

## **Visual Computing - Lecture notes week 6**

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# 8. Optical Flow

## 8.1 Brightness Constancy

We define **optical flow** as the apparent motion of brightness patterns. Ideally, the optical flow is the *projection* of the three-dimensional velocity vectors of the image.

$I(x, y, t)$  is the brightness at  $(x, y)$  at time  $t$ . This way we can define:

*Brightness constancy assumption:*

$$I\left(x + \frac{dx}{dt}\delta t, y + \frac{dy}{dt}\delta t, t + \delta t\right) = I(x, y, t)$$

*Optical flow constraint equation:*

$$\frac{dI}{dt} = \frac{\partial I}{\partial x} \frac{dx}{dt} + \frac{\partial I}{\partial y} \frac{dy}{dt} + \frac{\partial I}{\partial t} = 0$$

## 8.2 The Aperture Problem

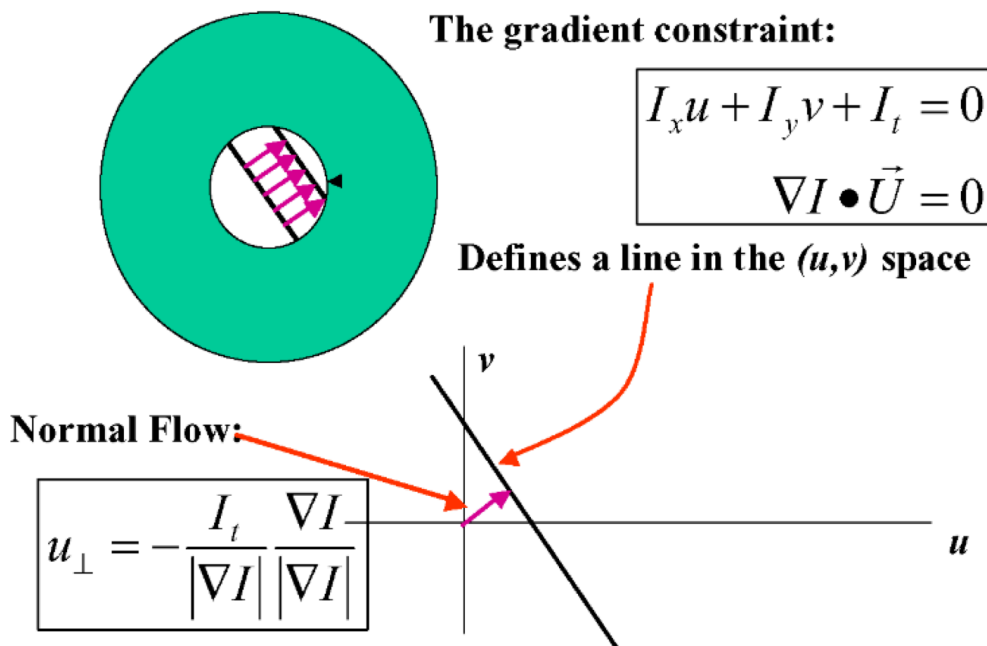
If we do the following substitution:

$$u = \frac{dx}{dt}, v = \frac{dy}{dt}, I_x = \frac{\partial I}{\partial x}, I_y = \frac{\partial I}{\partial y}, I_t = \frac{\partial I}{\partial t}$$

we arrive at the following equation:

$$I_x u + I_y v + I_t = 0$$

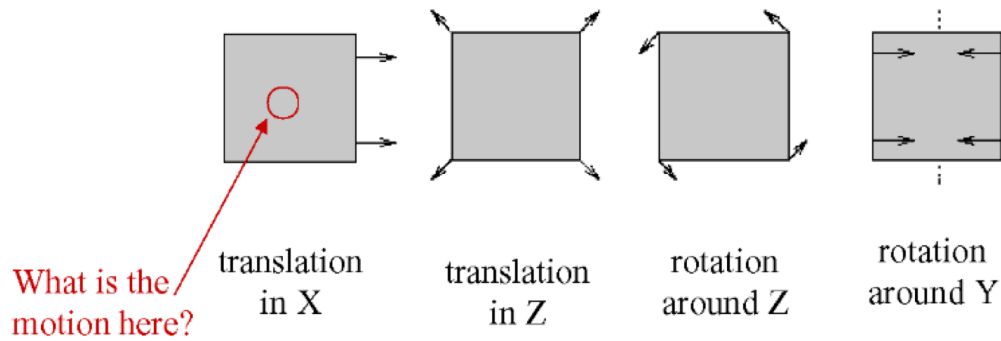
However, this is one equation in two unknowns, which is known as the **Aperture Problem**.



Optical flow is not always well-defined! We can compare the different kinds of flows:

- *Motion Field*: Projection of 3D motion field

- *Normal Flow*: Observed tangent motion
- *Optical Flow*: Apparent motion of the brightness pattern, hopefully equal to motion field



Scene Flow: →  
 Normal Flow: undef  
 Optic Flow: ?, probably 0

## 8.3 Regularization

We might use the **Horn & Schuck algorithm**, described as follows:

- We have an additional smoothness constraint

$$e_s = \int \int ((u_x^2 + u_y^2) + (v_x^2 + v_y^2)) dx dy$$

- besides our Optical Flow constraint equation term:

$$e_c = \int \int (I_x u + I_y v + I_t)^2 dx dy$$

The goal is to *minimize*  $e_s + \lambda e_c$  !s

## 8.4 Lucas-Kanade

Assume a single velocity for all pixels within an image patch

$$E(u, v) = \sum_{x, y \in \Omega} (I_x(x, y)u + I_y(x, y)v + I_t)^2$$

$$\frac{dE(u, v)}{du} = \sum 2I_x(I_x u + I_y v + I_t) = 0$$

$$\frac{dE(u, v)}{dv} = \sum 2I_y(I_x u + I_y v + I_t) = 0$$

Solve with:

$$\begin{bmatrix} \sum I_x^2 & \sum I_x I_y \\ \sum I_x I_y & \sum I_y^2 \end{bmatrix} \begin{pmatrix} u \\ v \end{pmatrix} = - \begin{pmatrix} \sum I_x I_t \\ \sum I_y I_t \end{pmatrix}$$

On the LHS: sum of the 2x2 outer product tensor of the gradient vector

$$(\sum \nabla I \nabla I^T) \vec{U} = - \sum \nabla I I_t$$

With respect to singularities and the aperture problem, we proceed as follows. Let:

$$M = \sum (\nabla I)(\nabla I)^T \quad \text{and} \quad b = \begin{bmatrix} - \sum I_x I_t \\ - \sum I_y I_t \end{bmatrix}$$

Then, the algorithm is as simple as solving at each pixel for  $U$  in  $MU = b$ .  $M$  is *singular* if all gradient vectors point in the same direction.

## 8.5 Coarse To Fine

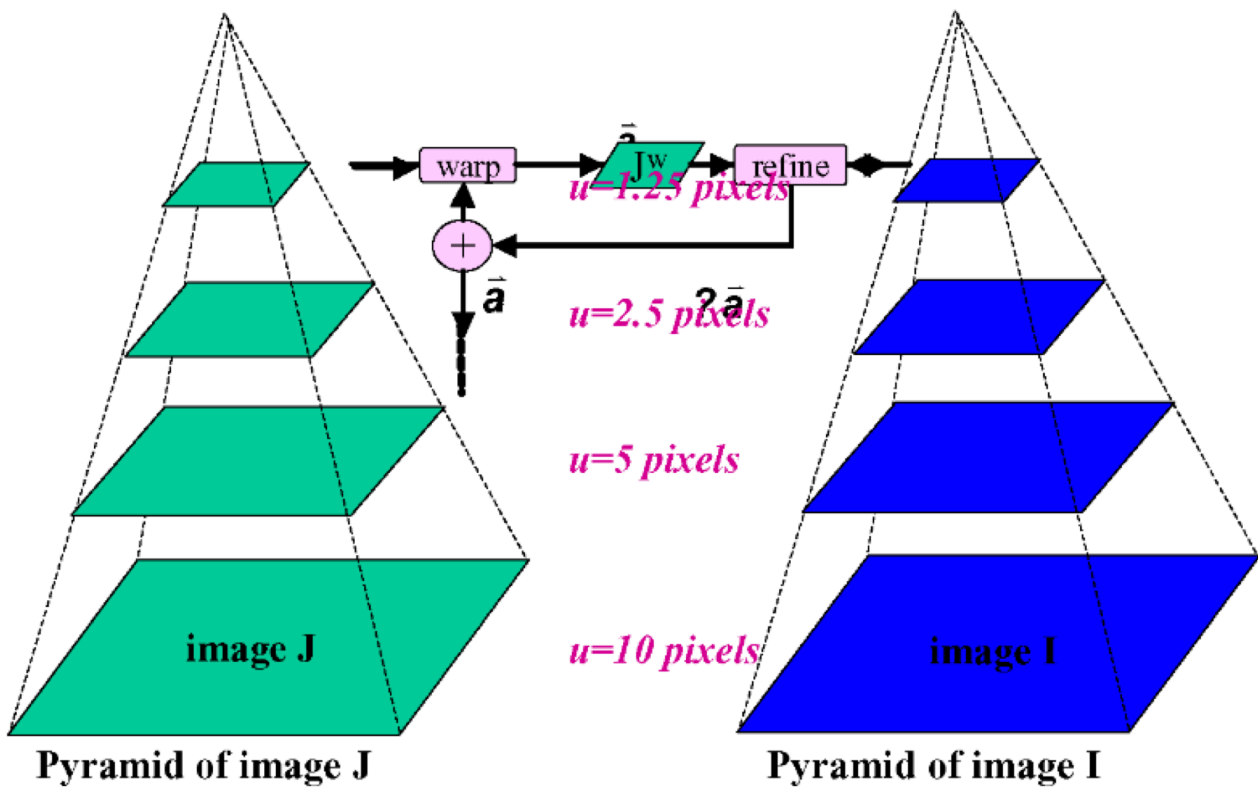
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The local gradient method has some limitation:

- Fails when the intensity structure within the window is poor
- Fails when the displacement is large (typical operating range is motion of 1 pixel per iteration!)

We can combat this with a Pyramid or "*Coarse-to-fine*" estimation:

$$I_x \cdot u + I_y \cdot v + I_t \approx 0 \implies \text{small } u \text{ and } v \dots$$



## 8.6 Parametric Motion Models

Global motion models offer:

- more constrained solutions than smoothness models (Horn-Schunck)
- integration over a larger area than a translation-only model can accommodate (Lucas-Kanade)