

# Physics on the Celestial

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# Overview

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# Lorentz Group & Poincaré group

## Definition

Lorentz group ( $O(1,3)$ ) is the isometric transformations in Minkowski spacetime

Four disconnected parts:

- Proper orthochronous lorentz group:  $SO(1,3)^\uparrow$
- Discret Transformation:  $\mathcal{P}, \mathcal{T}$

We just need to care about  $SO(1,3)^\uparrow$ .

## Definition

$$\text{Poincaré Group} = SO(1,3)^\uparrow \ltimes \mathbb{R}^{1,3}$$

It's the biggest symmetry group of 1+3 spacetime.

# Double Cover

As we all know:

$$SO(3) \cong SU(2)/\mathbb{Z}_2 \quad (1)$$

Similarly:

$$SO(3,1)^\uparrow \cong SL(2, \mathbb{C})/\mathbb{Z}_2 \quad (2)$$

More fun with identities:

$$SO(2,1)^\uparrow \cong SL(2, \mathbb{R})/\mathbb{Z}_2, \quad SO(5,1)^\uparrow \cong SL(2, \mathbb{H})/\mathbb{Z}_2, \quad SO(9,1)^\uparrow \cong SL(2, \mathbb{O})/\mathbb{Z}_2 \quad (3)$$

Because

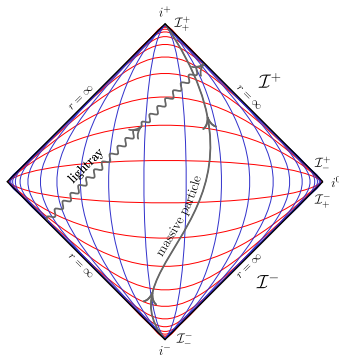
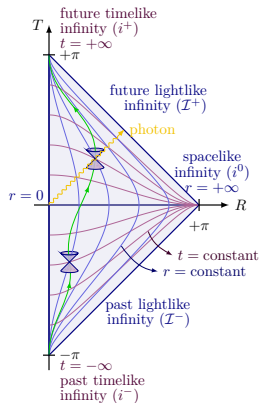
## Theorem (Hurwitz)

*The normed division algebras over  $\mathbb{R}$  are precisely  $\mathbb{R}, \mathbb{C}, \mathbb{H}$  and  $\mathbb{O}$ .*

So we can never find ternary number!

# Penrose Diagrams ( $\text{Mink}_4$ )

Penrose diagram conformally compactify the spacetime, represent it by a finite scale diagram.



## Definition

Every point of  $\mathcal{I}^\pm$  represent a sphere of infinite radius, called **Celestial Sphere** ( $\mathcal{CS}^2$ )

- Because  $\mathcal{CS}^2$  can be seen as a complex plane after single-point compacting, we can use  $(z, \bar{z})$  to label its angular coordinates ( $\mathcal{S}^2 \cong \mathbb{C} \cup \{\infty\}$ ). We prefer to use retarded coordinates  $(u, r, z, \bar{z})$  parametrize  $\mathcal{I}^+$ , and advanced coordinates  $(v, r, z, \bar{z})$  on  $\mathcal{I}^-$ . Remember, The angle parameters  $(z, \bar{z})$  are antipodally identity between  $\mathcal{I}^\pm$ .
- $SO(1, 3)^\uparrow \cong SL(2, \mathbb{C})/\mathbb{Z}_2$ , Lorentz transformations generate conformal transformations on the celestial sphere!

$$z \mapsto \frac{az + b}{cz + d}, \quad \bar{z} \mapsto \frac{\bar{a}\bar{z} + \bar{b}}{\bar{c}\bar{z} + \bar{d}} \quad (4)$$

We can use infinite symmetry on  $\text{CFT}_2$  to analyse scattering process in  $\text{Mink}_4$

# Conformal Basis

In QFT, we usually use fourier transformation to work in moment space, which can manifest the translational symmetry of  $\text{Mink}_d$ . To use the conformal symmetry of  $\text{CFT}_{d-2}$ , we can work with **Conformal Basis**:

- Photons:

$$A_{\mu a}^{\Delta, \pm}(X^\mu; \vec{w}) = -\frac{1}{(-q \cdot X \mp i\epsilon)^{\Delta-1}} \frac{\partial}{\partial X^\mu} \frac{\partial}{\partial w^a} \log(-q \cdot X \mp i\epsilon) \quad (5)$$

- Gravitons:

$$h_{\mu_1 \mu_2; a_1 a_2}^{\Delta, \pm}(X; \vec{w}) = P_{a_1 a_2}^{b_1 b_2} \frac{1}{(-q \cdot X \mp i\epsilon)^{\Delta-2}} \partial_{b_1} \partial_{\mu_1} \log(-q \cdot X \mp i\epsilon) \partial_{b_2} \partial_{\mu_2} \log(-q \cdot X \mp i\epsilon) \quad (6)$$

where  $P_{a_1 a_2}^{b_1 b_2} \equiv \delta_{(a_1}^{b_1} \delta_{a_2)}^{b_2} - \frac{1}{d} \delta_{a_1 a_2} \delta^{b_1 b_2}$ ,  $\Delta \in \frac{d}{2} + i\mathbb{R}$ .

However, in the soft region, which forced  $\Delta = 1$ , the conformal wave functions need to be redefined. (arXiv: 1810.05219)

For example, we can use Mellin transformation to transform momentum space amplitudes to conformal space amplitudes:

$$\tilde{\mathcal{A}}(\Delta_i, \vec{w}_i) \equiv \prod_{k=1}^n \int_{\mathbb{H}_{d+1}} [d\hat{p}_k] G_{\Delta_k}(\hat{p}_k; \vec{w}_k) \mathcal{A}(\pm m_i \hat{p}_i^\mu) \quad (7)$$

3 massive scalar scattering ( $\phi^3$ -theory) at tree-level ( $m_{\text{in}} = (2 + \epsilon)m_{\text{out}}$ ):

$$\tilde{\mathcal{A}} = \frac{i 2^{\frac{9}{2}} \pi^6 \lambda \Gamma(\frac{\Delta_1 + \Delta_2 + \Delta_3 - 2}{2}) \Gamma(\frac{\Delta_1 + \Delta_2 - \Delta_3}{2}) \Gamma(\frac{\Delta_1 - \Delta_2 + \Delta_3}{2}) \Gamma(\frac{\Delta_1 - \Delta_2 + \Delta_3}{2}) \Gamma(\frac{-\Delta_1 + \Delta_2 + \Delta_3}{2}) \sqrt{\epsilon}}{m^4 \Gamma(\Delta_1) \Gamma(\Delta_2) \Gamma(s_3) |w_1 - w_2|^{\Delta_1 + \Delta_2 - \Delta_3} |w_2 - w_3|^{\Delta_2 + \Delta_3 - \Delta_1} |w_3 - w_1|^{\Delta_3 + \Delta_1 - \Delta_2}} + \mathcal{O} \quad (8)$$

But we need a more systematic approach, CCFT is coming!



# Infrared Triangle

Surprisingly, three seemingly unrelated areas can be linked by some kind of relationship!

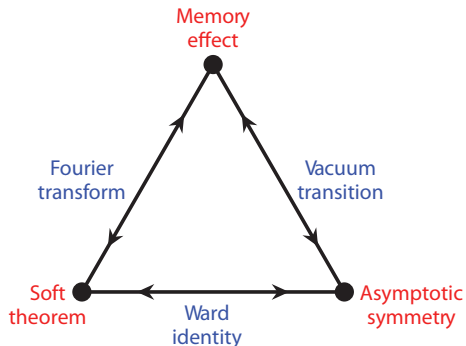


Figure: The infrared triangle

# Large gauge symmetry $\not\subseteq$ Gauge symmetry

## Definition

$$\text{Asymptotic symmetries} = \frac{\text{Allowed symmetries}}{\text{Trivial symmetries}}$$

We can use Asymptotic analyse on QED and perturbative gravity to find asymptotic symmetries, called Large gauge/diffeomorphic symmetry. Noether theorem tells us there must be some conserved charges correspond to asymptotic symmetries, in fact, there are infinite number of conserved charges! These symmetries will spontaneous breaking and bring soft photons and soft gravitons to our world.

$$\langle \text{out} | (Q_{\epsilon}^{+} \mathcal{S} - \mathcal{S} Q_{\epsilon}^{-}) | \text{in} \rangle = 0 \iff \sum_{k=1}^m \frac{e Q_k^{\text{out}} p_k^{\text{out}} \cdot \epsilon}{p_k^{\text{out}} \cdot q} - \sum_{k=1}^n \frac{e Q_k^{\text{in}} p_k^{\text{in}} \cdot \epsilon}{p_k^{\text{in}} \cdot q} \quad (9)$$

Asymptotic symmetries  $\overset{\text{Ward Identity}}{\iff}$  Soft theorem!

# Open Questions

There are some open questions:

- CCFT
- Non-Abelian promotion
- Supersymmetry on the celestial
- BCFW, CHY, KLT relations of the celestial Amplitudes
- ...

There's a lot of new physics on the celestial waiting to be discovered !

## Lectures on the Infrared Structure of Gravity and Gauge Theory

Andrew Strominger (Harvard U.)

Mar 15, 2017

158 pages

ISBN: 9780691179735

e-Print: [1703.05448](#) [hep-th]

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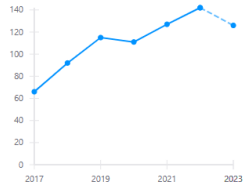


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779 citations

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## Conformal basis for flat space amplitudes

Sabrina Pasterski (Harvard U.), Shu-Heng Shao (Princeton, Inst. Advanced Study)

May 2, 2017

17 pages

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