

Summary part I

- *Nature of Light (chapter 1)*
- *Geometrical Optics (chapter 2)*
- *Optical Instrumentation (chapter 3)*

Nature of Light

- Light behaves like waves in its propagation and in the phenomena of interference and diffraction; however, it exhibits particle-like behavior when exchanging energy with matter, as in the Compton and photoelectric effects

- Photon energy:

$$E = h\nu \quad h : \text{Planck constant (} 6.626 \times 10^{-34} \text{ J}\cdot\text{s) }$$

- Energy unit eV: $1\text{eV} = 1.6 \times 10^{-19} \text{ J}$

- Relations between wavelength, frequency, and speed:

In air: $c = \lambda \nu \quad \text{or} \quad \lambda = c / \nu$

In media: $\nu = c / n, \quad \lambda = \frac{\nu}{\nu} = \frac{c}{n\nu} = \frac{\lambda_o}{n}$

Geometrical Optics

Law of reflection: $\theta_i = \theta_r$

Law of refraction (Snell's Law): $n_i \cdot \sin \theta_i = n_t \cdot \sin \theta_t$

- Wavefront : The surfaces joining all points of equal phase are known as wavefronts.
- Huygens's Principle : every point on a propagating wavefront serves as the source of spherical secondary wavelets, such that the wavefront at some later time is the envelope of these wavelets.
- Principle of Least Time: The actual path between two points taken by a beam of light is the one that is traversed in the least time
- Re-state Fermat's Principle : $t = OPL / c$
Light, in going from point S to P, traverses the route having the smallest *OPL*
- Prove Snell's law

Geometrical Optics

- Total Internal Reflection:

$$n_1 > n_2 \quad \theta_c = \sin^{-1} \left(\frac{n_2}{n_1} \right)$$

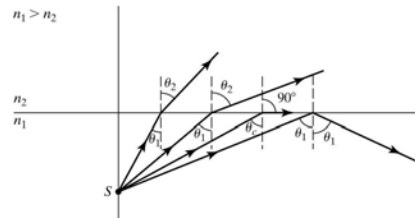
Critical angle: is the incident angle when refraction angle is equal to 90°

- Refraction through Plane Surfaces:

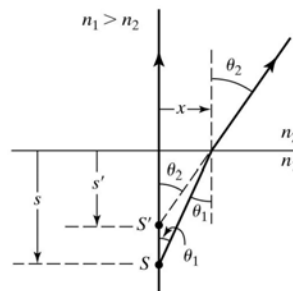
Paraxial rays: $\sin \theta \cong \tan \theta \cong \theta$

$$n_1 \sin \theta_1 = n_2 \sin \theta_2$$

$$\Rightarrow s' = \left(\frac{n_2}{n_1} \right) s$$



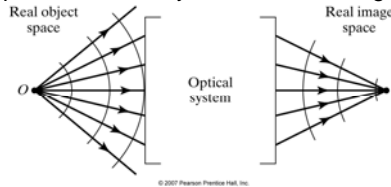
(c)



(b)

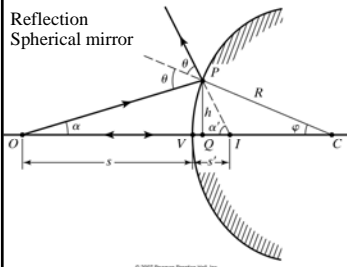
Geometrical Optics

- O and I are conjugate points – any pair of object-image points
.... by the principle of reversibility can be interchanged



1. Reflection at a spherical surface

Reflection
Spherical mirror



- Mirror equation:

$$\frac{1}{s} + \frac{1}{s'} = \frac{1}{f}$$

with $f = -\frac{R}{2}$
Convex mirror: $f < 0$
Concave mirror: $f > 0$

- Lateral magnification:

$$m = \frac{h_i}{h_o} = -\frac{s'}{s}$$

Sign convention

s	+	O is to the left of V
s'	+	I is to the left of V
R	+	C is to the right of V
f	+	F is to the left of V

Geometrical Optics

2. Refraction at a spherical surface:

- Refraction equation:

$$\frac{n_1}{s} + \frac{n_2}{s'} = \frac{n_2 - n_1}{R}$$

- Special case: plane refracting surface

$$R \rightarrow \infty$$

$$s' = -\left(\frac{n_2}{n_1}\right)s$$

s' : apparent depth

- Lateral magnification:

$$m = \frac{h_i}{h_o} = -\frac{n_1 s'}{n_2 s}$$

Sign convention

s	+	O is to the left of V
s'	+	I is to the right of V
R	+	C is to the right of V

Describe images:

- Type: Virtual or real
- Orientation: erect or inverted
- Relative size

Meanings associated with the sign

Quantity	Sign	
	+	-
s	real object	virtual object
s'	real image	virtual image
m	erect image	inverted image

Geometrical Optics

3. Thin Lenses:

➤ Thin-lens equation: $\frac{1}{s} + \frac{1}{s'} = \frac{1}{f}$

➤ Lateral magnification: $m = \frac{h_i}{h_o} = -\frac{s'}{s}$

Lensmaker's equation:

$$\frac{1}{f} = \left(\frac{n_2}{n_1} - 1 \right) \left(\frac{1}{R_1} - \frac{1}{R_2} \right)$$

Sign convention		
s	+	O is to the left of V
s'	+	I is to the <u>right</u> of V
$R_{1,2}$	+	C is to the right of V

Meanings associated with the sign		
Quantity	Sign	
	+	-
s	real object	virtual object
s'	real image	virtual image
m	erect image	inverted image

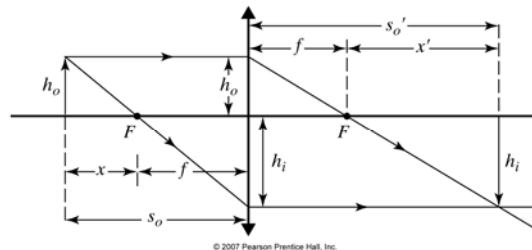
➤ Rules for ray tracing:

- Ray 1 parallel to central axis will pass through the right focal point of a converging lens (or the left focal point of the diverging lens)
- Ray 2 passing through the left focal point of a converging lens (or the right focal point of the diverging lens) will emerge from lens parallel to central axis
- Ray 3 is the undeviated ray through center of lens (Ray 3 is also called central ray)

Geometrical Optics

➤ Newtonian Equation for Thin Lens: $xx' = f^2$

➤ Lateral magnification: $m = \frac{h_i}{h_o} = -\frac{s'}{s} = -\frac{x'}{f} = -\frac{f}{x}$



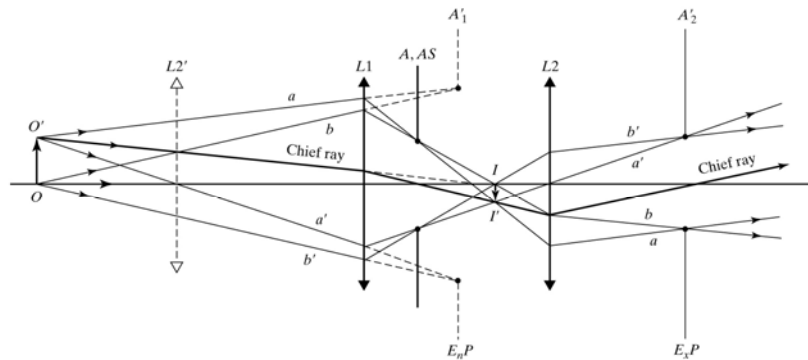
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Optical Instrumentation

3.1 Stops, Pupils, and Windows:

- Apertures are used to control Image Brightness and limit Field of view.
- Apertures for controlling Image Brightness : Aperture stops and Pupils
 - Aperture stop AS: The real element in an optical system that limits the size of the cone of rays accepted by the system from an axial object point
 - Entrance pupil (E_nP): The image of the aperture stop formed by the optical elements (if any) that precede it
 - Exit pupil (E_xP): The image of the aperture stop formed by the optical elements (if any) that follow it
- Chief Ray: The chief is a meridional ray that start at the edge of the object, and passes through the center of E_nP , AS, and E_xP
- Apertures for Limiting Field of view : Field stops and Windows
 - Field stop FS: The real element that limits the angular field of view formed by an optical system
 - Entrance Window (E_nW): The image of the field stop formed by the optical elements that precede it
 - Exit Window (E_xW): The image of the field stop formed by the optical elements that follow it

Optical Instrumentation

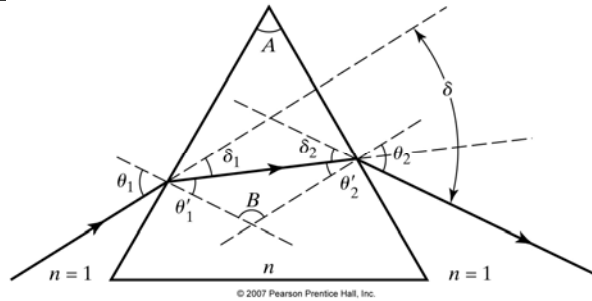


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- Aperture stop AS
- Entrance pupil (E_nP)
- Exit pupil (E_xP)
- Chief Ray: The chief is a meridional ray that start at the edge of the object, and passes through the center of E_nP , AS, and E_xP

Optical Instrumentation

3.3 Prisms:



- The deviation angle is: $\delta = \theta_1 + \theta_2 - (\theta'_1 + \theta'_2) = \theta_1 + \theta_2 - A$

$$\delta = \theta_1 + \sin^{-1} \left(\sqrt{n^2 - \sin^2 \theta_1} \sin A - \sin \theta_1 \cos A \right) - A$$

- Condition of Minimum deviations: The ray for which the deviation is a minimum traverses the prism symmetrically, that is, parallel to its base.
- Applications : Determination of refractive index n of transparent material

$$n = \frac{\sin \left[(\delta_m + A)/2 \right]}{\sin (A/2)} \quad \delta_m: \text{Minimum angular deviation}$$

Optical Instrumentation

- Empirical relation ---Cauchy Relation:

$$n_\lambda = A + \frac{B}{\lambda^2} + \frac{C}{\lambda^4} + \dots \quad A, B, C \dots \text{ are Cauchy constants}$$

- Dispersion: $\frac{dn}{d\lambda} \approx -\frac{2B}{\lambda^3}$

- Resolving power: $\mathcal{R} = \frac{\lambda}{(\Delta\lambda)_{\min}} = b \frac{dn}{d\lambda}$

3.4 The Camera:

$$s_1 = \frac{s_o f (f + Ad)}{f^2 + Ads_o}$$

$$s_2 = \frac{s_o f (f - Ad)}{f^2 - Ads_o}$$

$$\text{depth of field} = \frac{2Ads_o(s_o - f)f^2}{f^4 - A^2d^2s_o^2}$$

Optical Instrumentation

3.5 Simple Magnifiers and Eyepieces:

- **Angular magnification M :** the ratio of the retinal image as seen through the instrument over the size of the retinal image as seen by the unaided eye at normal viewing distance.

1. Image viewed at infinity: $M = \frac{25}{f}$ (Object is moved to the focal point)
2. Image viewed at normal near point: $M = \frac{25}{f} + 1$ (Object is just inside the focal point)
 $s' = -25 \text{ cm}$

- **Eyepieces or Oculars:** To reduce transverse chromatic aberration, two lens are most often used. The effective focal length f of the two thin lenses separated by a distance L is given:

$$\frac{1}{f} = \frac{1}{f_1} + \frac{1}{f_2} - \frac{L}{f_1 f_2}$$

with

$$\frac{1}{f_1} = (n-1) \left(\frac{1}{R_{11}} - \frac{1}{R_{12}} \right) = (n-1) K_1$$

$$\frac{1}{f_2} = (n-1) \left(\frac{1}{R_{21}} - \frac{1}{R_{22}} \right) = (n-1) K_2$$

To correct transverse chromatic aberration:

$$\frac{d(1/f)}{dn} = 0$$



Optimum separation:

$$L = \frac{1}{2}(f_1 + f_2)$$

Optical Instrumentation

3.6 Microscopes:

- **Compound Microscope :** Objective + Eyepiece (it may provide high MP)
- Objective: form a real, inverted, magnified image → Lateral Magnification - M_o
- Eyepiece: further magnify the intermediate image → Angular Magnification - M_e
- The intermediate image is at or just inside of the 1st focal point f_e of the eyepiece.

- The separation of the two lens: $d = f_o + L + f_e$

- **Magnification:** $M = M_o M_e = - \left(\frac{L}{f_o} \right) \left(\frac{25}{f_e} \right)$ Unit: cm

or $M = \frac{25}{f_{eff}}$ with $\frac{1}{f_{eff}} = \frac{1}{f_1} + \frac{1}{f_2} - \frac{d}{f_1 f_2}$

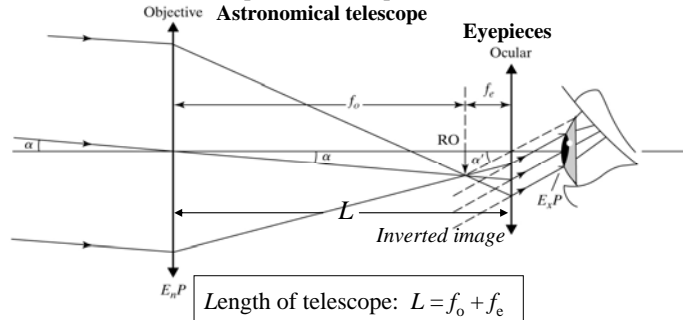
d : Separation of the two lenses

L : Distance between the 2nd focal point and the objective image

Optical Instrumentation

3.7 Telescopes:

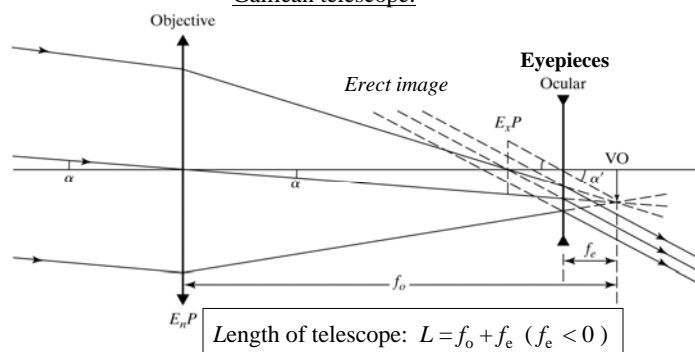
Keplerian telescope: Astronomical telescope



- Eyepiece is located so that 1st focus overlaps 2nd focus of Objective
- Nearly parallel rays of light from a distant object are collected by a positive lens--Objective, formed a real inverted image in its focal plane
- Then the intermediate image, located at or near the focal point of eyepiece, serves as a real object for the Ocular
- Angular magnification (when image is viewed at infinity): $M = \frac{\alpha'}{\alpha} = -\frac{f_o}{f_e}$

Optical Instrumentation

Galilean telescope:



- Angular magnification: $M = \frac{\alpha'}{\alpha} = -\frac{f_o}{f_e}$