Radio-Quiet AGN

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Outline

Definition of Radio-quiet AGN

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Definition of Radio-quiet AGN

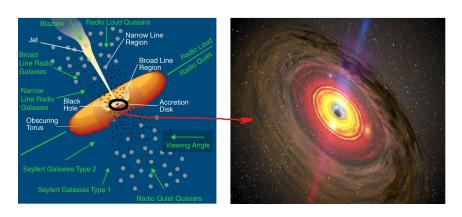


Figure 1: Model of AGN and the animation of accretion disc.

Definition of Radio-quiet AGN - Cont.

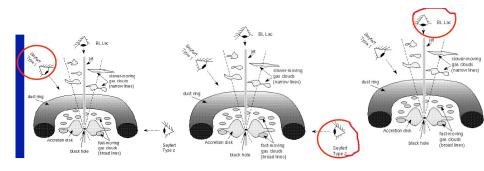
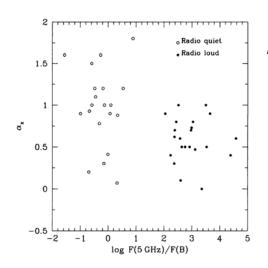


Figure 2: different lines of viewing AGN

- Through obscuring AGN ⇒ broad lines emitted by hot fast moving gas around the BH.
- Through a dust torus ⇒ narrow lines emitted by slower moving clouds farther from the BH.

Definition of Radio-quiet AGN - Cont.



$$R = 1.36 \times 10^5 \times \frac{L(5 GHz)}{L(4400 \text{Å})} (1)$$

Figure 3: Spectral index VS Radio loudness [Labiano ,2006]



Definition of Radio-quiet AGN - Cont.

Radio-quiet AGN \Longrightarrow are those with a lack of emitted relativistic jets and with radio loudness lower than ten (R < 10).

Classification of Radio-quiet AGN

- Seyfert-I;
- Seyfert-II;
- Quasars and
- 4 Liners.

Classification of Radio-quiet AGN - Cont.

Seyfert galaxies have **bright compact cores** which spit out lots of strong **infrared emission**.

- Type-I ⇒ faster-moving hot gas → broad emission lines (strong continuum emission).
- Type-II ⇒ slower moving clouds → narrow emission lines (very weak continuum emission).

Classification of Radio-quiet AGN - Cont.



Figure 4: Seyfert-I: NGC 1068 [https://en.wikipedia.org/wiki/Seyfert-galaxy]



Figure 5: Seyfert-II: NGC 7742 [https://www.spaceanswers.com/deepspace/what-is-a-seyfert-galaxy/]

Quasars — are more luminous versions of Seyfert-I — always show strong optical and X-ray continuum emissions — broad and narrow optical emission lines. The host galaxies of quasars can be spirals, irregulars or ellipticals.



Figure 6: Quasar [https://www.nasa.gov/mission-pages/webb/news/dusty-quasar.html]

Final Considerations

The Studies of the physical processes that take place in the nuclear region of radio-quiet AGN continue to be a principal role of Scientists as, once all those processes are well understood, not only will be a great leap in most of the matter related to the **high energy particle** accelerators but also useful information related to the possible link between **AGN** and the star formation will be known.

The End!

Khanimambo!

Thank you!