NuWro - neutrino MC event generator

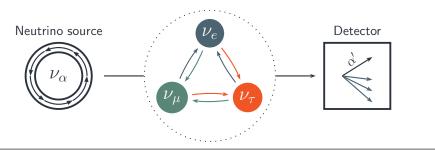
Tomasz Golan

28.04.2017, UW HEP Seminar





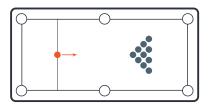
1. Motivation



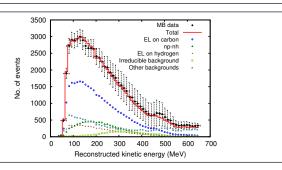
2. NuWro event generator



3. Final State Interactions



4. MB NCEL data analysis



Introduction



Basic properties of neutrinos

Introduction

Neutrino properties

PMNS matrix
Probability of oscillation
Neutrino oscillation
Measurement idea
Example: T2K
Energy reconstruction

MC generators

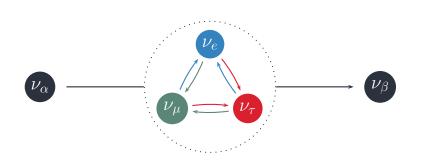
NuWro

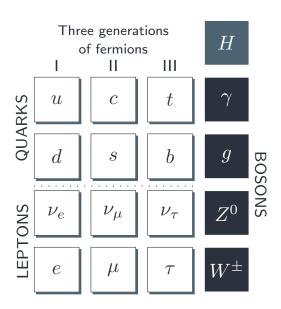
Final state interactions

MB NCEL analysis

Backup slides

- $\frac{1}{2}$ -spin, no electric charge, small mass (extremely hard to detect)
- Interactions with elementary particles
 - → electroweak theory (Standard Model)
- Interactions with nucleons
 - \rightarrow form factors
 - → parton distribution functions
- lacksquare Interactions with nuclei ightarrow nuclear effects
- The neutrino flavor state is a superposition of the mass states:





 $|\nu_{\alpha}\rangle = \sum_{i} U_{\alpha i} |\nu_{i}\rangle$

The neutrino produced in α state can be measured in β state. The phenomenon is called neutrino oscillation.



Pontecorvo-Maki-Nakagawa-Sakata matrix

The PMNS matrix defines the mass mixing in the lepton sector:

$$U = \begin{bmatrix} 1 & 0 & 0 \\ 0 & c_{23} & s_{23} \\ 0 & -s_{23} & c_{23} \end{bmatrix} \begin{bmatrix} c_{13} & 0 & s_{13}e^{-i\delta} \\ 0 & 1 & 0 \\ -s_{13}e^{i\delta} & 0 & c_{13} \end{bmatrix} \begin{bmatrix} c_{12} & s_{12} & 0 \\ -s_{12} & c_{12} & 0 \\ 0 & 0 & 1 \end{bmatrix}$$

$$c_{ij} = \cos \theta_{ij}$$

$$s_{ij} = \sin \theta_{ij}$$

$$\theta_{ij}$$
 - mixing angels

$$c_{ij} = \cos \theta_{ij}$$
 $s_{ij} = \sin \theta_{ij}$ θ_{ij} - mixing angels δ - CP phase factor

Measurements:

$$\sin^2(2\theta_{12}) = 0.846 \pm 0.021$$

$$\sin^2(2\theta_{23}) > 0.92$$

$$\sin^2(2\theta_{13}) = 0.093 \pm 0.008$$

 δ can be measured if and only if all mixing angles are nonzero. If $\delta \neq 0$ CP is violated.



Probability of oscillation

Introduction

Neutrino properties PMNS matrix

Probability of oscillation

Neutrino oscillation Measurement idea Example: T2K Energy reconstruction

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Backup slides

$$P(\nu_{\alpha} \to \nu_{\beta}) = |\langle \nu_{\beta}(x) | \nu_{\alpha}(y) \rangle|^{2} = \delta_{\alpha\beta}$$

$$- 4 \sum_{i>j} \operatorname{Re} \left[U_{\alpha i}^{*} U_{i\beta} U_{\alpha_{j}} U_{j\beta}^{*} \right] \sin^{2} \left(\frac{\Delta m_{ij}^{2} L}{4E} \right)$$

$$+ 2 \sum_{i>j} \operatorname{Im} \left[U_{\alpha i}^{*} U_{i\beta} U_{\alpha_{j}} U_{j\beta}^{*} \right] \sin^{2} \left(\frac{\Delta m_{ij}^{2} L}{2E} \right)$$

L - traveled distance (fixed), E - neutrino energy (will be discussed)

Measurements:

Solar neutrinos:
$$\Delta m_{21}^2 = (7.53 \pm 0.18) \cdot 10^{-5} eV^2$$

■ Atmospheric neutrinos:
$$|\Delta m_{32}^2| = (2.44 \pm 0.06) \cdot 10^{-3} eV^2$$

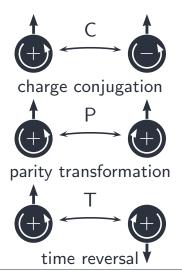


Neutrino oscillation

The probability of neutrino oscillation depends on:

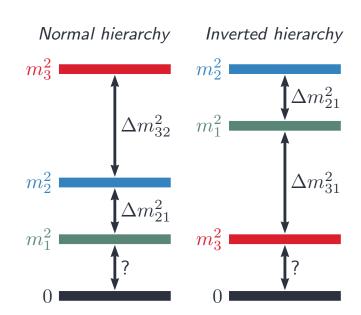
- What we want to measure:
 - three mixing angles already measured
 - the difference between squared mass of the mass states the sign of Δm^2_{32} is still unknown, what is the mass hierarchy?
 - the CP phase factor is the CP symmetry broken?

$$P(\bar{\nu}_{\alpha} \to \bar{\nu}_{\beta}) = P(\nu_{\beta} \to \nu_{\alpha}) \stackrel{?}{=} P(\nu_{\alpha} \to \nu_{\beta})$$





- the distance traveled by a neutrino
- neutrino energy this is a problem





How to measure neutrino oscillation?

Introduction

Neutrino properties PMNS matrix Probability of oscillation Neutrino oscillation

Measurement idea

Example: T2K

Energy reconstruction

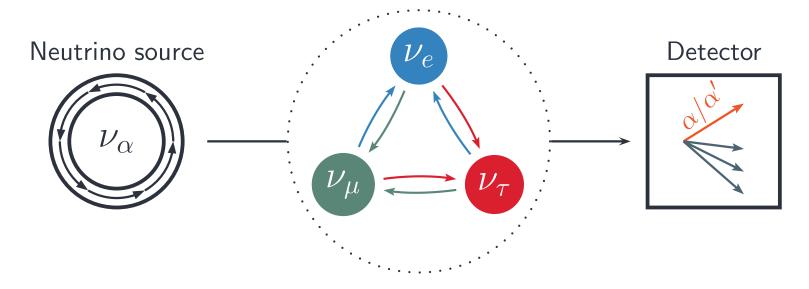
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Backup slides



Disappearance method

Number of α -flavor neutrinos

 \sim Number of lpha-flavor charged leptons

Appearance method

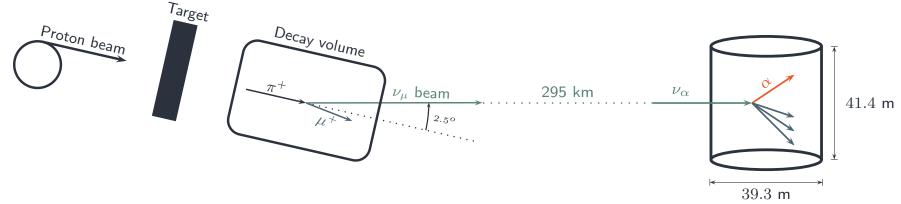
Number of α -flavor neutrinos

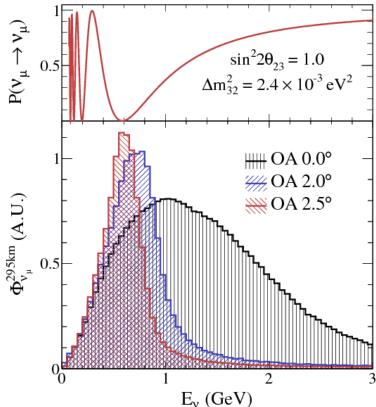
←

Number of α' -flavor charged leptons



Example: T2K





- Off-axis beam
- Cherenkov detector (Super-Kamiokande)
- 50 000 tons of ultra-pure water
- Only charged leptons and final state charged hadrons are visible
- The neutrino energy is unknown



The problem with neutrino energy reconstruction

Introduction

Neutrino properties PMNS matrix Probability of oscillation Neutrino oscillation Measurement idea Example: T2K

Energy reconstruction

MC generators

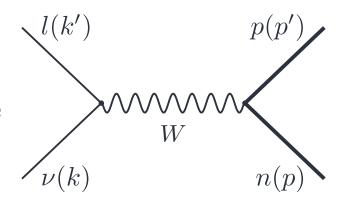
NuWro

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Backup slides

In the case of neutrino scattering off nucleon at rest neutrino energy can be calculated from lepton kinematics:



$$M_p^2 = p'^2 = (p + k - k')^2$$

$$M_p^2 = m_l^2 + M_n^2 - 2M_n E_l + 2E_{\nu}(M_n - E_l + |\vec{p}_l| \cos \theta_l)$$

$$E_{\nu} = \frac{M_p^2 - M_n^2 - m_l^2 + 2M_n E_l}{2(M_n - E_l + |\vec{p}_l| \cos \theta_l)}$$

Quasi-elastic scattering

$$k = (E_{\nu}, 0, 0, E_{\nu})$$

$$k' = (E_{l}, \vec{p}_{l})$$

$$p = (M_{p}, 0, 0, 0)$$

$$p' = p + q$$

$$q = k - k'$$



The problem with neutrino energy reconstruction

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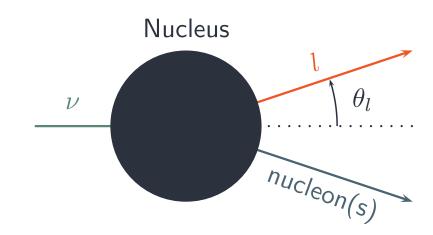
■ Usually, the energy reconstruction procedure is based on quasi-elastic neutrino-nucleon scattering:

$$E_{\nu}^{REC} = \frac{M_p^2 - (M_n - E_B)^2 - m_l^2 + 2(M_n - E_B)E_l}{2(M_n - E_B - E_l + |\vec{p_l}|\cos\theta_l)}$$

Impulse Approximation

No Fermi motion

Binding Potential





The problem with neutrino energy reconstruction

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Neutrino properties PMNS matrix Probability of oscillation Neutrino oscillation Measurement idea Example: T2K

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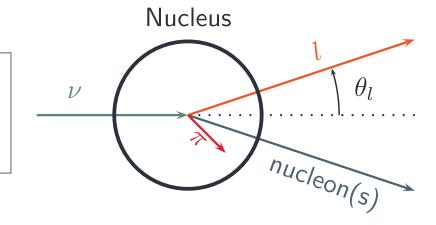
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■ Usually, the energy reconstruction procedure is based on quasi-elastic neutrino-nucleon scattering:

$$E_{\nu}^{REC} = \frac{M_p^2 - (M_n - E_B)^2 - m_l^2 + 2(M_n - E_B)E_l}{2(M_n - E_B - E_l + |\vec{p_l}|\cos\theta_l)}$$

Never judge an event by its final state particles!



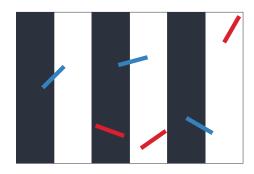
Monte Carlo generators



Buffon's needle problem

Suppose we have a floor made of parallel strips of wood, each the same width, and we drop a needle onto the floor. What is the probability that the needle will lie across a line between two strips?

Georges-Louis Leclerc, Comte de Buffon 18th century



blue are good red are bad

Monte Carlo without computers

If needle length (l) < lines width (t):

$$P = \frac{2l}{t\pi}$$

which can be used to estimate π :

$$\pi = \frac{2l}{tP}$$

MC experiment was performed by Mario Lazzarini in 1901 by throwing 3408 needles:

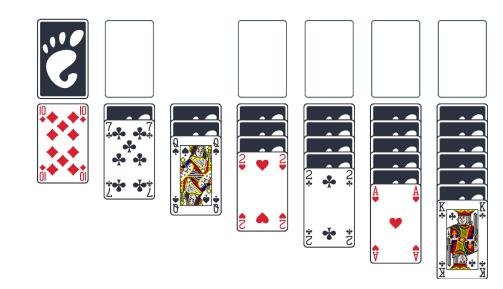
$$\pi = \frac{2l \cdot 3408}{t \cdot \#red} = \frac{355}{113} = 3.14159292$$



From Solitaire to Monte Carlo method

- Stanisław Ulam was a Polish mathematician
- He invented the Monte Carlo method while playing solitaire
- The method was used in Los Alamos, performed by ENIAC computer





- What is a probability of success in solitaire?
 - Too complex for an analytical calculations
 - Lets try N = 100 times and count wins
 - lacktriangle With $N \to \infty$ we are getting closer to correct result



MC integration (hit-or-miss method)

Introduction

MC generators

Buffon's needle problem From Solitaire to MC

Hit-or-miss method

Crude method
Methods comparison
Accept-reject
MC generators
Why do we need them?
The main problem
Cooking generator

NuWro

Final state interactions

MB NCEL analysis

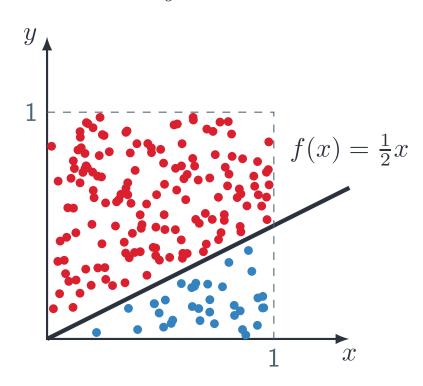
Backup slides

Lets do the following integration using MC method:

$$\int_0^1 f(x)dx = \int_0^1 \left(\frac{1}{2}x\right)dx = \left.\frac{1}{2}\frac{x^2}{2}\right|_0^1 = \frac{1}{4}$$

- take a random point from the $[0,1] \times [0,1]$ square
- lacktriangle compare it to your f(x)
- lacktriangle repeat N times
- lacktriangle count n points below the function
- you results is given by

$$\int_0^1 f(x)dx = P_{\square} \cdot \frac{n}{N} = \frac{n}{N}$$





MC integration (crude method)

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Buffon's needle problem From Solitaire to MC Hit-or-miss method

Crude method

Methods comparison Accept-reject MC generators Why do we need them? The main problem Cooking generator

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Backup slides

Lets do the following integration using MC method once again:

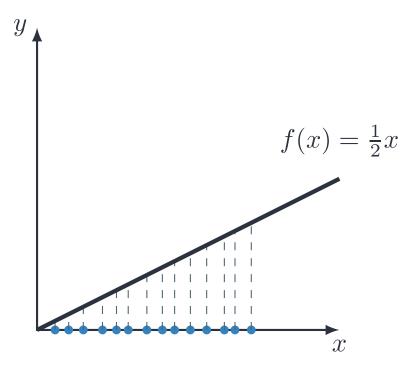
$$\int_0^1 f(x)dx = \int_0^1 \left(\frac{1}{2}x\right)dx = \left.\frac{1}{2}\frac{x^2}{2}\right|_0^1 = \frac{1}{4}$$

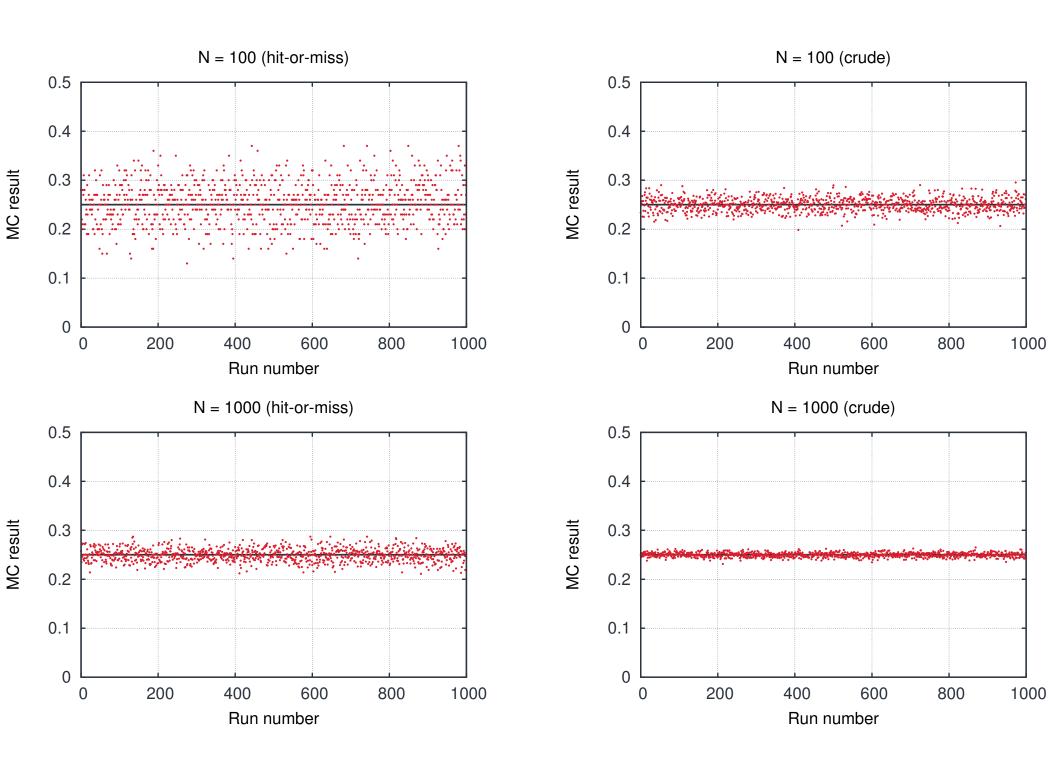
One can approximate integral

$$\int_{a}^{b} f(x)dx \approx \frac{b-a}{N} \sum_{i=1}^{N} f(x_i)$$

where x_i is a random number from [a, b]

- It can be shown that crude method is more accurate than hit-or-miss
- We will skip the math and look at some comparisons





Acceptance-rejection method

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Buffon's needle problem From Solitaire to MC Hit-or-miss method Crude method Methods comparison

Accept-reject

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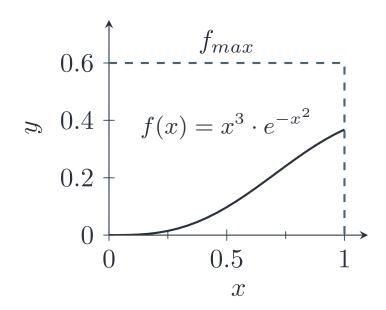
Backup slides

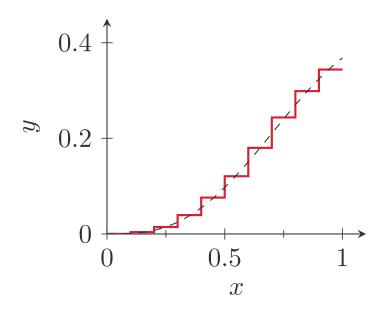
For generating events according to a distribution.

■ Evaluate $f_{max} \ge \max(f)$

Note: $f_{max} > max(f)$ will affect performance, but the result will be still correct

- \blacksquare Generate random x
- Accept x with $P = \frac{f(x)}{f_{max}}$
 - generate a random u from $[0, f_{max}]$
 - lack accept if u < f(x)







Monte Carlo event generators

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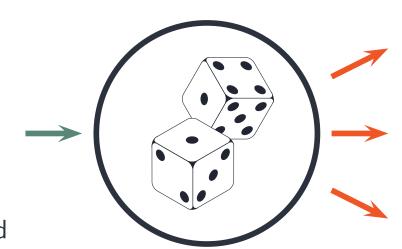
Final state interactions

MB NCEL analysis

Backup slides

- Monte Carlo generators basically do two things:
 - integrate cross section formulas
 - generate events using accept-reject method

with many optimization tricks



- Physicists have been using them since ENIAC
- Some common generators used in neutrino community:
 - transport of particles through matter: Geant4, FLUKA
 - neutrino interactions: GENIE, GIBUU, NEUT,
 NUANCE, NuWro



Why do we need them?

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Buffon's needle problem From Solitaire to MC Hit-or-miss method Crude method Methods comparison Accept-reject MC generators

Why do we need them?

The main problem Cooking generator

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Backup slides



- Monte Carlo event generators connect experiment (what we see) and theory (what we think we should see)
- Any neutrino analysis relies on MC generators
- From neutrino beam simulations, through neutrino interactions, to detector simulations
- Used to evaluate systematic uncertainties, backgrounds, acceptances...



What is the main problem?

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From Solitaire to MC
Hit-or-miss method
Crude method
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Why do we need them?

The main problem

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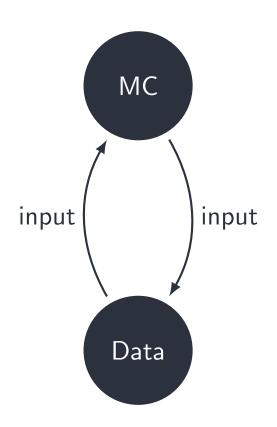
MB NCEL analysis

Backup slides

"You use Monte Carlo until you understand the problem"

Mark Kac

- In perfect world MC generators would contain "pure" theoretical models
- In real world theory does not cover everything
- Neutrino and non-neutrino data are used to tune generators

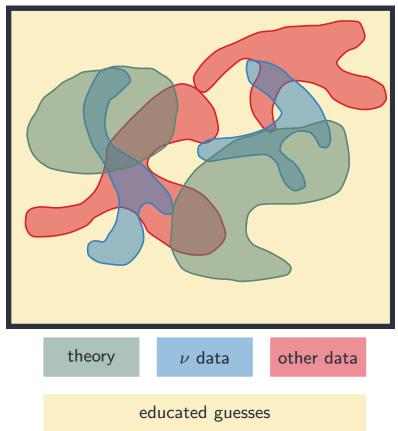




How to build generator

INGREDIENTS:

Phase space



RECIPE:



NuWro



NuWro MC event generator

Introduction

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NuWro

NuWro MC

Dynamics
(Q)EL scattering
RES pion production
Deep Inelastic Scattering
π production
Impulse approximation
Fermi gas
Spectral function
Two-body current
COH pion production
Summary

Final state interactions

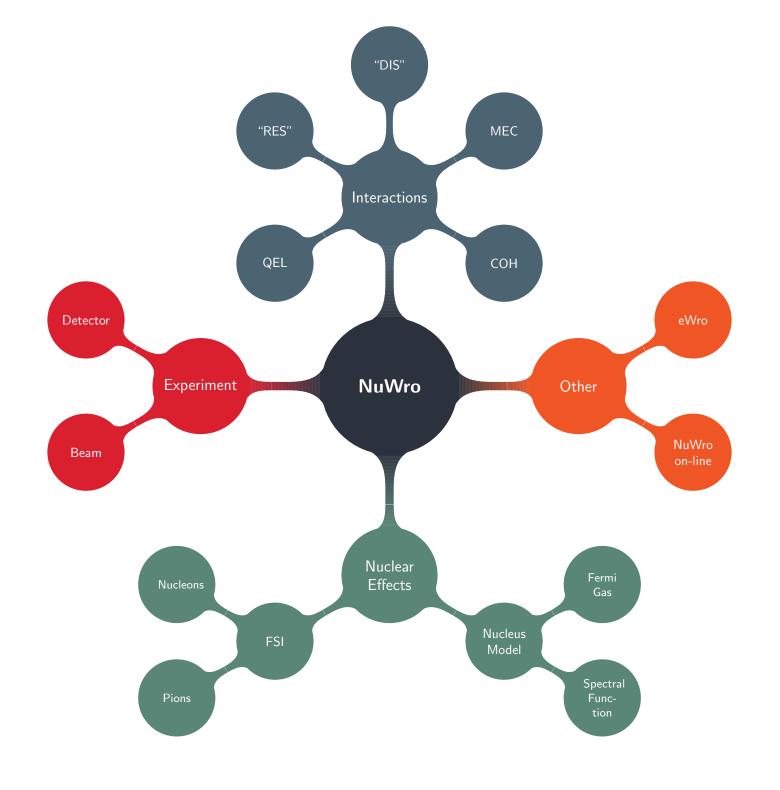
MB NCEL analysis

Backup slides



formerly known as WroNG

- It has been developed at Wroclaw University since 2006
 - Currenly involved: Jan Sobczyk, Cezary Juszczak,
 Krzysztof Graczyk, Tomasz Golan, Kajetan Niewczas
 - Significant contribution: Jarosław Nowak, Jakub Żmuda, Artur Kobyliński, Maciej Tabiszewski, Paweł Przewłocki, Patrick Stowell, Luke Pickering
- The authors were encouraged by prof. Danuta Kiełczewska
- The open source code: https://github.com/nuwro/





Implemented dynamics

All major interaction channels are implemented, for charged and neutral current, covering neutrino energy region from a few hundreds MeV (Impulse Approximation limit) to several TeV:

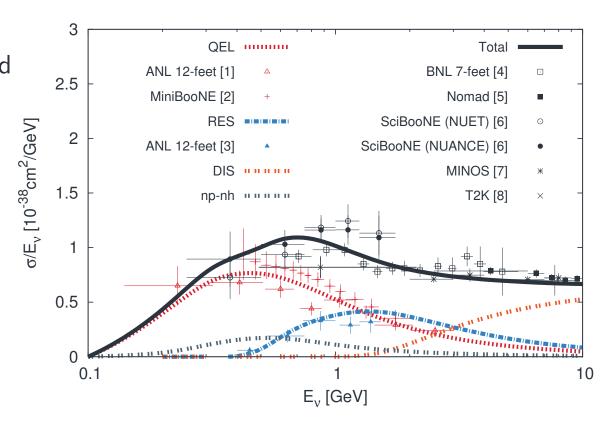
QEL (quasi-)elastic scattering

RES pion production through a Δ resonance excitation

DIS more inelastic processes

COH coherent pion production

np-nh two body current contribution



[1] PRD 19 (1979) 2521 [5] PLB 660 (2008) 19

[2] PRD 81 (2010) 092005 [6] PRD 83 (2011) 012005

[3] PRD 16 (1977) 3103 [7] PRD 81 (2011) 072002

[4] PRD 25 (1982) 617 [8] PRD 87 (2013) 092003



(Quasi-)elastic scattering

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NuWro MC Dynamics

(Q)EL scattering

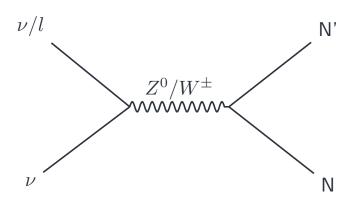
RES pion production
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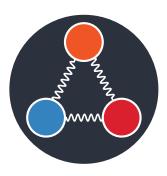
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MB NCEL analysis

Backup slides

- Llewellyn-Smith model is used for charged current quasi-elastic scattering
- Not much difference here between generators (but default parameters)





 Nucleon structure is parametrized by form factors

- Vector → Conserved Vector Current (CVC)
- Pseudo-scalar → Partially Conserved Axial Current (PCAC)
- lacktriangle Axial ightarrow dipole form with one free parameter (axial mass, M_A)



Resonance pion production

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NuWro MC Dynamics

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RES pion production

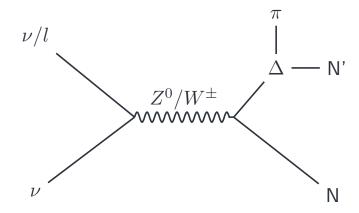
Deep Inelastic Scattering π production Impulse approximation Fermi gas Spectral function Two-body current COH pion production Summary

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MB NCEL analysis

Backup slides

- Most of generators (like NEUT and GENIE) uses
 Rein-Sehgal model
- RS model describes single pion production through baryon resonances below $W=2~{\rm GeV}$



- In NuWro Adler-Rarita-Schwinger formalism is used to calculate Δ resonance explicitly
- Non-resonant background is estimated using quark-parton model



Deep inelastic scattering [DIS]

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NuWro MC Dynamics (Q)EL scattering RES pion production

Deep Inelastic Scattering

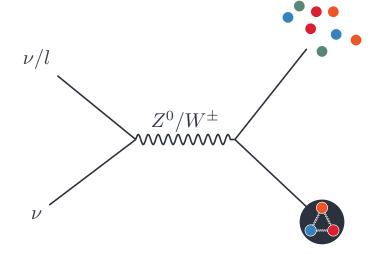
π production
Impulse approximation
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COH pion production
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Backup slides

- Quark-parton model is used for deep inelastic scattering
- Bodek-Young modification to the parton distributions at low Q^2 is included by most generators



Hadronization



- Hadronization is the process of formation hadrons from quarks
- Pythia is widely used at high invariant masses



Pion production in NuWro

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(Q)EL scattering
RES pion production
Deep Inelastic Scattering

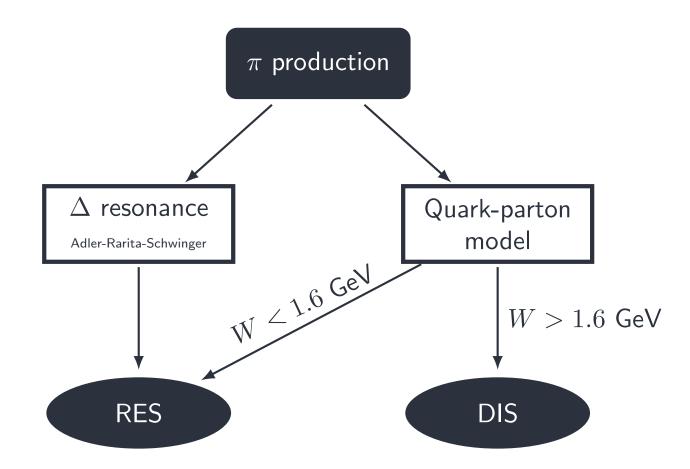
π production

Impulse approximation Fermi gas Spectral function Two-body current COH pion production Summary

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Backup slides



RES/DIS distinguish is arbitrary for each MC generator!



Impulse approximation

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NuWro MC Dynamics (Q)EL scattering RES pion production Deep Inelastic Scattering π production

Impulse approximation

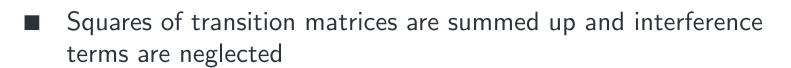
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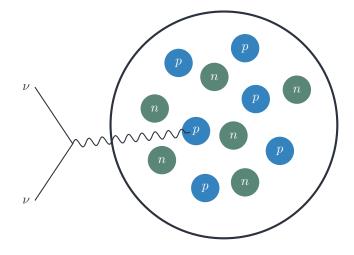
Backup slides

- In impulse approximation neutrino interacts with a single nucleon
- If $|\vec{q}|$ is low the impact area usually includes many nucleons
- For high $|\vec{q}|$ IA is justified



$$\sigma^A = \sum_{i=1}^{Z} \sigma_p + \sum_{i=1}^{A-Z} \sigma_n$$

High $|\vec{q}|$ means more than 400 MeV. However, IA is always assumed





Fermi gas

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NuWro MC Dynamics (Q)EL scattering RES pion production Deep Inelastic Scattering π production

Impulse approximation

Fermi gas

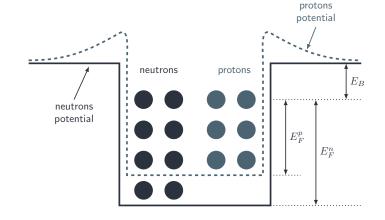
Spectral function Two-body current COH pion production Summary

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Backup slides

Nucleons move freely within the nuclear volume in constant binding potential.

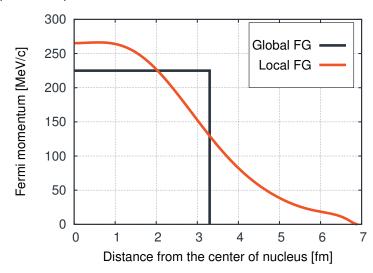


Global Fermi Gas

$$p_F = \frac{\hbar}{r_0} \left(\frac{9\pi N}{4A} \right)^{1/3}$$

Local Fermi Gas

$$p_F(r) = \hbar \left(3\pi^2 \rho(r) \frac{N}{A} \right)^{1/3}$$

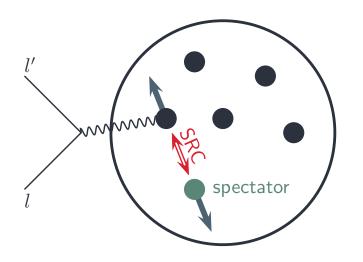


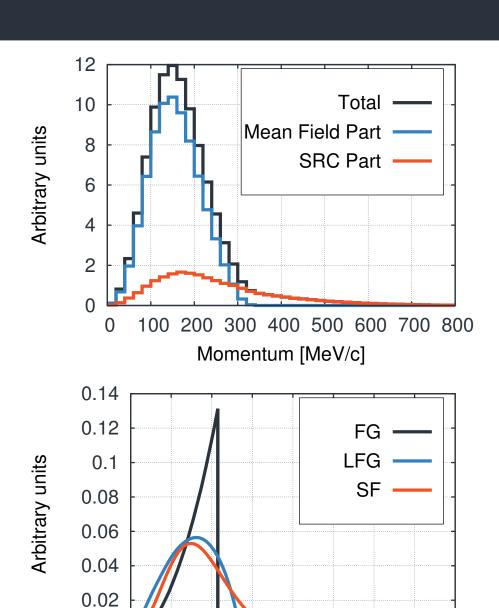


Spectral function

The probability of removing of a nucleon with momentum \vec{p} and leaving residual nucleus with excitation energy E.

$$P(\vec{p}, E) = P_{MF}(\vec{p}, E) + P_{corr}(\vec{p}, E)$$





0.2 0.3 0.4 0.5 0.6 0.7 0.8

Nucleon momentum [GeV/c]



Two-body current interactions

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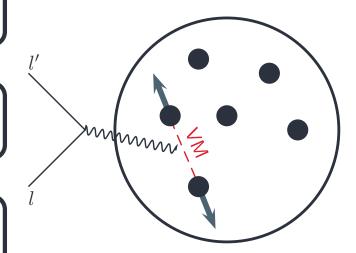
MB NCEL analysis

Backup slides

Two Body Current

2 particles - 2 holes (2p-2h)

Meson Exchange Current (MEC)



Models in generators

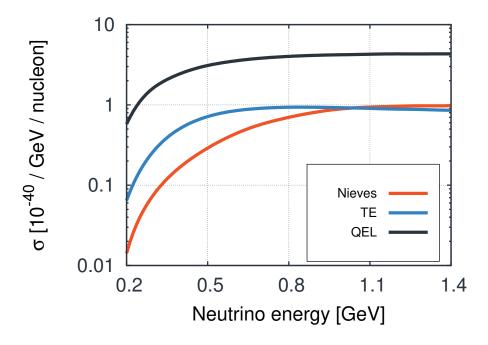
- Nieves model (GENIE, NEUT, NuWro) CC only
- Transverse Enhancement model (NuWro) both CC and NC



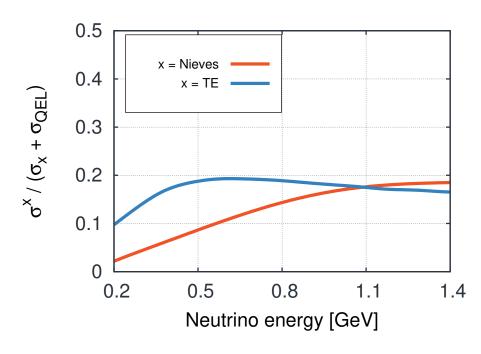
Two-body current interactions

- Nieves model is microscopic calculation
- TE model introduce 2p-2h contribution by modification of the vector magnetic form factors

Total MEC cross section



MEC / (QEL + MEC)





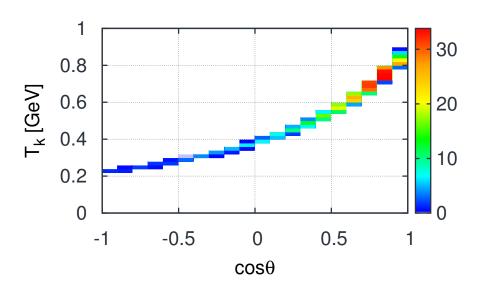
Two-body current interactions

- Both models provide only the inclusive double differential cross section for the final state lepton
- Final nucleons momenta are set isotropically in CMS

Nieves

$\begin{array}{c} 1 \\ 0.8 \\ \hline 0.6 \\ 0.4 \\ 0.2 \\ 0 \end{array}$ -1 -0.5 0 0.5 1 $\cos \theta$

Transverse Enhancement





Coherent pion production

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NuWro

NuWro MC
Dynamics
(Q)EL scattering
RES pion production
Deep Inelastic Scattering
π production
Impulse approximation
Fermi gas
Spectral function
Two-body current

COH pion production

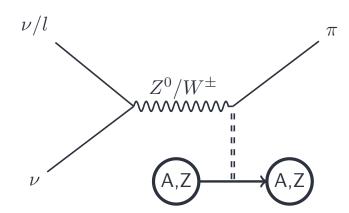
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- Rein-Sehgal model is commonly used for coherent pion production
- Note: it is different model than for RES
- Berger-Sehgal model replaces RS (NuWro, GENIE - coming soon)



Comments

- In COH the residual nucleus is left in the same state (not excited)
- The interaction occurs on a whole nucleus no final state interactions



Neutrino interactions - summary

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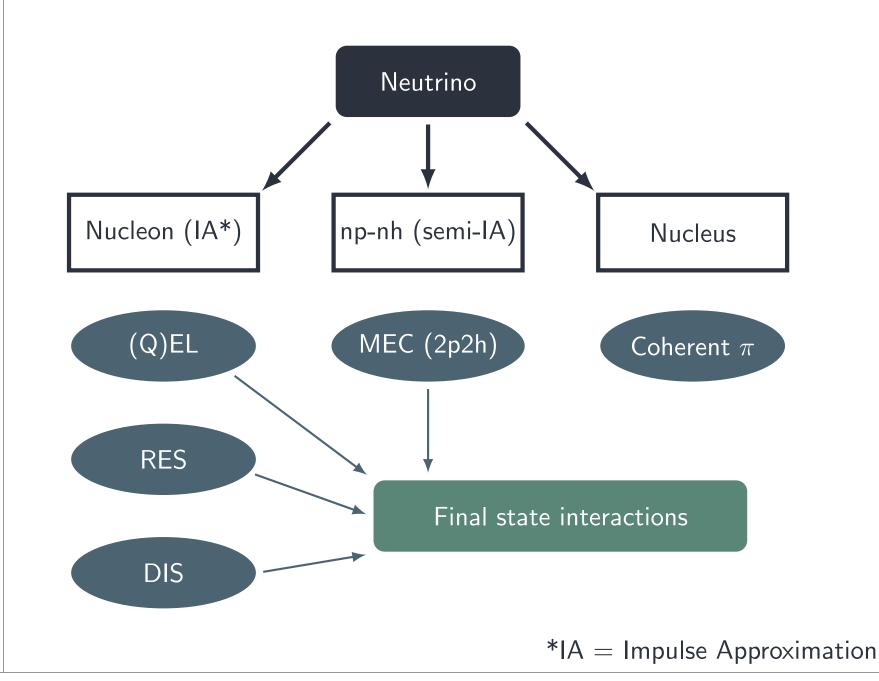
NuWro MC
Dynamics
(Q)EL scattering
RES pion production
Deep Inelastic Scattering
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Final state interactions



Final state interactions

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FSI

Intranuclear cascade LP effect

Formation time

ronnation time

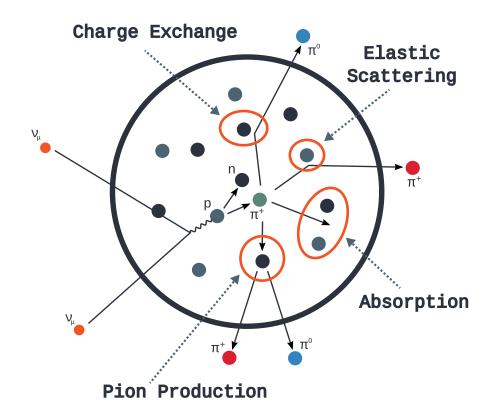
NOMAD

NC π Summary

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FSI describe the propagation of particles created in a primary neutrino interaction through nucleus



All MC generators (but GIBUU) use intranuclear cascade model



Intranuclear cascade

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FSI

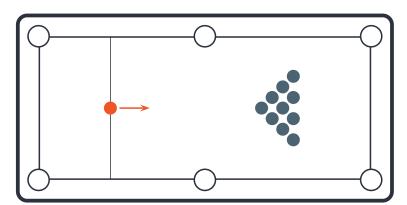
Intranuclear cascade

LP effect Formation time NOMAD NC π Summary

MB NCEL analysis

Backup slides

■ In INC model particles are assumed to be classical and move along the straight line.



The probability of passing a distance λ (small enough to assume constant nuclear density) without any interaction is given by:

$$P(\lambda) = e^{-\lambda/\tilde{\lambda}}$$

$$\tilde{\lambda} = (\sigma
ho)^{-1}$$
 - mean free path

 σ - cross section

ho - nuclear density

Can be easily handled with MC methods.



Landau Pomeranchuk effect

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LP effect

Formation time NOMAD

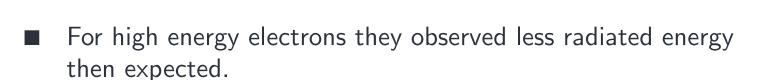
NC π

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The concept of formation time was introduced by Landau and Pomeranchuk in the context of electrons passing through a layer of material.



■ The energy radiated in such process is given by:

$$\frac{\mathrm{d}I}{\mathrm{d}^3k} \sim \left| \int_{-\infty}^{\infty} \vec{j}(\vec{x},t) e^{i(\omega t - \vec{k} \cdot \vec{x}(t))} \mathrm{d}^3x \mathrm{d}t \right|^2$$

 $\vec{x}(t)$ describes the trajectory of the electron.

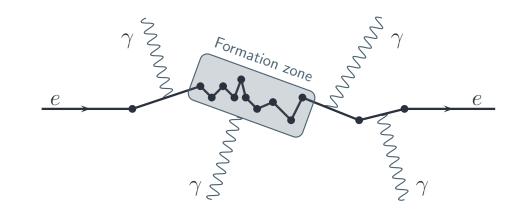
 ω , \vec{k} are energy and momentum of the emitted photon.



Landau Pomeranchuk effect

Assuming the trajectory to be a series of straight lines (the current density $j \sim \delta^3(\vec{x} - \vec{v}t)$) the radiation integral is:

$$\sim \int_{path} e^{i(\vec{k}\vec{v}-\omega)t} dt$$



Formation time is defined as:

$$t_f \equiv \frac{1}{\omega - \vec{k}\vec{v}} = \frac{E}{kp} = \frac{E}{m_e} \frac{1}{\omega_{r.f.}} = \gamma T_{r.f.}$$

k, p - photon, electron four-momenta $\omega_{r.f.}$ - photon frequency in the rest frame of the electron

■ Formation time can be interpreted as the "birth time" of photon.

- If time between collisions $t >> t_f$, there is no interference and total radiated energy is just the average emitted in one collision multiplied by the number of collisions.
- If $t << t_f$, a photon is produced coherently over entire length of formation zone, which reduces the bremsstrahlung.



Formation time in INC

- One may expect a similar effect in hadron-nucleus scattering.
- In terms of INC it means that particles produced in primary vertex travel some distance, before they can interact.





$$t_f = \tau_0 \frac{E \cdot M}{\mu_T^2}$$

where E, M - nucleon energy and mass, $\mu_T^2 = M^2 + p_T^2$ - transverse mass

- SKAT parametrization (similar but with $p_T = 0$)
- NEUT and GENIE use SKAT parametrization
- lacktriangle NuWro uses Ranft parametrization for DIS and a model based on Δ lifetime for RES



Comparison with NOMAD data

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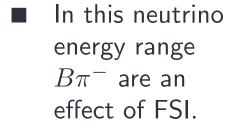
NOMAD

 $\begin{array}{c} {\rm NC} \ \pi \\ {\rm Summary} \end{array}$

MB NCEL analysis

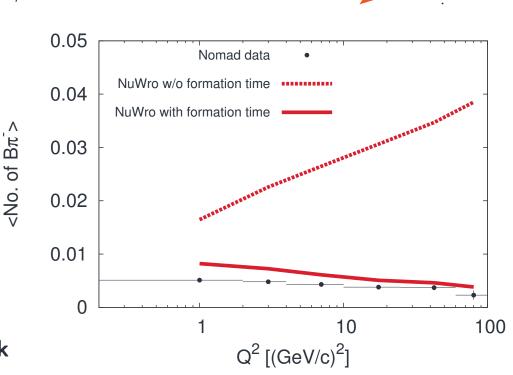
Backup slides

- Nomad data from Nucl. Phys. B609 (2001) 255.
- The average number of backward going negative pions with the momentum from 350 to 800 MeV/c.



The observable is very sensitive to formation time effect.

Golan, Juszczak, Sobczyk PRC86 (2012) 015505



 $\langle E_{\nu} \rangle \sim 24 \text{ GeV}$



Comparison with data for NC π^0 production

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Formation time NOMAD

NC π

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Backup slides

- The cross section for π^0 production through neutral current is measured by K2K [1], MiniBooNE [2] and SciBooNE [3] experiments.
- The signal is defined as: no charged leptons nor charged pions and one neutral pion (or at least one for ScibooNE) in the final state.
- The result depends on primary vertex and FSI, as π can be:
 - produced in primary vertex;
 - produced in FSI;
 - affected by charge exchange;
 - absorbed.

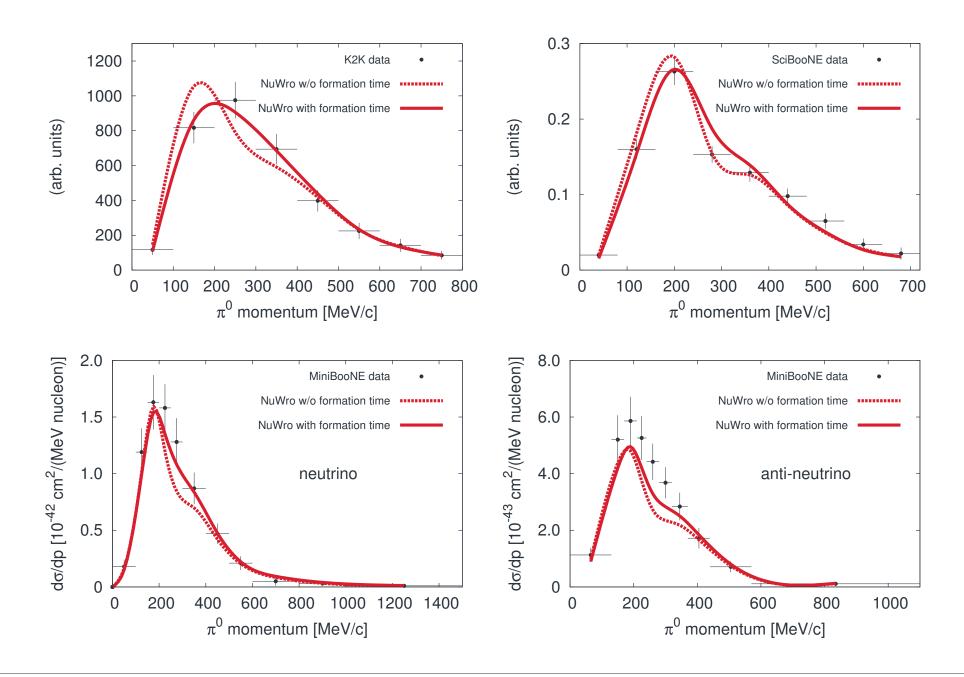
- [1] Phys. Lett. B619 (2005) 255
- [2] Phys. Rev. D81 (2010) 013005

 0μ

[3] Phys. Rev. D81 (2009) 033004



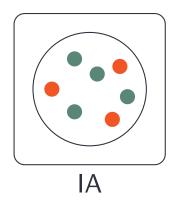
Comparison with data for NC π^0 production

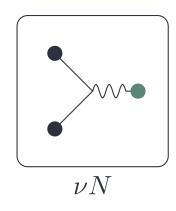




Neutrino-nucleus interactions

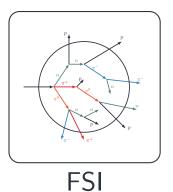
For all channels (but coherent) neutrino interactions are factorized in the following way











- Is the physics really factorized this way?
- This factorization is common for all generators
- However, some pieces are done in different way

The analysis of MB NCEL data



(Quasi-)elastic neutrino-nucleon scattering

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QEL formalism

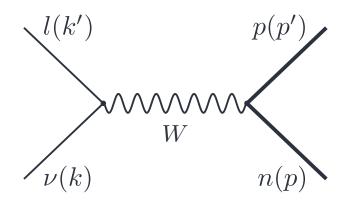
Form factors MiniBooNE data np-nh Results with np-nh The ratio issue

Backup slides

Summary

■ For (quasi-)elastic neutrino scattering off nucleon the cross section is given by:

$$\sigma \sim |j_{\mu}h^{\mu}|^2$$



Quasi-elastic scattering

$$j_{\mu} = \bar{u}(k')\gamma^{\mu} (1 - \gamma_5) u(k) \rightarrow \text{lepton current}$$

$$h_{\mu} = \bar{u}(p')\Gamma^{\mu}u(p) \longrightarrow \mathsf{hadron} \; \mathsf{current}$$

- Leptonic vertex can be calculated from the basis.
- However, due to the complex structure of nucleon, hadronic vertex needs a phenomenological input.
- lacksquare Γ^{μ} can be parametrized by the functions of Q^2 , called form factors.



Form factors

■ Hadronic vertex can be expressed in terms of form factors:

$$\Gamma^{\mu}_{NC,p(n)} = \gamma^{\mu} F_1^{NC,p(n)}(Q^2) + \frac{i\sigma^{\mu\nu}q_{\nu}}{2M} F_2^{NC,p(n)}(Q^2) - \gamma^{\mu}\gamma_5 G_A^{NC,p(n)}(Q^2)$$

Vector form factors are expressed by electromagnetic form factors (Conserved Vector Current - CVC):

$$F_{1,2}^{NC,p(n)}(Q^2) = \pm \frac{1}{2} \left(F_{1,2}^p(Q^2) - F_{1,2}^n(Q^2) \right) - 2\sin^2\theta_W F_{1,2}^{p(n)}(Q^2) - \frac{1}{2} F_{1,2}^s(Q^2)$$

Axial form factor is assumed to have a dipole form:

 $G_A^{NC,p(n)}(Q^2) = \pm \frac{1}{2}G_A(Q^2) - \frac{1}{2}G_A^s(Q^2) = (\pm g_A - g_A^s) \left(1 + Q^2/M_A^2\right)^{-2}$ strangeness

axial mass

different sign for proton/neutron g_A^s sensitive to $\sigma(\nu p)/\sigma(\nu n)$



MiniBoonE data for NCEL

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■ The cross section for NCEL is measured by MiniBooNE - PRD82 (2010) 092005.



- The signal is defined as: no charged leptons nor any kind of pions in the final state.
- The detector measures the Cherenkov and scintillation light.
- lacktriangle Protons with the kinetic energy of order tens MeV are detectable.
- Neutrons are visible as an effect of interactions with protons (inside or outside the nucleus).
- The M_A and g_A^s is extracted from the measurement.



MiniBoonE data for NCEL

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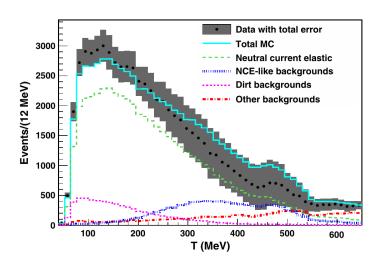
MB NCEL analysis

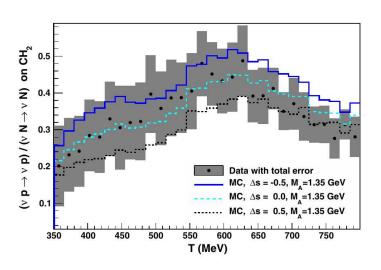
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 ${\cal T}$ stands for the total reconstructed energy of all nucleons in the final state.

- $\blacksquare \quad \text{Here assume } g_A^s = 0.$
- Best fit for:

$$M_A = 1.39 \pm 0.11 \; \text{GeV}$$

Inconsistency with older measurements $M_A \sim 1 \text{ GeV}.$

- Here assume $M_A = 1.35$ GeV.
- Best fit for:

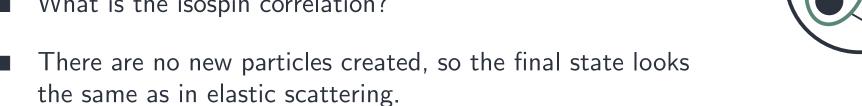
$$g_A^s = 0.08 \pm 0.26$$

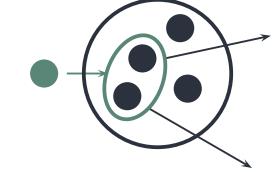
" $\nu p \rightarrow \nu p$ " stands for events with proton above Cherenkov treshold.



Two body current (or np - nh) contribution

- The np-nh interactions occur on at least two nucleons.
- What is the isospin correlation?

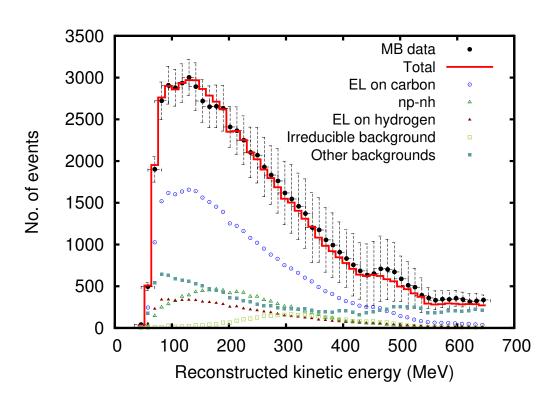




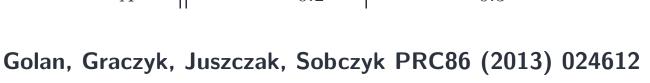
- np-nh is not taken into account in MiniBooNE analysis, which may cause the discrepancy with previous measurements of axial mass.
- There are two theoretical models of np-nh (IFIC, Lyon), which take care about proper energy transfer distribution. Unfortunately, both are available only for CC channel.
- The phenomenological Transverse Enhancement (TE) model is used in the calculation. The np-nh contribution is introduced by the modification of the vector magnetic form factors. Lepton kinematics is the same as for elastic scattering.
- The goal is to repeat the analysis with the inclusion if the np-nh contribution.

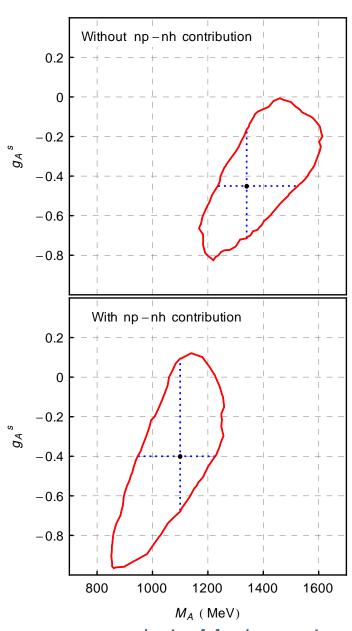


Simultaneous extraction of M_A and g_A^s



	\mid w/o $np-nh$	with $np - nh$
M_A [GeV]	$1.34^{+0.18}_{-0.10}$	$1.10^{+0.13}_{-0.15}$
g_A^s	$-0.5^{+0.2}_{-0.2}$	$-0.4^{+0.5}_{-0.3}$





made in Mathematica

Tomasz Golan NuWro @ HEP UW 53 / 61



The ratio issue

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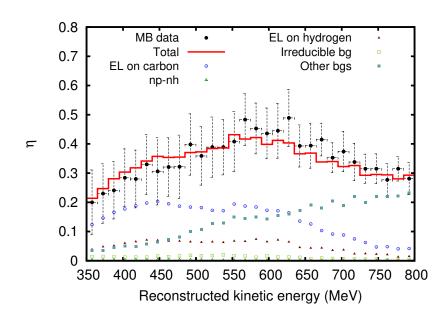
Results with np - nh

The ratio issue

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- The plot is made using best fit values from the last slide.
- It turns out that the ratio is very sensitive on the energy transfer distribution (No. of protons above threshold).



- The Transverse Enhancement does not take care properly of lepton kinematics, which affects the energy distribution of final state nucleons. It is not a proper model to analyze this data.
- This data, however, has a potential to discriminate between IFIC and Lyon models...
- ... when they appear for neutral current channel.



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Summary

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- MC generators are irreplaceable tools in high-energy physics
- People use them before experiment exists (feasibility studies, requirements ...)
- And during data analysis (systematics uncertainties, backgrounds ...)
- NuWro contains all major neutrino-nucleon interaction channels, reliable FSI model and realistic nucleus description it is a ready tool to use in the investigation of neutrino data

Thank you for the attention!

Backup slides



Formation time in MC generators

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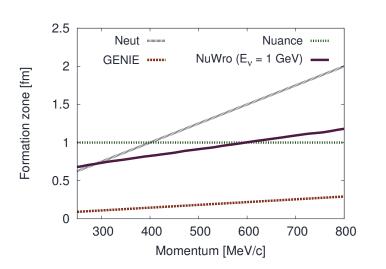
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FT in MC generators

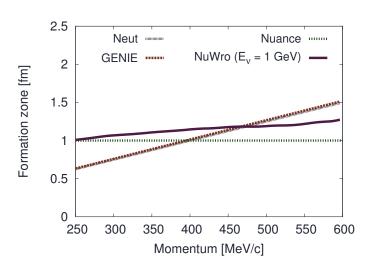
Form factors TE model

Cascade algorithm

Formation time for nucleons



Formation time for pions



- NUANCE uses constant formation length x = 1 fm.
- NEUT uses SKAT parametrization ($\mu^2 = 0.08 \pm 0.04 \text{ GeV}^2$):

$$x = \frac{|\vec{p}|}{\mu^2}$$

GENIE uses Rantf parametrization (assuming transverse momentum $p_T = 0$):

$$x = \tau_0 \frac{E}{M} = \tau_0 \gamma$$



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TE model

Cascade algorithm

$$\Gamma^{\mu}_{CC}(q) = \gamma^{\mu} F_1^V(Q^2) + \frac{i\sigma^{\mu\nu}q_{\nu}}{2M} F_2^V(Q^2) - \gamma^{\mu}\gamma_5 G_A(Q^2) - q^{\mu}\gamma_5 \frac{F_P(Q^2)}{2M}$$

■ Vector form factors are expressed by electromagnetic form factors (Conserved Vector Current - CVC):

$$F_{1,2}^V(Q^2) = F_{1,2}^p(Q^2) - F_{1,2}^n(Q^2)$$

Axial form factor is assumed to have a dipole form:

$$G_A(Q^2) = \frac{g_A}{(1 + Q^2/M_A^2)^2}$$

■ Pseudoscalar form factor is related to the axial one (Partially Conserved Axial Current - PCAC):

$$F_P(Q^2) = \frac{4M^2}{m_\pi^2 + Q^2} G_A(Q^2)$$



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$$\Gamma^{\mu}_{NC,p(n)} = \gamma^{\mu} F_1^{NC,p(n)}(Q^2) + \frac{i\sigma^{\mu\nu}q_{\nu}}{2M} F_2^{NC,p(n)}(Q^2) - \gamma^{\mu}\gamma_5 G_A^{NC,p(n)}(Q^2)$$

■ Vector form factors are expressed by electromagnetic form factors (Conserved Vector Current - CVC):

$$F_{1,2}^{NC,p(n)}(Q^2) = \pm \frac{1}{2} \left(F_{1,2}^p(Q^2) - F_{1,2}^n(Q^2) \right) - 2\sin^2\theta_W F_{1,2}^{p(n)}(Q^2) - \frac{1}{2} F_{1,2}^s(Q^2)$$

Axial form factor is assumed to have a dipole form:

$$G_A^{NC,p(n)}(Q^2) = \pm \frac{1}{2}G_A(Q^2) - \frac{1}{2}G_A^s(Q^2)$$

■ The axial strange form factor is assumed to have a dipole form:

$$G_A^s(Q^2) = \frac{g_A^s}{(1 + Q^2/M_A^2)^2}$$



Transverse Enhancement model

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Cascade algorithm

■ In the Transverse Enhancement model the two body current contribution is introduced by the modification of the vector magnetic form factors:

$$G_M^{p,n} \to \tilde{G}_M^{p,n} = \sqrt{1 + AQ^2 \exp\left(-\frac{Q^2}{B}\right)} G_M^{p,n}(Q^2)$$

- \blacksquare A, B are established from the electron data.
- The cross section for np nh can be obtained by taking the difference:

$$\frac{\mathrm{d}^2 \sigma^{np-nh}}{\mathrm{d}q \mathrm{d}\omega} \equiv \frac{\mathrm{d}^2 \sigma^{QEL}}{\mathrm{d}q \mathrm{d}\omega} (\tilde{G}_M^{p,n}) - \frac{\mathrm{d}^2 \sigma^{QEL}}{\mathrm{d}q \mathrm{d}\omega} (G_M^{p,n})$$

■ The disadvantage of the model is lepton kinematics ("copied" from the QEL scattering).





$$\tilde{\lambda}(r) = [\sigma \rho(r)]^{-1}$$

