



Suppose that the position of Sun is fixed and earth is orbiting around it.

$$F = -G \frac{M_e M_s}{r^2} \hat{r}$$

$$F_x = -F \cos \theta = m_e a_x = m_e \ddot{x}$$

$$F_y = -F \sin \theta = m_e a_y = m_e \ddot{y}$$

$$\Rightarrow \begin{cases} \frac{d^2 x}{dt^2} = \frac{-G M_s}{r^2} \cos \theta \\ \frac{d^2 y}{dt^2} = \frac{-G M_s}{r^2} \sin \theta \end{cases}$$

Now we want to turn dimensions into those ones that

I) Large/Small number doesn't appear

II) Length - time - mass are comparable with  
Characteristic length, time, and mass of

the problem.

a good starting point is to set  $G M_s = 1$

$$G M_s = \underbrace{[G M_s]}_{\text{value}} m^3 \text{kg}^{-1} \text{s}^{-2} = 1 m^*{}^3 \text{kg}^{*-1} \text{s}^{*-2}$$

to satisfy that we should have:

$$\left( \frac{m^*}{m} \right)^3 \left( \frac{\text{kg}^*}{\text{kg}} \right)^{-1} \left( \frac{\text{s}^*}{\text{s}} \right)^{-2} = [G M_s]$$

we all know that a reasonable:

time scaler  $\rightarrow 1 \text{ year}$

Length scale  $\rightarrow 1 \text{ AU} \xrightarrow{\text{Sun-Earth distance}}$

$$S^* = 1 \text{ year} \rightarrow \left( \frac{\text{s}^*}{\text{s}} \right) = 31536000 = 1 [\text{year}]$$

$$m^* = 1 \text{ AU} \rightarrow \left( \frac{m^*}{m} \right) = 149597870700 = 1 [\text{AU}]$$

$$\Rightarrow \left( \frac{\text{kg}^*}{\text{kg}} \right)^{-1} = [G M_s] [\text{AU}]^{-3} [\text{year}]^2 \simeq 10^6 \text{ kg}$$

