

# ASTR 531 - Stellar Interiors and Evolution

Exam 1

TOM WAGG

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## Thriving under high pressure

*Dearest Grad Student,*

*My name is Professor Daniel Eyemund and I'm reaching out to you with an exciting business idea - diamonds! I've realised that I could get around applying for grants by generating a large number of diamonds and discretely selling them! They might be rare on Earth but with the right conditions perhaps we could create a **bunch** of them out in space. The plan is as follows:*

- 1. Collect a big cloud of carbon*
- 2. Squish it down and inject it into a white dwarf*
- 3. Create diamonds*
- 4. ...*
- 5. Profit?*

*Your mission, should you choose to accept it (... which you have to because this is an exam), is to find out whether a white dwarf could produce diamonds, if so, what kind of white dwarf we need! I feel your experiences with stellar interiors, not to mention high pressure environments, would make you ideally suited to this task. Please note your timely response would be appreciated as I will be meeting with Elon to pitch the idea in 30 minutes.*

*Yours sincerely,  
Prof. D. Eyemund*

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### **Part a - Complete degeneracy pressure for electrons**

*Given that, for complete degeneracy, the  $n_e(p)$  distribution is rectangular for  $p < p_F$  such that  $n_e(p) = 2/h^3$ , show that the pressure scales as  $P_e \propto n_e^{5/3}$ .*

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### **Part b - Polytrope central pressure**

*By showing that part a is true, you've proven that we can use an  $n = 3/2$  polytrope! It'll now be useful to know an expression for the central pressure of a polytrope star.*

*Use the fact that  $P_c = K\rho_c^{1+1/n}$  to show that  $P_c = A\frac{GM^2}{R^4}$  (and find the constant  $A$  in terms of  $D_n, \pi, n, M_n$ ). Hint: You can find  $K$  by comparing an expression for  $\alpha$  and an expression for  $M_n$ . Additionally note that  $\rho_c = \bar{\rho}D_n$ .*

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To start, we need to find an expression for  $K$ . We have the definition of  $\alpha$  from the Lane-Emden

equation (see Eq. 11.5).

$$\alpha = \left[ \frac{(n+1)K}{4\pi G \rho_c^{1-1/n}} \right]^{1/2} \quad (1)$$

Next we can use Equation 11.10a to find another expression for  $\alpha$  from the definition of  $M_n$

$$M = 4\pi a^3 \rho_c M_n \quad (2)$$

$$\alpha = \left( \frac{M}{M_n 4\pi \rho_c} \right)^{1/3} \quad (3)$$

Combining these yields an expression for  $K$

$$\left[ \frac{(n+1)K}{4\pi G \rho_c^{1-1/n}} \right]^{1/2} = \left( \frac{M}{M_n 4\pi \rho_c} \right)^{1/3} \quad (4)$$

$$K = \left( \frac{M}{M_n 4\pi \rho_c} \right)^{2/3} \frac{4\pi G \rho_c^{1-1/n}}{(n+1)} \quad (5)$$

$$K = \left( \frac{M}{M_n} \right)^{2/3} \frac{(4\pi)^{1/3} G \rho_c^{1/3-1/n}}{(n+1)} \quad (6)$$

Now, given that  $P_c = K \rho_c^{1+1/n}$ , this means that we can write the central pressure as

$$P_c = \left( \frac{M}{M_n} \right)^{2/3} \frac{(4\pi)^{1/3} G \rho_c^{1/3-1/n}}{(n+1)} \rho_c^{1+1/n} \quad (7)$$

$$= \left( \frac{M}{M_n} \right)^{2/3} \frac{(4\pi)^{1/3} G \rho_c^{4/3}}{(n+1)} \quad (8)$$

$$= \left[ \frac{(4\pi)^{1/3} G}{(n+1) M_n^{2/3}} \right] M^{2/3} \rho_c^{4/3} \quad (9)$$

We also know that the average density is given by  $\bar{\rho} = \frac{M}{\frac{4}{3}\pi R^3}$  and it is related to the central density as  $\rho_c = \bar{\rho} D_n$ . Therefore we can express the central pressure

$$P_c = \left[ \frac{(4\pi)^{1/3} G}{(n+1) M_n^{2/3}} \right] M^{2/3} \left( \frac{M D_n}{\frac{4}{3}\pi R^3} \right)^{4/3} \quad (10)$$

$$= \left[ \frac{(4\pi)^{1/3} G}{(n+1) M_n^{2/3}} \right] \left( \frac{D_n}{\frac{4}{3}\pi} \right)^{4/3} \frac{M^2}{R^4} \quad (11)$$

$$P_c = \underbrace{\left[ \frac{(3D_n)^{4/3}}{4\pi(n+1) M_n^{2/3}} \right]}_{=A} \frac{GM^2}{R^4} \quad (12)$$

### **Part c - White dwarf M-P<sub>c</sub> relation**

Now let's take that and create a relation for the central pressure of a white dwarf as only a function of mass. Apply the M-R relation for WDs (assuming  $\mu_e = 2$ ) and that  $n = 3/2$  to show that  $P_c = \frac{AG}{B^4} M^C$  where  $B$  and  $C$  are constants you should find.

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The WD  $M$ - $R$  relation with  $\mu_e = 2$  gives that

$$R = 0.012 \, \text{R}_\odot \left( \frac{M}{\text{M}_\odot} \right)^{-1/3} \quad (13)$$

This gives the final result of the central pressure as a function of WD mass

$$\boxed{P_{\text{c,WD}} = \frac{AG}{B^4} M^C} \quad (14)$$

where  $B = 0.012 \, \text{R}_\odot \, \text{M}_\odot^{1/3}$  and  $C = 10/3$ .