a rough range of neutrino masses.

I should say at the outset that I do not have any real novelty to offer today. The study of neutrino masses and neutrino oscillations is an area where experiment is definitely ahead of theory. That may be even more true by the end of this meeting. All I can do is to review some ideas about neutrino masses that are not novel and are undoubtedly familiar to many of you but that remain fresh and important.¹

In the 1950's, the discovery of the “two component” nature of the neutrino – neutrinos have left-handed spin around their direction of motion, and antineutrinos have right-handed spin – was generally understood to show that the neutrino must be exactly massless. Indeed, a massive spin one-half particle has two helicity states (“spin up” and “spin down”), while the neutrino only has one.

There is a potential fallacy here, which was understood at an early stage by a few physicists,² but came to be widely appreciated only in the 1970's in the course of the development of unified gauge theories of the strong, weak, and electromagnetic interactions, which are often called GUT's. In the course of this work, particle theorists came to widely suspect that neutrinos have mass and to guess a likely mass range that overlaps with the range explored in present experi-

ments.

The fallacy is as follows. To argue, given that the neutrino only has one helicity state, that it should be exactly massless, one has to assume that *lepton number* is an exactly conserved quantity. If lepton number is not conserved, one can treat the left-handed neutrino and right-handed antineutrino as two different helicity states of one particle, and combine them to make a massive spin 1/2 particle. (Such a particle is called a massive Majorana fermion; this just means that it is its own antiparticle and in particular carries no conserved charge.) If lepton number is an exact symmetry, then the neutrino and antineutrino, because they have different lepton number, cannot be combined in this way as two different states of one particle.

In fact, observed particle interactions conserve the baryon number B and the three lepton numbers L_e , L_μ , and L_τ . I will write L for the total lepton number $L = L_e + L_\mu + L_\tau$. In the 1970's, theorists generally became convinced that none of these would be true symmetries, and (hence) that neutrinos would be massive.

I'll try to list, in the rough historical order in which they had their impact (I depart from the historical presentation at the end), the principal arguments that suggested that the baryon and lepton symmetries would be violated. (Some of the arguments are more compelling for baryon number.)

(1) All attempts to unify the particles and forces in nature have required that physicists postulate symmetry-violation, as a result of putting fermions of different types in the same gauge mul-

¹Because the material is so standard, I have omitted references. I will just mention one original source where some of the issues are treated: S. Weinberg, “Expectations For Baryon And Lepton Nonconservation,” First Workshop on Grand Unification (Durham, NH 1980).

²In the talk after mine, John Bahcall reminded us of the prescient early work of Bruno Pontecorvo.

triplets. This was true in the Pati-Salam model, and in the grand unified models based on groups such as $SU(5)$, $SO(10)$, E_6 , and it is certainly true in unified string theories. To be more exact, the $SU(5)$ model in its minimal form conserves only one linear combination of the four symmetries, namely $B - L$, while in $SO(10)$ and E_6 (and the string theories), it is natural to break all of them. For example, in the case of the $SU(5)$ model, one combines antiquarks and leptons into a single gauge multiplet

$$\begin{pmatrix} \bar{q} \\ \bar{q} \\ \bar{q} \\ \nu \\ e \end{pmatrix} \quad (1)$$

with a somewhat similar structure for the other fermions. As a result, transitions inside the gauge multiplets can violate the baryon and lepton symmetries.

$B - L$ symmetry is enough to make the neutrinos exactly massless (assuming that the usual left-handed neutrinos and right-handed antineutrinos are the only relevant light helicity states). This is because the neutrino and anti-neutrino have different values of $B - L$, so the naive argument for a massless neutrino is valid if $B - L$ is a symmetry. Hence in practice the minimal $SU(5)$ model has massless neutrinos, while most of its more elaborate cousins (such as the $SO(10)$ and E_6 models in which all fermions in a “generation” are unified in a single gauge multiplet) have massive neutrinos.

(2) At first sight, it might seem that if B , L_e , L_μ , and L_τ are not fundamental symmetries of nature, there is a mystery to explain: Why do ordinary particle interactions respect these symmetries?

It turns out that this question has a completely natural answer. Using the fields of the standard model, it is impossible at the classical level to violate the baryon and lepton number symmetries by renormalizable interactions.³ Asking why the

³As I explain shortly, B and L violation in the standard model is induced at an incredibly low level, much too low to be detectable, by a quantum anomaly. The point at the moment is to explain why ordinary standard model

standard model conserves B and L is thus analogous to asking why QED conserves parity even though other forces do not. The answer seems to be that QED conserves parity because with only the fields of QED (photons and charged leptons), it is impossible to violate parity by renormalizable interactions. The same argument explains conservation of B , L_e , L_μ , and L_τ by ordinary standard model processes.

The distinguished role of renormalizable interactions comes because they have dimensionless coupling constants that are pure numbers. In contrast, unrenormalizable interactions are expected to be suppressed by a power of $1/M$, where M is a mass scale at which the standard model breaks down, perhaps a mass 10^{15} GeV suggested by the Georgi-Quinn-Weinberg computation of running of coupling constants (and hence characteristic of many grand unified theories), or the Planck mass, roughly 10^{19} GeV, characteristic of gravity.

In the standard model, the lowest dimension operator that violates baryon number is

$$\frac{1}{M^2} QQQ L \quad (2)$$

(where Q and L are standard model gauge multiplets containing quark and lepton fields, respectively) while the lowest dimension operator that violates lepton number is

$$\frac{1}{M} H H L L, \quad (3)$$

where H is a multiplet of Higgs fields.

If we assume that 10^{15} GeV $\leq M \leq 10^{19}$ GeV, we get a proton lifetime in the range

$$10^{30} \text{ years} \leq \tau_p \leq 10^{45} \text{ years} \quad (4)$$

and a neutrino mass in the range

$$10^{-5} \text{ eV} \leq m_\nu \leq 10^{-1} \text{ eV}. \quad (5)$$

For the proton lifetime, there was an attractive model (the minimal $SU(5)$ model without supersymmetry) that predicted that the proton lifetime would be quite close to the lower end of couplings conserve B and L .

the suggested range. This prediction has been excluded experimentally, and since then, though I think most particle theorists still believe that the above estimate is on the right track, there has been no convincing quantitative prediction of τ_p . For neutrino masses, the considerations have always been qualitative, and, despite some interesting attempts, there has never been a convincing quantitative model of the neutrino masses.

What I have explained above is a slightly abstract way to estimate the neutrino mass. An elegant mechanism that fits into this general scheme (and which enables one to see how to make the neutrino mass bigger or smaller than the above range, if desired) is the “see-saw mechanism.” Here one assumes that a right-handed neutrino ν_R exists but is very heavy. In fact, we assume that ν_R is a singlet of the standard model gauge group (as the $V - A$ theory of weak interactions suggests). One can then have an ordinary Dirac neutrino mass

$$m\bar{\nu}_R\nu_L + c.c. \quad (6)$$

that need not be small. It originates from a standard model coupling

$$H\bar{\nu}_R \begin{pmatrix} \nu \\ e \end{pmatrix}, \quad (7)$$

similar to the couplings that give bare masses to quarks and charged leptons. Hence one might expect m to be comparable to the quark and charged lepton masses. But ν_R can also have a very large Majorana mass

$$M\bar{\nu}_R\nu_R + c.c. \quad (8)$$

(this violates lepton number conservation, since it couples two $\bar{\nu}$'s rather than a ν and a $\bar{\nu}$). When we combine these two terms, we get a 2×2 mass matrix

$$\begin{pmatrix} 0 & m \\ m & M \end{pmatrix} \quad (9)$$

for the states $\begin{pmatrix} \nu_L \\ \bar{\nu}_R \end{pmatrix}$. Supposing that $m \ll M$, the result, after “integrating out” $\bar{\nu}_R$, is a Majorana mass term

$$\frac{m^2}{M}\nu_L\nu_L + c.c., \quad (10)$$

that is, the ordinary neutrino gets a nonzero L -violating mass m^2/M .

If we set m equal to the scale of electroweak symmetry breaking, and use the above range for M , this leads to the range of neutrino masses suggested above. Actually, since quark and charged lepton masses vary over many orders of magnitude (from the electron to the top quark), we can make the neutrino mass quite a lot less by making m smaller. Or we can make the neutrino masses bigger by making M smaller than the usual GUT scale. For example, in conventional GUT theories, this will occur if ν_R is naturally massless at tree level.

Whether we take the abstract approach or rely on the see-saw mechanism, we would expect neutrino oscillations, since it is hard to see why the neutrino mass matrix would be diagonal in the basis

$$\begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{pmatrix}. \quad (11)$$

However, large mixing angles seem a bit surprising, at least to me.

(3) So far we have considered arguments that suggest that baryon and lepton number violation, and hence neutrino masses, are natural. Let us now consider arguments that tend to show that B and L violation are inevitable.

The first such argument involves electroweak instantons. 't Hooft showed in 1976 that via instantons of the $SU(2) \times U(1)$ gauge group, the standard model actually has a mechanism that violates most of the baryon and lepton number symmetries. In fact, the only linear combination of those symmetries that is conserved additively is $B - L$; in addition, B is conserved mod three (because the standard model has three generations). The 't Hooft process will cause the deuteron to decay to an antiproton plus antileptons at a fantastically slow rate (the lifetime is far more than 10^{100} years). Despite being unmeasurably small (and not generating neutrino masses because it conserves $B - L$), this process shows that, given our present knowledge, it is all but impossible to interpret B and L as fundamental symmetries of nature.

Indeed, back in 1976, one could imagine adding additional massive fields to the standard model in such a way as to cancel the B and L violation by instantons. With the high precision tests of the standard model that we now have, this is virtually impossible.

(4) Now I will consider black holes and quantum gravity. This is the one point at which I will consider issues that, while certainly not at all novel, are perhaps not always considered in relation to neutrino masses.

Classically, a black hole absorbs matter and radiation and does not emit. Quantum mechanically, this is impossible as it violates the hermiticity of the Hamiltonian H . Indeed, if $\langle f|H|i\rangle$ is a nonzero matrix element describing absorption, then its complex conjugate $\langle i|H|f\rangle$ is a nonzero matrix element describing emission. In fact, Hawking showed in 1974 that quantum black holes do emit approximately thermal radiation,⁴ at a temperature that is fantastically small for black holes of astronomical masses. Such a massive black hole, placed in empty space, will slowly lose mass to Hawking radiation. As it loses mass, it shrinks. Ultimately, the hole shrinks down to the Planck scale. At that point, Hawking's approximations break down. By this time, the great bulk of the hole's mass has been radiated away. The radiation emitted after Hawking's approximations break down is presumably highly non-thermal, but can carry away only a very small fraction of the rest mass of the black hole, since the bulk of the rest mass has been radiated away during the period when Hawking's approximations were valid.

Now, let us consider making a black hole from a collapsing star, or from infalling tables and chairs – anything with large baryon and lepton number. One does not get the baryons back in the outgoing Hawking radiation, since for the great bulk of the black hole evaporation process, the Hawking temperature is much too small compared to the proton mass to enable the emission of any significant number of baryons or antibaryons. At the

endpoint of the black hole evaporation, we may get back some baryons, but only a small number since only a small mass remains in the black hole. So in sum, the formation and evaporation of a black hole does not conserve baryon number, and hence this cannot be an exactly conserved quantity in nature.

What about lepton number? Here we have to be a little more careful, because it may be that some of the neutrino species are massless, and if so neutrinos and antineutrinos make up (along with photons, gravitons, and any other massless particles that may exist in nature) a large part of the black hole emission. However, because the black hole emission is approximately thermal, no significant net lepton number is emitted during the bulk of the black hole history. In case there may exist massless neutrinos in nature, carrying lepton number, we should worry that perhaps the final pulse of radiation from a black hole could carry away the original lepton number. (If so, this “pulse” would last a very long time, since Fermi statistics limit how many neutrinos carrying a bounded energy can be emitted in a given space in a given time.) This hypothesis would apparently contradict subtler aspects of the black hole evaporation story. Consider a black hole with a mass M_S comparable to that of the sun. If the lepton number of a black hole is a well-defined quantum number, such a black hole might have a lepton number of 10^{57} (if it formed from collapse of a sun-like star) or any larger number (if the black hole formed from collapse of a larger mass and then emitted Hawking radiation to reduce its mass to M_S). So black holes with mass M_S have infinitely many quantum states, labeled by the lepton number. This contradicts the claim that the number of quantum states of a black hole of given mass is the exponential of the Bekenstein-Hawking entropy (which is $1/4$ of the area of the horizon in Planck units) and in particular is finite.

I do not claim that arguments such as these are absolutely air-tight, but they do show that it will be hard to reconcile a claim of absolute lepton number conservation in nature with what we know about black hole physics.

In contrast to global symmetries such as baryon and lepton number, gauge symmetries such as

⁴The corrections to the thermal description of the outgoing radiation are, for a macroscopic black hole, so subtle and small that they have never been successfully computed.

electric charge conservation cause no trouble for black hole physics. First of all, electric charge definitely is a well-defined quantum number of black holes. One sees it in the classical (Reissner-Nordstrom) black hole solution; it can be measured exterior to the black hole horizon, by a flux integral for the electric field on a large sphere at infinity. The electric field E of the hole contributes an energy $\int d^3x \sqrt{g} E^2 / 8\pi$ to the black hole mass, so black holes of bounded mass also have bounded electric charge. It is *global* additive symmetries that cause difficulties for black hole physics.

Since we have only used general properties of black holes and their Hawking emission (which is deduced semi-classically and not on the basis of a detailed microscopic understanding of black holes), the conclusion that baryon and lepton number are not symmetries of black hole physics should apply in any consistent theory of quantum gravity. Hence, we can test it out in string theory, which is our only real candidate at the moment for such a theory. We do not need to assume that string theory describes nature; the arguments above indicate that global additive conservation laws such as baryon and lepton number conservation cannot hold in any consistent quantum gravity theory, whether it describes nature or not.

We indeed find that in string theory, we never get an additive global conservation law. At least in known string vacua, the apparent additive global symmetries turn out – sometimes in a tricky fashion – to be either gauged or explicitly broken. For example, one can add Chan-Paton factors to open strings, apparently introducing a global symmetry, but upon further examination this turns out to be a gauge symmetry. Similarly, in the heterotic string, the global symmetries on the world sheet turn out to be gauge symmetries (not global symmetries) in space time. As an example of the opposite kind, in compactification of the Type II or heterotic string theories from ten to four dimensions on a compact manifold K , one gets approximately massless “axions” associated with the second Betti number $b_2(K)$. The shift symmetries of these axions appear to be continuous global symmetries (albeit spontaneously

broken), but on closer examination, one finds that these symmetries are explicitly violated by world-sheet instantons.

In summary, these arguments seem to show that B and L cannot be conserved quantities in any theory that includes quantum gravity. This is a strong hint of neutrino masses, but there is a loophole. Although L , for example, cannot be an exactly conserved quantity, it might be conserved mod 7, or mod n for any integer n . For $n > 2$, this would suffice to prevent neutrino masses. If one wishes in string theory to make a model in which neutrino masses are exactly zero, one would arrange to have a discrete \mathbf{Z}_n symmetry producing such a mod n conservation law. Discrete symmetries that prevent neutrino masses could well prevent quark or charged lepton masses, so it actually would take some care to arrange a model in which the neutrinos are exactly massless and the other fermions have mass.

(5) Finally, the last hint that B and L are not exact symmetries in nature is simply that the real universe has a small but nonzero B and L density relative to the entropy density. (In the case of L , we do not know this for sure, since the observed net density of electrons might conceivably be canceled by not yet observed relic antineutrinos.) If B and L are not exact symmetries, then making use of the C and CP violation in particle interactions, and the CPT violation in cosmology that comes from the fact that the universe is expanding, it is possible to spontaneously generate nonzero B and L densities. There is not a convincing quantitative theory, but the observed B and L densities (about 10^{-9} of the entropy density) are in a natural range. If B and L are exactly conserved quantities, they must be postulated as initial conditions in the Big Bang; this is certainly less attractive.

I hope that today I have at least managed, albeit without saying anything that is really new, to convey a sense of the exciting and wide-ranging theoretical ideas relevant to the neutrino mass question. Hopefully, experimentalists are beginning to come back with answers. I am sure we will hear more in the next few days.

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