

# Homework Of Superconductivity and Magnetism

Vo Chau Duc Phuong <sup>1</sup>

<sup>1</sup>the Abdus Salam International Center for Theoretical Physics

April 2025

## 1 Superconductivity

- Considering the 2-D potential have the form:

$$U(r) = \begin{cases} -U_0, & r < a \\ 0, & r > a \end{cases} \quad (1)$$

Find the shallow energy level when  $U_0 \ll \hbar^2/ma^2$  for the case  $M_z = 0$ , where  $M_z$  is the projection of the orbital moment on the z-axis.

In the polar coordinate, the kinetic have the form:

$$K(r, \theta)\psi(r, \theta) = \frac{1}{r} \frac{\partial}{\partial r} \left( r \frac{\partial \psi(r, \theta)}{\partial r} \right) + \frac{1}{r} \frac{\partial^2 \psi(r, \theta)}{\partial \theta^2} \quad (2)$$

Since  $M_z = 0$ , we have the ansatz:

$$\psi(r, \theta) = \frac{1}{\sqrt{4\pi}} R(r) \quad (3)$$

Therefore, the second term of (2) vanished. The stationary Schrödinger equation have the form:

$$-\frac{\hbar^2}{2m} \left( \frac{1}{r} \frac{\partial R}{\partial r} + \frac{\partial^2 R}{\partial r^2} \right) + V(r)R = ER$$

Or in the form of Bessel equation:

$$r^2 \frac{\partial^2 R}{\partial r^2} + r \frac{\partial R}{\partial r} + \frac{2m(E - V)}{\hbar^2} R = 0 \quad (4)$$

Inside the circle, we have the equation:

$$r^2 \frac{\partial^2 R}{\partial r^2} + r \frac{\partial R}{\partial r} + \frac{2m(E + U_0)}{\hbar^2} R = 0 \quad (5)$$

Which have the general solution:

$$R(r) = C_1 J_0 \left( \sqrt{\frac{2m(E + U_0)}{\hbar^2}} r \right) + C_2 Y_0 \left( \sqrt{\frac{2m(E + U_0)}{\hbar^2}} r \right) \quad (6)$$

But since the radius have to be continuous at  $r = 0$ , where  $Y_0(r)$  is not, therefore:  $C_2 = 0$ . Hence

$$R(r) = C_1 J_0 \left( \sqrt{\frac{2m(E + U_0)}{\hbar^2}} r \right) \quad (7)$$

Outside the circle, we have:

$$r^2 \frac{\partial^2 R}{\partial r^2} + r \frac{\partial R}{\partial r} - \frac{2m|E|}{\hbar^2} R = 0 \quad (8)$$

Which yields the Bessel modified function as the solution:

$$R(r) = C_3 I_0 \left( \sqrt{\frac{2m|E|}{\hbar^2}} r \right) + C_4 K_0 \left( \sqrt{\frac{2m|E|}{\hbar^2}} r \right) \quad (9)$$

The constrain condition now is  $R(r)$  have to vanished at  $r \rightarrow \infty$ , which  $I_0\left(\sqrt{\frac{2m|E|}{\hbar^2}}r\right)$  is not, therefore:

$$R(r) = C_2 K_0\left(\sqrt{\frac{2m|E|}{\hbar^2}}r\right) \quad (10)$$

At  $r = a$ , the function have to be continuous:

$$C_2 = C_1 \frac{J_0\left(\sqrt{\frac{2m(E+U_0)}{\hbar^2}}a\right)}{K_0\left(\sqrt{\frac{2m|E|}{\hbar^2}}a\right)} \quad (11)$$

And also their derivative, we using the properties of Bessel function:

$$\frac{d}{dx} J_0(x) = -J_1(x); \quad \frac{d}{dx} K_0(x) = -K_1(x) \quad (12)$$

to get:

$$\sqrt{\frac{E+U_0}{|E|}} C_1 J_1\left(\sqrt{\frac{2m(E+U_0)}{\hbar^2}}a\right) = C_2 K_1\left(\sqrt{\frac{2m|E|}{\hbar^2}}r\right) \quad (13)$$

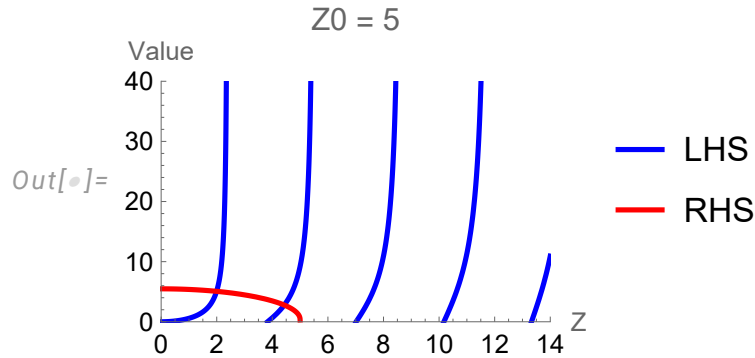
Or:

$$\sqrt{\frac{E+U_0}{|E|}} = \frac{K_1\left(\sqrt{\frac{2m|E|}{\hbar^2}}a\right)}{J_1\left(\sqrt{\frac{2m(E+U_0)}{\hbar^2}}a\right)} \frac{J_0\left(\sqrt{\frac{2m(E+U_0)}{\hbar^2}}a\right)}{K_0\left(\sqrt{\frac{2m|E|}{\hbar^2}}a\right)} \quad (14)$$

set  $Z = \sqrt{2m(E+U_0)}a/\hbar$ ,  $Z_0 = \sqrt{2mU_0}a/\hbar$ , equation become:

$$\begin{aligned} \frac{Z}{\sqrt{Z^2 - Z_0^2}} &= \frac{K_1(\sqrt{Z_0^2 - Z^2})}{J_1(Z)} \frac{J_0(Z)}{K_0(\sqrt{Z_0^2 - Z^2})} \\ Z \frac{J_1(Z)}{J_0(Z)} &= \sqrt{Z^2 - Z_0^2} \frac{K_1(\sqrt{Z_0^2 - Z^2})}{K_0(\sqrt{Z_0^2 - Z^2})} \end{aligned} \quad (15)$$

Plot these two, we have: According to the LHS, we seeing that however the  $U_0$  small, always exist a solution  $Z$  for



both side to match. At very shallow case:  $U_0 \ll \frac{\hbar^2}{2ma^2}$  or  $Z \ll 1$ , we need to approximate the

## 2 Magnetism

Consider Hamiltonian:

$$\mathcal{H}_{int.} = -\vec{M} \cdot \vec{H} \quad (16)$$

Choosing the axis to be along the magnetic field  $\vec{H}$ .

- Using semi-classical:

$$Z = 2\pi \int_0^\pi e^{\beta m H \cos(\theta)} \sin \theta d\theta = \frac{2\pi}{\beta m H} 2 \sinh(\beta m H) \quad (17)$$

Calculate:  $\langle \vec{M} \cdot \vec{a} \rangle$ ,  $\vec{a} = (a_x, a_y, a_z)$ . We have:

$$\langle \vec{M} \cdot \vec{a} \rangle = \langle a_x M_x + a_y M_y + a_z M_z \rangle = a_x \langle M_x \rangle + a_y \langle M_y \rangle + a_z \langle M_z \rangle \quad (18)$$

$$\langle M_x \rangle \propto m \int \cos(\phi) \sin(\theta) e^{\beta m H \cos \theta} \sin(\theta) d\phi d\theta = 0 \quad (19)$$

$$\langle M_y \rangle \propto m \int \sin(\phi) \sin(\theta) e^{\beta m H \cos \theta} \sin(\theta) d\phi d\theta = 0 \quad (20)$$

$$\begin{aligned} \langle M_z \rangle &= \frac{m}{Z} \int_0^{2\pi} \int_0^\pi \cos \theta e^{\beta m H \cos(\theta)} \sin \theta d\theta d\phi = \frac{2\pi m}{Z} \partial_{\beta m H} \int_0^\pi e^{\beta m H \cos \theta} \sin \theta d\theta \\ &= m \left( \coth \beta m H - \frac{1}{\beta m H} \right) = mL(\beta m H), \end{aligned} \quad (21)$$

in which  $L(x)$  is Langevin function. Therefore:

$$\langle \vec{M} \cdot \vec{a} \rangle = a_z mL(\beta m H) \quad (22)$$

Consider:  $\langle M_z^2 \rangle - (\langle M_z \rangle)^2$ .

$$\begin{aligned} \langle M_z^2 \rangle &= \frac{m^2}{Z} \frac{\partial^2 Z}{\partial (\beta m H)^2} = m^2 \frac{\beta m H}{\sinh \beta m H} \left( \frac{\sinh(\beta m H)}{\beta m H} - 2 \cosh(\beta m H) \frac{1}{(\beta m H)^2} + \frac{2 \sinh(\beta m H)}{(\beta m H)^3} \right) \\ &= m^2 \left( 1 - 2 \coth(\beta m H) \frac{1}{\beta m H} + \frac{2}{(\beta m H)^2} \right) \\ (\langle M_z \rangle)^2 &= m^2 \left( 1 + \frac{1}{\sinh(\beta m H)^2} - \frac{2 \coth \beta m H}{\beta m H} + \frac{1}{(\beta m H)^2} \right) \end{aligned} \quad (23)$$

$$\langle M_z^2 \rangle - (\langle M_z \rangle)^2 = m^2 \left( \frac{1}{(\beta m H)^2} - \frac{1}{\sinh(\beta m H)^2} \right) \quad (24)$$

At  $T \rightarrow 0, \beta \rightarrow \infty$ :

$$\langle M_z^2 \rangle - (\langle M_z \rangle)^2 \rightarrow 0; \langle M_z \rangle = m \left( \coth(\beta m H) - \frac{1}{\beta m H} \right) \rightarrow m \quad (25)$$

In this case, all spin aligned with the external field, give no fluctuation in the distribution of the  $\langle M_z \rangle$ . At  $T \rightarrow \infty, \beta \rightarrow 0$ :

$$\langle M_z^2 \rangle - (\langle M_z \rangle)^2 \rightarrow \frac{m^2}{3}; \langle M_z \rangle = m \left( \coth(\beta m H) - \frac{1}{\beta m H} \right) \rightarrow 0 \quad (26)$$

In this case, the spin align in both side with the same number. And the fluctuation difference than 0, proportional to  $m^2$ .

• Using quantum theory, know that  $\vec{L} = 0, S = 1/2, \vec{M} = 2\mu_B \vec{S}$ :

$$Z = \sum_{m_z=-1/2}^{1/2} e^{\beta 2\mu_B H m_z} = 2 \cosh(\beta \mu_B H) \quad (27)$$

Which in turn give:

$$\langle S_z \rangle = \frac{1}{Z} \sum_{m_z=-1/2}^{1/2} m_z e^{\beta 2\mu_B H m_z} = \frac{1}{2} \tanh(\beta \mu_B H) \quad (28)$$

$$\langle S_z^2 \rangle = \frac{1}{Z} \sum_{m_z=-1/2}^{1/2} m_z^2 e^{\beta 2\mu_B H m_z} = \frac{1}{4} \quad (29)$$

$$\langle S_z^2 \rangle - \langle S_z \rangle^2 = \frac{1}{4} (1 - \tanh^2(\beta \mu_B H)) = \frac{1}{4 \cosh(\beta \mu_B H)} \quad (30)$$

At  $T \rightarrow 0, \beta \rightarrow \infty$ :

$$\langle S_z^2 \rangle - \langle S_z \rangle^2 \rightarrow 0, \langle S_z \rangle \rightarrow \frac{1}{2} \quad (31)$$

Like the classical results, all the spin align with the external field and no fluctuation with the distribution of  $S_z$ , all aligned.

At  $T \rightarrow \infty, \beta \rightarrow 0$ :

$$\langle S_z^2 \rangle - \langle S_z \rangle^2 \rightarrow \frac{1}{4}, \langle S_z \rangle \rightarrow 0 \quad (32)$$

All the spin equally divided into up and down, make the mean  $S_z$  to 0, their distribution make the variation of the spin to be half of the magnitude 1/2, indicate the distribution.

• Considering the dynamic of magnetization:

$$\frac{d\vec{M}}{dt} = -\gamma[\vec{M} \times \vec{H}] \quad (33)$$

With  $\vec{H} = H_0\hat{z} + H_1e^{-\kappa t}\hat{x}$ , we have the set of equation:

$$d_t M_x = -\gamma M_y H_0 \quad (34)$$

$$d_t M_y = \gamma M_x H_0 - \gamma M_z H_1 e^{-\kappa t} \quad (35)$$

$$d_t M_z = \gamma M_y H_1 e^{-\kappa t} \quad (36)$$

Since we have the condition  $H_1 \ll H_0$ , we can approximate the final equation to  $d_t M_z = 0 \rightarrow M_z = m_0$ , the  $m_0$  also be the total magnitude of the system due to the initial condition  $M_x(0) = M_y(0) = 0$ . The set of equations become:

$$d_t M_x = -\gamma H_0 M_y \quad (37)$$

$$d_t M_y = \gamma M_x H_0 - \gamma m_0 H_1 e^{-\kappa t} \quad (38)$$

Take derivative the first equation and substituting the second one into it to find:

$$d_t^2 M_x + (\gamma H_0)^2 M_x = \gamma m_0 H_1 e^{-\kappa t} \quad (39)$$

Homogeneous solution yields:

$$M_{xh} = C_1 \cos(\gamma H_0 t) + C_2 \sin(\gamma H_0 t) \quad (40)$$

Particular solution can be achieved using the ansatz:

$$M_{xp} = A e^{-\kappa t} \quad (41)$$

Substituting it to get the constant  $A$ :

$$M_{xp}(t) = \frac{m_0 \gamma H_1 e^{-\kappa t}}{(\gamma H_0)^2 + \kappa^2} \quad (42)$$

The general solution is:

$$M_x(t) = C_2 \sin(\gamma H_0 t) + \frac{m_0 \gamma H_1}{(\gamma H_0)^2 + \kappa^2} e^{-\kappa t} + C_1 \cos(\gamma H_0 t) \quad (43)$$

At  $t = 0$ :

$$M_x(0) = 0 \rightarrow C_1 = -\frac{m_0 \gamma H_1}{(\gamma H_0)^2 + \kappa^2} \quad (44)$$

Also at  $t = 0$ :

$$M_y(t) = d_t M_x(t) = C_2 \gamma H_0 \cos(\gamma H_0 t) - \kappa \frac{m_0 \gamma H_1}{(\gamma H_0)^2 + \kappa^2} e^{-\kappa t} - C_1 \gamma H_0 \sin(\gamma H_0 t) \quad (45)$$

$$M_y(0) = 0 \rightarrow C_2 = \kappa m_0 \frac{H_1}{H_0} \frac{1}{(\gamma H_0)^2 + \kappa^2} \quad (46)$$

So, the solution is:

$$M_x(t) = \kappa m_0 \frac{H_1}{H_0} \frac{\sin(\gamma H_0 t)}{(\gamma H_0)^2 + \kappa^2} + \frac{m_0 \gamma H_1}{(\gamma H_0)^2 + \kappa^2} \left( e^{-\kappa t} + \cos(\gamma H_0 t) \right) \quad (47)$$