

Ay 20 PS1 - Solutions

Q1

- a) The tidal height is proportional to the effective tidal force (F_T), which scales as $\frac{M_s}{D_s^3}$ ($M_s \equiv$ satellite mass, $D_s \equiv$ satellite distance).

$$\text{So, } h_{\text{sun}} = \frac{2 \times 10^{33} \text{ g}}{7 \times 10^{25} \text{ g}} \cdot \frac{(3.8 \times 10^{10} \text{ cm})^3}{(1.5 \times 10^{13} \text{ cm})^3} \times 1 \text{ m} \\ = 0.46 \text{ m.} \quad (1 \text{ point})$$

- b) Applying the formula from lectures,

$$\frac{d\nu}{dt} = - \frac{5 h g G M_s R_E^2 \sin 2\alpha}{M_E D_s^3},$$

we find a rate of $\sim 3 \times 10^{-13} \text{ s s}^{-1}$
or $-4 \times 10^{-23} \text{ Hz s}^{-1}$.

(1 point)

c) Difference between solar & sidereal day
today : $\Delta t_{\text{now}} = \frac{24 \text{ hr}}{365.25} = 237 \text{ s.}$

(1 point)

Considering both the Moon & Sun Tides,

$$\frac{d\nu}{dt} = -2.4 \times 10^{-22} \text{ Hz s}^{-1}$$

In 10^8 yr , $\Delta \nu = +7.6 \times 10^{-7} \text{ Hz}$,

$$\text{and } \nu = \frac{1}{86400 - 237} + 7.6 \times 10^{-7} = 1.24 \times 10^{-5} \text{ Hz}$$

or 22.5 hours. This is a sidereal day.

There will always be one more sidereal days per year than solar days. So

$$\Delta t_{10^8} = 208 \text{ s.} \quad (1 \text{ point})$$

d) The Moon moves further away. A simple calculation (assuming constant velocity of the Moon) suggests a distance increase of $\sim 40\%$.

Q3 As the primary explodes symmetrically, the secondary will not be subject to the primary's gravitational pull (see e.g. Purcell 3rd Ed. Fig 1-22, the field is 0 inside a spherical shell of charge/mass).
[1 pt]

The orbital velocity of the reduced mass in the CM frame is given by

$$V_{\mu} = \sqrt{\frac{GM}{R}} = \sqrt{\frac{GM\omega}{V_{\mu}}} \Rightarrow V_{\mu} = (GM\omega)^{1/3}$$

The orbital velocity of the secondary is then given by

$$V_2 = \frac{\mu}{M_2} V_{\mu} = \frac{M_1}{M_1 + M_2} [G(M_1 + M_2)\omega]^{1/3}$$

With $M_1 = 0.61 M_{\odot}$, $M_2 = 0.21 M_{\odot}$,

$$\omega = \frac{2\pi}{415} \text{ rad s}^{-1},$$

$$V_2 = 873 \text{ km s}^{-1} \quad [1 \text{ pt}]$$

Escape velocity

$$V_{\text{esc}} = \sqrt{\frac{2GM_{\text{tot}}}{R}} = 1272 \text{ km s}^{-1} \text{ with radius}$$

from question.

$$\text{So } V_2 \leq V_{\text{esc}}$$