

The Non-Astronomer's Discerning Guide to Tom's Defence

A quick cheatsheet for a variety of astro jargon

Stellar evolution



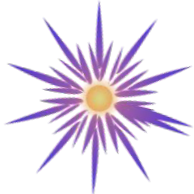
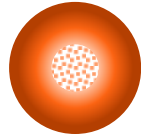
Massive stars for my purposes are those that end their lives in a supernova, so at least $\sim 7x$ more massive than the sun

More massive stars evolve more quickly because of higher pressures/temperatures in their cores, accelerating nuclear fusion



The **metallicity** of a star is the fraction that is “metals”, meaning not hydrogen or helium. The *earlier* in the Universe a star is born, generally the *lower* its metallicity

Massive stars have **convective cores** (which mix up all of the elements), and **radiative envelopes** (which don't)



A **supernova** is a huge explosion when the core of a massive star collapses – they can obliterate entire solar systems and outshine galaxies!



After a supernova, the star leaves behind a **neutron star** or a **black hole** (approximately depending on the star's mass) – they are extremely dense and massive objects

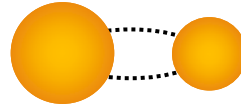


Supernovae aren't *completely* symmetrical; more material goes in one direction. So, by Newton's third law, the resulting objects get a **natal kick**, which can cause them to fly off at 100s of km/s!

These kicks can **unbind** a binary, causing the star to also be ejected at high velocities, potentially producing **runaway stars** (which move faster than ~ 30 km/s)

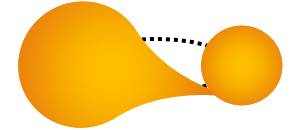


Binary stellar interactions



Most massive stars have ≥ 1 companion that they orbit in a **binary**

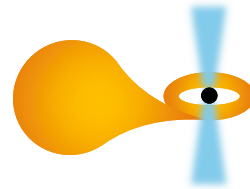
Stars in binaries can interact through **mass transfer**, where the **donor** star pours material onto the **accretor** star



If this becomes unstable, the two stellar cores can orbit in a single **common envelope**, which can shrink the orbit, or lead to a **stellar merger**



You can make all sorts of exotic binaries, like **X-ray binaries** where a star is pouring material onto a compact object, forming a disc of material, with huge polar jets – all emitting lots of X-rays!



If you get really lucky, you can form a binary with two compact objects (e.g. two black holes). We can't see these with telescopes, but when they collide and merge, they shake the fabric of spacetime so violently that we can detect the **gravitational waves** they emit!

Miscellaneous terms

Parsec: ~ 3 light-years, $\sim 20,000$ billion miles

Asteroseismology: The study of stellar pulsations

High-redshift galaxies: Galaxies from the early universe

LISA: Future space-based gravitational-wave detector

Hydrodynamical simulations: Encapsulate gas physics, heating, gravity & more

Compact object: A white dwarf, neutron star, or black hole

Double compact object: As above, but two bound together in a binary

O B A F G K M

Stellar classifications, most massive are O, least are M