

PHY319ObservingProject

May 8, 2018

Python module to align different frames: 1. Take all R-band images. 2. Identify x,y positions of the same bright stars in all images (either automatically or by-eye). 3. Work out the x and y shift between the images. 4. Shift images so the stars align (Have the same x-y positions). 5. Mean-combine the images to increase the S/N. (Also to try median combine) 6. Do the same for Halpha images. The galaxy will start to appear in Halpha after adding more images. This will increase the S/N. 7. Subtract the R-band. Before that, the R-band image should be normalised. On normalising, the stars in the Halpha image disappear after subtraction. This leaves with the galaxy in Halpha.

```
In [1]: #Reduction of images of M61
        from astropy.io import fits
        import os
        import numpy as np
        from astropy.visualization import (MinMaxInterval, SqrtStretch, ImageNormalize)
        # display plots in notebook
        %matplotlib inline
        import matplotlib.pyplot as plt
        import ccdproc
        from astropy import units as u

        #assigning directories
        data_dir = '../..../assignments/Session5/M61/raw_data'
        destination_dir = '../..../assignments/Session5/M61/calib_data'

        images = ccdproc.ImageFileCollection(location=data_dir,keywords=['type','exposure','filt

In [2]: print(images.summary) #the image types of the images are fine
        #hdu_list_b1 = fits.open('../..../assignments/Session5/M61/raw_data/m61_h_150_2.fits')
        #header_b1 = hdu_list_b1[0].header
        #print(header_b1['type'])
```

file	type	exposure	filter	xbinning
m61_h_150_1.fits	SCIENCE	150.0	H	2
m61_h_150_10.fits	SCIENCE	150.0	H	2
m61_h_150_11.fits	SCIENCE	150.0	H	2
m61_h_150_12.fits	SCIENCE	150.0	H	2
m61_h_150_13.fits	SCIENCE	150.0	H	2

m61_h_150_14.fits	SCIENCE	150.0	H	2
m61_h_150_15.fits	SCIENCE	150.0	H	2
m61_h_150_16.fits	SCIENCE	150.0	H	2
m61_h_150_17.fits	SCIENCE	150.0	H	2
m61_h_150_18.fits	SCIENCE	150.0	H	2
m61_h_150_19.fits	SCIENCE	150.0	H	2
m61_h_150_2.fits	SCIENCE	150.0	H	2
m61_h_150_20.fits	SCIENCE	150.0	H	2
m61_h_150_3.fits	SCIENCE	150.0	H	2
m61_h_150_4.fits	SCIENCE	150.0	H	2
m61_h_150_5.fits	SCIENCE	150.0	H	2
m61_h_150_6.fits	SCIENCE	150.0	H	2
m61_h_150_7.fits	SCIENCE	150.0	H	2
m61_h_150_8.fits	SCIENCE	150.0	H	2
m61_h_150_9.fits	SCIENCE	150.0	H	2
m61_r_240_1.fits	SCIENCE	240.0	R	2
m61_r_240_2.fits	SCIENCE	240.0	R	2

```
In [3]: calib_frames = ccdproc.ImageFileCollection(location=destination_dir, keywords=['type', 'e
```

```
print(calib_frames.summary)
```

file	type	exposure	filter	xbinning
-----	-----	-----	-----	-----
Bflatnorm1x1.fits	SKY	2.54617976223	B	1
Bflatnorm2x2.fits	SKY	2.54617976223	B	2
Bflatnorm3x3.fits	SKY	2.54617976223	B	3
Hflatnorm1x1.fits	SKY	21.1649156994	H	1
Hflatnorm2x2.fits	SKY	21.1649156994	H	2
Hflatnorm3x3.fits	SKY	21.1649156994	H	3
Iflatnorm1x1.fits	SKY	3.75051017342	I	1
Iflatnorm2x2.fits	SKY	3.75051017342	I	2
Iflatnorm3x3.fits	SKY	3.75051017342	I	3
Rflatnorm1x1.fits	SKY	2.20479702354	R	1
...
m61_h_150_20_reduced.fits	SCIENCE	150.0	H	2
m61_h_150_2_reduced.fits	SCIENCE	150.0	H	2
m61_h_150_3_reduced.fits	SCIENCE	150.0	H	2
m61_h_150_4_reduced.fits	SCIENCE	150.0	H	2
m61_h_150_5_reduced.fits	SCIENCE	150.0	H	2
m61_h_150_6_reduced.fits	SCIENCE	150.0	H	2
m61_h_150_7_reduced.fits	SCIENCE	150.0	H	2
m61_h_150_8_reduced.fits	SCIENCE	150.0	H	2
m61_h_150_9_reduced.fits	SCIENCE	150.0	H	2
m61_r_240_1_reduced.fits	SCIENCE	240.0	R	2
m61_r_240_2_reduced.fits	SCIENCE	240.0	R	2

Length = 43 rows

```
In [3]: def getCalibs(hdu):
        filename = hdu.header['filter']
        xbin = hdu.header['xbinning']
        dark, bias, flat = None, None, None
        for hdu in calib_3184_3.ccds(xbinning=2, ccd_kwargs=dict(unit=u.adu)):
            if hdu.header['type'] == 'DARK':
                dark = hdu
            elif hdu.header['type'] == 'BIAS':
                bias = hdu
            elif hdu.header['type'] == 'SKY' and hdu.header['filter'] == filename:
                flat = hdu
        return dark, bias, flat
```

```
In [7]: for ccd, fname in images.ccds(return_fname=True, ccd_kwargs=dict(unit=u.adu)):
        dark, bias, flat = getCalibs(ccd)
        bias_subtracted = ccdproc.subtract_bias(ccd, bias)
        dark_subtracted = ccdproc.subtract_dark(bias_subtracted, dark, exposure_time='exposure')
        reduced_image = ccdproc.flat_correct(dark_subtracted, flat)
        root, ext = os.path.splitext(fname)
        ofname = os.path.join(destination_dir, root) + '_reduced.fits'
        reduced_image.write(ofname)
```

WARNING: FITSFixedWarning: RADECSYS= 'FK5 '

the RADECSYS keyword is deprecated, use RADESYSa. [astropy.wcs.wcs]

WARNING:astropy:FITSFixedWarning: RADECSYS= 'FK5 '

the RADECSYS keyword is deprecated, use RADESYSa.

INFO:astropy:using the unit adu passed to the FITS reader instead of the unit adu in the FITS file.

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INFO:astropy:using the unitadu passed to the FITS reader instead of the unitadu in the FITS file. [astropy.io.fits] using the unitadu passed to the FITS reader instead of the unitadu in the FITS file. [astropy.io.fits] INFO:astropy:using the unitadu passed to the FITS reader instead of the unitadu in the FITS file. [astropy.io.fits] INFO: using the unitadu passed to the FITS reader instead of the unitadu in the FITS file. [astropy.io.fits] INFO:astropy:using the unitadu passed to the FITS reader instead of the unitadu in the FITS file. [astropy.io.fits] INFO: using the unitadu passed to the FITS reader instead of the unitadu in the FITS file. [astropy.io.fits] INFO:astropy:using the unitadu passed to the FITS reader instead of the unitadu in the FITS file. [astropy.io.fits] INFO: using the unitadu passed to the FITS reader instead of the unitadu in the FITS file. [astropy.io.fits] INFO:astropy:using the unitadu passed to the FITS reader instead of the unitadu in the FITS file. [astropy.io.fits] INFO: using the unitadu passed to the FITS reader instead of the unitadu in the FITS file. [astropy.io.fits]

```
In [5]: os.listdir(destination_dir)
```

```
Out[5]: ['m61_h_150_20_reduced.fits',  
        'Iflatnorm3x3.fits',  
        'Rflatnorm1x1.fits',  
        'bias3x3.fits',  
        'Bflatnorm2x2.fits',  
        'dark3x3.fits',  
        'Hflatnorm3x3.fits',  
        'Vflatnorm3x3.fits',  
        'Rflatnorm2x2.fits',  
        'M61_calibratedimages.zip',  
        'Bflatnorm1x1.fits',  
        'Hflatnorm2x2.fits',  
        'm61_h_150_10_reduced.fits',  
        'm61_h_150_11_reduced.fits',  
        'm61_h_150_13_reduced.fits',  
        'Bflatnorm3x3.fits',  
        'dark2x2.fits',  
        'm61_h_150_12_reduced.fits',
```

```

'bias2x2.fits',
'Vflatnorm1x1.fits',
'm61_r_240_1_reduced.fits',
'm61_h_150_17_reduced.fits',
'm61_h_150_8_reduced.fits',
'm61_h_150_9_reduced.fits',
'm61_h_150_16_reduced.fits',
'Iflatnorm2x2.fits',
'm61_h_150_14_reduced.fits',
'm61_r_240_2_reduced.fits',
'm61_h_150_15_reduced.fits',
'dark1x1.fits',
'm61_h_150_1_reduced.fits',
'm61_h_150_2_reduced.fits',
'Hflatnorm1x1.fits',
'm61_h_150_3_reduced.fits',
'm61_h_150_19_reduced.fits',
'm61_h_150_6_reduced.fits',
'm61_h_150_7_reduced.fits',
'm61_h_150_18_reduced.fits',
'm61_h_150_5_reduced.fits',
'Rflatnorm3x3.fits',
'Iflatnorm1x1.fits',
'm61_h_150_4_reduced.fits',
'Vflatnorm2x2.fits',
'bias1x1.fits']

```

In [3]: *#reduction of images of ngc3184_1*

#assigning directories

data_dir_3184_1 = '../..'/assignments/Session5/NGC3184_1/raw_data_1'

destination_dir_3184_1 = '../..'/assignments/Session5/NGC3184_1/calib_data_1'

images_3184_1 = ccdproc.ImageFileCollection(location=data_dir_3184_1,keywords=['type','e

#print(images_3184_1.summary)

calib_3184_1 = ccdproc.ImageFileCollection(location=destination_dir_3184_1,keywords=['ty

#print(calib_3184_1.summary)

file	type	exposure	filter	xbinning
Bflatnorm1x1.fits	SKY	2.71539647751	B	1
Bflatnorm2x2.fits	SKY	2.71539647751	B	2
Bflatnorm3x3.fits	SKY	2.71539647751	B	3
Hflatnorm1x1.fits	SKY	21.2001404396	H	1
Hflatnorm2x2.fits	SKY	21.2001404396	H	2
Hflatnorm3x3.fits	SKY	21.2001404396	H	3
Iflatnorm1x1.fits	SKY	4.24281279695	I	1
Iflatnorm2x2.fits	SKY	4.24281279695	I	2
Iflatnorm3x3.fits	SKY	4.24281279695	I	3

Rflatnorm1x1.fits	SKY	2.51441818049	R	1
Rflatnorm2x2.fits	SKY	2.51441818049	R	2
Rflatnorm3x3.fits	SKY	2.51441818049	R	3
Vflatnorm1x1.fits	SKY	3.39509675375	V	1
Vflatnorm2x2.fits	SKY	3.39509675375	V	2
Vflatnorm3x3.fits	SKY	3.39509675375	V	3
bias1x1.fits	BIAS	0.03	H	1
bias2x2.fits	BIAS	0.03	H	2
bias3x3.fits	BIAS	0.03	H	3
dark1x1.fits	DARK	60.0	H	1
dark2x2.fits	DARK	60.0	H	2
dark3x3.fits	DARK	60.0	H	3

```
In [8]: for ccd, fname in images_3184_1.ccds(return_fname=True, ccd_kwargs=dict(unit=u.adu)):
        dark, bias, flat = getCalibs(ccd)
        bias_subtracted = ccdproc.subtract_bias(ccd, bias)
        dark_subtracted = ccdproc.subtract_dark(bias_subtracted, dark, exposure_time='exposure')
        reduced_image = ccdproc.flat_correct(dark_subtracted, flat)
        root, ext = os.path.splitext(fname)
        ofname = os.path.join(destination_dir_3184_1, root) + '_reduced.fits'
        reduced_image.write(ofname)
```

WARNING: FITSFixedWarning: RADECSYS= 'FK5 '
the RADECSYS keyword is deprecated, use RADESYSa. [astropy.wcs.wcs]
WARNING:astropy:FITSFixedWarning: RADECSYS= 'FK5 '
the RADECSYS keyword is deprecated, use RADESYSa.

```
-----

OSError                                Traceback (most recent call last)

<ipython-input-8-a7af521dd2af> in <module>()
      6     root, ext = os.path.splitext(fname)
      7     ofname = os.path.join(destination_dir_3184_1, root) + '_reduced.fits'
----> 8     reduced_image.write(ofname)
      9 os.listdir(destination_dir_3184_1)

/ext/anaconda3/lib/python3.5/site-packages/astropy/nddata/mixins/ndio.py in write(self,
    35         formats.
    36         """
----> 37         io_registry.write(self, *args, **kwargs)

/ext/anaconda3/lib/python3.5/site-packages/astropy/io/registry.py in write(data, format,
```

```

558
559     writer = get_writer(format, data.__class__)
--> 560     writer(data, *args, **kwargs)
561
562

/ext/anaconda3/lib/python3.5/site-packages/astropy/nddata/ccddata.py in fits_ccddata_wri
569     hdu = ccd_data.to_hdu(hdu_mask=hdu_mask, hdu_uncertainty=hdu_uncertainty,
570                           hdu_flags=hdu_flags)
--> 571     hdu.writeto(filename, **kwd)
572
573

/ext/anaconda3/lib/python3.5/site-packages/astropy/utils/decorators.py in wrapper(*args,
485         # one with the name of the new argument to the function
486         kwargs[new_name[i]] = value
--> 487     return function(*args, **kwargs)
488
489     return wrapper

/ext/anaconda3/lib/python3.5/site-packages/astropy/io/fits/hdu/hdulist.py in writeto(sel
865     # file object that's open to write only, or in append/update modes
866     # but only if the file doesn't exist.
--> 867     fileobj = _File(fileobj, mode='ostream', overwrite=overwrite)
868     hdulist = self.fromfile(fileobj)
869     try:

/ext/anaconda3/lib/python3.5/site-packages/astropy/utils/decorators.py in wrapper(*args,
485         # one with the name of the new argument to the function
486         kwargs[new_name[i]] = value
--> 487     return function(*args, **kwargs)
488
489     return wrapper

/ext/anaconda3/lib/python3.5/site-packages/astropy/io/fits/file.py in __init__(self, fil
173     self._open_fileobj(fileobj, mode, overwrite)
174     elif isinstance(fileobj, str):
--> 175     self._open_filename(fileobj, mode, overwrite)
176     else:
177     self._open_filelike(fileobj, mode, overwrite)

/ext/anaconda3/lib/python3.5/site-packages/astropy/io/fits/file.py in _open_filename(sel

```



```

515
516         if mode == 'ostream':
--> 517             self._overwrite_existing(overwrite, None, True)
518
519         if os.path.exists(self.name):

/ext/anaconda3/lib/python3.5/site-packages/astropy/io/fits/file.py in _overwrite_existing
412             os.remove(self.name)
413         else:
--> 414             raise OSError("File {!r} already exists.".format(self.name))
415
416     def _try_read_compressed(self, obj_or_name, magic, mode, ext=''):

```

OSError: File '../..../assignments/Session5/NGC3184_1/calib_data_1/ngc3184_h_240_1_reduced

In [9]: os.listdir(destination_dir_3184_1)

```

Out[9]: ['Bflatnorm1x1.fits',
        'Vflatnorm3x3.fits',
        'Rflatnorm2x2.fits',
        'bias3x3.fits',
        'Hflatnorm3x3.fits',
        'Bflatnorm2x2.fits',
        'Rflatnorm1x1.fits',
        'Iflatnorm3x3.fits',
        'dark3x3.fits',
        'Iflatnorm1x1.fits',
        'Rflatnorm3x3.fits',
        'Vflatnorm2x2.fits',
        'dark1x1.fits',
        'bias1x1.fits',
        'Hflatnorm1x1.fits',
        'dark2x2.fits',
        'Vflatnorm1x1.fits',
        'Iflatnorm2x2.fits',
        'Bflatnorm3x3.fits',
        'Hflatnorm2x2.fits',
        'ngc3184_h_240_1_reduced.fits',
        'bias2x2.fits']

```

```

In [2]: #reduction of images of ngc3184_2
        #assigning directories
        data_dir_3184_2 = '../..../assignments/Session5/NGC3184_2/raw_data_2'
        destination_dir_3184_2 = '../..../assignments/Session5/NGC3184_2/calib_data_2'

```

```

images_3184_2 = ccdproc.ImageFileCollection(location=data_dir_3184_2,keywords=['type','e
print(images_3184_2.summary)
calib_3184_2 = ccdproc.ImageFileCollection(location=destination_dir_3184_2,keywords=['ty
print(calib_3184_2.summary)

```

file	type	exposure	filter	xbinning
ngc3184_h_240_2.fits	SCIENCE	240.0	H	3
file	type	exposure	filter	xbinning
bias1x1.fits	BIAS	0.03	H	1
bias2x2.fits	BIAS	0.03	H	2
bias3x3.fits	BIAS	0.03	H	3
dark1x1.fits	DARK	60.0	H	1
dark2x2.fits	DARK	60.0	H	2
dark3x3.fits	DARK	60.0	H	3

```

In [7]: for ccd, fname in images_3184_2.ccds(return_fname=True, ccd_kwargs=dict(unit=u.adu)):
        dark, bias, flat = getCalibs(ccd)
        bias_subtracted = ccdproc.subtract_bias(ccd, bias)
        dark_subtracted = ccdproc.subtract_dark(bias_subtracted, dark, exposure_time='exposu
        reduced_image = ccdproc.flat_correct(dark_subtracted, flat)
        root, ext = os.path.splitext(fname)
        ofname = os.path.join(destination_dir_3184_2, root) + '_reduced.fits'
        reduced_image.write(ofname)

```

WARNING: FITSFixedWarning: RADECSYS= 'FK5 '

the RADECSYS keyword is deprecated, use RADESYSa. [astropy.wcs.wcs]

WARNING:astropy:FITSFixedWarning: RADECSYS= 'FK5 '

the RADECSYS keyword is deprecated, use RADESYSa.

```

-----
ValueError                                Traceback (most recent call last)

```

```

<ipython-input-7-652e23d1a9e0> in <module>()
    1 for ccd, fname in images_3184_2.ccds(return_fname=True, ccd_kwargs=dict(unit=u.adu))
    2     dark, bias, flat = getCalibs(ccd)
----> 3     bias_subtracted = ccdproc.subtract_bias(ccd, bias)
    4     dark_subtracted = ccdproc.subtract_dark(bias_subtracted, dark, exposure_time='ex
    5     reduced_image = ccdproc.flat_correct(dark_subtracted, flat)

```

```

/ext/anaconda3/lib/python3.5/site-packages/ccdproc/log_meta.py in wrapper(*args, **kwd)
    94     # Grab the logging keyword, if it is present.
    95     log_result = kwd.pop(_LOG_ARGUMENT, True)

```

```

---> 96         result = func(*args, **kwd)
      97
      98         if not log_result:

/ext/anaconda3/lib/python3.5/site-packages/ccdproc/core.py in subtract_bias(ccd, master)
572             "image".format(ccd.unit, master.unit))
573         else:
--> 574             raise e
      575
      576         result.meta = ccd.meta.copy()

/ext/anaconda3/lib/python3.5/site-packages/ccdproc/core.py in subtract_bias(ccd, master)
565
566     try:
--> 567         result = ccd.subtract(master)
      568     except ValueError as e:
      569         if 'operand units' in str(e):

/ext/anaconda3/lib/python3.5/site-packages/astropy/nddata/ccddata.py in inner(self, operand, operation, operand2, **kwargs)
47         _config_ccd_requires_unit = False
48         result = self._prepare_then_do_arithmetic(op, operand,
---> 49                                     operand2, **kwargs)
      50         # Wrap it again as CCDDData so it checks the final unit.
      51         _config_ccd_requires_unit = True

/ext/anaconda3/lib/python3.5/site-packages/astropy/nddata/mixins/ndarithmetic.py in _prepare_then_do_arithmetic(self, operation, operand, operand2, **kwargs)
612
613         # Now call the _arithmetics method to do the arithmetics.
--> 614         result, init_kwds = operand._arithmetic(operation, operand2, **kwargs)
      615
      616         # Return a new class based on the result

/ext/anaconda3/lib/python3.5/site-packages/astropy/nddata/mixins/ndarithmetic.py in _arithmetic(self, operation, operand, **kwargs)
244         # Then calculate the resulting data (which can but not needs to be a
245         # quantity)
--> 246         result = self._arithmetic_data(operation, operand, **kwargs2['data'])
      247
      248         # Determine the other properties

/ext/anaconda3/lib/python3.5/site-packages/astropy/nddata/mixins/ndarithmetic.py in _arithmetic_data(self, operation, operand, **kwargs)
320         else:
321             result = operation(self.data * self.unit,

```

```

--> 322                                operand.data * operand.unit)
    323
    324         return result

```

ValueError: operands could not be broadcast together with shapes (490,728) (736,1092)

```

In [2]: #reduction of images of ngc3184_3
        #assigning directories
        data_dir_3184_3 = '../..'/assignments/Session5/NGC3184_3/raw_data_3'
        destination_dir_3184_3 = '../..'/assignments/Session5/NGC3184_3/calib_data_3'

        images_3184_3 = ccdproc.ImageFileCollection(location=data_dir_3184_3,keywords=['type','e
        print(images_3184_3.summary)
        calib_3184_3 = ccdproc.ImageFileCollection(location=destination_dir_3184_3,keywords=['ty
        print(calib_3184_3.summary)

```

file	type	exposure	filter	xbinning
ngc3184_r_240_1.fits	SCIENCE	240.0	R	2
ngc3184_r_240_2.fits	SCIENCE	240.0	R	2
ngc3184_r_240_3.fits	SCIENCE	240.0	R	2

file	type	exposure	filter	xbinning
Bflatnorm1x1.fits	SKY	2.54617976223	B	1
Bflatnorm2x2.fits	SKY	2.54617976223	B	2
Bflatnorm3x3.fits	SKY	2.54617976223	B	3
Hflatnorm1x1.fits	SKY	21.1649156994	H	1
Hflatnorm2x2.fits	SKY	21.1649156994	H	2
Hflatnorm3x3.fits	SKY	21.1649156994	H	3
Iflatnorm1x1.fits	SKY	3.75051017342	I	1
Iflatnorm2x2.fits	SKY	3.75051017342	I	2
Iflatnorm3x3.fits	SKY	3.75051017342	I	3
Rflatnorm1x1.fits	SKY	2.20479702354	R	1
Rflatnorm2x2.fits	SKY	2.20479702354	R	2
Rflatnorm3x3.fits	SKY	2.20479702354	R	3
Vflatnorm1x1.fits	SKY	3.00486467614	V	1
Vflatnorm2x2.fits	SKY	3.00486467614	V	2
Vflatnorm3x3.fits	SKY	3.00486467614	V	3
bias1x1.fits	BIAS	0.03	H	1
bias2x2.fits	BIAS	0.03	H	2
bias3x3.fits	BIAS	0.03	H	3
dark1x1.fits	DARK	60.0	H	1
dark2x2.fits	DARK	60.0	H	2
dark3x3.fits	DARK	60.0	H	3

```

In [4]: for ccd, fname in images_3184_3.ccds(return_fname=True, ccd_kwargs=dict(unit=u.adu)):

```

```

dark, bias, flat = getCalibs(ccd)
bias_subtracted = ccdproc.subtract_bias(ccd, bias)
dark_subtracted = ccdproc.subtract_dark(bias_subtracted, dark, exposure_time='exposu
reduced_image = ccdproc.flat_correct(dark_subtracted, flat)
root, ext = os.path.splitext(fname)
ofname = os.path.join(destination_dir_3184_3, root) + '_reduced.fits'
reduced_image.write(ofname)

```

WARNING: FITSFixedWarning: RADECSYS= 'FK5 '
the RADECSYS keyword is deprecated, use RADESYSa. [astropy.wcs.wcs]
WARNING:astropy:FITSFixedWarning: RADECSYS= 'FK5 '
the RADECSYS keyword is deprecated, use RADESYSa.

```

In [11]: #loading the calibration object NGC6720
ccd_ngc6720 = fits.open('../assignments/Session5/NGC6720/ngc6720_h_30_1.fits')
ccd_data6720 = ccd_ngc6720[0].data
print(ccd_data6720.mean())

#fig, axes = plt.subplots(figsize=(8,8))
#Bins = 3000
#axes.hist(ccd_data6720.ravel(), Bins)
#axes.set_xlim(0,700)
#plt.show()

```

290.118368271

NameError Traceback (most recent call last)

```

<ipython-input-11-ddbc50053616> in <module>()
    3 ccd_data6720 = ccd_ngc6720[0].data
    4 print(ccd_data6720.mean())
----> 5 print(median(ccd_data6720))
    6 #fig, axes = plt.subplots(figsize=(8,8))
    7 #Bins = 3000

```

NameError: name 'median' is not defined

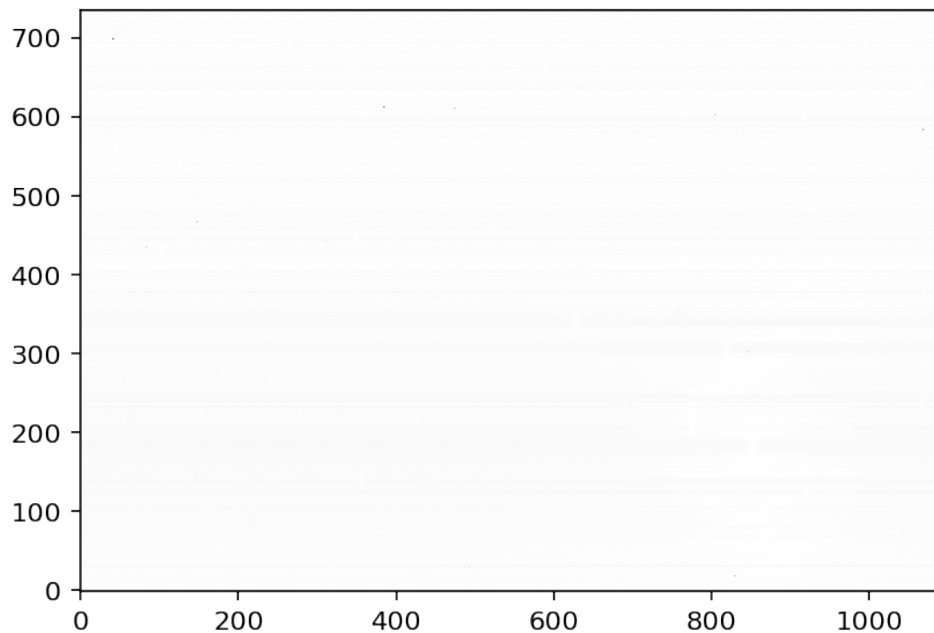
```

In [20]: #Scikit learn module used
from sklearn.preprocessing import normalize
#?normalize
normed_data = normalize(mean_r_data, axis=1, norm='max')
print(normed_data)

```

```
[[ 0.27124596  0.30585245  0.52528778 ...,  0.64014297  0.54018972
  0.50692661]
 [ 0.20176851  0.34586829  0.47074685 ...,  0.71502778  0.63944739
  0.41248747]
 [ 0.21111121  0.35209332  0.55122093 ...,  0.64131794  0.73883172
  0.61950779]
 ...,
 [ 0.31939159  0.31181076  0.55997743 ...,  0.85751808  0.67299692
  0.67654348]
 [ 0.41664443  0.35075512  0.64554666 ...,  0.64457353  0.88232864
  0.78651874]
 [ 0.40494603  0.3099271   0.76234955 ...,  0.58667189  0.53818709
  0.62189351]]
```

Out [20]:



```
In [4]: #Normalize the rows
def normalize_rows(a):
    row_sums = a.sum(axis=1)[:,np.newaxis]
    return a/row_sums

mean_normed_r_data = normalize_rows(mean_r_data)
print(mean_normed_r_data)

[[ 0.0003847  0.00043378  0.00074499 ...,  0.00090789  0.00076613
  0.00071895]
```

```
[ 0.0002778  0.0004762  0.00064814 ...,  0.00098447  0.00088041
 0.00056792]
[ 0.00030275  0.00050492  0.00079048 ...,  0.00091969  0.00105953
 0.00088841]
...,
[ 0.00045213  0.0004414  0.00079271 ...,  0.00121391  0.0009527
 0.00095772]
[ 0.00053014  0.0004463  0.0008214 ...,  0.00082016  0.00112268
 0.00100077]
[ 0.00057377  0.00043914  0.00108019 ...,  0.00083127  0.00076257
 0.00088117]]
```

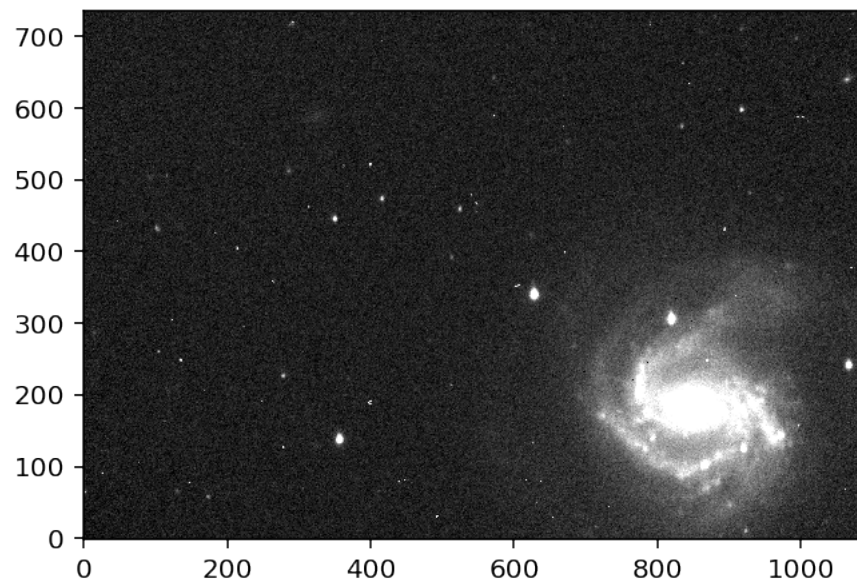
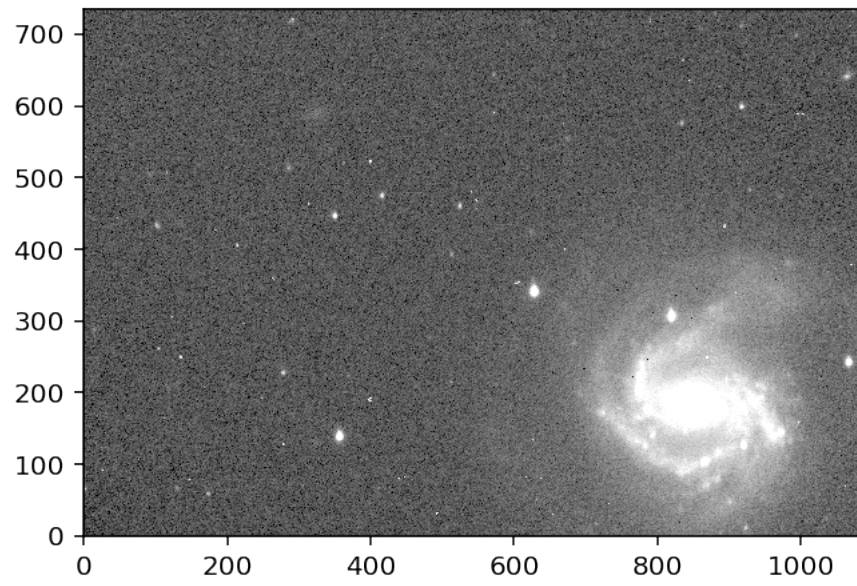
```
In [9]: #Normalizing the R-band image
        #matplotlib normalization
        mean_r_ccd = fits.open('../assignments/Session5/M61/calib_data/m61_mean_r.fits')
        #print(mean_r_ccd[0].data)
        mean_r_data = mean_r_ccd[0].data
        mean_r_header = mean_r_ccd[0].header

        #create a ImageNormalize object
        norm = ImageNormalize(mean_r_data, interval=MinMaxInterval(),stretch=SqrtStretch())

        fig,ax = plt.subplots(figsize=(8,8), nrows=2)
        vmin, vmax = np.percentile(mean_r_data,[1,99])
        implot = ax[0].imshow(mean_r_data, cmap='gray', origin='lower',vmin=vmin, vmax=vmax, nor
        implot1 = ax[1].imshow(mean_r_data, cmap='gray', origin='lower', vmin=vmin, vmax=vmax)
        plt.show()
        print(type(implot))

/ext/anaconda3/lib/python3.5/site-packages/astropy/visualization/stretch.py:112: RuntimeWarning:
  np.sqrt(values, out=values)
/ext/anaconda3/lib/python3.5/site-packages/matplotlib/colors.py:504: RuntimeWarning: invalid val
  xa[xa < 0] = -1
```

Out [9]:



```
<class 'matplotlib.image.AxesImage'>
```

```
AttributeError
```

```
Traceback (most recent call last)
```

```
<ipython-input-9-3765cbc575b4> in <module>()
    15 plt.show()
```



```

16 print(type(implot))
--> 17 print(implot.get_data())

```

AttributeError: 'AxesImage' object has no attribute 'get_data'

```

In [74]: #combining the images for R-band and h-alpha
         #destination_dir_align = '../assignments/Session5/M61/calib_aligned'
         #calib_images = ccdproc.ImageFileCollection(location=destination_dir_align,keywords=['e

         #R-band
         #hdu_r_1 = fits.open('../assignments/Session5/M61/calib_aligned/aligned_m61_r_240_1_
         #ccd_r_data_1 = fits.getdata('../assignments/Session5/M61/calib_aligned/aligned_m61_
         #ccd_r_data_2 = fits.getdata('../assignments/Session5/M61/calib_aligned/aligned_m61_

         """The mean combined r-band and h-alpha images are loaded.
            The r-band image is normalised using the maximum value of the counts in the image."""
         #os.listdir(destination_dir)
         #create a function to normalize each column of the data
         def normalize_data(data):
             rows,cols = data.shape
             for j in range(cols):
                 data[:,j] /= abs(data[:,j]).max()
                 j=j+1
             new_data = data
             return new_data
         mean_r_ccd = fits.open('../assignments/Session5/M61/calib_data/m61_mean_r.fits')
         #print(mean_r_ccd[0].data)
         mean_r_data = mean_r_ccd[0].data
         mean_r_header = mean_r_ccd[0].header

         print(mean_r_data)
         #filepath = os.path.join(destination_dir, 'm61_mean_r_norm.fits')
         #mean_r_ccd1 = fits.open('../assignments/Session5/M61/calib_data/m61_mean_r_norm.fits')
         #print(mean_r_ccd1[0].data)
         #mean_r_data1 = fits.getdata(filepath)

         #mean_r_norm = normalize_data(mean_r_data)
         #print(mean_r_norm)
         #hdu = fits.PrimaryHDU(data=mean_r_norm, header=mean_r_header)
         #hdu.writeto('m61_mean_r_norm.fits',clobber=True)

         fig,axis = plt.subplots(figsize=(8,8))
         Nbins=800
         #axis.hist(mean_r_data.ravel(),Nbins)
         #axis.set_xlim(-2,5)
         #axis.set_ylim(0,5000)

```

```

#imshow = axis[0].imshow(mean_r_norm,cmap='gray',origin='lower')
vmin, vmax = np.percentile(mean_r_data,[1,99])
imshow1 = axis.imshow(mean_r_data,cmap='gray',origin='lower',vmin=vmin,vmax=vmax)
axis.grid(False)
plt.show()

[[ 77.41653442  87.29360199 149.92282104 ..., 182.70373535
 154.17599487 144.68234253]
 [ 55.77303314  95.60522461 130.12426758 ..., 197.64862061
 176.75662231 114.02015686]
 [ 60.9186554 101.60072327 159.06137085 ..., 185.05993652
 213.19869995 178.76635742]
 ...,
 [ 85.8229599  83.78593445 150.47021484 ..., 230.42166138 180.8394165
 181.79240417]
 [101.39308167  85.35849762 157.09790039 ..., 156.86108398 214.7203064
 191.40435791]
 [109.49563599  83.80293274 206.13598633 ..., 158.6335144 145.5234375
 168.1572876 ]]

```

Out[74]:

