# Special and General Relativity for the Educated Pedestrian

Virgil Elings, pedestrian

My introduction to special relativity was as an undergraduate engineering student. We were told that the speed of light was constant for all observers traveling at different velocities with respect to one another and then spent the rest of our time manipulating Lorentz transformations. Due to a combination of bad teaching and bad learning, I thought that moving clocks ran slow because it took time for the information to get back to the observer. Of course, this idea was wrong, but it was what I was left with. I will try to stay away from manipulating formulas but try to derive things as we go along (learning by thinking, I call it). I will try to do almost everything in special and general relativity using time dilation and length contraction. The aim is to get a good feel for relativity, without the complexity of tensors and all that fancy stuff the experts use. The serious reader needs to get out a pen and pencil and fill in some of the math I have left out.

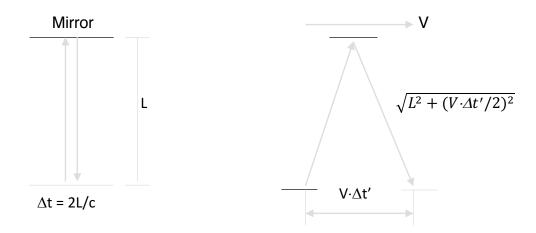
Einstein's assumptions for special relativity were that in systems traveling with constant velocity with respect to each other (no acceleration), they are equal and the physics inside each of the systems is the same. He assumed that the speed of light in free space, with respect all observers, is a constant. This assumption was not well received because it really screws up space and time. This assumption came from Maxwell's equations in which one solution showed a traveling electromagnetic wave traveling at a velocity which seemed to be close to that of the speed of light, but it didn't say in what system this speed was measured. There were also experiments which showed that the velocity of light from a star was the same whether you were traveling, in the earth orbit, toward the star or away from the star.

The assumption in general relativity was that acceleration and gravity are the same. We will get to that later.

Other than some basic classical mechanics, everything will be done using only these assumptions. No proper time, no geodesics, no tensors, and no Schwartzchild metric. If you don't know what these are, good, you won't need them. You are a pedestrian like me. Relativity is not my field and so I'm afraid there's a lot of approaches that I haven't looked at. I'm sure all this has been done before, probably in the same way. I apologize for not giving credit to people, I'm afraid, I don't know about.

### Special relativity

Let's start with the standard optical clock. This clock consists of an emitter/detector which emits pulses which go up to a mirror, spaced a distance L from the emitter, and then back down to the emitter/ detector where a new pulse is sent out, making one tick of the clock. There is an identical clock which is moving off to the right at a velocity V in another reference frame, which we will call the primed frame. So why the mirror? It turns out in relativity if you're going to measure a time interval, it is best to measure the start and end of the interval at the same place in whatever reference frame you are in. Time and space get mixed up so it is better if you can keep space out of it. We will see an example of this shortly.



What we want to calculate is the ratio of the time interval in the primed frame, as seen by an observer in the stationary frame, to the time interval in the stationary frame for one "tick" of the clocks. Now the velocity of light, c, is the same,  $3x10^8$  m/sec, in both systems whether the system is moving with a velocity to the right or not. In our stationary system, one tick is just 2L divided by c. In the primed frame one tick is clearly going to be longer because the path is now longer than 2L. During the time interval  $\Delta$  t' (one tick of the moving clock) the detector has moved a

distance V· $\Delta t'$ . The distance the pulse must travel is, therefore,  $2\sqrt{L^2 + (V\cdot \Delta t'/2)^2}$ .

So, 
$$\Delta t' = 2/c\sqrt{L^2 + (V \cdot \Delta t'/2)^2}$$
. With  $\Delta t = 2L/c$ , we get:

$$\Delta t' = \gamma \cdot \Delta t$$

 $\Delta t' = \gamma \cdot \Delta t$  where  $\gamma = 1/\sqrt{1-V^2/c^2}$  a parameter greater than or equal to 1. In our everyday world where velocities are not much greater than a few hundred meters per second, this parameter is essentially 1 and so we rarely see any of this time dilation (GPS is the glaring exception-more on this later).

The takeaway from this is that the clock in the primed (moving) system is ticking slower than the stationary clock; Moving clocks run slow! So, then we take a leap of

faith; time is slower in the moving system. By the way, if we were in the primed frame looking at the clock in the unprimed system, we would see it running slow by the same factor. Clocks moving by us run slower than our clock. Now Einstein said that without acceleration or gravity, these reference frames are equal. So, who is moving? The other guy, of course.

Back in my deep dark past I was part of an experiment at the Berkeley Bevatron to measure the lifetimes of the pi plus and pi minus mesons. If we could find a difference in the lifetimes, it would mean that a fundamental theory, CPT, was violated and therefore it would be a big deal. Alas (for us, but not for physics), we measured the lifetimes to be equal. The point of all this is that the pi mesons, which have an average lifetime at rest of  $2.6\times10^{-8}$  sec, traveling at the speed of light, without relativity, would only travel about 8 m on the average. But our spectrometer was about 30 m in length over which we measured the decay of the mesons. They were living about twice their "at rest" lifetime as they moved through the lab. Why? Moving clocks run slow and in our case the clock was the meson's lifetime and  $\gamma$  was about 2.

Now let's say that a particular pi meson decays halfway down the spectrometer. People in the lab frame and the meson in the meson's frame both agree where this happens, halfway down the spectrometer. But in the meson's frame the clock is not running slow and therefore the lifetime is the at-rest lifetime. In this case the meson can only see 8 m of spectrometer go by before it decays. How did it get to the 15 meter mark? It must be that the lab, which is moving by the meson, is contracted in length, otherwise the meson would decay when only 8 m of spectrometer had passed by. From the meson's point of view, the spectrometer is contracted by about a factor of 2 so that in the meson's normal at-rest lifetime, half of the spectrometer passes by. Objects in a reference frame moving by a stationary observer are length contracted in the direction of motion by  $\gamma$  i.e. L'= L/ $\gamma$ . This length contraction is only along the direction of motion. There is no transverse length contraction, as opposed to time dilation which occurs everywhere equally in the moving frame. So, the meson sees a distorted moving space where dimensions along the direction of the moving lab are contracted but the transverse dimensions are not. Time dilation and length contraction go hand-in-hand and the magnitude of both relativistic effects is the same.

With this length contraction idea, let's go back and revisit our optical clock, but this time we will lay it along the X axis, the direction of motion, so that the pulses travels along the X axis. One might think offhand that, due to the length contraction of the clock, it will actually take less time for the pulses to go to the mirror and back than it does in the stationary frame. If that is true, we are in deep trouble. The rate time passes in the primed frame should not depend on the orientation of our clock.

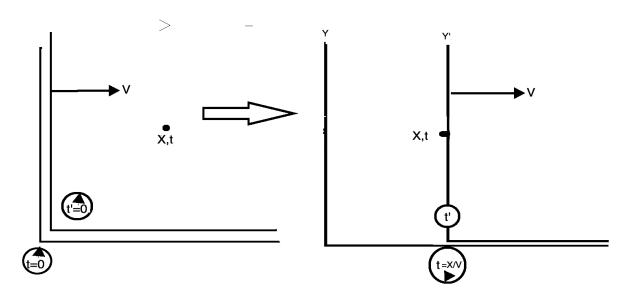
Let's break the round-trip of the path into two parts. The first part being from the emitter/detector to the mirror (time t1') and the second part from the mirror back to the emitter/detector (t2'). In the first part the mirror is moving away from the pulse, which is travelling at velocity c, so the distance traveled is  $(L/\gamma + V \cdot t1') = t1' \cdot c$  where the length of the clock has been length contracted to  $L/\gamma$ . We solve for t1' and get t1' =  $L/\gamma \cdot (c - V)$ . In the second part the emitter/detector is moving toward the pulse and so the distance traveled is  $(L/\gamma - V \cdot t2') = t2' \cdot c$  and so we get  $t2' = L/\gamma \cdot (c + V)$ . The total time for one tick of the moving clock is t1' + t2' which is  $t1' \cdot t2' \cdot$ 

(moving) frame, spaced a distance L apart, by placing a pulse emitter halfway between the clocks and sending out pulses of light which start the clocks running when the pulses arrive. In the primed frame the clocks are now synchronized. As seen from the stationary frame, the left clock is running into the pulse and therefore gets set before the right clock, which is moving away from the pulse, and so the left clock measures time which is ahead of the right clock as seen from the stationary frame. If we calculate how much it is ahead, it involves the same calculation we did for our clock which was laying down.

Lorentz Transformations Again, we consider two reference frames, one the primed frame moving with a velocity V to the right of the stationary frame. We have a point in the stationary frame with coordinates X and t and we now want to calculate the coordinates X' and t' in the primed frame. At time zero all clocks are synchronized in each of the respective frames and set to zero and the origins of both coordinate systems are superimposed.

Length Contraction First we will derive the Lorentz transformations, for transforming the space and time coordinates from a stationary frame to a moving frame, by simply using only length contraction. Classically we have, as seen from the stationary frame, X' = X - Vt. With relativity, the moving frame is length contracted and, therefore, we see, from the stationary frame, that X' is now  $X'/\gamma$ . So,  $X' = \gamma(X - Vt)$ , and looking from the primed frame we see an equal relation,  $X = \gamma(X' + Vt')$ , except that V is now in the other direction. Eliminating X' between these relations we get;  $t' = \gamma(t - X \cdot V/c^2)$  and, similarly,  $t = \gamma(t' + X' \cdot V/c^2)$ . The term  $X \cdot V/c^2$  means that the time on the X' axis, as seen from the stationary frame, depends on the position in the stationary frame. Therefore, the Lorenz transformation for transferring the coordinates in one frame to a moving frame are the following.

$$X' = \gamma(X - Vt)$$
 and  $t' = \gamma(t - X \cdot V/c^2)$   
 $X = \gamma(X' + Vt)$  and  $t = \gamma(t' + X' \cdot V/c^2)$ 



**Time Dilation** We will now derive the Lorentz transformations using only time dilation. We already found that it's dangerous to compare time on two clocks in the moving frame that are separated in space, so we will use only one clock in the moving frame. The clocks in the stationary frame all read the same time.

Consider a clock at the origin of the moving frame and count the time on this clock, as viewed from the stationary frame, from time zero until the Y' axis of the moving frame is coincident with the point of interest, X,t. We have,  $t' = t/\gamma$  (the moving clock is running slow).

Multiplying the top and bottom by  $\gamma$ , we get:

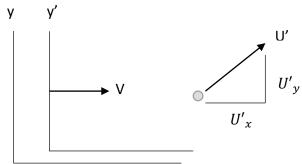
$$t' = \gamma/\gamma^2 \cdot t = \gamma t \cdot (1 - V^2/c^2)$$
, but  $t = X/V$  (V = X/t) so:  $t' = \gamma(t - X \cdot V/c^2)$ 

As before, the rest of the transformations follow from this. The Lorentz transformations can be derived from either length contraction or time dilation. Time and distance, they're like birds of a feather that flock together, or is it just one bird and we have double vision? Sort of, they are both results of the same thing, the bold assumption of Einstein's that the velocity of light is the same for all observers. A big difference is that length contraction is directional and time dilation is not. In general relativity, we will see that we have to use both to get the right answers.

# Velocity Addition

We now will calculate how to add velocities in relativity. We will do this two ways, one by thinking, and one using the Lorenz transformations. Let's do thinking first. Consider a particle moving in the primed frame with a velocity U' and with the frame moving with a velocity V in the X direction. What we want to calculate is the velocity U of the particle in the stationary frame ie. We want to add U' and V. This velocity U must satisfy several boundary conditions. Two are that

when U' or V increases to the velocity of light(c), U also goes to c and cannot go above it. Also, if V is very small compared to the velocity of light, we should get the classical result, which is just U' plus V. Whatever relation we derive, it must be symmetrical in U' and V.



Consider first only the x component of U', the component in the direction of V. In the low velocity case we have:

 $U_x = {U'}_x + V$ , but as V $\Rightarrow$ c  $U_x = \alpha \cdot ({U'}_x + c) \Rightarrow$ c where  $\alpha$  is some factor.  $\alpha$  therefore must have a form something like  $c/({U'}_x + c)$  but similarly must have the form c/(V + c) when  ${U'}_x \Rightarrow$ c. To include all these conditions and the symmetry between U and V, we wind up with  $\alpha = c/(c + V \cdot {U'}_x/c) = 1/(1 + V \cdot {U'}_x/c^2)$ .

So: 
$$U_x$$
 =(  ${U'}_x$ +V) /(1+V· ${U'}_x/c^2$ ).  $\Rightarrow$ c when V,  ${U'}_x$   $\Rightarrow$ c and  $\Rightarrow$  ${U'}_x$ +V when V/c  $\Rightarrow$ 0

Another way to do this is to calculate dX/dt from the two Lorentz transformations;  $X = \gamma(X' + Vt')$  and  $t = \gamma(t' + X' \cdot V/c^2)$  dX/dt =  $(dX' + Vdt')/(dt' + dX'V/c^2)$  Dividing by dt' top and bottom we get, as before

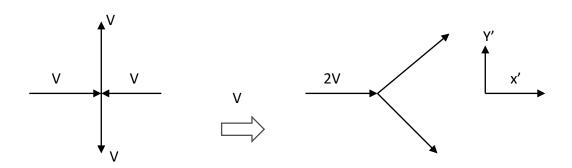
$$dX/dt = U_x = (U'_x + V) / (1 + V \cdot U'_x/c^2).$$

Now consider the Y component of U. You might assume that the velocity in the Y direction should not be affected by a transformation between two coordinate systems which have a relative velocity in the X direction. Let's see how that works out. Consider a particle travelling only along the Y 'axis. When transformed to the stationary system, we will take the Y velocity to be the same and the X velocity will just be V ( $U'_x$ =0). The resultant velocity will be  $\sqrt{U'_y{}^2 + V^2}$ . Well, you can see the problem. As V  $\Rightarrow$ c, this resultant velocity is clearly going to be larger than the velocity of light, and that is a no-no. The Y velocity measured in the stationary frame,  $U_y$ , must be smaller than the Y' velocity measured in the primed frame,  $U'_y{}$ , to avoid this problem. Let's take  $U_y{}$  to be  $\kappa \cdot U'_y{}$  and see if we can figure out  $\kappa$ . We have  $U = \sqrt{(\kappa \cdot U'_y{})^2 + V^2}$ . If this is to be c as V goes to c, we have  $U^2 = \kappa^2 \cdot U'_y{}^2 + c^2 = c^2$ . So  $\kappa$  must go to zero as V goes to c. Let  $U'_y{}$  go to c and the result must be the same.  $\kappa^2 c^2 + V^2 = c^2$ . Or:  $\kappa^2 = 1 - V^2/c^2 = 1/\gamma^2$ . Now if time were running slower in the primed frame, by the factor  $\gamma{}$ , then dY/dt would be slower than

dY'/dt' (velocity is inverse time). But, this is true because we already knew that time runs slower in the moving frame. Apparently, you can derive time dilation by only insisting that velocities do not exceed c, and therefore derive all of the Lorentz transformations as we did before. We assumed that Y = Y' but we did that for the optical clock that we used at the beginning.

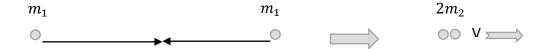
## Conservation of Momentum

So far, special relativity has made a mess out of space and time. Let's see if we can save Newton's most famous law of motion, action equals reaction, also known as conservation of momentum. Some people think that F = ma Is his most famous law of motion but to me it is just the definition of force, and besides the force is measured in Newtons! Consider the collision of two particles of mass m and velocity V travelling in opposite directions, colliding, and then going off at 90° with velocity V. In this frame, momentum is conserved because both the X and Y components of momentum are zero both before and after the collision. Call this the primed frame. We now let this frame move to the right with velocity V and observe the collision in a stationary frame. Is momentum still conserved?



Let's consider the classical case first. In this case we have a particle moving with the velocity 2V striking a stationary particle and then the particles go off at a 45° angle with each particle having an X component of velocity V . We have an incoming X momentum of  $2V \cdot m$  and outgoing momentum of  $mV \cdot 2$ . The total Y momentum is still zero, and so classically momentum is conserved in the collision. Let's now use our velocity addition formula to see what relativity thinks of this. The incoming velocity is not 2V but rather  $2V/(1+V^2/c^2)$  which is less than 2. The X component of momentum after the collision is still  $2 \cdot mV$ , and so momentum is  $\underline{NOT}$  conserved. To make this work we need to have the mass of the incoming particle be larger than either of the masses of the 2 outgoing particles. The likely suspect is, of course, to multiply the rest mass of the particles by  $\gamma$  so that particles with a higher velocity would have a larger mass. This would also be convenient because it would mean that the transverse(y) momentum would stay constant when we transform between the moving frame and the stationary frame, the larger mass making up for the smaller transverse velocity. If you do this fudge to the mass by multiplying the rest masses by  $\gamma$ , then you'll find out, if you do the algebra, that in this example

the momentum is conserved. Try not to make any mistakes because the algebra is sort of a mess. And be sure that the mass fudge factor involves the TOTAL velocity of the particle.



Let's take a simpler example, an inelastic collision where we have two particles of mass  $m_1$ travelling with equal and opposite velocities V, colliding with each other and sticking together, making a mass  $2 \cdot m_2$ . We call the masses different, because we know where this is headed. Momentum is conserved in this frame because both the X and Y total momentum are zero before and after the collision. Let's again move this frame off to the left with a velocity V so that in a stationary frame we now have only one particle moving with a velocity given by relativity of  $2V/(1+V^2/c^2)$  which we call  $V_{in}$ . After the collision, we have a particle of mass  $2 \cdot m_2$  moving to the right with a velocity V. Conserving momentum, we have:

$$2 \cdot m_1 \cdot V_{in} / \sqrt{1 - V_{in}^2} / c^2 = 2m_2 \cdot V / \sqrt{1 - V^2} / c^2$$

you can do the algebra (there's a lot of cancellation) and what you end up with is that the <u>rest</u> mass  $m_2$  is NOT equal to the rest mass  $m_1$  but rather:

 $m_2=m_1/\sqrt{1-V^2}/c^2$  this is the REST MASS of  $m_2$ , not the moving mass! We now have a problem. We either give up our idea of conservation of momentum or we need to do something about this thing we call mass. We will stick with conservation of momentum. Classically the loss in kinetic energy goes into heat of the final particles that are stuck together. Let's see what this has to do with the increase in rest mass. In the stationary system the kinetic energy before the collision is  $1/2 \ m \cdot (2 \ V)^2$  and after the collision is  $2 \cdot 1/2 m V^2$  so the loss in kinetic energy in the collision is, classically,  $1/2 \ m \cdot (2 \ V)^2 - 2 \cdot 1/2 m V^2 = m V^2$  since the input and output masses are the same. To first order in  $V^2/c^2$  we have, from above,  $2m_2 = 2m_1(1+1/2(V^2/c^2)) = 2m_1 + m_1 V^2/c^2$ , or in English; the outgoing particles (which are stuck together) have a total rest mass equal to the two incoming particle's rest mass + the loss in kinetic energy in the collision divided by  $c^2$ . This collision has turned kinetic energy into mass! Let's see what it looks like in the moving frame where it is easier to analyze things. Here we have two particles headed toward each other with the kinetic Energy of  $1/2 \cdot m V^2$ . The final kinetic energy is zero. If this is turned into mass, then we get an increase in the rest mass/ $c^2$  of  $m V^2$ , the same as see from the stationary frame.

We need to be careful. We may be adding apples and oranges. We combined "mass is energy", a relativistic concept, with kinetic energy, a Newtonian concept. It looks like the fudge factor gamma which we multiplied the mass with to make conservation of momentum work, may be a way to include in the mass, it's energy of motion. Rest energy  $(mc^2)$ + energy of motion =  $\gamma mc^2 = E$ . This certainly works where V/c is much less than one. We get E=  $\gamma mc^2 \cong (1+1/2 \cdot V^2/c^2) mc^2 = mc^2 + 1/2 \cdot mV^2$ . It also works at higher velocities.

With total energy  $E = \gamma mc^2$  and momentum  $P = \gamma mV$ , we have that  $E = Pc^2/V$ . For a particle travelling at the speed of light (a photon) this gives E = Pc. In the limit of V going to 0 (P=mV), we have  $E = mc^2$ , the rest energy. Everything is consistent.

# Work, energy and conservation of energy

We can derive the mass-energy relationship by using the age-old concept of work (force times distance) equals the increase in energy. We have  $dE = Fdx = dP/dt \cdot dx = dP/dV \cdot VdV$ . Integrating from 0 to V we have:

Change in Energy =  $\Delta E = \int_0^V dP/dV \cdot V dV$ .

The difference with relativity is that P involves  $\gamma$  and so the integral becomes more involved. You can convince yourself that dP/dV is just  $\gamma^3 m$  and that the integral is  $\gamma mc^2$ ]  $_0^V = \gamma mc^2 - mc^2$ .

The increase in energy from rest to velocity V is the total energy  $(\gamma mc^2)$  minus the rest energy  $(mc^2)$ .

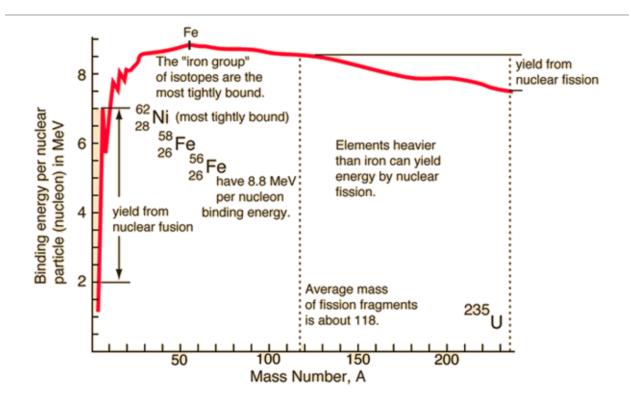
Since energy is now mass, then our concept of conservation of energy is the conservation of total mass,  $\gamma mc^2$ . Let's try this in our inelastic collision example. In the moving frame, the total mass before the collision is just  $2\gamma m$  and therefore that is the mass after the collision, which is just the rest mass of the final particles stuck together. This is the same answer we got when we used conservation of momentum.

Are there other examples of turning energy into mass? Of course, that's what they do at particle accelerators. I mentioned before the measurement of  $\pi^+$  and  $\pi^-$  meson lifetimes. But if you want to measure these lifetimes, first you need to make the mesons. This is done with a proton-proton inelastic collision,  $p+p \Longrightarrow p+p+\pi^++\pi^-$ , where the first proton comes out of a proton accelerator and the second proton is stationary in a liquid hydrogen target. In the center of mass system, it is clear that the threshhold kinetic Energy of each proton must be such to create the rest energy of one pi meson, about 135 Mev (million electron volts). So, in the center of mass we have  $\gamma m_p = m_p + m_\pi$  for each proton. Transforming to the laboratory frame where one of the protons is stationary, using our velocity transformation formula and transforming to the laboratory system, we get that the total energy of the proton from the accelerator is  $E = \gamma' m_p = m_p + 4m_\pi + 2m_\pi^2/m_p$ . The kinetic energy required,  $4m_\pi + 2m_\pi^2/m_p$ , is over twice that needed in the center of mass frame. Not very efficient. If we were making a proton- antiproton pair, each with a rest energy of 1000 Mev, instead of a pi meson pair in the collision, then the kinetic energy required in the lab frame would be three times that required in the center of mass. This ratio gets bigger the larger the masses that are created in the collision.

In order to create more massive exotic particles, the problem was solved by making the center of mass frame and the lab frame the same using colliding beam accelerators. The first proton colliding beam accelerator was the Intersecting Storage Rings (ISR) built in Geneva in 1971. This machine stored protons from a synchrotron into two counter-rotating beams and then steered them with magnets to collide with each other, making the lab frame the center of mass frame. The current(2017) most powerful colliding beam accelerator is the large hadron collider (LHC), also at the CERN lab in Geneva. This has two counter rotating beams each with kinetic energy of

6.5 TeV, yes, that is with a T. This accelerator is huge and has a diameter of 27 km. Yes, that's kilometers.

We've shown how mass can be created with kinetic energy, so how about the reverse where we create kinetic energy from mass. Well, it happens every day in your automobile. In this case carbon and oxygen combine to make  $CO_2$  plus 4.1 eV of energy. The reactants contain a total of 44 nuclei which have a total rest energy of about 44 billion electron volts. It would take a very good scale to notice that 4.1 eV of rest energy is missing in the final product. This reaction doesn't convert much of the rest energy into kinetic energy. These atomic reactions involve energies of the order of electron volts, the binding energy of electrons in atoms. Nuclear reactions, on the other hand, involve energies of millions of electron volts, the binding energy of nucleons in the nucleus. The method would be to take one or more nuclei with a small binding energy and somehow convert them into one or more nuclei that have a much higher binding energy. This binding energy might be thought of as a negative energy. If two things bind together then it takes energy to separate them so therefore when they bind together, they give off kinetic energy equal to the binding energy. From the diagram below we see that the most tightly bound nuclei are the iron group and so we could create kinetic energy by taking something very heavy, like uranium, and splitting it to make smaller nuclei or we could take very light nuclei like hydrogen and deuterium and fuse them together to make heavier nuclei like helium. The first is easier to do but unfortunately some of the products in the fission reaction are very radioactive and last for tens of decades. The second fusion process is very hard to do because you need to get the light nuclei together to react, but they are repelled by the Coulomb force between the positively charged nuclei. I spent one summer at the Oak Ridge National Laboratory and had a chance to tour the fusion facility. They told me that they were just turning the corner to success. The problem is, that was 1961. It has been a big corner but progress is still being made.



This brings special relativity to an end. Let's see what we have learned based on just two bold assumptions, one that the speed of light is the same for all observers and two, that reference frames travelling at a constant velocity with respect one another are equal.

- 1. Clocks moving with respect to us run slow
- 2. Objects moving with respect to us are length contracted.
- 3. Velocities don't add as easily as they used to.
- 4. Momentum can only be conserved if mass increases as a function of velocity.
- 5. The mass of a body contains both its rest energy and kinetic energy.
- 6. Conservation of energy is conservation of mass.
- 7. Energy can be turned into mass.
- 8. Mass can be turned into energy.

Quite a bit from simple assumptions. If this all seems strange, stranger things are yet to come.

# General relativity

In special relativity, we dealt with reference frames which we call inertial frames, they travelled with constant velocity and had no acceleration. Einstein wanted to include acceleration.

One of the results of classical mechanics was that the mass attracted by gravity (weight) was exactly the same as the mass involved in collisions, that is, the mass involved in acceleration (F=ma, if you will). Could it be that gravity and acceleration are the same thing? It is a little hard to see how gravity, this mystical force from a massive body, could be the same as acceleration which has to do with motion. They don't seem to be the same. Einstein said they are the same, or I think that is what he said. Actually, this is not exactly a new concept. In classical mechanics, we go into a rotating coordinate system, such as a merry-go-round, and then replace what was centripetal acceleration in the stationary frame with a mystical force called centrifugal force in the opposite direction. Acceleration replaced with a force. For gravity, we are going to replace a force with acceleration.

The question is, does gravity affect time and distance like velocity does in special relativity? Our approach to general relativity is going to be quite simple so that even I can understand it. Our approach is robust enough that we can reproduce Einstein's most famous calculations, the bending of light by gravity and the gravitational effect on the precession of the orbit of mercury. Our approach avoids a lot of the formalism in general relativity and so therefore our approach is not the end all, but is simple enough that we can get a good feel for general relativity. Oddly enough, one of the problems is that It gives the right answer when it shouldn't.

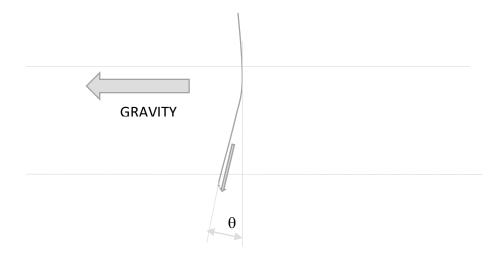
We are going to analyze a reference frame which has gravity from a massive body, like the sun, and then cancel gravity with acceleration so we can analyze the frame from the viewpoint of motion. Getting rid of gravity is not that hard, it's done every day in the space station. The occupants accelerate toward the earth just the right amount to eliminate the force of gravity. They are in freefall. Einstein's idea was an elevator in freefall, but that was before the space station. A reference frame in freefall is an inertial frame, there is no gravity and no acceleration, so we can use special relativity to analyze other frames like the "gravity" frame. From the standpoint of this freefall frame, it is the rest of the world that is accelerating past it, moving faster and faster as time goes by. But remember, we have simply replaced gravity with acceleration. As the world moves by, we see that three things happen in that world. The first is that if there were a row of clocks stretching to the sun, as the clocks go by they are running slow with respect our clock, but the closer the sun gets to us (or we get to the sun) the slower they run. Time dilation causes clocks nearer the source of gravity (the sun) to run slower. The second thing is that if there was a line of rulers stretching to the sun, we would notice that the rulers would get shorter and shorter the closer we got to the sun due to length contraction. There is no length contraction perpendicular to the direction of motion and therefore space becomes warped. We now have that the circumference around the source of gravity is no longer two pi times the radius, but is greater than that. Space is no longer "flat". The third thing, which is disturbing to me, is that as the sun gets closer and closer to us, its' mass increases because its' velocity is greater and greater. If mass is energy, where did the energy come from?

Sorry to disappoint, but that's about all there is to our simple approach to general relativity. Let's now get on to the calculations using our simple approach and see how they agree with measurements. We, of course, will have to get into the details and that requires some grinding.

# The bending of light by gravity

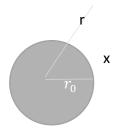
### First a classical calculation

Before we get into the relativity calculation, let's calculate classically the deflection of a massive particle, travelling at the speed of light, passing near the surface of the sun. This is a classical calculation, so a particle travelling at the speed of light is OK. This is a calculation that could've been made by Newton, because he thought light was a particle, but apparently, he never made it. Consider the imaginary gravitational field shown below which is directed toward the left. As a particle passes through this field, the angle of deflection, for small angles, is simply the amount of horizontal momentum given to the particle by the gravitational force, divided by the total momentum of the particle.  $\theta = \Delta P_t/P$  where  $\Delta P_t$  is the change in transverse momentum.



We will now calculate this deflection for a particle passing close to the surface of the sun at a distance  $r_0$  from its center. Since the deflection is small, there is no reason to worry about the shape of the path. We will use just a straight path and calculate the horizontal momentum given to the particle. You can pretend the deflection is there, but it's so small you can't see it in the diagram. We integrate half the path.

$$\Delta P_t = \int_0^\infty dP = \int_0^\infty dP/dt \cdot dt/dx \cdot dx = \int_0^\infty F_t/\text{c-dx}$$
 
$$F_t = \text{mMG}/r^2 \cdot \cos\theta \text{ where } \cos\theta = r_0/r$$
 and 
$$r = \sqrt{x^2 + r_0^2}$$
 m is the mass of the particle, M is the mass of



the sun and G is the gravitational constant.

$$\Delta P_t/\mathrm{P} = \mathrm{mMG/m}c^2 \cdot \int_0^\infty r_0/r^3 \cdot dx$$
 =  $\mathrm{MG/c^2} \int_0^\infty r_0/\sqrt[3/2]{x^2 + r_0^2} = \mathrm{MG/r_0}c^2$  We must multiply by 2 since we only integrated half of the path, so the total deflection is  $2\theta = 2\mathrm{MG/r_0}c^2$ 

The product 2MG/ $c^2$  is called the Swartzchild radius,  $r_s$ ,

and it will come up again and again. The deflection, then, is  $r_{\rm s}/r_0$ , a dimensionless number. This angle, which is only 0.87 arc seconds, or about  $4\cdot 10^{-6}$  radians, is half the correct angle for deflection of light by the sun. So, we need to do better.

#### A word of caution

Given the parameters at hand; the mass of the sun M, the gravitational constant G, the radius of the sun  $r_0$ , and the speed of light c, we need to construct a dimensionless parameter, because the angle of deflection is dimensionless. The only dimensionless parameter that you can create is  $\mathrm{MG}/r_0c^2$ , or powers of it. So, no matter what your theory, even if it's a crackpot theory (maybe like this one), but you keep track of units, you are going to get "something" times this parameter. The important point is, what is the "something"?

Einstein's first calculation of the bending of light - the dropping of a photon Einstein's first idea for the bending of light by gravity was in 1911 and the idea was to give the photon a mass and then see how this would affect the bending of light. From special relativity, we know that energy is equal to mass and so we can set the mass equal to  $E/c^2$ . The easy way out in the above calculation would be to replace the mass of the particle with  $E/c^2$  and replace the momentum P with E/c.

$$\Delta P_t/P = E/c^2MG/E \cdot \int_0^\infty r_0/r^3 \cdot dx = MG/c^2 \cdot \int_0^\infty r_0/r^3 \cdot dx = MG/r_0c^2$$

The energy cancels out and you, of course, get the same answer as the classical calculation. But Einstein took a different approach. He took the energy of the photon to be hf (why not, he discovered that) where h is Plank's constant and f is the frequency of the photon, and the mass therefore is  $m = E/c^2 = hf/c^2$ . As the photon falls in the gravitational field, the energy would increase and therefore the frequency would increase. Since the speed of light is constant, this means that the wavelength would decrease. If the frequency is  $f_0$  at infinity then, as the photon falls, we get:

 $E = E_0 + mMG/r = hf_0 + hf_0 MG/rc^2 = hf_0 (1 + MG/rc^2)$  where r is the distance to the center of the sun. Similarly we have:  $\lambda = \lambda_0 (1 - MG/rc^2)$  where we have used, and will use a lot,  $(1 + x)^n = 1 + nx$  for  $x \ll 1$ , for all values of n. Remember,  $MG/rc^2$  is very small, even for r as small as the radius of the sun. Einstein then used the concept that light travels from one point to another on the path of the minimum number of wavelengths and proceeded to calculate

the angle of bending. If you have not heard about the path of minimum number of wavelengths, let me give you my thoughts and use some crude diagrams

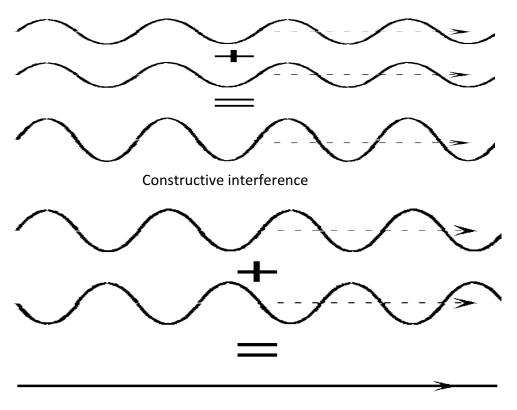
### The assumptions:

### 1. Light is a wave

2. Light can take any and all paths from A to B.

I am going to blame this second assumption on Richard Feynman, which is where I think I ran across it (summing over paths). Maybe, though, it was Huygens several centuries ago. By the way, you can do essentially all of optics using only these two assumptions. That's a whole different writeup.

Now people usually think light travels in a straight line, and so what is this business about taking any path? But light is a wave and, as you know, two waves out of phase can cancel each other, so it may turn out that a lot of these paths cancel each other.

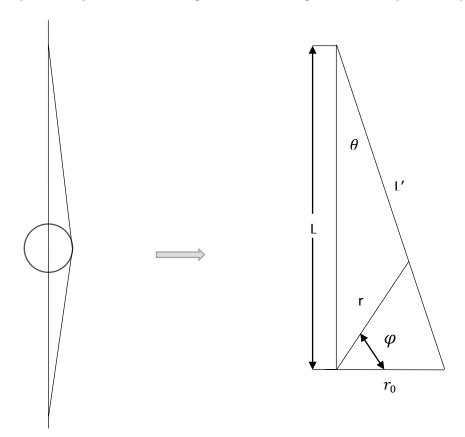


Destructive interference

If we take the long path from A to B, there is always a shorter path, shorter by one half a wavelength, that cancels that path and so on down until we get to the path of the minimum number of wavelengths from A to B. When we get to that point, the paths can no longer cancel each other and around one quarter of a wavelength of the path of minimum wavelengths is the path on which the light travels. Absent other things, this path will be the shortest distance between A and B, a straight line. If the wavelength varies in space, as in this case, then the path

will not be a straight line. A similar thing happens when light goes through a medium with an index of refraction that varies in space and the light does not follow a straight line. Graded index lenses work this way.

The path is shown below. The bending is small, so there is no reason to worry about the exact shape of the path. We just use two straight lines but integrate over only half the path.



We now calculate the average wavelength over the path L' by integrating over  $\varphi$  from 0 to  $\pi/2$ , with dL' =  $r/\cos(\varphi - \theta)$  d $\varphi$ . Remember,  $\lambda = \lambda_0(1 - MG/rc^2)$ 

$$\lambda_{avg} = \lambda_0 / L \cdot \int_0^{\pi/2} (1 - a/r) dL'$$
 where a is MG/ $c^2$  this gives:

$$\lambda_{avg} = \lambda_0 - \lambda_0 \int_0^{\pi/2} 1/\cos(\varphi - \theta) \mathrm{d}\varphi = \lambda_0 (1 - \mathrm{a/L'} \cdot \ln(\tan(\varphi/2 - \theta/2 + \pi/4))_0^{\pi/2}$$

We need to be careful at the upper limit because the tangent is very large. For small  $\theta$ ,  $\tan(\pi/2-\theta/2)$  is just  $2/\theta$ . So, for  $\lambda_{avg}$  we get:

$$\lambda_{avg} = \lambda_0 (1 - a/L' \cdot (ln2/\theta - ln(tan(\pi/4 - \theta/2)))$$

the first In term is large and sensitive to  $\theta$  whereas the second In term is small, essentially In1 which is zero. Notice that all of the "action" is at the upper limit of the integral. To first order we have:

$$\lambda_{avg} = \lambda_0 (1 - a \cdot \ln(2/\theta)/L')$$

This average wavelength is less than the wavelength far from the sun,  $\lambda_0$ , because the closer we get to the sun, the higher the frequency of the photon, therefore the shorter the wavelength. Notice that  $\lambda_{ava}$  increases as  $\theta$  increases.

We now want to vary  $\theta$  to find the value of  $\theta$  when the path is the minimum number of wavelengths. This path will favor distances away from the sun because there the average wavelength is longer. The number of wavelengths on the path L' is L'/ $\lambda_{avg}$ . We minimize this by taking the derivative wrt  $\theta$  and set it equal to zero.

 $\#\lambda = L'/\lambda_{avg}$  and  $d\#\lambda/d\theta = dL'/d\theta/\lambda_{avg} - \lambda_0 a/\theta/\lambda_{avg}^2 = 0$ . Notice that the 2 in the ln makes no difference in the derivative.

Now L' = L/cos( $\theta$ ) = L(1+ $\theta^2$ /2) so, dL'/d $\theta$  = L $\theta$ : to first order we can use  $\lambda_0$  for  $\lambda_{avg}$  in the denominators

So: 
$$L\theta - a/\theta = 0$$
 or, since  $L\theta = r_0$ ,  $\theta = a/r_0$ 

total deflection = 
$$2\theta = 2a/r_0$$
 = 2 MG/  $r_0c^2 = r_s/r_0$ 

Seems like a lot of work to get the same answer, but a step in the right direction. No time dilation or length contraction though. The dropping of a photon was basically a shot at time dilation. The effect is the same but what is really going on is that the frequency of the photon is increasing as it drops toward the sun because it is entering regions of slower and slower time. The longer the seconds, the more cycles per second. This dropping of a photon is used to describe things like the gravitational red shift and the rate of clocks that are orbiting the earth. In the red shift, one says that the photon loses energy, and therefore gains wavelength, as it comes out of the gravitational pull of a distant star where it was created. The real explanation is that the photon was created in a region of slow time, that is, in the gravitational field of the distant star, and has now entered a region of faster time, where we see it. Shorter seconds means lower frequency.

Let me relay a personal story. I was having lunch with my friend Ray Sawyer, who is a theoretical physicist (as you can see, I am not), and somehow general relativity came up in the discussion. Ray said he wished people would quit teaching general relativity as the "phony" dropping of a photon. I was embarrassed to tell Ray what I had been teaching in a course, called "Thinking out of the Box", because it was exactly that (too far out of the box, apparently). This made me curious

to understand more about general relativity but avoiding the formalization, which I would have trouble understanding. So, what is phony? In the dropping of a photon the clocks all run locally at the same rate, but if you compare a clock higher up in the gravitational field with yours, it "appears" to be running faster because photons coming from it have a higher frequency than your similar photons. The real reason is just that time runs faster "up there" (time dilation). Now what about length contraction? If we include that, you can guess where the correct answer for the bending of light is going to come from. So, let's do this thing with time dilation and length contraction.

## General relativity and the bending of light

In our freefall exercise we found three things going on in the "gravity frame" compared to our inertial (freefall) frame. 1. The time intervals were expanded by  $\gamma$  (time dilation), 2. Lengths in the radial (but not tangential) direction were contracted by  $\gamma$  and, 3. the mass of the sun was increased by  $\gamma$ , where  $\gamma = 1/\sqrt{1-V^2/c^2}$  and V is the freefall velocity.

For V 
$$\ll$$
 c, we have  $\gamma = 1 + 1/2 \cdot V^2/c^2$ .

The freefall velocity, by the time we get to the surface of the sun (or it gets to us), will be equal to the escape velocity from the surface of the sun (freefall run backwards) which is about 620 km/sec, 1/500 the speed of light, so we can use the approximation that  $V^2/c^2 \ll 1$ . Our freefall velocity squared, starting with zero at infinity, will be 2MG/r at a distance r from the center of the sun. The relativistic corrections, therefore, to time, radial length, and mass of the sun, as seen in the gravitational field are, proportionately,

Relativistic corrections = $\pm MG/rc^2$ .

Let's see how small the radius r can be before our approximations start to fall apart. If we take, arbitrarily,  $V^2/c^2$  to be less than 1/3, then we should keep the radius larger than 6MG/ $c^2$ =  $3r_s$ . We want to look at the effects of relativity as seen by someone stationary wrt the sun but at a distance where there is very little gravitational field, an inertial frame. We call this the distant observer (like us here on earth looking at the bending of light passing near the sun).

We will look at the effect of time dilation with no length contraction, and then we will look at the effect of length contraction with no time dilation, and then add the results. You may say that we should multiply the results, but for a small effect, multiplying and adding are the same. For a, b,  $c \ll 1$ ,  $(1+a)\cdot(1+b)/(1+c) \cong 1+a+b-c$ . We will do a lot of this when we get to the orbit of mercury.

#### Time Dilation

With time dilation, the frequency of the light will be greater as it passes nearer the sun. Since the velocity of light is constant, the wavelength of light will be shorter near the sun. We have:  $\lambda = \lambda_0 (1 - \text{MG/rc}^2)$  where  $\lambda_0$  is the wavelength at infinity, just as in the case of dropping a photon. The rest of the calculation to find the path of the minimum number of wavelengths is the same as what we have already calculated and therefore the answer is that the angle of deflection is

 $2\theta = 2MG/r_0c^2 = r_s/r_0$ , deflection caused by time dilation

### Length Contraction

Our distant observer now sees dimensions near the sun to be distorted somewhat like the map of the northern hemisphere shown on a flat page. Let's say you had such a map with distorted grid of lines 100 km apart and were asked to calculate the shortest path from say Paris to New York. You would calculate, in your flat world, a curved path which favors the northern latitudes where the kilometers in the distorted grid are quite large. Using the dimensions distorted by gravity, we want to calculate the path of minimum length(distance), which is the path that light will take (more on that later). This path will not be a straight line, somewhat like the curved airline routes shown on a flat map of the northern hemisphere. This calculation will be like our time dilation integral, calculating the path of the minimum number of wavelengths, except with one difference, only the radial dimensions are distorted, not the tangential dimensions. The bending is quite small and so we choose the "curved" path to be two straight lines as shown in the previous figure with only  $\theta$  as a variable. The integrations will be done in the flat space of the distant observer.

Consider a meter stick laying along L'. The radial component of this meter stick is  $m_0 \cdot \sin(\varphi - \theta)$  and the tangential component is  $m_0 \cdot \cos(\varphi - \theta)$ , essential all radial component at large r and all tangential component small r.  $m_0$  is the length of a meter stick with no gravity around i.e. in the frame of the distant observer. Let's calculate the length of this meter stick with distorted dimensions. The length contraction is only in the radial direction, in the amount a/r where, as before, a=MG/ $c^2$ . Therefore:

$$m = m_0 \cdot \sqrt{\left\{\sin^2(\varphi - \theta) \cdot (1 - a/r)^2 + \cos^2(\varphi - \theta)\right\}}$$

where m is the length of the meter stick in the distorted frame. This all reduces, in first order, to  $m = m_0 \cdot (1-a/r \cdot sin^2(\varphi-\theta))$ . We now integrate along L' to get the average length of a meter along that path under the influence of gravity. The integral is similar to what we have already done for wavelength except for the  $sin^2(\varphi-\theta)$ .

$$\begin{split} & m_{avg} = m_0 / L' \int_0^{\pi/2} (1 - a/r \cdot sin^2(\varphi - \theta)) dL' \\ &= m_0 / L' \int_0^{L'} dL' - \int_0^{\pi/2} (a \ sin^2(\varphi - \theta)) d\varphi / cos(\varphi - \theta) \end{split}$$

The first term is L' and the second term is, after another trip to the Integral Tables,

a[ 
$$\sin(\varphi-\theta)$$
 -  $\ln(\tan(\varphi/2-\theta/2+\pi/4)) \frac{\pi/2}{0} =$  a[(1+  $\theta$ ) -  $\ln(2/\theta)$ -  $\ln 1$ ].  
So  $m_{avg} = m_0[1 - a/L' \cdot (-(1+\theta) + \ln(2/\theta))]$ .

#meters along L' = L'/
$$m_{avg}$$
 = L'/ $m_0$ [1 – a/L'· ( -(1+  $\theta$ ) + ln(2/ $\theta$ ))]

We minimize the number of meters in the grid distorted by gravity, by taking the derivative of # with respect to  $\theta$  and setting it equal to zero (it is probably easier to maximize 1/# and so we will do that).

$$\mathrm{d}/\mathrm{d}\theta(m_{avg}/L')=m_0d/d\theta[1-\mathrm{a/L'}\cdot(\,\text{-(1+}\,\theta)+\ln(2/\theta))]/\mathrm{L'}=0$$

as before, with dL'/d $\theta$  = L $\theta$  and L $\theta$  =  $L'\theta$  =  $r_0$  , we get, after some algebra

$$r_0$$
-a( 1/ $\theta$  - 1)=0

Since  $1/\theta$  is about  $10^5$ , the 1 is second order, so:  $\theta = a/r_0$ 

total deflection due to length contraction = 2 $\theta$  = 2 MG/  $r_0c^2$  =  $r_s/r_0$ 

The total deflection due to both effects is 4 MG/ $r_0c^2$ , 1.75 arcseconds

A few observations. The effects of time dilation and length contraction are long range effects, they drop off only as 1/r. We noticed in the integrals that the major contribution is from the high-end of the integrals whereas at the low-end, with small values of r, there is essentially no contribution. The  $sin^2(\phi-\theta)$  term, which kills off contributions only at the smaller values of r, made little or no difference. Also, for length contraction, I said that light follows the path of minimum number of wavelengths but then I said it follows the shortest path. Well, for length contraction, they are the same. Lengths in the radial direction are contracted and that includes the wavelength of light, as seen by the distant observer. So, as the meters are contracted, the wavelength of light is also contracted by the same amount, and therefore our calculation could've been the path of minimum number of wavelengths which, in this case, turns out to be the path of minimum number of meters.

This whole thing could've been done in one fell swoop by just doing the path of minimum number of wavelengths and taking into account both the effects of length contraction and time dilation on the wavelength of the light. The integral would've been

$$\begin{split} \lambda_{avg} &= \lambda_0 \ / \text{L'} \int_0^{\pi/2} (1 - a/r - a/r \cdot sin^2(\phi - \theta)) dL', \text{ where a is MG}/c^2 \text{ and dL'} = \text{r/cos}(\phi - \theta) \text{d}\phi. \\ \lambda_{avg} &= \lambda_0 \ (1 - a/L' \int_0^{\pi/2} \ (1 + \ sin^2(\phi - \theta)) / \text{cos}(\phi - \theta) \cdot \text{d}\phi) \\ &= \lambda_0 \ (1 - a/L' \int_0^{\pi/2} \ (2 - \ cos^2(\phi - \theta)) / \text{cos}(\phi - \theta) \ \text{d}\phi) \end{split}$$

The first term in the integral gives a factor of two times what we got for time dilation and the second factor is equal to  $1 + \theta$ , but then gives a second order contribution when we calculate the path of minimum number of wavelengths.

This is analogous to classical optics. In our flat world, the path of minimum number of wavelengths in free space turns out to be a straight line i.e. the shortest distance. But if there is something else going on, like a variable index of refraction (or time dilation), then the path is not a straight line. Time dilation, without length contraction, causes an increase in the wavelength of light with the distance from the sun, causing the path to bend, even in that flat world of no length contraction.

We have ALMOST beaten this light bending thing to death. In classic optics, there is something called Fermat's Principal that says light goes on the path of least time. In our case, with time

dilation, the path of minimum number of wavelengths prefers paths away from the sun where the wavelengths are longer (fewer wavelengths), whereas in Fermat's case the light prefers paths near the sun where the seconds are longer (fewer seconds). The answer would have the same magnitude but the opposite sign, so light would bend in the opposite direction for time dilation from what we get for the path of minimum number of wavelengths. In Einstein's first calculation of the bending of light, neither time nor light velocity varied in space and therefore the path of minimum time would've just been a straight path for the beam of light, no bending. Fermat and Einstein don't mix. Why is this? In classical optics, with a variable Index of refraction, the velocity of light changes but the frequency stays the same. If the frequency stays the same, then the wavelength varies and regions of higher velocity are regions of longer wavelength. Paths of least time (high velocity) are equal to paths of least number of wavelengths. In relativity, with time dilation, the speed of light is constant and so in areas of slower time (higher frequency) the wavelengths are shorter and therefore paths of least time are not paths of the least number of wavelengths. Before I repeat myself again, let's move on to something easier, the precession of the orbit of the planet mercury (or any other planet).

# Precession of the Perihelion of Mercury

Given what we know, we can treat this problem as a simple harmonic oscillator.

First, a little history. In the mid-1800's it was calculated that the precession of the orbit of mercury should be 532 arc seconds/century due to the gravitational force from the other planets, but unfortunately the measured value was 575 arc seconds/century. This leaves about 43 arc seconds to account for. Can relativity be the answer? Let's see. Now the angle of precession per revolution of the planet in its orbit is a dimensionless number. As we have already mentioned, the answer will be "something" times  $\mathrm{MG}/rc^2$  where r is now the radius of the orbit. We need to find the "something". If you were to guess 4, as in the bending of light, you would be off by about the same factor.

Let's consider the parameters involved in the orbital motion. When relevant, the parameters will be per unit mass of the planet. We want to consider their deviation from their non-relativistic values, as seen by our distant observer. The parameters will be something like A =  $A_0 \cdot (1 + \text{nMG/}rc^2)$  where n is a positive or negative integer. Remember that multiplying is the same as adding. We will list the n values for each parameter. For instance, r will be -1 because of length contraction in the radial direction. Also, 1/r is +1 and  $r^2$  is -2. The orbital velocity V is  $\ll$  c.

Parameter Design	nation	n	Comment
Radius	r	-1	length contraction
Time	t	+1	time dilation
Mass of the sun	M	+1	freefall analysis
Velocity in orbit	V	+1	time dilation, no tangential length contraction
Angular momentum	rV, h	0	a constant of motion that stays constant
Centripetal accel	$V^2/{ m r}, h^2/r^3$	+3	what causes this?
Gravitational force	$MG/r^2$	+3	looks like the mass increase might be needed

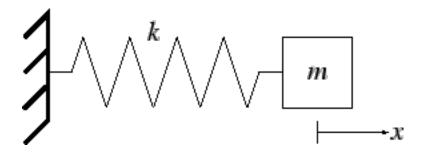
If one doesn't like the increase in M, then one can argue that the centripetal acceleration increases by three units and therefore there must be a gravitational force caused by "something" which causes this acceleration. Except for that pesky M, the angular momentum contains everything we (I) know about general relativity: length contraction in the radial direction, no length contraction in the tangential direction, and time dilation in all directions. Let's write our dimensionless number in terms of the constant of motion h (which does not depend on r), and the radius, r. Since  $h^2/r^3 = \text{MG}/r^2$ , then  $\text{MG}/rc^2 = h^2/r^2c^2$ . The gravitational force on the planet now becomes,

$$F = MG/r^2 \cdot (1 + 3 h^2/r^2c^2).$$

Our calculations will be done in the coordinate system of the distant observer, using the altered force of gravity. The calculations will be classical.

#### Orbital Mechanics and Harmonic Oscillators

Let's start very simple and work our way up. Consider the mass on a spring shown in the following diagram. We will do things per unit mass, that is, m = 1.



For small movements  $\Delta x$  from the equilibrium position, the restoring force F is  $-k \cdot \Delta x$  where k is the spring constant. The equation of motion is:

$$d^2(\Delta x)/dt^2 + k\Delta x = 0$$
 and the solution is  $\Delta x = a \cdot \sin(\omega t)$  where  $\omega = \sqrt{k} = \sqrt{dF/dx}$  and a is an amplitude

If k is a function of x, the analysis is still OK for small  $\Delta x$ . If k is just a constant, then the analysis is good for large motions of x. Our springs will be nonlinear and so we will consider only small deviations from the equilibrium position.

#### Classical Calculation

Consider a planet orbiting the sun. We go into the orbiting frame of the planet (you don't have to, but it is easier to visualize) and add a centrifugal force outward from the sun to replace the

centripetal acceleration. We now have two forces on the planet, a centrifugal force and a gravitational force as shown below.

$$MG/r^2$$
  $h^2/r^3$ 

A few things to notice. As r increases, we notice that the centrifugal force decreases faster than the gravitational force and therefore there is a restoring force to decrease r. If the gravitational force were a  $1/r^3$  force, then there would be no restoring force and the universe would consist of things floating around randomly.  $1/r^4$  is even worse. Luckily for us, the force is  $1/r^2$ .

The total force is  $F_t = h^2/r^3 - MG/r^2$ 

For oscillations in the radial direction:  $\mathbf{k}=\mathrm{d}F_t/\mathrm{d}\mathbf{r}=-3h^2/r^4+2\mathrm{MG}/r^3=-h^2/r^4$  (restoring) Therefore  $\omega_r=\sqrt{\mathbf{k}}=\mathrm{h}/r^2$ , but, because  $\mathbf{h}=\omega_o r^2$ , this is the same as the frequency of the planet orbiting the sun,  $\omega_o(\mathbf{r})$  for radial, o for orbital). This means that the point of maximum radius (perihelion) occurs at the same point every revolution of the planet around the sun. This slightly elliptical orbit is, therefore, stationary and has no precession.

Classically: radial frequency = orbital frequency

## **Relativity Calculation**

Relativity adds a small  $1/r^4$  term to the gravitational force.

$$MG/r^2 (1 + \delta/r^2) \longleftarrow h^2/r^3$$

Where  $\delta = 3h^2/c^2$ 

Before we calculate, let's think about the problem. The gravitational force, which provides the restoring force for increasing r, now has a small  $1/r^4$  term and therefore decreases faster, with increasing r, then it did before. Therefore, the restoring force will be less than classical, and the radial oscillation frequency will be less. Also notice that when  $\delta/r^2 = 3/2 \cdot r_s/r = 1$  (r=3/2 $r_s$ ) then the restoring force has equal parts  $1/r^2$  and  $1/r^4$ .

The spring constant is:  $k = dF_t/dr = d/dr(h^2/r^3 - MG/r^2(1 + \delta/r^2))$ 

After a little algebra, one gets,  $k = h^2/r^4(1-2\delta/r^2)$ . Notice that for radii smaller than  $3r_s$ , the restoring force goes negative and there would be no stable orbits at smaller radii. This is on the hairy edge of where our approximations are still reasonable.

So, the frequency of radial oscillations is, remembering that  $(1 + x)^n = 1 + nx$  for small x,

$$\omega_r = \sqrt{k} = h/r^2(1 - \delta/r^2) = \omega_o(1 - \delta/r^2)$$

This is less than the frequency of orbital rotations,  $\omega_o$  and, therefore the orbit has a precession and is not stationary. The point of maximum radius (perihelion) will occur later-and-later each revolution of the planet around the sun. The difference per revolution, in radians per second is:

 $\omega_o - \omega_r = \omega_o \delta/r^2$ . If we multiply by the time of one orbit,  $2\pi/\omega_0$ , we get the number of radians of precession per revolution.

$$\Delta\theta$$
/revolution =  $\omega_0 \delta/r^2 \cdot 2\pi/\omega_0 = 6\pi h^2/r^2c^2 = 6\pi MG/rc^2$ 

This is the answer.  $6\pi$  is the "something" we were looking for.

For light bending,  $4\,\mathrm{MG}/r_0c^2$  was equal to 1.75 arcseconds , so we replace  $4\,\mathrm{by}\,6\pi$  and the radius of the sun  $r_0$  ( $7\mathrm{x}10^5\,\mathrm{km}$ ) with the radius of mercury's orbit. It is a little sloppy, but we will just use the mean radius of  $57\mathrm{x}10^6\,\mathrm{km}$ . The answer is 0.102 arc seconds/ revolution. The period of mercury's orbit is 88 of our days, so we multiply by  $365/88\cdot100$  to get 42 arc seconds/century compared to the 43 that we needed. Treating the amplitude of oscillation of Mercury's orbit radius as being small is pushing it a little bit, so a more careful analysis would give a better answer. This whole analysis is for any planet going around a large body like the sun. If you do the calculation for Earth you get about 3.8 arc seconds per century and, it's not so sloppy because earth's orbit is fairly circular.

## Black Holes, the Photon Sphere, and Trouble (or lack of trouble)

So far, our simple approach to general relativity has done well as far as calculating goes. We have taken examples only where the radius from the large gravitational object is much larger than the Swartzschild radius, which is true for things except a black hole. So, let's consider a black hole. We revisit our free fall example, but this time falling towards a very small but very massive object. We want to calculate at what radius our free fall velocity with respect to this object is equal to the speed of light. The first calculation will be done assuming that the massive object accelerating towards us doesn't change its mass as it's velocity towards us increases.

$$dV = acc \cdot dt = acc \cdot dt/dr \cdot dr = MG/r^2 \cdot 1/V \cdot dr$$
 So  $VdV = MG/r^2 \cdot dr$  acc is acceleration

integrating with V equals zero at infinity, we get, for the radius where V equals c,

r =  $2MG/c^2$ , which is the Swartzschild radius  $r_s$ .

For our sun, this radius is about 3 kilometers compared to its actual radius of 0.7 million km. We will never get to the Schwartzschild radius by free-falling toward the sun. For its mass, it is not small enough. An object smaller than this radius is called a Swartzschild black hole.

In the gravity frame things are pretty grim. The clocks have slowed to a halt and the radial dimensions have now decreased to zero due to length contraction. This radius is called the event horizon. Anything that occurs inside of this event horizon will not make it out. Massive particles would have to be traveling at the velocity of light just to make it out. Photons would have to have infinite frequency and, even then, they would be red shifted down to zero frequency if they escaped.

If we are in freefall toward a massive body, and our reference frame is inertial, then it seems that the mass of the body should depend on its velocity with respect to us. If this is so, and special relativity says it is so, then the force of gravity will increase faster than we had previously calculated and the event horizon will be at a larger radius than the Swartzschild radius. Let's calculate, including the increase in mass of the body, at what radius the velocity will become equal to the speed of light. We now have:

$$dV = g \cdot dt = g \cdot dt/dr \cdot dr = MG\gamma/r^2 \cdot 1/V \cdot dr$$
 So  $VdV\sqrt{1 - V^2/c^2} = MG/r^2 \cdot dr$ 

integrating, again with V equals zero at infinity, we get, for the radius where V equals c,

r =  $3MG/c^2$  =  $3/2r_s$ . At this radius we would see, in our freefall, the mass of the black hole go to infinity.

## The Photon Sphere

In classical optics there is an interesting example I will call the light merry-go-round. Let's say we have a transparent disk which has an index of refraction which decreases with the radius. For instance, let  $n = n_0(1-r/r_0)$  where n is the index of refraction. Because of this index, the velocity and wavelength of light will increase with increasing radius. Hold onto your seats because here comes the path of minimum number of wavelengths again. If the index were constant, then for a circular path of light, an increase in the circumference of the path would always cause an increase in the number of wavelengths and so there would be no path of minimum number of wavelengths. But in our case the index decreases as the radius increases which means that the wavelength increases. There may be a radius at which the percent increase or decrease in circumference of the path and the increase or decrease in the wavelength, as the radius increases or decreases, are equal and therefore the number of wavelengths does not change. This would be the path of minimum number of wavelengths. Let's see if there is such a radius. The number of wavelengths is just the circumference divided by the wavelength.  $\lambda = \lambda_0/n = \lambda_0/n(1-r/r_0)$ 

# wavelengths =  $2\pi r/\lambda = 2\pi r \cdot (1 - r/r_0) n_0/\lambda_0$ .

Taking the derivative and setting it equal to zero gives  $r = r_0/2$ . One could make  $n_0$  equal to 3 and therefore the orbit would be where the index is 1.5. Such a thing would be interesting on the scale of several microns such as a section of a graded-index multi-mode optical fiber. This effect could cause a spiral mode down the length of a graded-index multi-mode fiber. The index

does not need to vary linearly with r. For any index n(r), all that is needed is that there be a radius where dn/dr = -n/r.

In the gravitational field around a black hole, we have a similar situation. Time dilation causes the wavelength of light to increase the further away you get from the black hole, giving

 $\lambda$ =(1-a/r) $\lambda_0$  where  $\lambda_0$  is the wavelength at infinity. The wavelength, in a circular orbit, is not affected by length contraction. The situation is a little more complicated than the previous example, because the circumference is <u>not</u> equal to  $2\pi$  times the contracted radius. The circumference is equal to two pi times the <u>uncontracted</u> radius, r/(1-a/r), because the circumference is not affected by length contraction as is the radius. We then have that the number of wavelengths on a path of radius r (contracted) is:

# wavelengths =  $2\pi r/(1-a/r)\lambda = 2\pi r/(1-a/r)^2\lambda_0$  where a = MG/ $c^2$ Here we must be careful and not use  $1/(1-a/r) \cong 1+a/r$  since a/r is not  $\ll 1$ . The constants drop out when we set the derivative equal to zero. d#/dr = d/dr( $r/(1-a/r)^2 = 0$ : This gives  $r^2 - 4ar + 3a^2 = 0$ . Solving for r:

 $r = 3a = 3/2r_s$ . This is the radius of what is called the photon sphere,

a radius where light "can" orbit around the black hole. For light to actually orbit the black hole continuously, the orbit would need to be stable to a wavelength. Maybe over a distance of several microns, but not over kilometers. The continuous orbiting is not stable. But, this radius is outside the bounds of where our approximations are reasonable.

Here is the trouble.  $3/2r_s$  is the correct answer when the experts calculate with all the formalism, but there is no reason (I guess there is one) why we should've got the correct answer. We made assumptions at the beginning that the radius was much larger than  $r_s$  (freefall velocity much less than c) but here we are doing an example where r is very close to  $r_s$  and we still get the right answer. The experts may know the reason, but at the present, I do not.

If you have been keeping track, you will notice that this is the third time that we have come across  $r = 3/2r_s$ .

- 1. Radius of the orbit where the  $1/r^2$  and  $1/r^4$  terms in the gravitational force are equal.
- 2. Radius of the event horizon when free falling toward a mass which increases with velocity.
- 3. Radius of the photon sphere

1 and 3 involve the approximations we made at the start(which are not good at this radius) but 2 has no approximations, just special relativity. Food for thought?

### The twin paradox

A famous paradox of special relativity is the twin paradox. Here we have two twins, one which stays home and the other who travels in a spaceship and then turns around and returns to earth. Both twins see the other twin's clock as running slow and so which one is younger when the second twin returns? The rule is to analyze the situation from the point of view of the inertial

frame. This is the frame of the twin that stays home, so the answer is that the twin who took the ride in the spaceship comes back younger. How does he see this? The explanation is in the turnaround. Here the twin in the spaceship experiences an acceleration which is equal to a gravitational field. This field will be in a direction from earth to the spaceship. During turnaround then, the clock on earth will be higher in the gravitational field and therefore the twin in the spaceship will see earth's clock running faster than his. This speed up in earth's clock is enough to more than compensate for the fact that the earth clock was running slow before and after turnaround and when the spaceship twin gets back to earth, he will find that he is younger than his earth twin. Let's calculate. Say that the travelling twin starts his turnaround at a distance L from earth. If he has been traveling with a velocity V. then they turn around acceleration, in a time T2, which we take to be much less than the travel time, will be g = 2V/T2. If you do the freefall exercise in this constant gravitational field, you will find the difference in the rate between two clocks that are spaced a distance X apart, is assuming  $V \ll c$ : % difference in clock rates =  $\gamma$ -1 = gX/ $c^2$ . You can throw in length contraction but it's a second-order effect. During the turnaround time T2, the earths clock, will gain  $gX/c^2 \cdot T2 = 2LV/c^2$  seconds on the traveler's clock. But, during the out trip and the back trip, the earth clock lagged behind the traveler's clock by  $(\gamma-1)$ -travel time =  $1/2V^2/c^2 \cdot 2L/V = LV/c^2$ , again assuming  $V \ll c$ . The net effect is that the earth clock, according to the traveler, advances  $LV/c^2$ ahead of his clock during the entire trip and so the traveler comes back younger. That is what we get when we look at the problem from the point of view of the earth twin. Paradox resolved.

A more realistic twin paradox would be one where the travelling twin accelerates from the earth up to a velocity V, turns around, and then decelerates down to zero velocity to land on earth. During the acceleration and deceleration near earth, the traveling twin sees the earth clock as running slow compared to his clock due to gravity. During turn around he sees the earth clock as running fast. These effects do not cancel. The effect of the acceleration on the clocks depends on the distance between the clocks and therefore the turnaround outweighs the effect of acceleration and deceleration near earth. The net effect is that the traveler comes back younger. For those who like to calculate, one might try the following exercise. Consider the traveler leaving earth with an acceleration a to a velocity V, and immediately decelerating and turning around until his velocity is V toward the earth. He then lands on the earth at the zero velocity. One can calculate to see whether the traveler and earth twins both agree on who is younger and by how much. A simpler exercises is to show that if a constant velocity section is added between the acceleration and deceleration sections, that they still agree, although, on a different number.

Do our orbiting astronauts come back younger? No, they come back older. More on that in the next section.

Another thing to think about with the twin paradox, when you have nothing else to do, is to consider the following. If the stationary twin looks at the moving twin's x axis, which is marked off in meters, he notices that it is length contracted with respect to his. During turn around, as the travelling twin slows down to zero velocity, his x-axis becomes the same scale as that of the stationary twin. Where did all those meter marks go that had passed by? Try to figure out how this all evolves.

#### Clocks and GPS

One occurrence of relativity in our everyday life is the Global Positioning System (GPS), although we probably don't know it (or care about it). GPS consists of 31 satellites (as of 2017) orbiting the earth at a distance of 20,000 km above the earth and with an angular velocity of twice that of the earth, so that the satellites pass over every 24 hours. The idea is to have every point on earth have a line of sight to at least 4 satellites at any given time. The satellites give out coded signals such that an Earth-based receiver, like your cell phone, can determine the time-of-flight of a radio signal from the satellite to your receiver. This time-of-flight gives the distance from the satellite to the receiver. There is a sphere around the satellite with this particular distance. If the receiver knows where the satellites are at any given time, then three spheres around three separate satellites will intersect at your position on earth (they also intersect at another point which is not near the surface of the earth). Radio waves (and light) travel at a speed of about 1 foot/ nano second. So, if you want to determine your position to a few feet, this timing needs to be done to a few nanoseconds.

There are two problems with the timing. The first one is that relativity causes clocks that are moving with respect to one another and clocks separated in a gravitational field, to run at different rates. To do the position calculations, the earth-based receiver must be synchronized to the satellites' atomic clocks to a few nanoseconds. The second is that your cell phone does not have an atomic clock in it and is not nearly a good enough clock to do the job. Let's consider the first.

We want to calculate the difference in rates between the satellite clock and our earth clock. If we calculate, say, in microseconds per day, this is a dimensionless number, time divided by time. As before, the answer will be "something" times  $MG/rc^2$ , where M now is the mass of the earth and r is the radius of the satellite orbit. Oddly enough, if you guessed 4, like the bending of light by gravity, you would be pretty close.

We will calculate things in terms of the radius of the satellite orbit,  $r_{sat}$ , the angular velocity of the satellite,  $\omega_{sat}$ , and the ratio of the satellite's orbit radius to the earth's radius, which we will call x. Remember that the angular velocity of the earth is half that of the satellite. We will calculate the difference in the rate of the two clocks as seen by a distant observer. First, the effect of velocity as seen on the distant observer's clock.

The % difference in rates due to the velocity with respect to the distant observer's clock is just  $\gamma$  -1, which, with V $\ll c$ , is equal to  $1/2 \cdot V^2/c^2$ . For the satellite, V =  $\omega_{sat} \cdot r_{sat}$  and for the earth clock,  $\omega_{sat} \cdot r_{sat}/2x$ . The difference in rate between the two clocks, therefore, is

 $\omega_{sat}^2 \cdot r_{sat}^2 / 2c^2 (1 - 0.25/x^2)$ . The centripetal acceleration of the satellite,  $\omega_{sat}^2 r_{sat}$  is equal to the gravitational force per unit mass on the satellite, MG/ $r_{sat}^2$ . Therefore,  $\omega_{sat}^2 \cdot r_{sat}^2 = \text{MG}/r_{sat}$ , and the rate difference of the clocks due to their velocity with respect to the distant observer is MG/ $2c^2r_{sat}$  (1 - 0.25/ $x^2$ ), with the satellite clock running slow.

We are not done because we still need to consider the effect of gravity. We showed before that in a central gravitational field, time dilation causes time to be a function of the radius, T =  $T_0$ (1-MG/ $rc^2$ ). The difference in the rate of time at the earth and the satellite due to gravity is, therefore equal to MG/ $c^2$  (1/ $r_e$ -1/ $r_{sat}$ ) = MG/ $r_{sat}$  $c^2$ ( $r_{sat}$ / $r_e$ -1) = MG/ $r_{sat}$ (x -1)

Putting this all together:

The net difference in clock rates is:  $MG/c^2r_{sat}(x-1.5+0.12/x^2)$ 

Now x = 26.4/6.4 = 4.1, so the difference in clock rates is MG/ $c^2r_{sat}$ ·2.6

And again, the answer is "something" times the dimensionless parameter MG/ $c^2$ r.

Putting in numbers, gravity causes 45  $\mu$ s per day and relative velocity, 6  $\mu$ s per day, the net difference in rate being 39  $\mu$ s per day with the satellite clock running fast.

So, that is the standard approach, with gravity causing an effect which is about 7 ½ times that of velocity, that is, "general" relativity contributes 7 ½ times that of "special" relativity. But are they really separate effects? It depends on your point of view. Our approach to general relativity is that gravity is equal to acceleration, acceleration leads to velocity, velocity leads too length contraction and time dilation. They seem to be all mixed together.

To show that, let's go into a frame rotating with the satellite with the earth at the center. Here, the satellite is at a point with no acceleration and no gravity because gravity at the satellite has been canceled by the centrifugal force. An inertial point. We must then subtract the effect of the centrifugal force,  $\omega_{sat}^2 r_{sat}$ , integrated all the way down to zero radius, from our previous calculation of the effect of gravity. We need to integrate all the way down to zero because the centrifugal force also affects the earth clock. If you like the concept of dropping a photon, then it is the dropping of a photon in a force field which contains the force of gravity and a centrifugal force in the opposite direction.

The effect of centripetal force =  $\int_0^{r_{sat}} F_{cen} dr = \omega_{sat}^2 \int_0^{r_{sat}} r dr = \omega_{sat}^2 \cdot r_{sat}^2/2$ .

As before,  $\omega_{sat}^2 \cdot r_{sat}^2 = \text{MG}/r_{sat}$ , so when we subtract this from our previous gravity calculation we get the result: net "gravity" effect = MG/ $c^2 r_{sat}$ ·(x-1.5)

There is a small velocity of the earth clock in this frame of  $\omega_{sat} \cdot r_{sat}/2x$ . In this case the effect now causes the earth clock to run slow with respect to the satellite clock. The effect on rates is  $MG/c^2r_{sat}\cdot 0.12/x^2$  and must be added because the effect is in the same direction as gravity.

The net difference in clock rates in the rotating frame is:  $MG/c^2r_{sat}(x-1.5+0.12/x^2)$ 

It is the same as before, but now, with x = 4.1, "general" relativity contributes 99.7% of the entire effect. It depends on how you look at it. One man's velocity is another man's gravity.

This difference in rates is not a big deal because the clocks in the satellites can be corrected with electronics. If they weren't corrected, the GPS system would lead you off course by about 11 km a day. As mentioned earlier, the other problem is that your cell phone doesn't have a clock good to nanoseconds and therefore the distances from all the satellites is probably wrong, but by the same amount, the same effect as relativity except fairly random. This, of course, gives the wrong position. If we include one more satellite, then we can dither the time on the cell phone until the distance from all 4 satellites coincide at some point. This is the correct position. One could also do this with uncorrected satellite clocks and just solve 4 equations and forget about the satellite clocks being off with respect to the cell phone. Oh well, it was an excuse to talk about relativity.

## The Expansion of the Universe and Dark Energy

In this section I will let my mind wander, probably off the edge. I make no pretenses; I don't know anything about dark energy, but I read on the internet that there seems to be a lot of it and it is responsible for the universe expanding faster than you would calculate from the gravitational attraction between stars. So where does the energy come from? Well, of course, we can change mass into energy, not through nuclear reactions but maybe in a different way. When we did the free fall analysis, falling toward the sun, we noticed that the mass of the sun increased the closer we got to it. This seemed a little strange but it was consistent when we analyzed the orbit of mercury. Consider the universe in which the distance between stars is expanding radially, that is, they are moving further apart from each other. The freefall analysis would say that the further the stars get from each other, the lower their mass as seen by other stars. By conserving energy, let's say this decrease in mass is turned into kinetic energy of the motion of the stars. Is this enough to combat the gravitational attraction of the stars to each other so that the universe expands more easily? We found that the increase in mass was  $MG/rc^2$ , which decreases as r increases. If we multiply by  $c^2$  we get the energy from this lost mass. This gain of energy is the same as the loss of kinetic energy due to the gain in potential energy in the gravitational field. This conversion of mass into energy makes up for the loss of kinetic energy by gravity as the stars expand away from each other. The answer is yes, there is enough energy. If this were true, the universe would never stop expanding.

On the other hand, if the universe were collapsing, this affect would slow down the collapse of the universe because, then, the increase in mass as the stars move closer will take away from the kinetic energy of the stars. So, those are my dark thoughts about dark energy. I will keep thinking about it. No harm, no foul.

Well, that's about all for relativity. This approach probably has a lot of rough edges. For instance, all those terms which I said were second order, maybe they shouldn't have been there at all in a more correct approach. On the other hand, maybe our approach is fairly correct. It's strength, if you want to call it that, is that no fancy formalism is required and so "any" undergraduate engineer or science student, if they wanted, could understand it fairly easily. To me, relativity, along with optics, are examples of where simple assumptions can get you a long way just by thinking. I will finish my write up on optics, using only the assumptions which I have already pointed out, and put it on this website when I'm done. There will be a lot on the path of minimum number of wavelengths in that writeup.