

Fig. 4. Left: VLT spectra of PG1159 stars in the vicinity of the Ne vII 3644Å line(indicated by the vertical shaded bar) and the O vI 3s–3p doublet at 3811/3834Å. Right: Expanded view centered around the Ne vII line. Overplotted are model spectra with Ne=2%, labeled with their $T_{\rm eff}$ (in kK) and log g. Smoothing of observed spectra with Gaussians (FWHM given): left panel: 0.4–0.7Å, right panel: no smoothing, except for Longmore 3 (0.2Å). The model spectra are smoothed accordingly. The NGC 246 model is also rotationally broadened (70 km/s).