

Project List

- Modeling Gene Regulatory Networks
- Planetary Orbits and Lightcurves from Transits
- Modeling Supernovae Remnants using Pyro
- Solving nuclear reactions in white dwarf interiors

Modelling gene regulatory networks

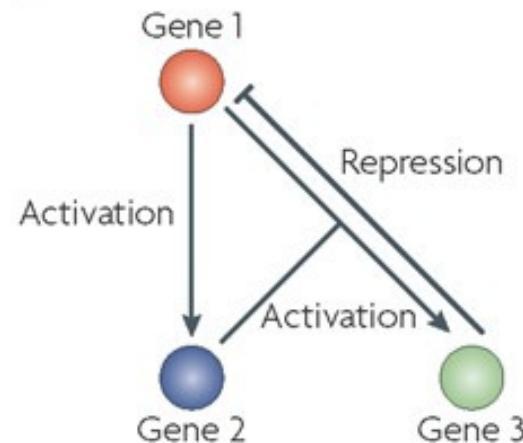
a

$$\frac{d(\text{gene}_1)}{dt} = k_{1,s} \cdot \frac{1}{1 + k_{1,3} \cdot \text{gene}_3} - k_{1,d} \cdot \text{gene}_1$$

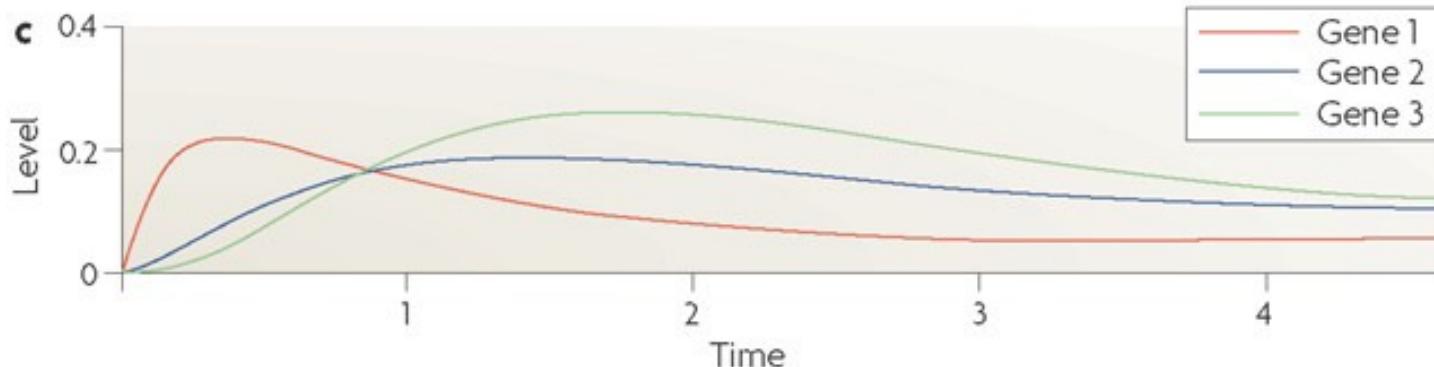
$$\frac{d(\text{gene}_2)}{dt} = k_{2,s} \cdot \frac{k_{2,1} \cdot \text{gene}_1}{1 + k_{2,1} \cdot \text{gene}_1} - k_{2,d} \cdot \text{gene}_2$$

$$\frac{d(\text{gene}_3)}{dt} = k_{3,s} \cdot \frac{k_{3,1} \cdot \text{gene}_1 \cdot k_{3,2} \cdot \text{gene}_2}{(1 + k_{3,1} \cdot \text{gene}_1) \cdot (1 + k_{3,2} \cdot \text{gene}_2)} - k_{3,d} \cdot \text{gene}_3$$

b



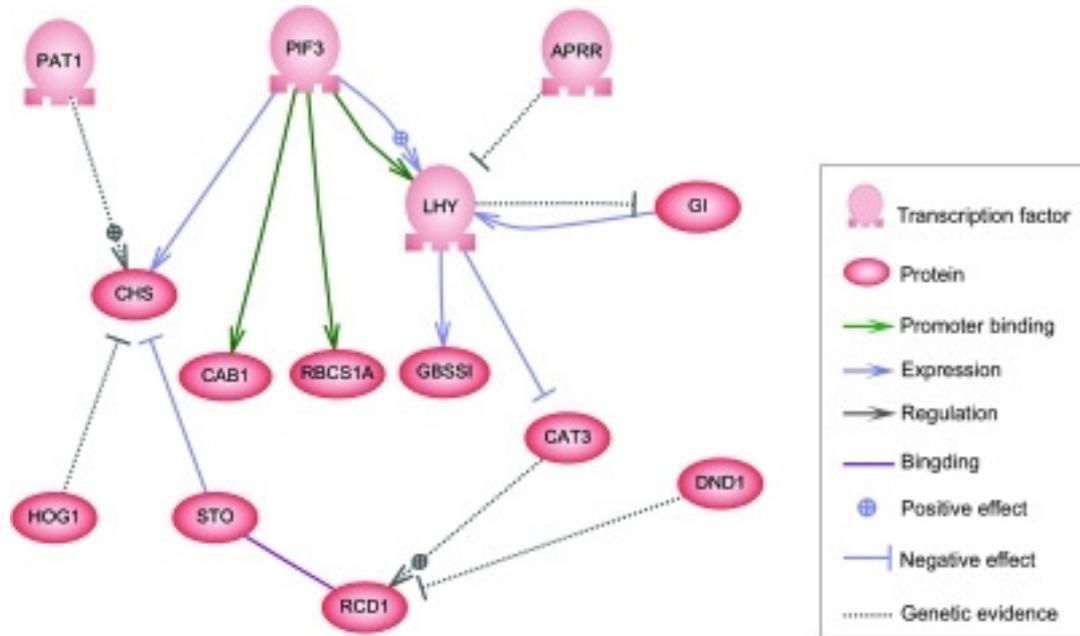
c



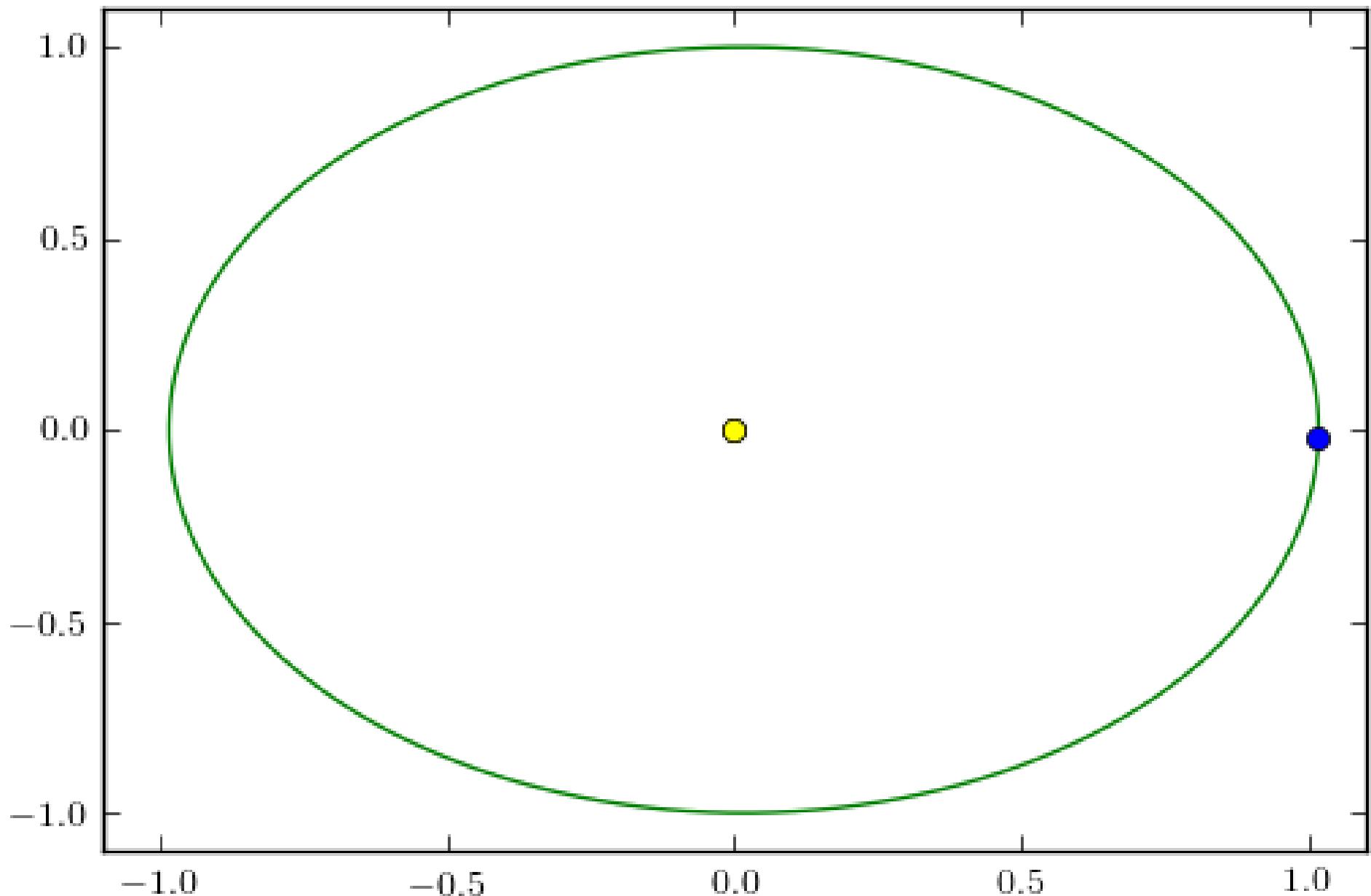
Nature Reviews | Molecular Cell Biology

- Important for understanding biological processes

- Student will build a model of gene regulations by solving ODE to model gene regulations
- Learn how to use gene databases to obtain real information
- Analyze information coming from the model



Modelling Planetary Orbits

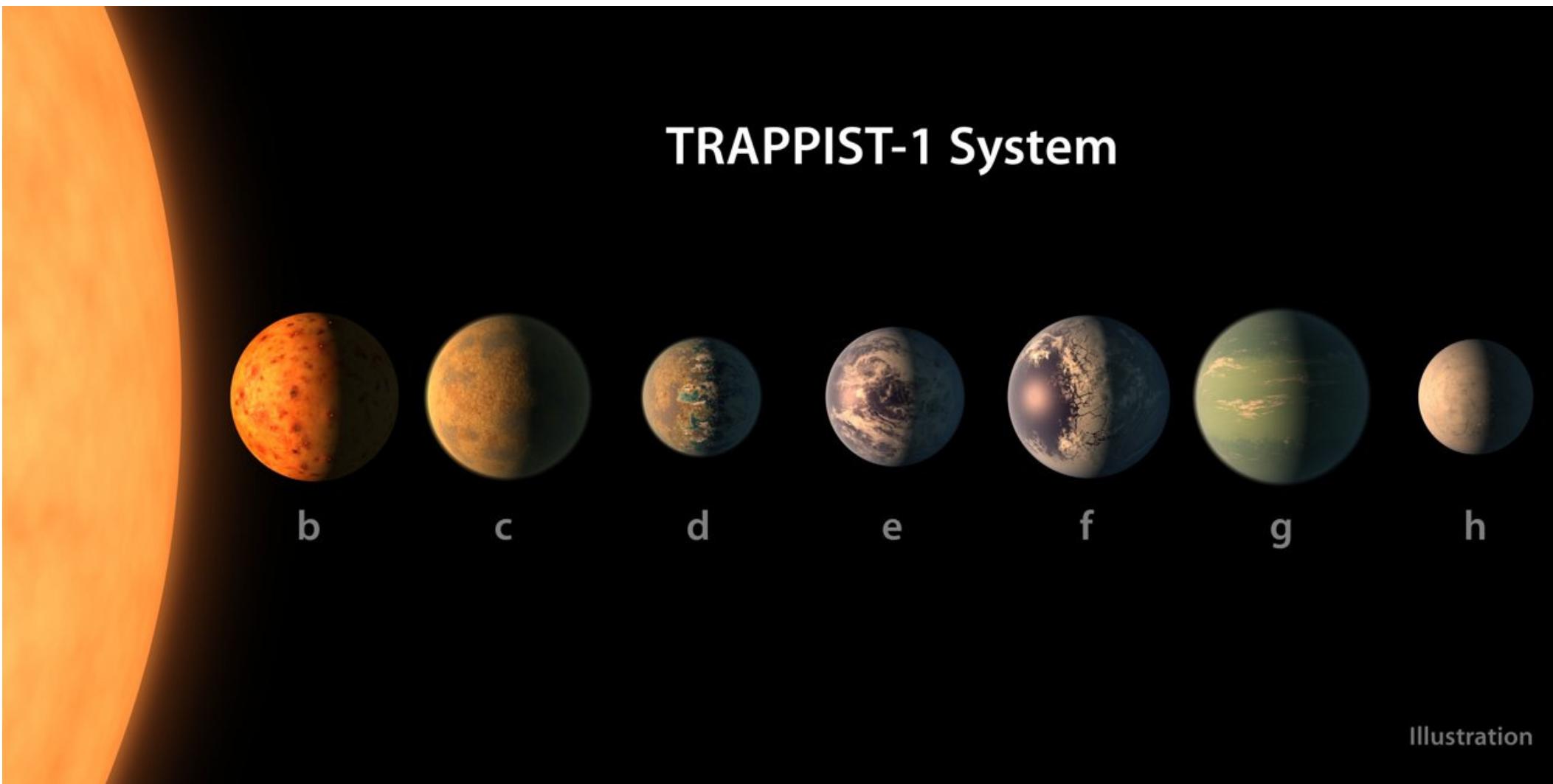


The Earth From Saturn – Cassini

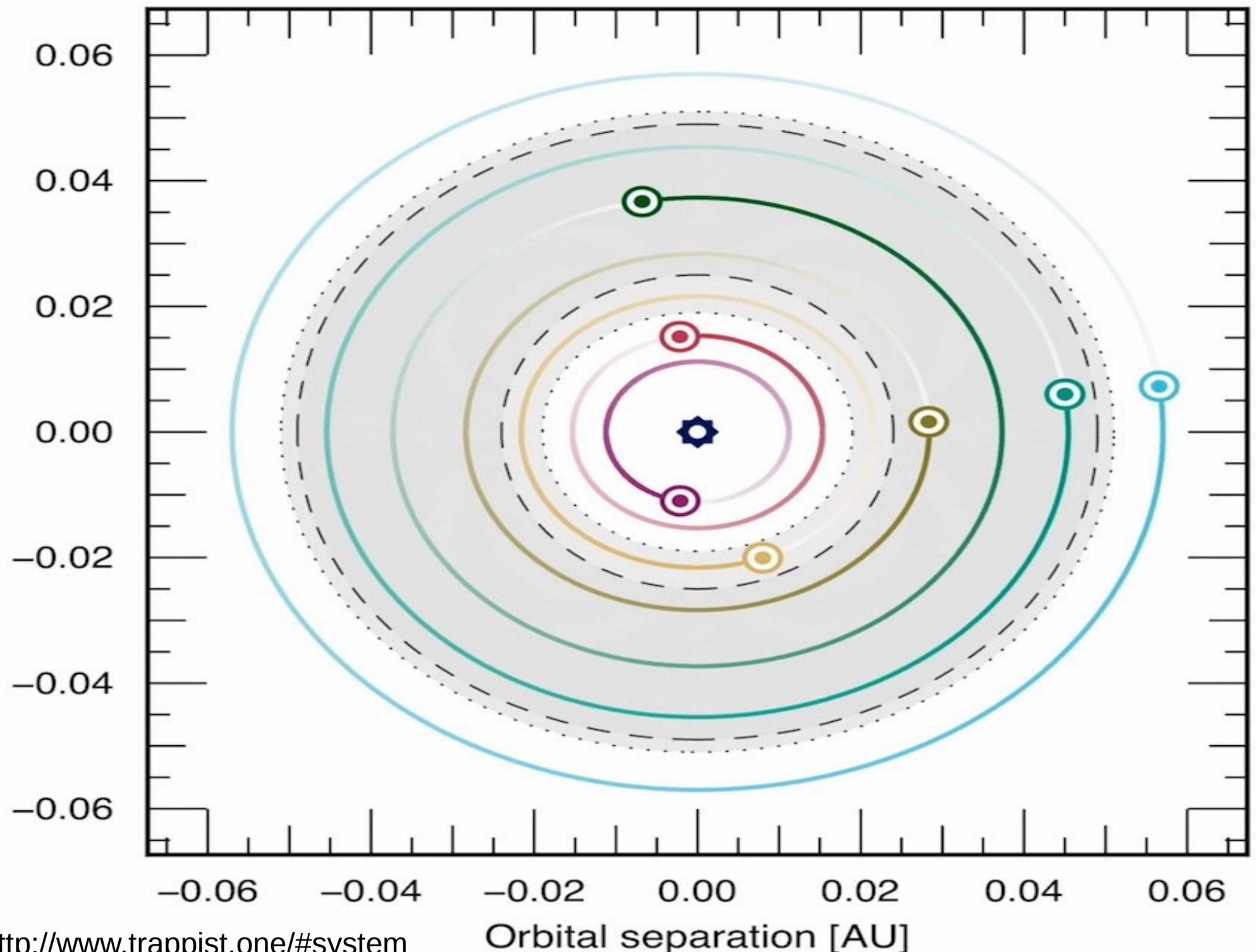


<https://saturn.jpl.nasa.gov/resources/5868/>

Newly Discovered – Trappist!



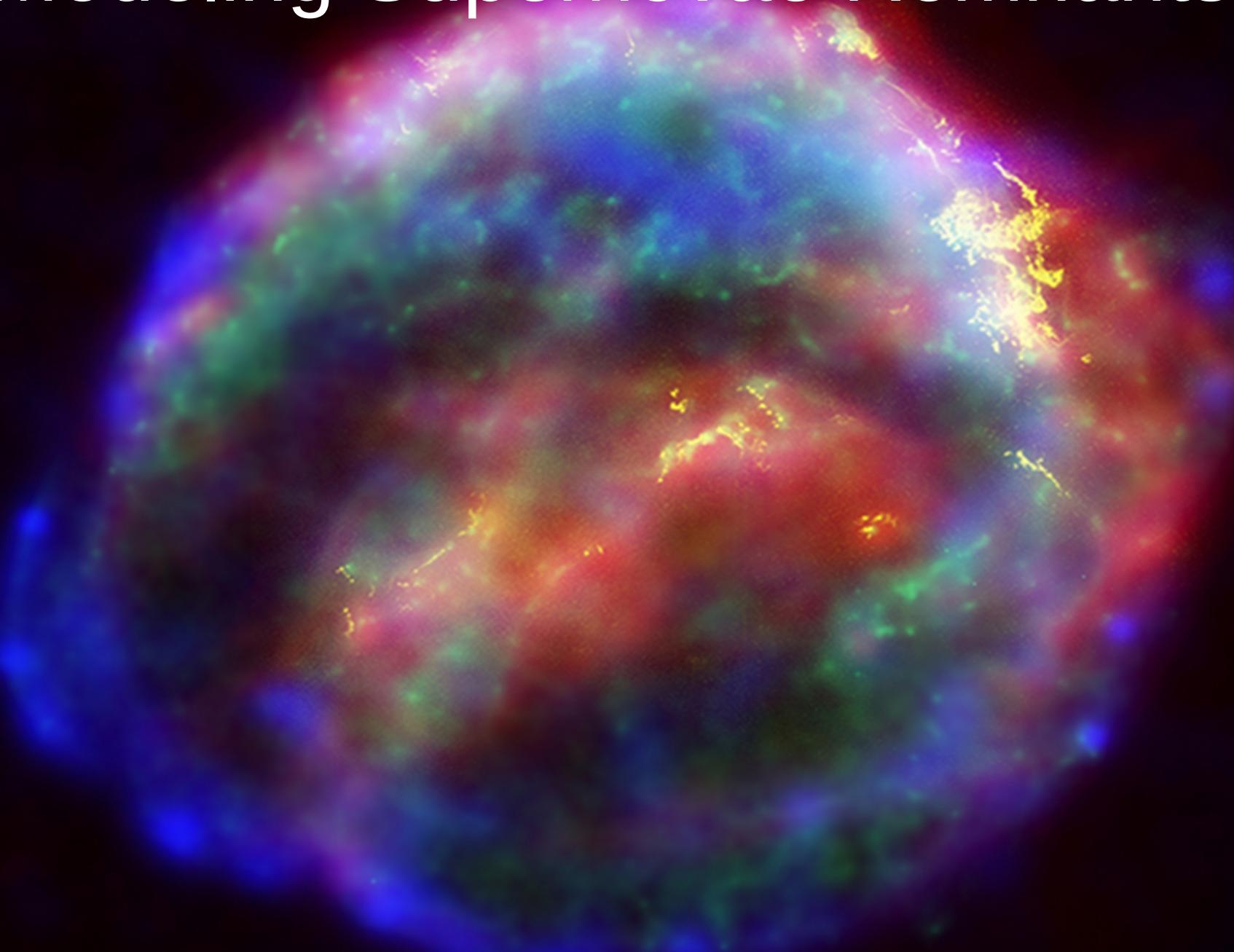
Illustration



Project

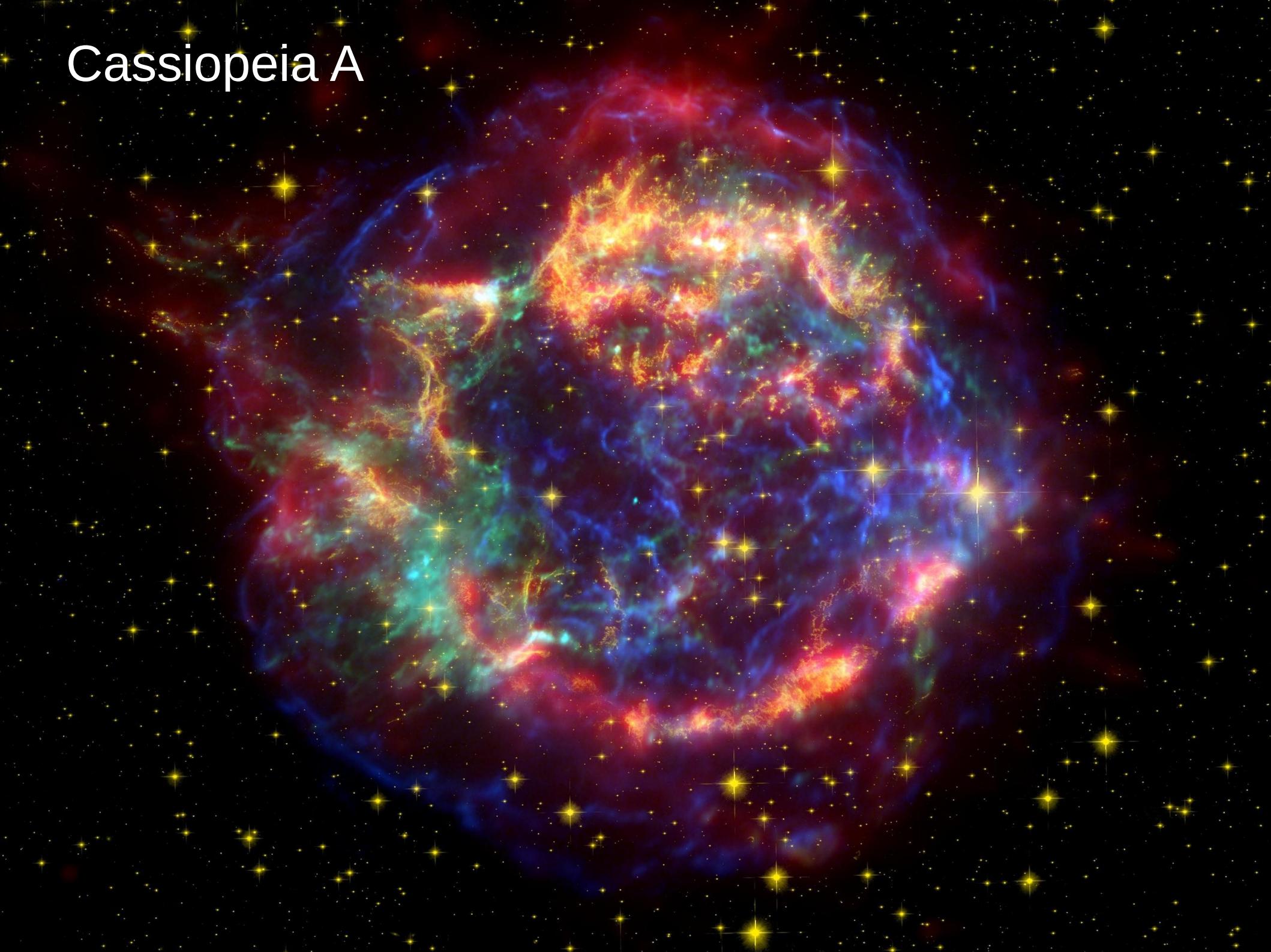
- For either our Solar System or Trappist-1:
 - Model the planetary orbits around the sun, accounting only for the sun's gravity
 - Study orbital perturbations arising from including gravitation from massive planets (e.g. Jupiter)
 - How would solar transits look for each planet from faraway?

Modeling Supernovae Remnants

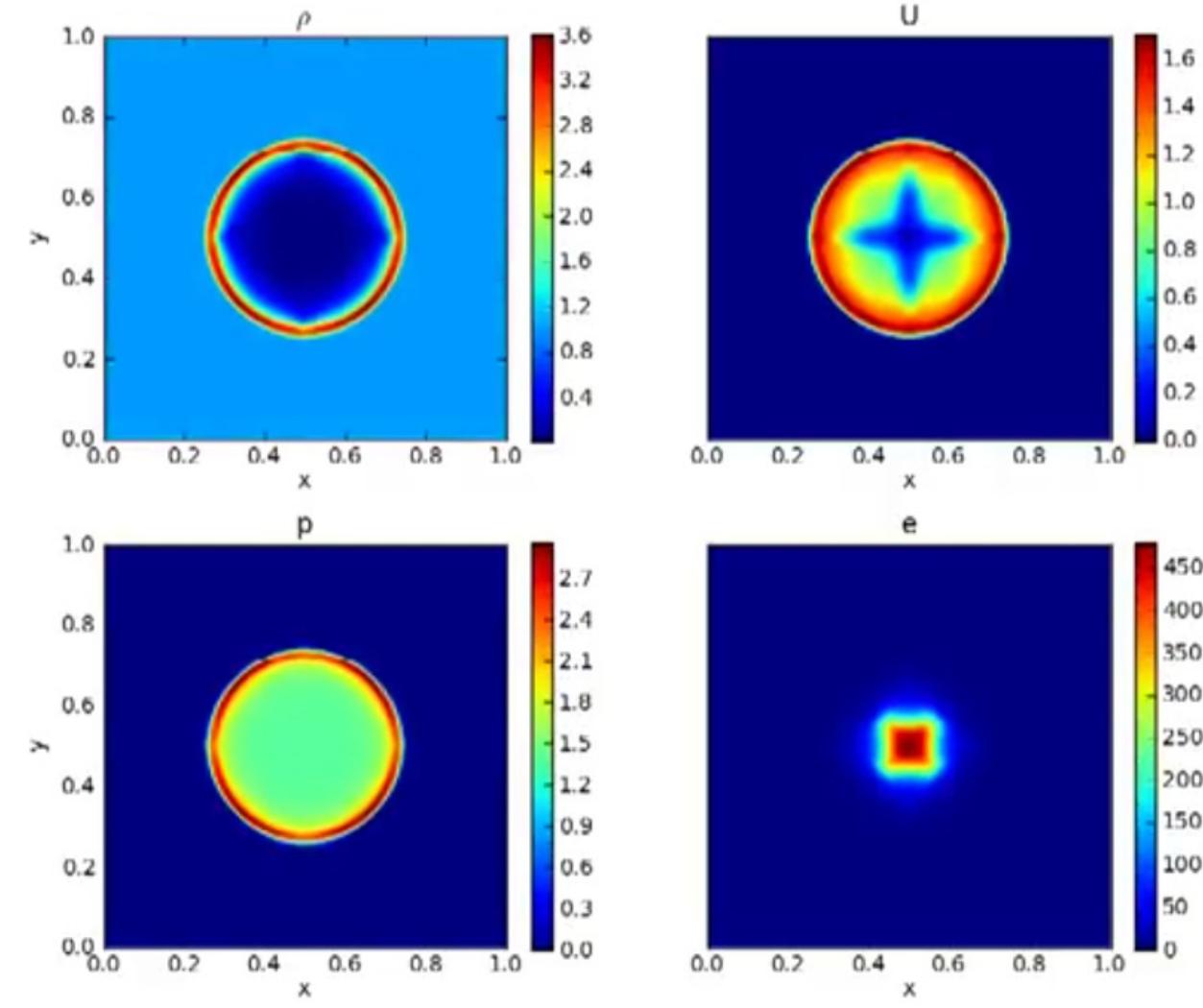


SN 1604 Kepler's SN

Cassiopeia A



Pyro – The Sedov Explosion

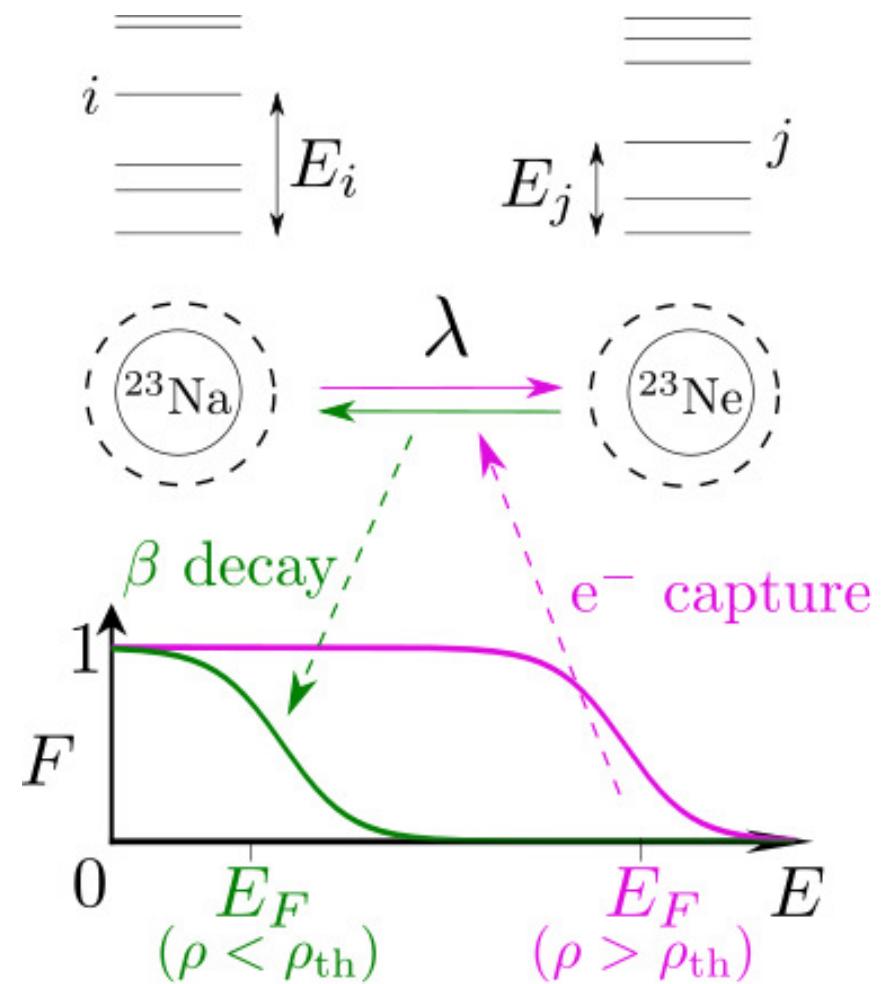
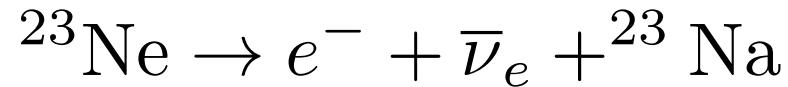
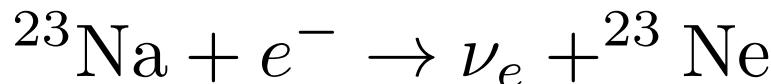
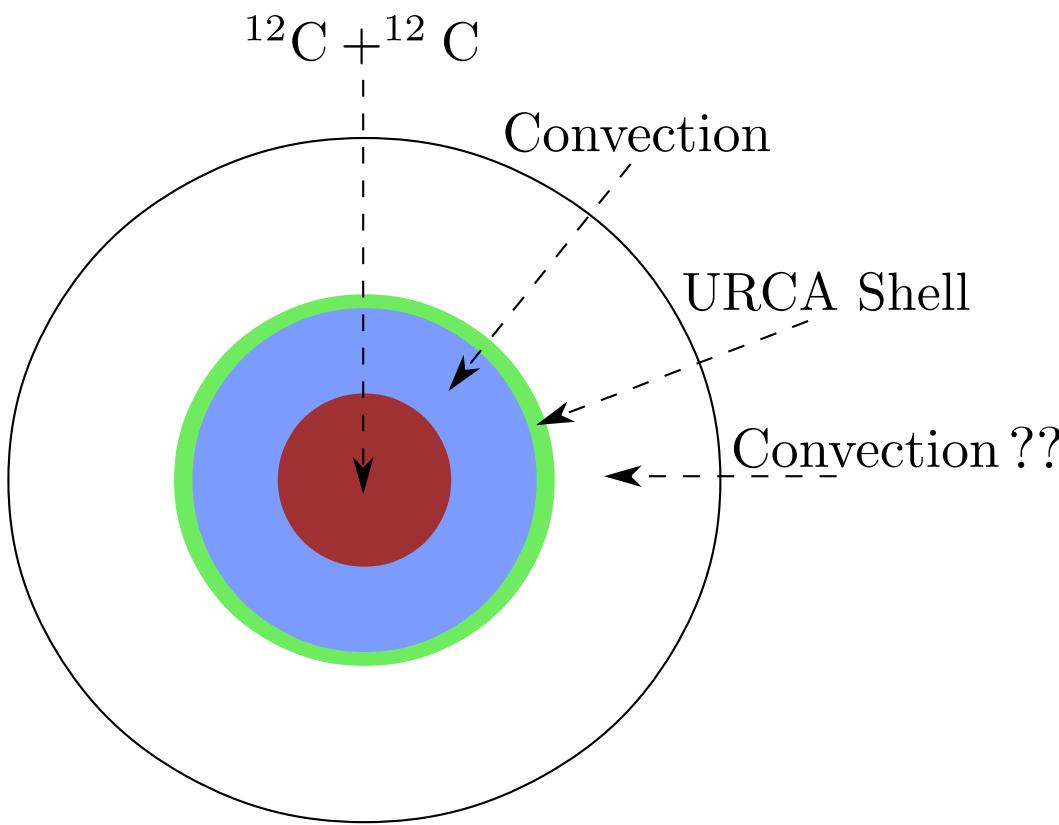


<https://zingale.github.io/pyro2/>

Project

- Assess which supernovae remnants we can model in the Sedov-Taylor adiabatic expansion phase.
- Experiment with Pyro using different blast initial conditions to determine how best to model various supernova remnants.
- Determine how shock speed and temperature depend on initial conditions.

Urca Process In White Dwarfs



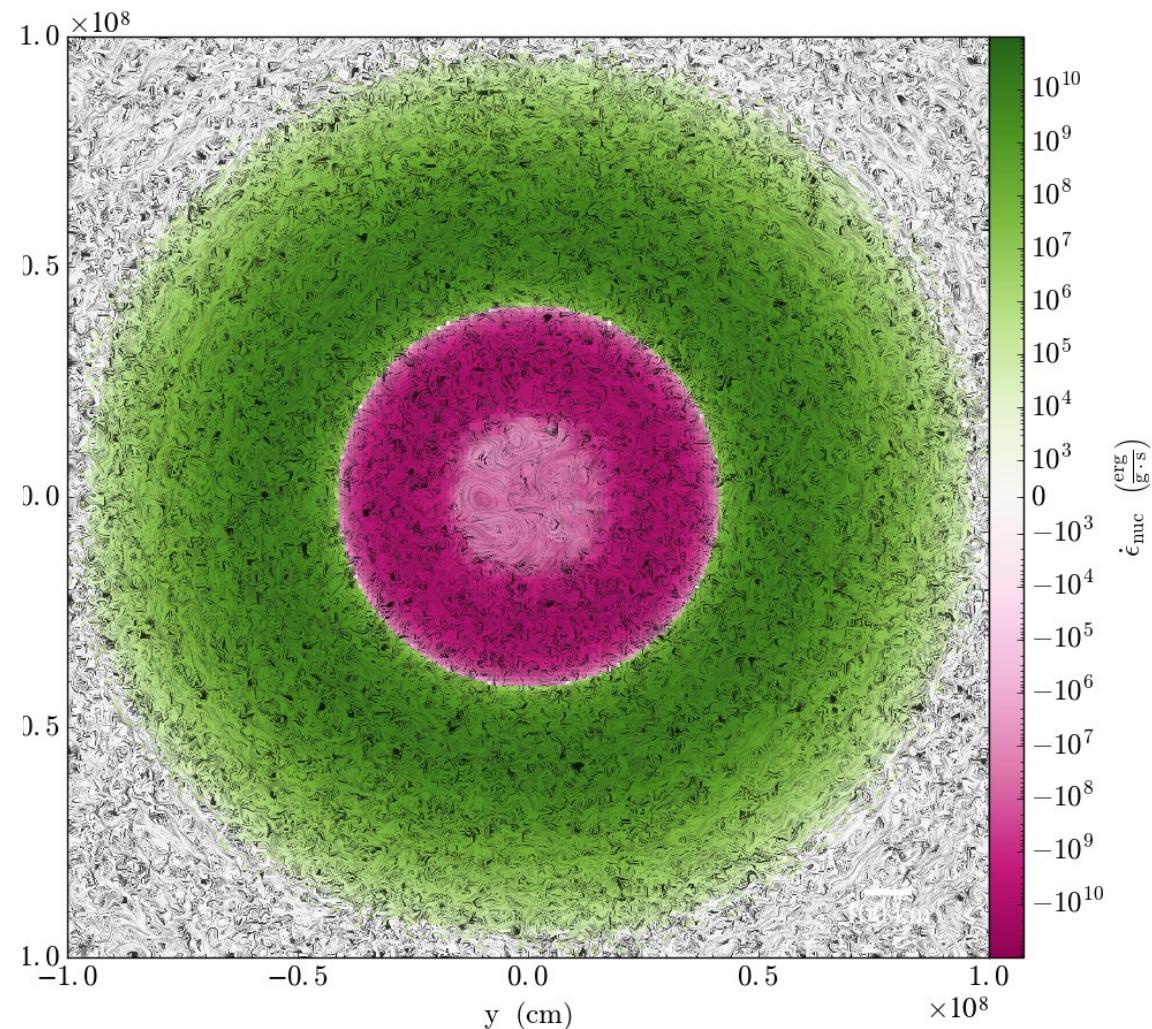
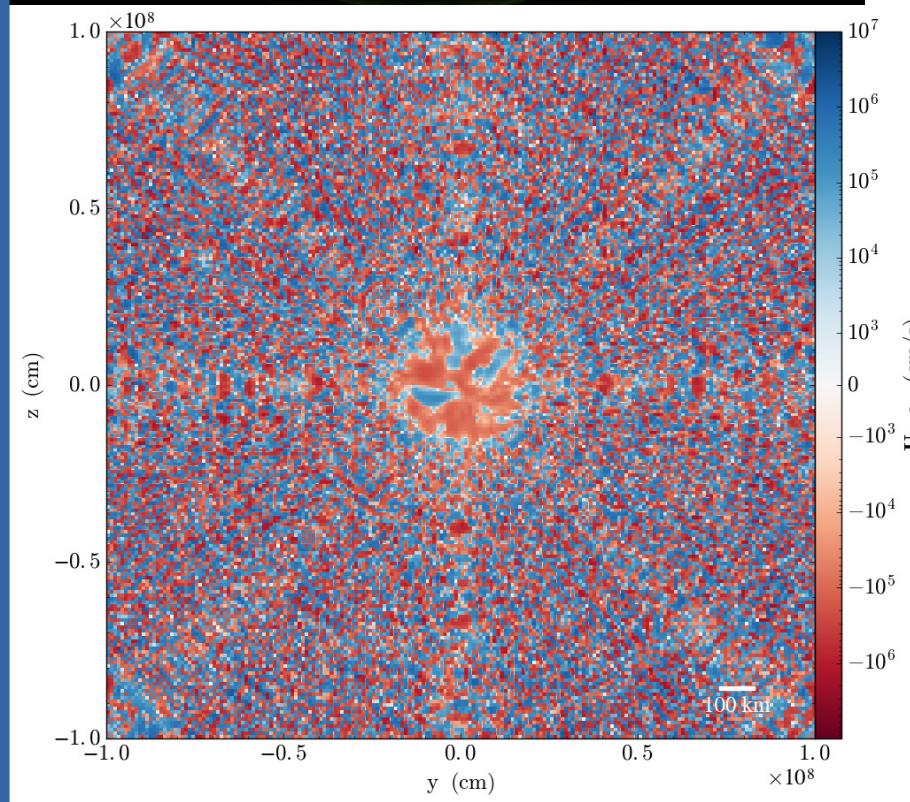
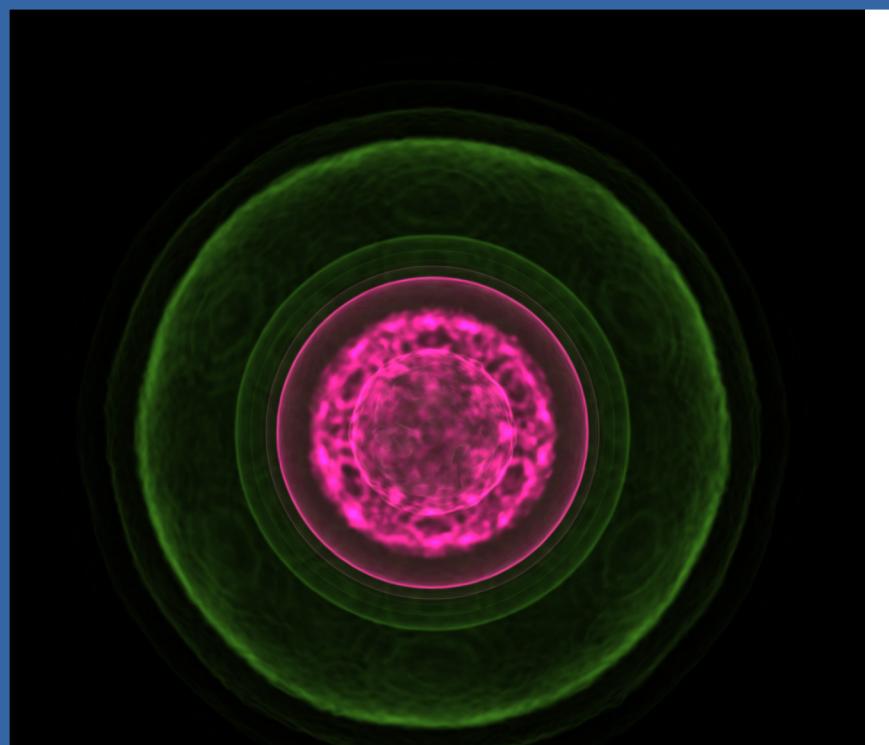
$$Q = 4.37590 \text{ MeV}$$

$$\log_{10}(\rho Y_e) = 8.92$$

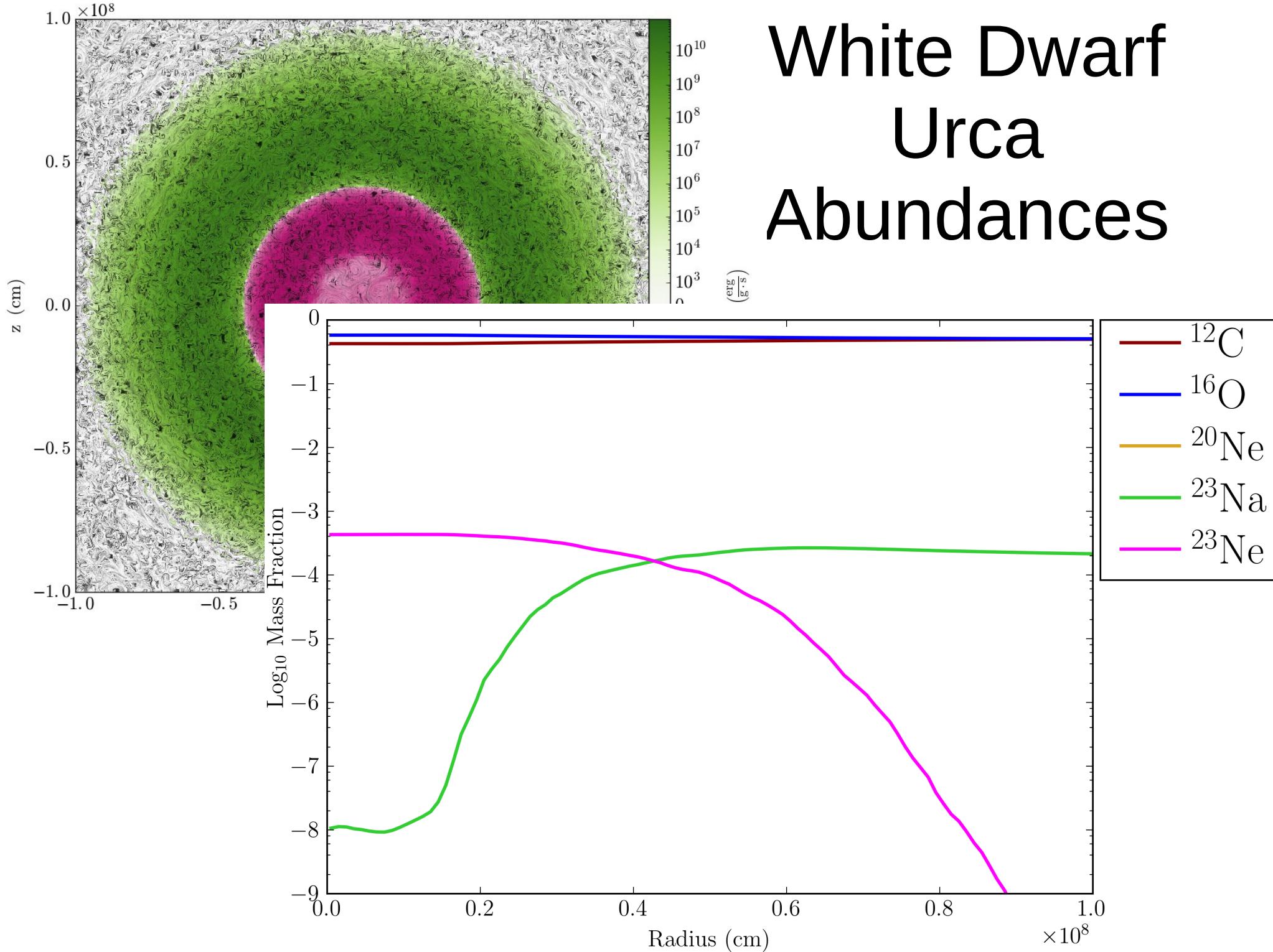
(Toki, Suzuki, Nomoto, Jones & Hirschi 2013)

White Dwarf Urca Process

$$\rho_c = 4.5 \times 10^9 \text{ g} \cdot \text{cm}^{-3}$$
$$T_c = 3.0 \times 10^8 \text{ K}$$



White Dwarf Urca Abundances



Project

- Use pyreaclib to construct a reaction network that links the nuclides that have a role in carbon fusion, electron capture, and beta decay.
- Determine the nucleosynthesis properties of the system and assess the dependence on:
 - Density
 - Temperature
 - Composition