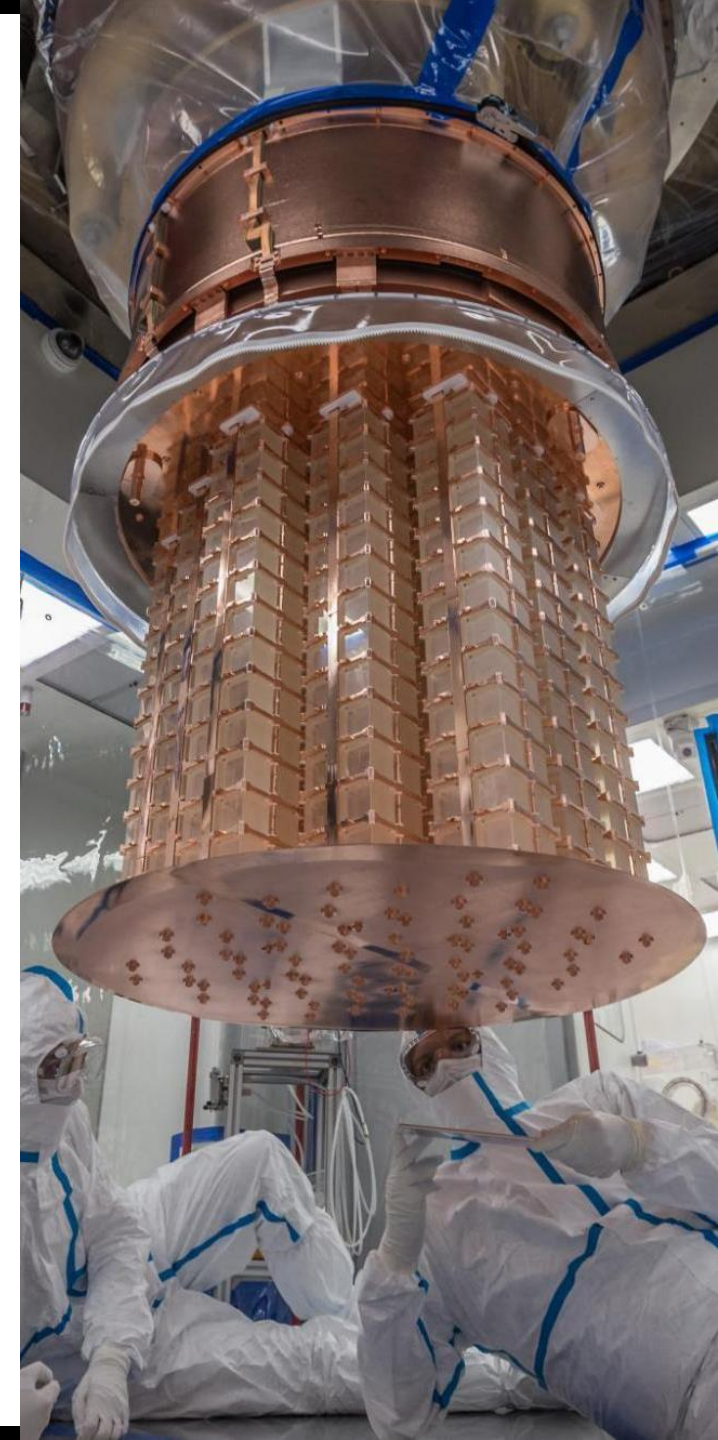


# SOMETHING IS (NOT) UP AT XENON1T

FEDERICO BATTISTI  
GRADUATE SYMPOSIUM 15/03/2021





# DARK MATTER

- Astrophysical and cosmological evidence suggests that most of the matter content in the Universe is made up of dark matter
- Several dark matter particle candidates have been proposed but all have thus far eluded direct detection.

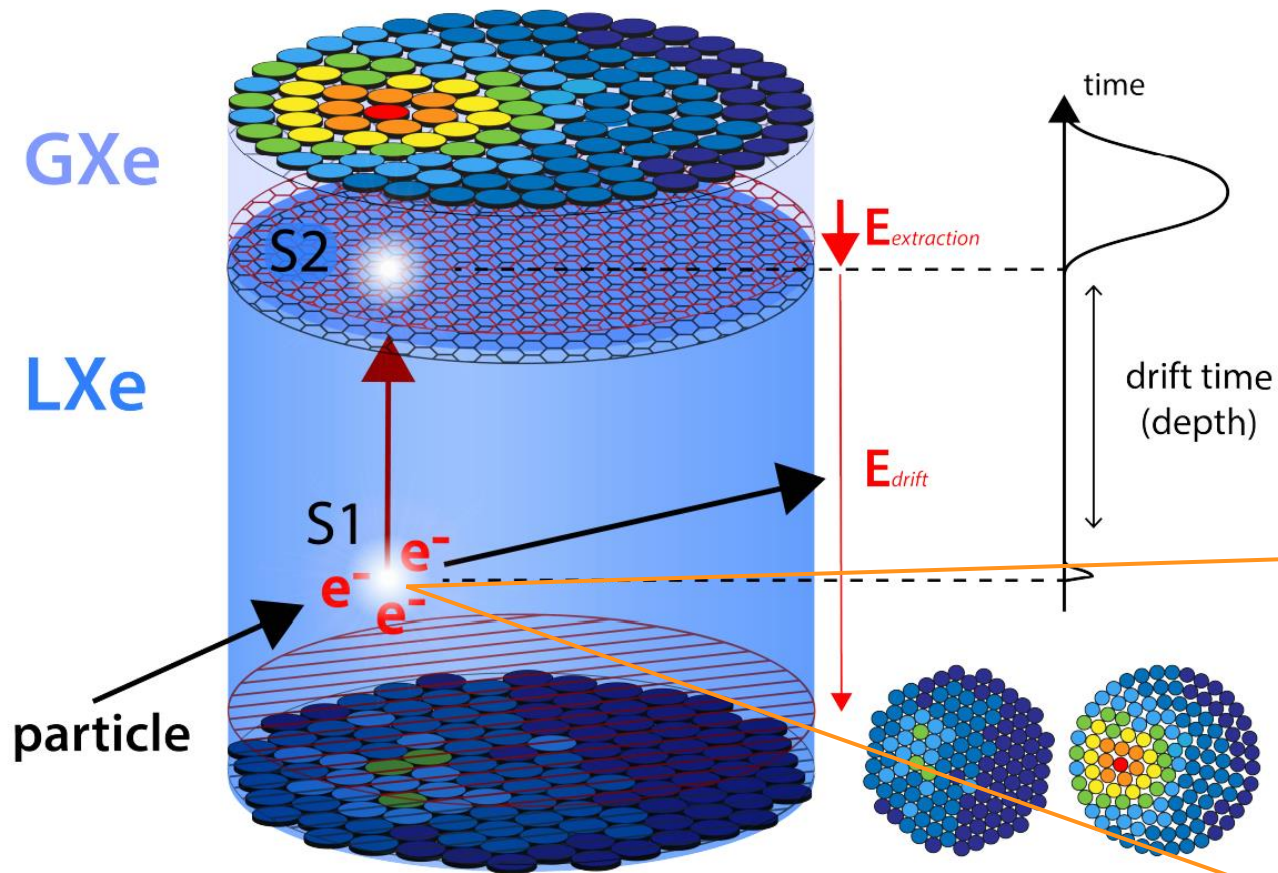


# WHAT IS XENON1T

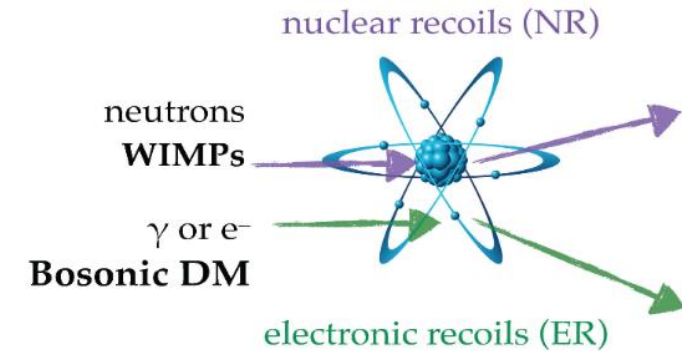


- XENON1T experiment operated underground at the INFN Laboratori Nazionali del Gran Sasso (LNGS) from 2016 to 2018
- Primarily designed to detect weakly interacting massive particle (WIMP) dark matter.
- Also sensitive to alternative dark matter candidates and other BSM

# WHAT XENON1T WAS BUILT FOR

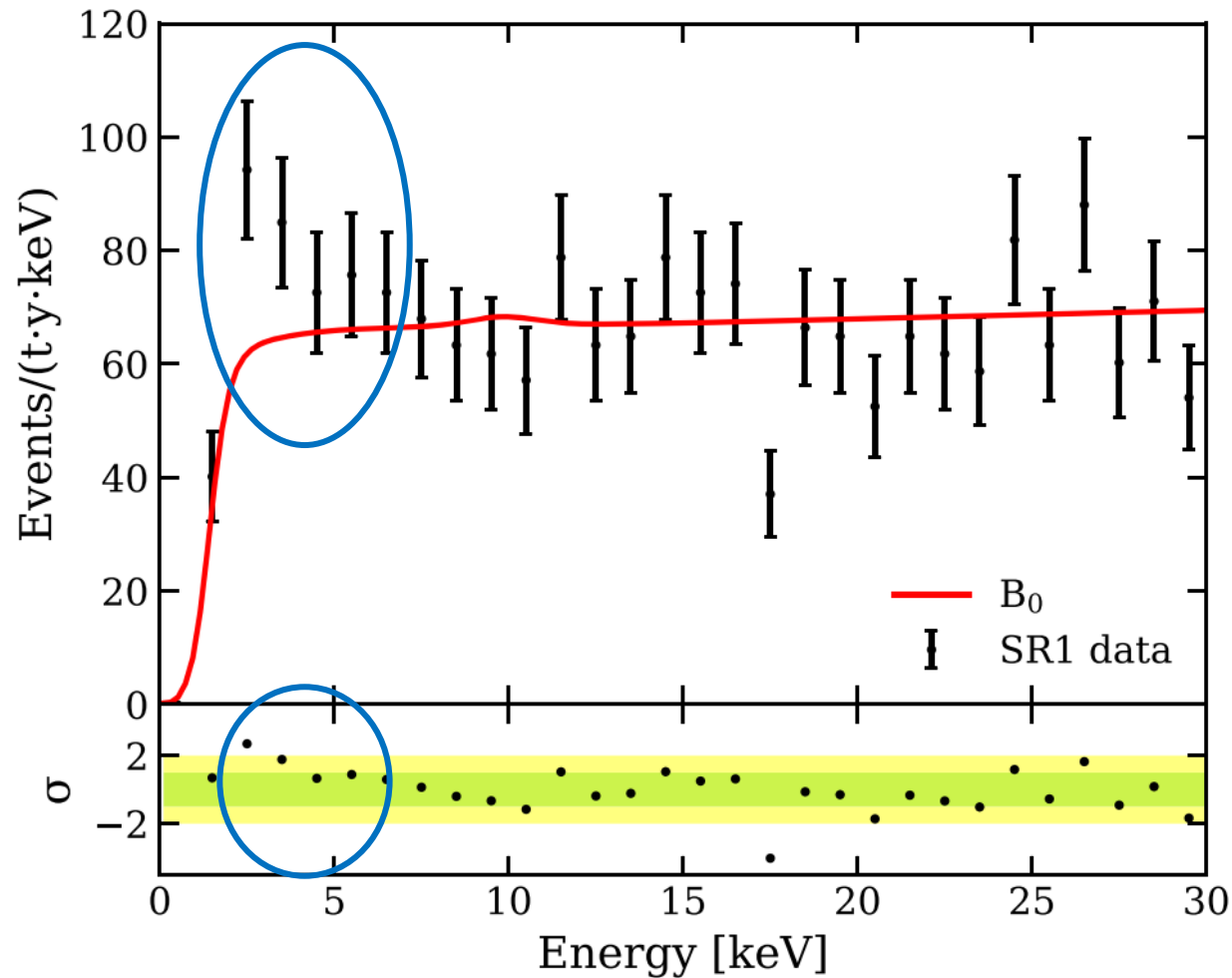


- Dual-phase TPC: liquid xenon 2.0-tonne active target with a gaseous phase on top + 2 arrays of PMTs
- Signal: prompt scintillation (S1) and delayed electroluminescence (S2)
- S2/S1 ratio used to distinguish nuclear recoils (NRs) from electronic recoils (ERs)





# AN ELECTRON RECOIL ECCESS



PHYSICAL REVIEW D **102**, 072004 (2020)

Featured in Physics

## Excess electronic recoil events in XENON1T

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(XENON Collaboration)<sup>‡‡</sup>

X. Mougeot<sup>29</sup>

- 1–7 keV: 285 ER events observed in the data VS expected  $232 \pm 15$  events from background only (**3.3 $\sigma$  Poisson fluctuation**)



# WHAT'S HAPPENING?

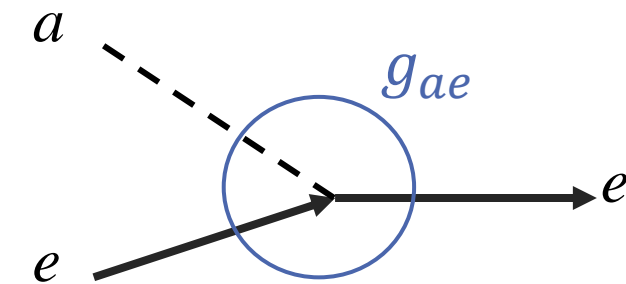
Four hypothesis consistent with signal type:

1. Solar Axions
2. Solar Neutrinos with anomalous magnetic moment
3. Dark matter bosons (Heavy axions or Dark photons)
4. Tritium background



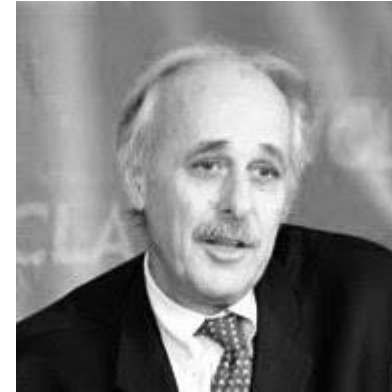
# SOLAR AXIONS

- 1977 Peccei and Quinn: postulate new field as solution to the strong CP problem
- Wilczek-Weinberg realize Peccei-Quinn field gives rise to new boson they call axion
- Axions produced in the Sun could produce electron recoil signal with energies in the keV range via axio-electric effect



AXIO-ELECTRIC EFFECT

Peccei



Quinn



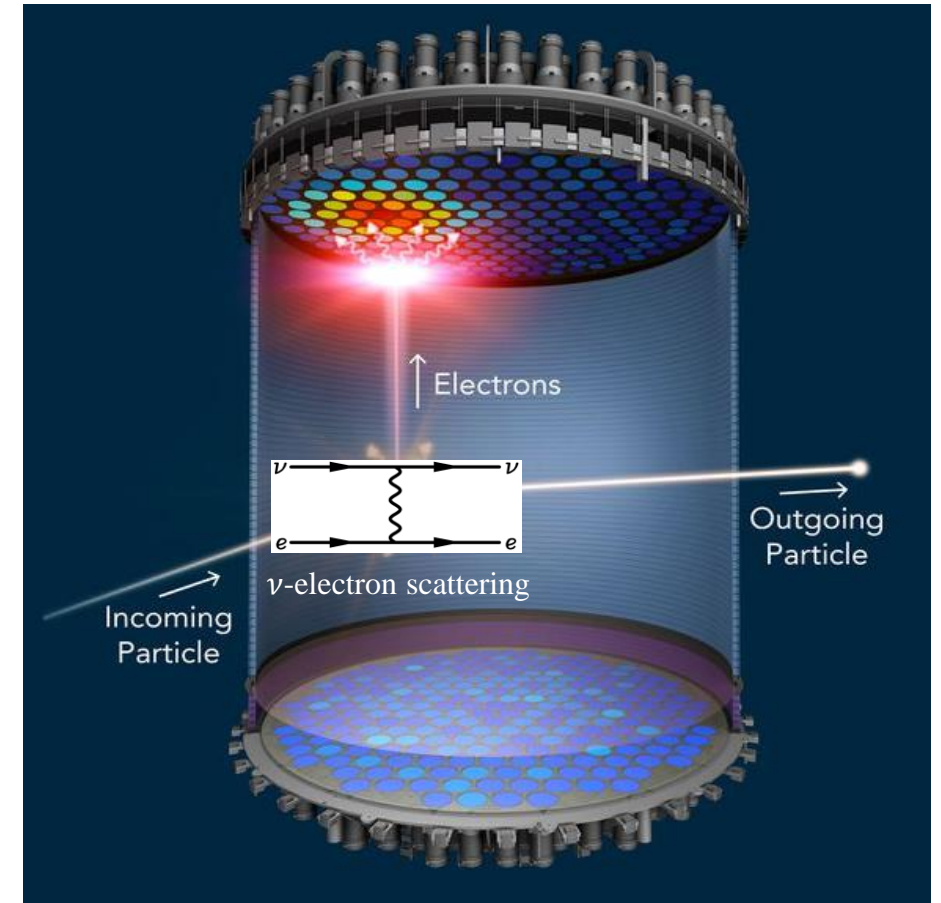
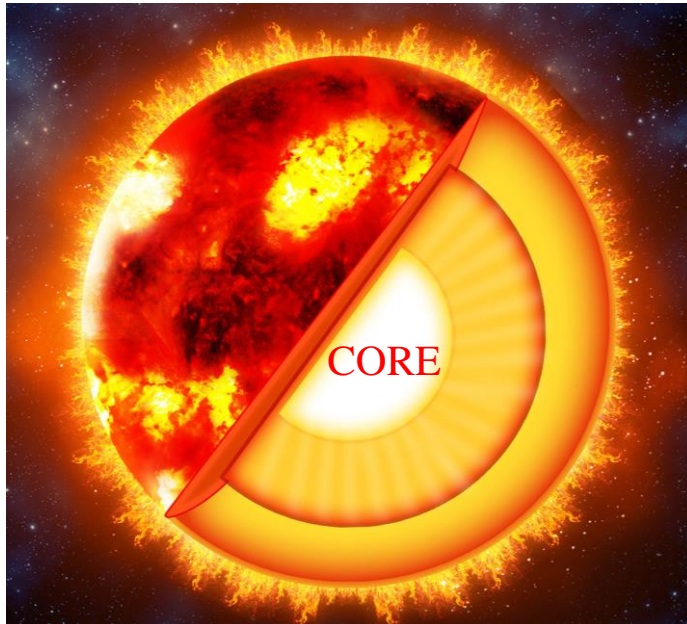
Wilczek

Weinberg



# ENHANCED MAGNETIC MOMENT SOLAR NEUTRINOS

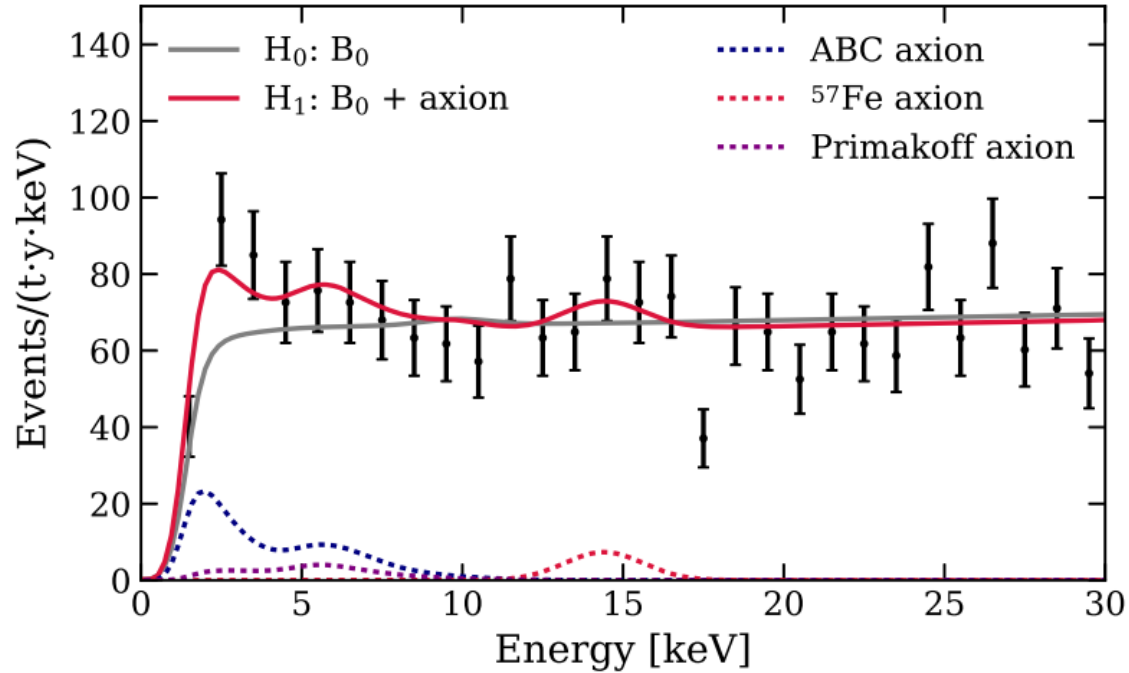
- Enhanced magnetic moment would increase the neutrino scattering cross sections at low energies
- $\mu_\nu \sim 7 \times 10^{-11} \times \mu_B$
- SM  $\mu_\nu \sim 10^{-20} \times \mu_B$  for massive neutrinos ( $10^{10}$  times smaller)





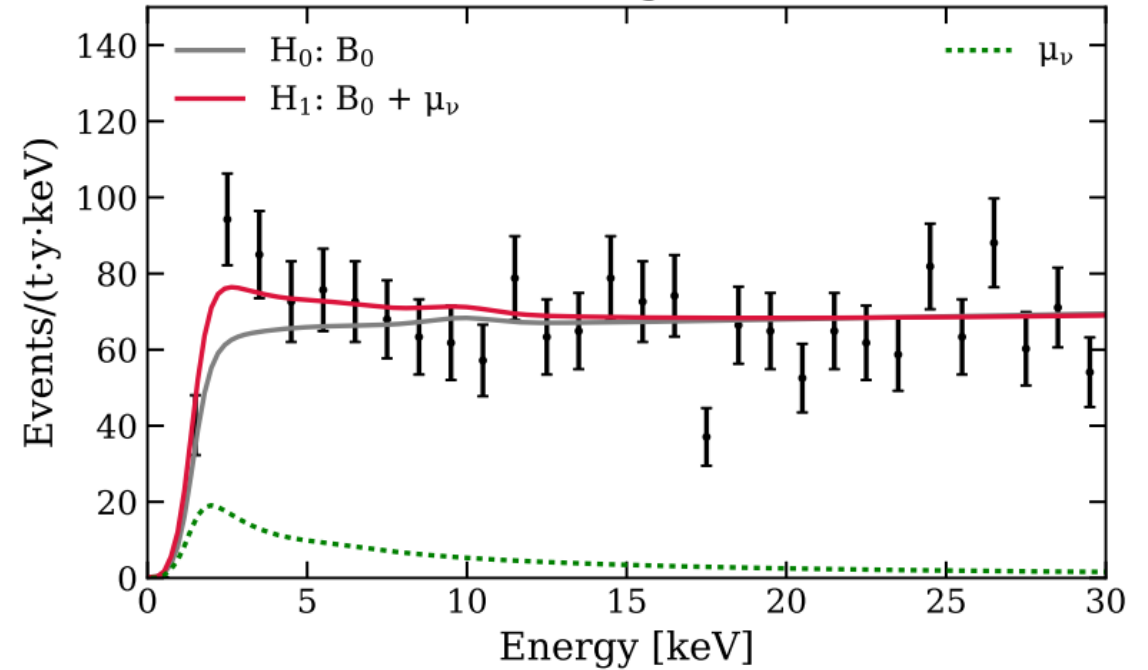
# SOLAR RESULTS

(b) Solar axion



- Fits over background:
  - **LEFT** Solar Axions: favored  $3.4\sigma$
  - **RIGHT** Neutrinos with enhanced magnetic moment: favored  $3.1\sigma$

(c) Neutrino magnetic moment



- **NOTE:** Result is in strong tension with indirect constraints from analyses of white dwarfs and globular clusters

A visualization of the cosmic web, showing a complex network of blue filaments and red nodes against a black background. The filaments represent the large-scale structure of the universe, while the red nodes represent galaxy clusters.

# BOSONIC DARK MATTER?

## 1. Axion-Like Particles (ALPs)

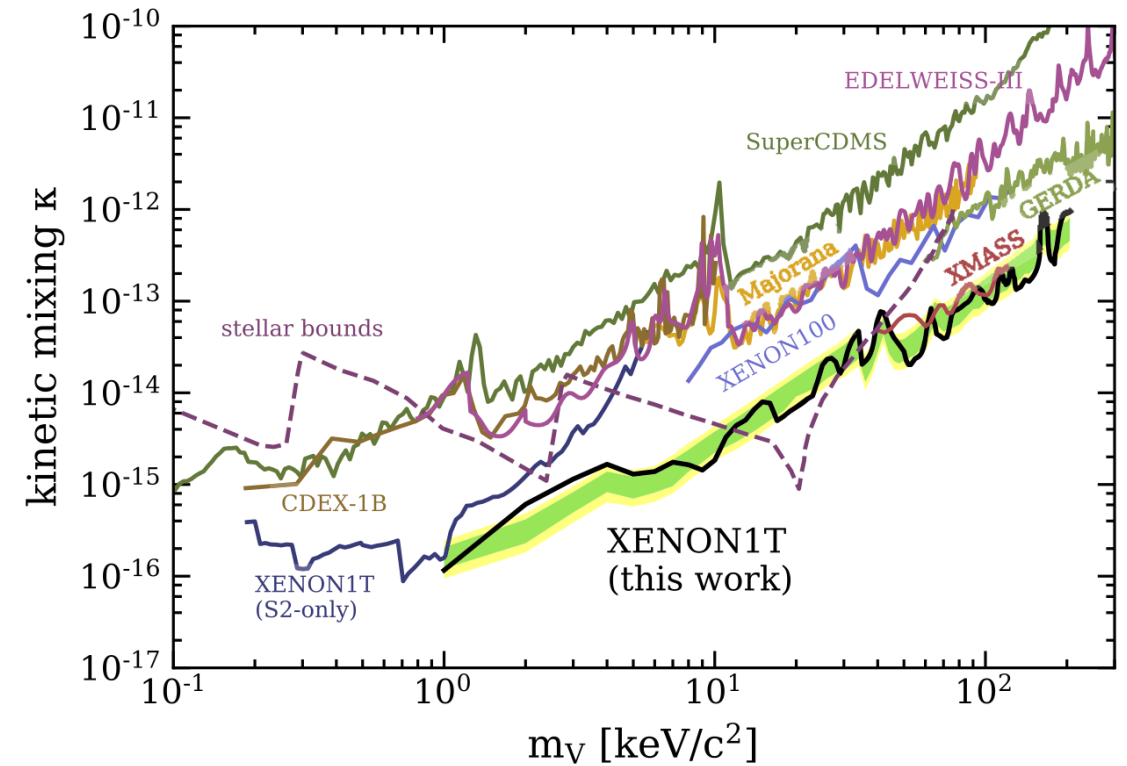
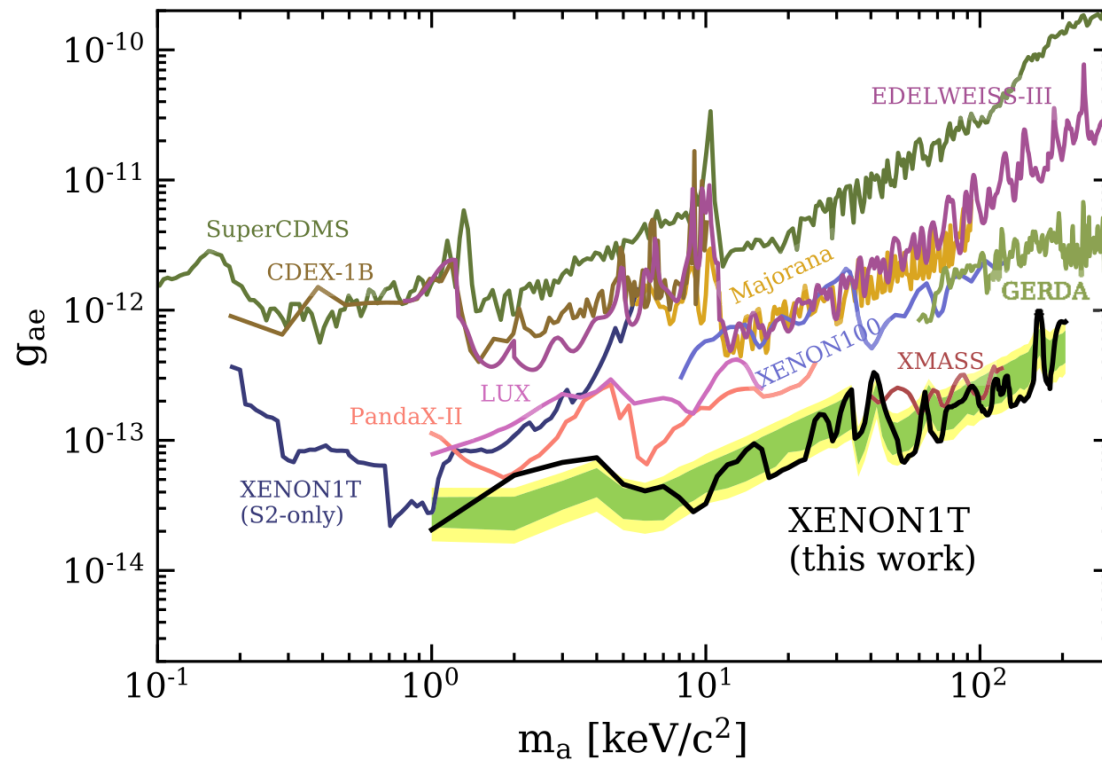
- Pseudoscalar bosons like QCD axions
- Higher masses, but do not solve the strong CP problem

## 2. Dark photons:

- Vector bosonic dark matter
  - Can couple weakly with SM photons through kinetic mixing
- 
- Expected signal: mono-energetic peak at the rest mass of the particle

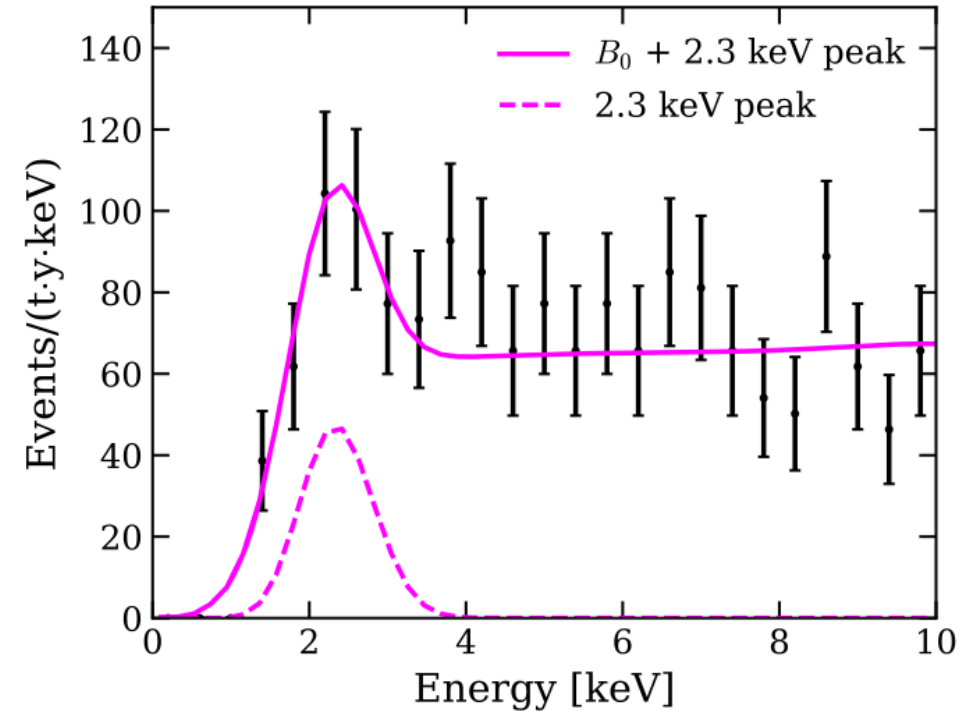
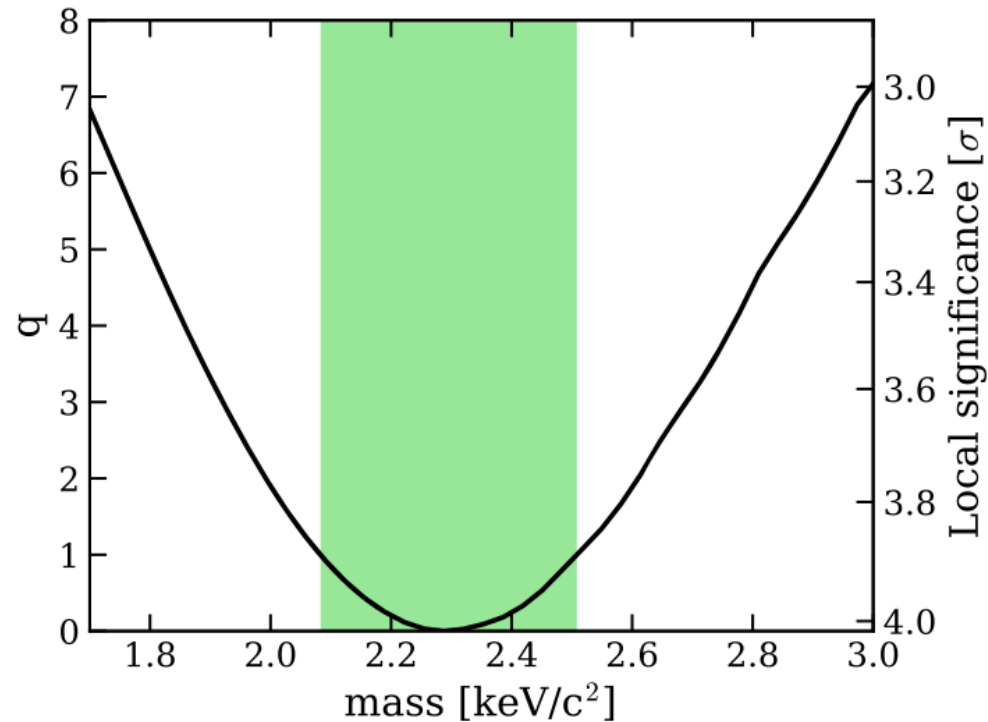


# DARK MATTER RESULTS



- Masses between 1 and 210  $\text{keV}/c^2$  : No global significance over  $3\sigma$  over the background model
- Set upper limit on the couplings  $g_{ae}$  (axion-electron coupling strength) and  $\kappa$  (kinetic mixing strength between photons and dark photons)
- Most stringent constraints for most mass values considered

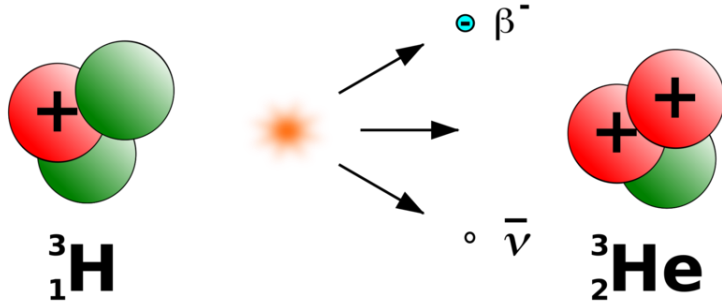
# DARK MATTER RESULTS



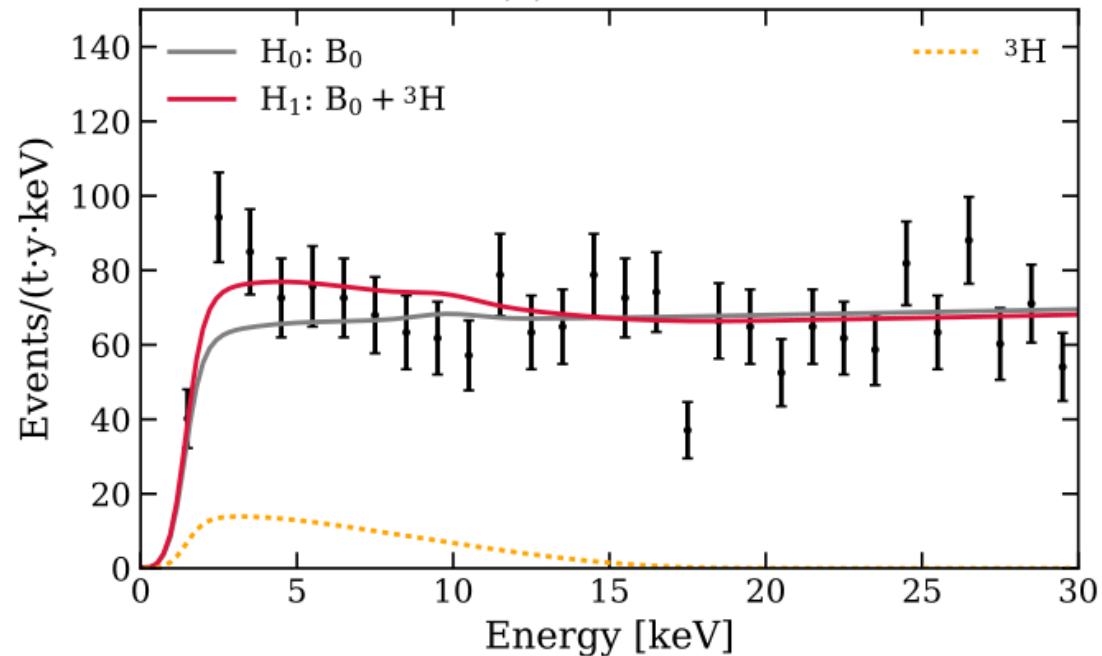
- Look for general peak not considering the type of particle
- Masses allowed to vary freely between 1.3 and 3.3  $\text{keV}/c^2$
- Favored mass value of  $2.3 \pm 0.2 \text{ keV}/c^2$  (68% C.L.) with a  $3.0\sigma$  global ( $4.0\sigma$  local) significance over background



# DON'T GET TOO EXCITED: IT MIGHT BE BORING TRITIUM



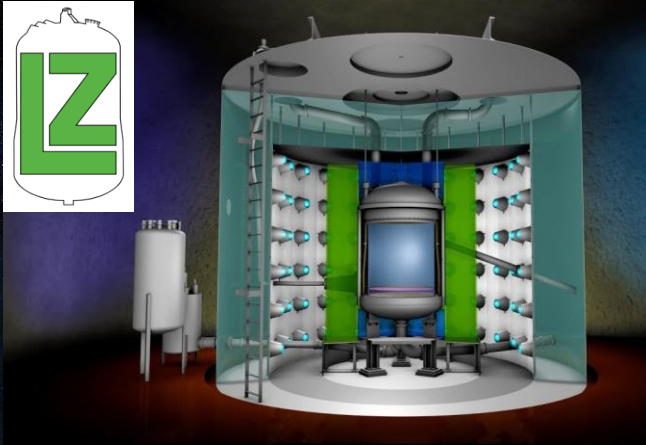
(a) Tritium



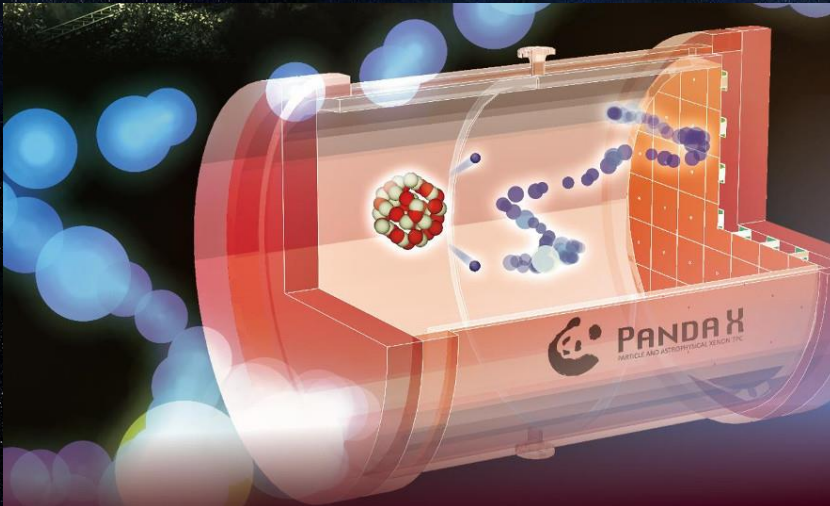
- Tritium expected to be present in the detector
- $\beta$  decay of tritium has similar spectrum to that observed
- Very low concentration required to account for the excess
- $O(10^{-25} \text{ mol/mol})$  of tritium  $3.2 \sigma$  favored over background
- Currently no way to discard Tritium hypothesis with current sensitivity levels



# CONCLUSIONS: TOO EARLY TO TELL



- $3.3\sigma$  anomaly is interesting but no hypothesis is currently sufficiently favored to claim a discovery
- Next-generation detectors will need to investigate







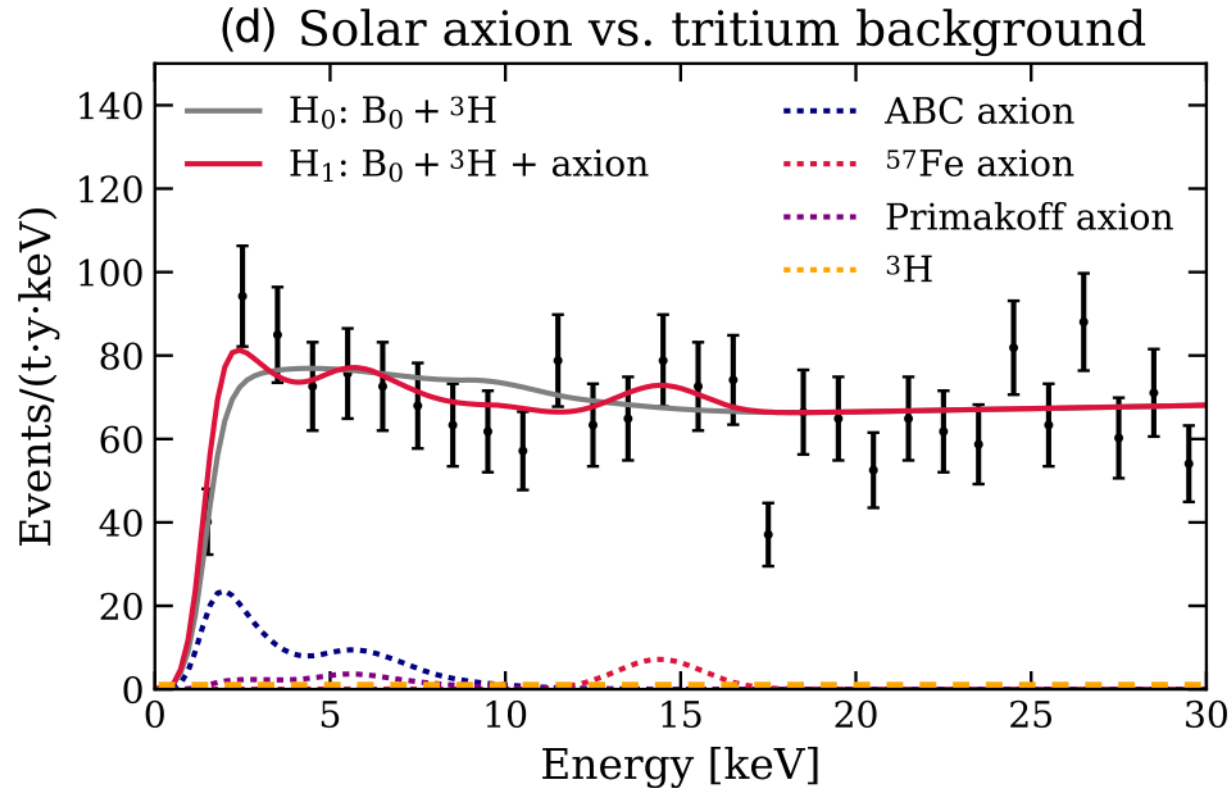
**THANK YOU!**



**BACK-UP**



# NEW RESULTS (CONSIDERING TRITIUM)



- Unconstrained tritium component was added to the background model: null hypothesis is  ${}^3\text{H} + B_0$
- Solar axion signal: still preferred in this test, but significance is reduced to  $2.0\sigma$
- Neutrino with anomalous magnetic moment: significance reduced to  $0.9\sigma$