

Searching for dark matter and exotic physics with atomic clocks and GPS

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Nature Comm. 8, 1195 (2017), arXiv:1704.06844
Phys. Review D (2018), arXiv:1803.10264
arXiv:1803.00617

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Outline:

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Ultralight DM +
TDs

GPS

Initial search/
first results

Bayesian search

Asymmetry &
ann. modulation

- Ultra light dark matter; “clumps”, e.g. Topological defects
- Transient signals: Global networks of precision devices
- GPS: 50,000km aperture sensor array
 - ~ 30 satellite clocks, > 15 years of archived data
- Initial search: domain walls
 - limits: orders of magnitude improvement for certain models
- Looking forward: Bayesian search technique + broader models
- Noise asymmetry & annual modulation signatures

Dark Matter: What is it?

- $\sim 25\%$ of Universe energy budget (cf $\sim 5\%$ for “normal” matter)
- We don't know what it is – but we have narrowed it down to a small 90 orders-of-magnitude window

Rough mass-range for various models:

- MACHOs: $10^{58} - 10^{68} \text{ eV}$
- WIMPs: $10^6 - 10^{12} \text{ eV}$
- I-WIMPS: $1 - 10^6 \text{ eV}$
- Axions: $10^{-10} - 10^{-4} \text{ eV}$
- Ultralight Q fields: $10^{-24} - 1 \text{ eV}$

(context: $m_{\text{Earth}} \sim 10^{60} \text{ eV}$ $m_{\text{electron}} \sim 10^6 \text{ eV}$)

- Even asserting that DM is a fundamental particle (i.e. ignore MACHOs) \implies 36 orders of magnitude range

Ultralight Dark Matter:

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WIMPs

- long-time “favourite” DM candidate
- Masses $\sim 10 - 1000$ GeV
- Many null WIMP results
- Increased interest in other forms of DM

Ultralight fields (e.g., axions)

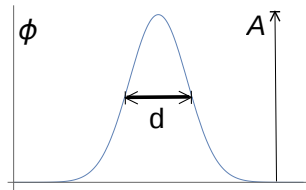
- Masses $\sim 10^{-24} - 1$ eV
- Oscillating field: $\phi = a \cos(m_a t)$ [*T. Kalaydzhyan: next talk*]
- Stable topological defects: monopoles, strings, walls
 - Also: Q-balls, solitons, “clumps”

- Peccei & Quinn '77, Weinberg '78, Dine & Fischler '82,...

Topological Defect DM

Topological Defects

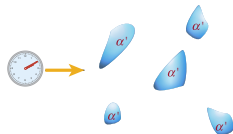
- monopoles, strings, walls,
- Defect width: $d \sim 1/m_\phi$
- Earth-scale object $\sim 10^{-14}$ eV



Inside: $\phi^2 \rightarrow A^2$, Outside: $\phi^2 \rightarrow 0$

Dark matter: Gas of defects

- DM: galactic speeds: $v_g \sim 10^{-3}c$
- A^2 , d , $\mathcal{T}_{b/w}$ collisions $\implies \rho_{DM}$



$$A^2 = \rho_{DM} v_g d \mathcal{T},$$

- Sikivie '82, Preskil '83, Vilekin '85,
Coleman '85, Lee '89, ...

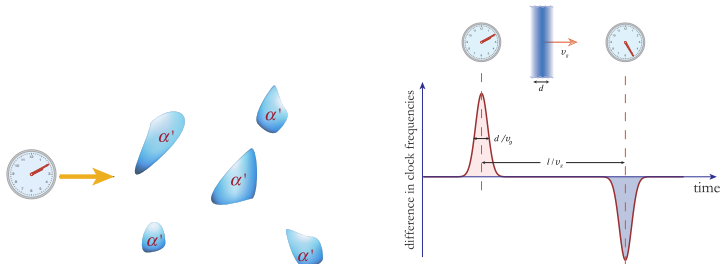
Shift in atomic clock frequencies

- DM may interact with: Photons, fermions
 \Rightarrow shifts in energy levels \Rightarrow shifts in clock frequencies

$$\frac{\delta\omega(r, t)}{\omega_0} = \phi^2(r, t) \sum_X K_X \Gamma_X$$

K_X sensitivity: Flambaum, Dzuba, Can. J. Phys. 87, 25 ('09).

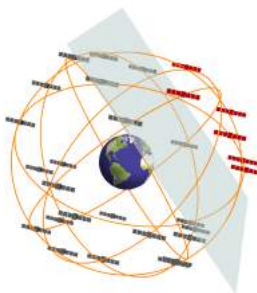
Monitor Atomic Clocks



- Derevianko, Pospelov, Nat. Phys. 10, 933 (2014).

GPS: 50,000 km DM observatory

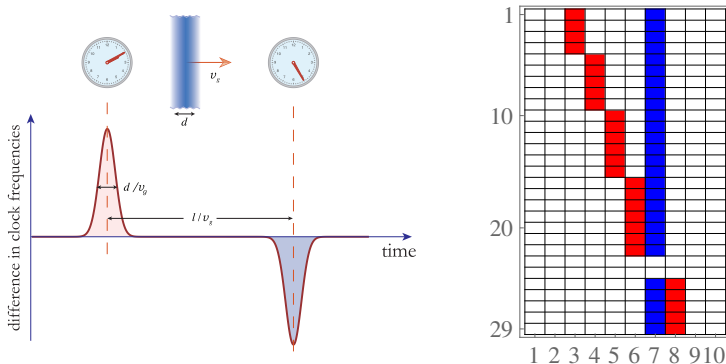
- 32 satellite clocks (Rb/Cs), ~ 16 years of high-quality data
- Also several H-maser ground-based clocks.
- Data from JPL: (sideshow.jpl.nasa.gov/pub/jpligsac/)
 - 30s sampled data; 0.01–0.1 ns precision
- Correlated, directional signal, with $v_g \sim 300$ km/s



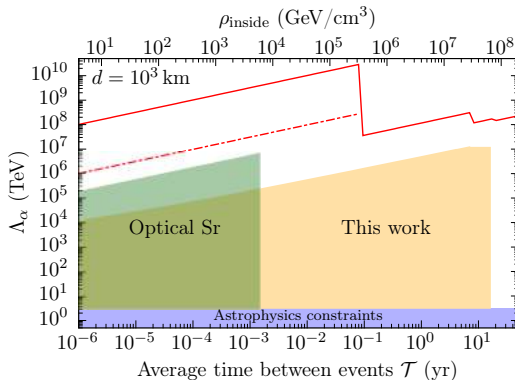
- Derevianko, Pospelov, Nat. Phys. 10, 933 (2014).
- & GNOME: Pospelov, Pustelny, Ledbetter, Kimball, Gawlik, Budker, PRL110, 21803 ('13).

DM Walls: Initial search/limits

- Thin wall [easiest case]: brief (< 30 s) frequency excursion



- \vec{v} encoded in time-delay and signal ordering: $\Delta t \sim$ minutes
- BMR, Blewitt, Dailey, Pospelov, Rollings, Sherman, Williams, Derevianko, Nature Comm. **8**, 1195 (2017). [1704.06844]

Results: Limits - Λ_α (photon)

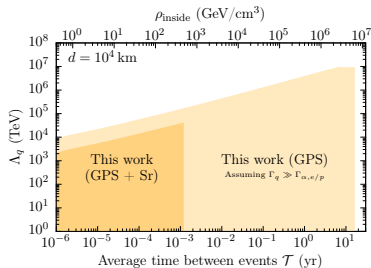
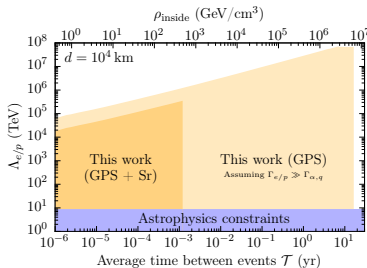
Sr: Wcislo, Morzynski, Bober, Cygan, Lisak, Ciurylo, Zawada, Nat. Astron. 1, 9 (2016).
Astro: Olive, Pospelov, Phys. Rev. D. 77, 43524 (2008).

- BMR, Blewitt, Dailey, Pospelov, Rollings, Sherman, Williams, Derevianko, Nature Comm. 8, 1195 (2017). [1704.06844]

Results: Limits - fermion masses

Combine Rb, Cs, and Sr (optical)

- Three different combo's of three couplings



Sr: Wcislo, Morzynski, Bober, Cygan, Lisak, Ciurylo, Zawada, Nat. Astron. 1, 9 (2016).
Astro: Olive, Pospelov, Phys. Rev. D. 77, 43524 (2008).

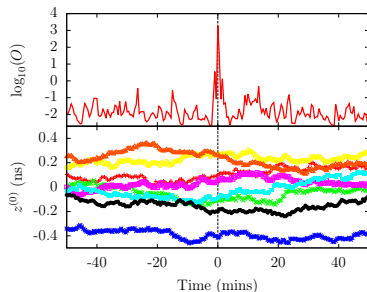
- BMR, Blewitt, Dailey, Pospelov, Rollings, Sherman, Williams, Derevianko, Nature Comm. 8, 1195 (2017). [1704.06844]

How to improve upon this?

- There may be events “hiding” below the noise.
- Other geometries: monopoles, strings, thicker walls

Bayesian Analysis

- Marginalise (integrate) all parameters (In-built Occam's Razor)
 - Time, velocity, object size, impact parameter
- Form odds ratios



- Events as small as:

$$s \approx \sigma / \sqrt{N} \approx 0.001 \text{ ns}$$

(for the best GPS clocks)

Bayes: True + false positive rates

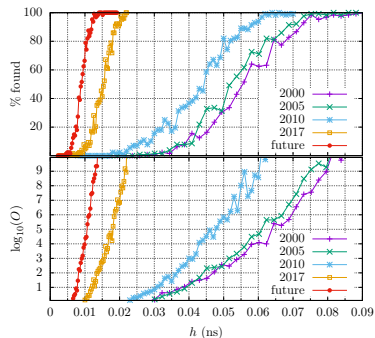
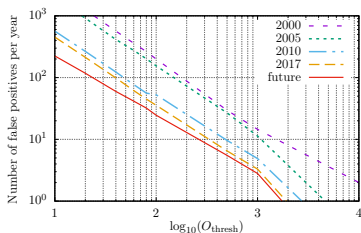
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Bayes: Projected sensitivity

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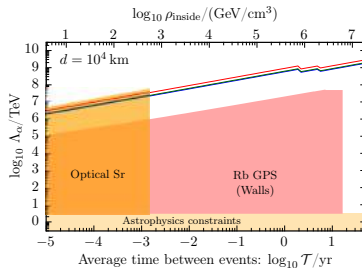
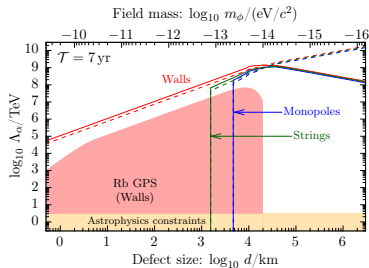
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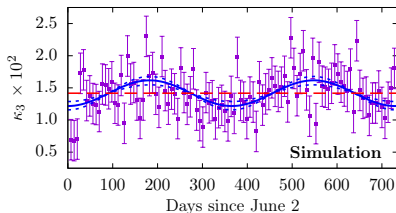
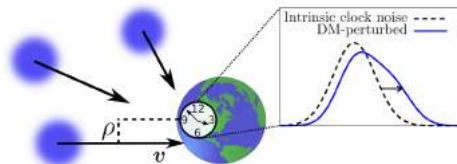
Large time between events (low number density), large objects:



- BMR, Blewitt, Dailey, Derevianko, Phys.Rev.D (2018). [1803.10264]

Asymmetry + an. modulation

Small objects, large event rate: asymmetry in noise distribution



- Yearly change in event rate:
- Sun + Earth velocities add
- $R(t) = R_0 + R_m \cos(\omega t + \phi_{\text{June2}})$

• BMR, Derevianko, arXiv:1803.00617

Asymmetry + an. modulation

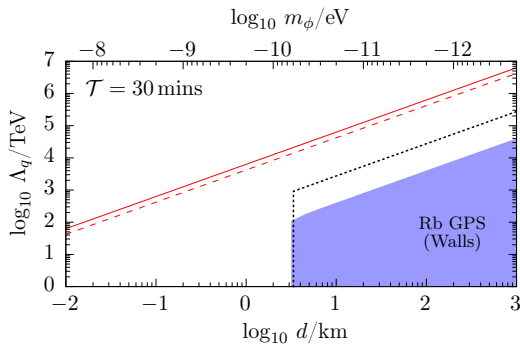
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Conclusion:

GPS: 50,000km aperture DM observatory

- Topological defect dark matter/transient exotic physics
 - GPS: 50,000km aperture sensor array
 - ~ 30 satellite clocks, many earth clocks, > 15 years of clock data
 - DM walls: Orders of magnitude improvement for certain models
 - Looking forward: Bayesian search technique
 - Monopoles, strings, signals below σ_{clock}
 - General technique: archived, time-stamped data
-
- Nature Comm. **8**, 1195 (2017). BMR¹, G. Blewitt¹, C. Dailey¹, M. Pospelov^{2,3}, A. Rollings¹, J. Sherman⁴, W. Williams¹, A. Derevianko¹ [1704.06844].
 - Phys. Rev. D (2018). BMR, G. Blewitt, C. Dailey, A. Derevianko [1803.10264].
 - BMR, A. Derevianko [1803.00617].

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Some references:

Axions, ultralight scalar DM:

- R. D. Peccei and H. R. Quinn, Phys. Rev. Lett. 38, 1440 (1977).
- P. Sikivie, Phys. Rev. Lett. 51, 1415 (1983); Phys. Rev. Lett. 48, 1156 (1982).

Topological defect DM:

- T. W. B. Kibble, Phys. Rep. 67, 183 (1980).
- A. Vilenkin, Phys. Rep. 121, 263 (1985).

non-topological solitons, Q-balls:

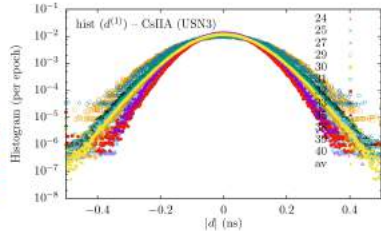
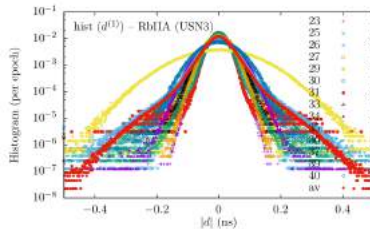
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- A. Kusenko and P. J. Steinhardt, Phys. Rev. Lett. 87, 141301 (2001).

Other non-gravitational TD searches:

- M. Pospelov, S. Pustelny, M. P. Ledbetter, D. F. J. Kimball, W. Gawlik, and D. Budker, Phys. Rev. Lett. 110, 21803 (2013).
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- E. D. Hall, T. Callister, V. V. Frolov, H. Muller, M. Pospelov, and R. X. Adhikari, arXiv:1605.01103.
- P. Wcisło, Morzynski, Bober, Cygan, Lisak, Ciuryło, M. Zawada, Nat. Astron. 1, 9 (2016).

Aside: challenges of re-purposed data

data from JPL: Histogram



- Possible that some clocks mis-identified
(Here, one of the “Rb” clocks is probably Cs).
- Same discrepancy in autocorrelation function, Allan variance etc.