Computer Practical: Saturn's Hexagon

Paul Connolly, November 2020

1 Overview

In this computer practical, a shallow water model on a rotating sphere, implemented in Fortran, is used to simulate Saturn's Hexagon.

2 Model formulation

The equations are largely the same as for the practical on Shallow Water Modelling; however, the circumference of the model domain now changes with latitude as it is mapped onto a sphere. In addition, the model can be split into sections and the equations in these sections can be solved on separate computer processors in parallel. This is key to how weather predictions are completed in time: the model is sped up by performing the calculations in parallel.

3 Downloading and compiling the code

Log into the server computer with an SSH window. Also, log into the server computer with an SFTP window. Space the windows out on your screen so that you can easily move between them.

If you have not done this already, download the code you will be using today from GitHub by typing the following:

```
git clone ssh://git@github.com/EnvModelling/shallow-water-model-on-sphere
```

this should download the code to your working directory.

In order to access the files you will need to change the working directory to where the files are. Type the following: cd shallow-water-model-on-

followed by the tab key and the command should auto-complete. Then press the enter key.

Type 1s followed by the enter key. The screen should list the files in this directory.

Finally, type make followed by the enter key. This will compile the FORTRAN code into machine code that the computer can run.

4 Viewing and Editing the code

We will use the nano text editor to view some files. In this practical there is an input file called namelist.in. To see what variables we can change, type:

```
nano -l namelist.in
```

the -1 means to show line numbers. The code can be edited in the text editor and saved by typing Ctrl-X at the same time and pressing Y to save the file. The important variable is u_jet on line 11 of this file.

5 Running the model

After the code has been compiled (see Section 3), the procedure for running the model is to:

- 1. Edit the file namelist.in to configure the model.
- 2. Type ./run.sh at the command line, followed by enter to run the model.

On Saturn the Hexagon appears around 77° N. Scientists have been able to determine the average wind speed by tracking the motion of clouds over time. However, the speed of these features also includes the speed of the planet's rotation. How do we determine the planet's rotation when there is no solid ground? There are radio emissions that emanate from the core of Saturn. These radio waves have a periodic nature, which repeats every 10 and a half hours. This is usually taken to be the rotation period of Saturn.

After running the model a NetCDF file will be generated at the location /tmp/<username>/output.nc. You can then run a python script by typing:

```
python3 python/height_and_streamlines.py
```

and a plot of the model simulation will be shown in /tmp/<username>/frame.png, which you can download using SFTP by typing:

```
get /tmp/<username>/frame.png
```

Additionally, we can make an animation of the output by typing:

```
python3 python/animate_output.py
```

and a plot of the model simulation will be shown in /tmp/<username>/animation.gif, which you can download using SFTP by typing:

```
get /tmp/<username>/animation.gif
```

6 Experiments

FOR THESE EXPERIMENTS WE DO NOT HAVE ENOUGH COMPUTE POWER FOR EVERYONE TO DO THE SIMULATIONS AT ONCE. THEREFORE YOU WILL WORK IN GROUPS ASSIGNED BY THE INSTRUCTOR. ONE PERSON FROM EACH GROUP WILL RUN THE SIMULATION, AND POST THE FILES FOR EVERYONE ELSE IN THE CLASS.

You will notice that the model takes far too long to run on a single processor. for this reason we will run the model on 12 processors. To do this we will add an argument to the ./run.sh command described in Section 5, which is the number of processors to run on as follows:

```
./run.sh 12
```

YOU MUST CHOOSE NO MORE THAN 12!

6.1 Simulations to do

There will be 5 groups. Simulations to do are to set the value u jet to the following values:

- 1. 45 m s^{-1}
- 2.60 m s^{-1}
- 3. 150 m s^{-1}
- 4. 250 m s^{-1}
- 5. 310 m s^{-1}

Your groups will be assigned in the class. See Section 5 for details of how the run the model and generate the output (with the update in Section 6). Download the file and then explain what you see in the class (and upload the output for the rest of the class).

Exercise: Where do the cyclones and anticyclones form?

Exercise: What do you notice happens as the wind speed in the jet is increased?