# **Computer Vision**

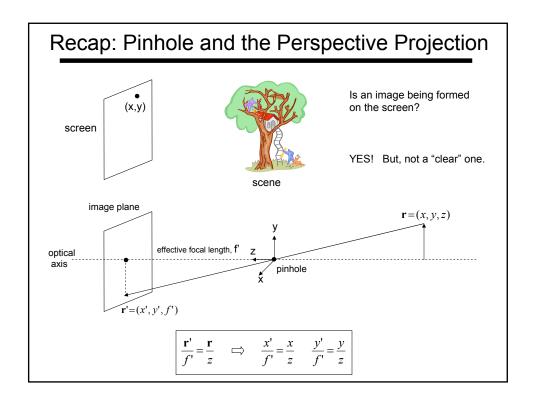
Spring 2006 15-385,-685

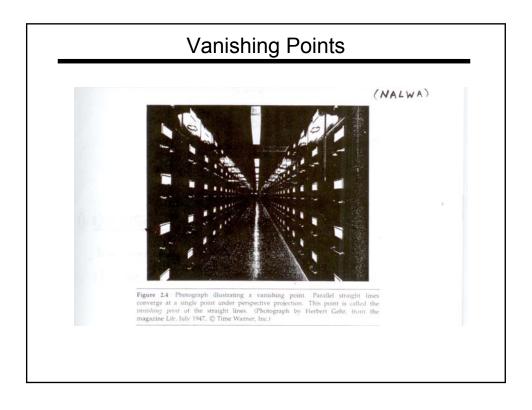
Instructor: S. Narasimhan

Wean 5403 T-R 3:00pm – 4:20pm

**Image Sensing** 

Lecture #3





# Vanishing Points

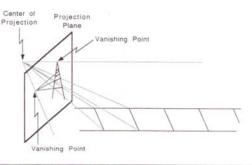


Figure 2.5 The vanishing point. The vanishing point of a straight line under perspective projection is that point on the projection surface at which the line would appear to "vanish" if the line were infinitely long in space. The location of the vanishing point of a straight line depends only on the orientation of the straight line in space, and not on the line's position: For any given spatial orientation, the vanishing point is located at that point on the projection surface where a straight line passing through the center of projection with the given orientation would intersect the projection surface.

# Pinhole Images





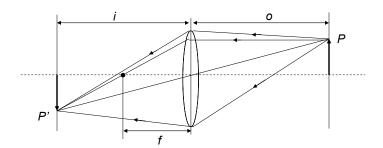
Exposure 4 seconds

Exposure 96 minutes

Images copyright © 2000 Zero Image Co.

# Image Formation using Lenses

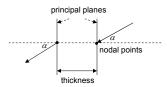
- · Lenses are used to avoid problems with pinholes.
- Ideal Lens: Same projection as pinhole but gathers more light!



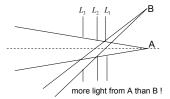
- Gaussian Thin Lens Formula:  $\frac{1}{i} + \frac{1}{o} = \frac{1}{f}$
- f is the focal length of the lens determines the lens's ability to refract light
- f different from the effective focal length f' discussed before!

#### Common Lens Related Issues - Summary

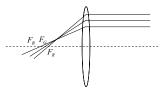
Compound (Thick) Lens



Vignetting



**Chromatic Abberation** 



Lens has different refractive indices for different wavelengths.

Radial and Tangential Distortion



image plane

#### Lens Glare

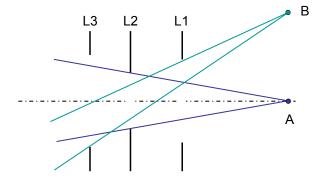




- Stray interreflections of light within the optical lens system.
- Happens when very bright sources are present in the scene.

Reading: http://www.dpreview.com

# Vignetting



More light passes through lens L3 for scene point A than scene point B

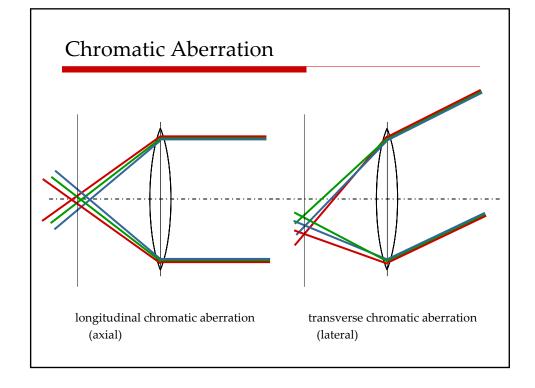
Results in spatially non-uniform brightness (in the periphery of the image)

# Vignetting





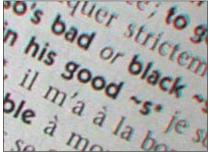
photo by Robert Johnes



#### **Chromatic Aberrations**

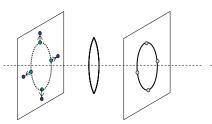


longitudinal chromatic aberration (axial)

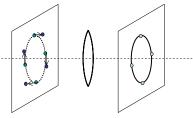


transverse chromatic aberration (lateral)

#### Geometric Lens Distortions



Radial distortion



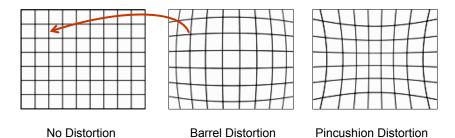
Tangential distortion



Photo by Helmut Dersch

Both due to lens imperfection Rectify with geometric camera calibration

#### **Radial Lens Distortions**



• Radial distance from Image Center:

$$r_u = r_d + k_I r_d^3$$

# **Correcting Radial Lens Distortions**







After

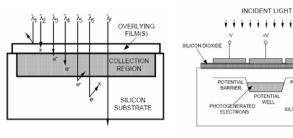
# Topics to be Covered

- Image Sensors
- Sensing Brightness
- Sensing Color
- Our Eyes

# Image Sensors Considerations Serial Register Considerations Speed Resolution Signal / Noise Ratio Cost Ftg. 4. Typical 512 × 512 CCD.



Convert light into electric charge



■CCD (charge coupled device)

Higher dynamic range

High uniformity Lower noise



■CMOS (complementary metal Oxide semiconductor)

Lower voltage

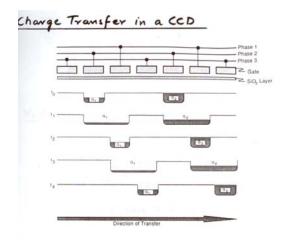
Higher speed

Lower system complexity



#### **Sensor Readout**

CCD Bucket Brigade



# Sensor Readout CCD Bucket Brigade To computer To computer To computer

Images Copyright © 2000 TWI Press, Inc.

CCD Performance Characteristics

Resolution:

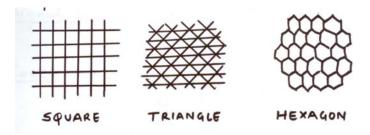
1k x 1k packed in 1-2 cm

No space between Pixels

No Photons wasted

#### **CCD Performance Characteristics**

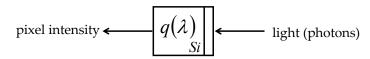
- Pixels must have same area
- Only 3 tessellations possible:



#### **CCD Performance Characteristics**

- Linearity Principle: Incoming photon flux vs. Output Signal
  - Sometimes cameras are made non-linear on purpose.
  - · Calibration must be done (using reflectance charts)---covered later
- Dark Current Noise: Non-zero output signal when incoming light is zero
- Sensitivity: Minimum detectable signal produced by camera

# **Sensing Brightness**



#### **Quantum Efficiency**

$$q(\lambda) = \frac{\text{generated electron flux}}{\text{photon flux of wavelength } \lambda}$$

Pixel intensity: I = k (generated electron flux)

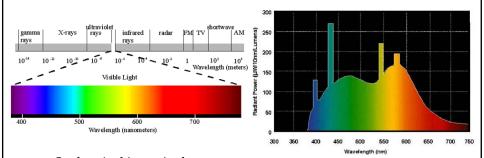
For monochromatic light  $(\lambda = \lambda_i)$  with flux  $P_i$ :

$$I = kq(\lambda_i)P_i$$

However, incoming light can vary in wavelength  $\lambda$ 

# **Sensing Brightness**

Incoming light has a **spectral distribution**  $p(\lambda)$ 



So the pixel intensity becomes

$$I = k \int_{-\infty}^{\infty} q(\lambda) p(\lambda) d\lambda$$

# **Sensing Color**

- Assume we have an image
- $I = k \int_{-\infty}^{\infty} q(\lambda) p(\lambda) d\lambda$
- $\square$  We know the pixel value I
- $\square$  We know our camera parameters  $k, q(\lambda)$

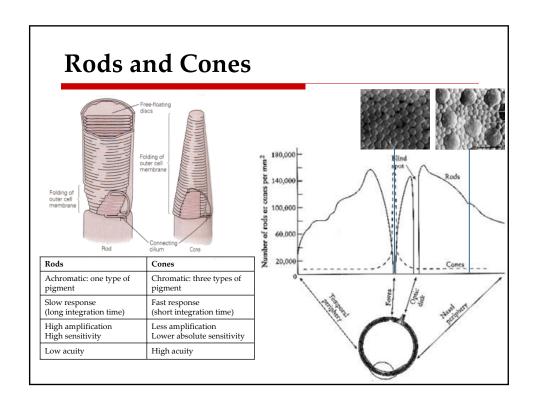
#### Can we tell the color of the scene?

(Can we recover the spectral distribution  $p(\lambda)$ )

$$I \longleftarrow q(\lambda)$$
Use a filter  $f_i(\lambda)$  Where  $f_i(\lambda) = \delta(\lambda - \lambda_i) = \begin{cases} 1 & \lambda = \lambda_i \\ 0 & \text{otherwise} \end{cases}$ 

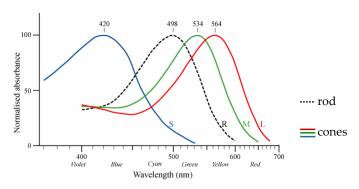
then

$$I = k \int_{-\infty}^{\infty} q(\lambda) p(\lambda) f(\lambda_i) d\lambda = kq(\lambda_i) p(\lambda_i)$$



#### How do we sense color?

Do we have infinite number of filters?

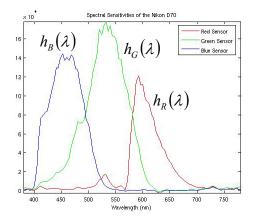


Three filters of different spectral responses

# **Sensing Color**

Tristimulus (trichromatic) values  $(I_R, I_G, I_B)$ 

Camera's spectral response functions:  $h_R(\lambda), h_G(\lambda), h_B(\lambda)$ 



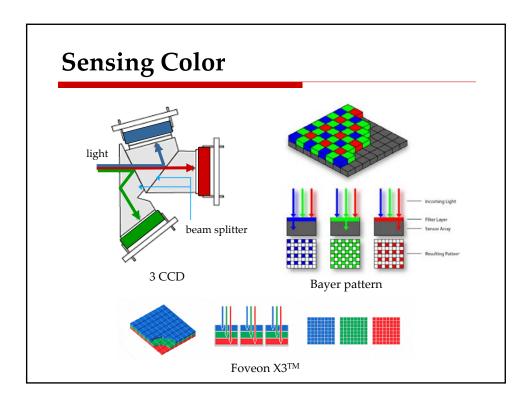
$$I_R = k \int_{-\infty}^{\infty} h_R(\lambda) p(\lambda) d\lambda$$

$$I_G = k \int_{-\infty}^{\infty} h_G(\lambda) p(\lambda) d\lambda$$

$$I_{R} = k \int_{-\infty}^{\infty} h_{R}(\lambda) p(\lambda) d\lambda$$

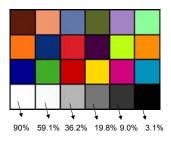
$$I_{G} = k \int_{-\infty}^{\infty} h_{G}(\lambda) p(\lambda) d\lambda$$

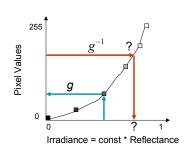
$$I_{B} = k \int_{-\infty}^{\infty} h_{B}(\lambda) p(\lambda) d\lambda$$



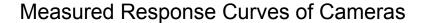
#### **Color Chart Calibration**

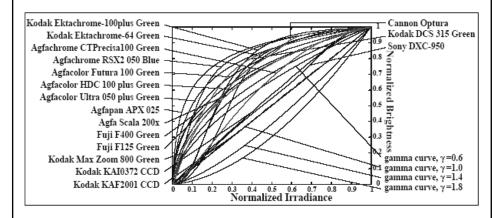
- Important preprocessing step for many vision and graphics algorithms
- Use a color chart with precisely known reflectances.





- Use more camera exposures to fill up the curve.
- Method assumes constant lighting on all patches and works best when source is far away (example sunlight).
- ullet Unique inverse exists because  $oldsymbol{\mathcal{G}}$  is monotonic and smooth for all cameras.

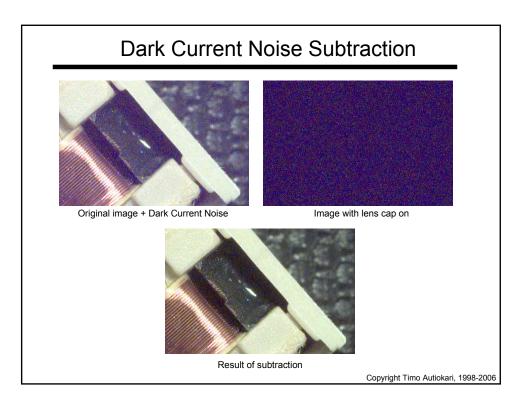


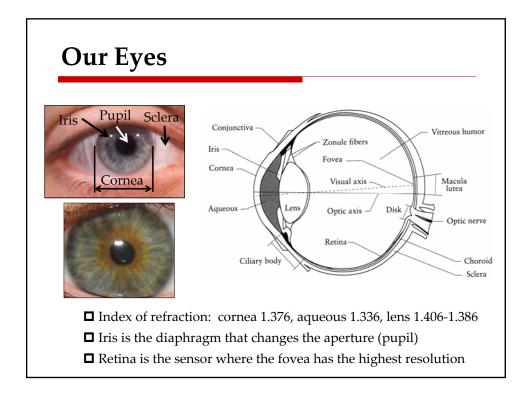


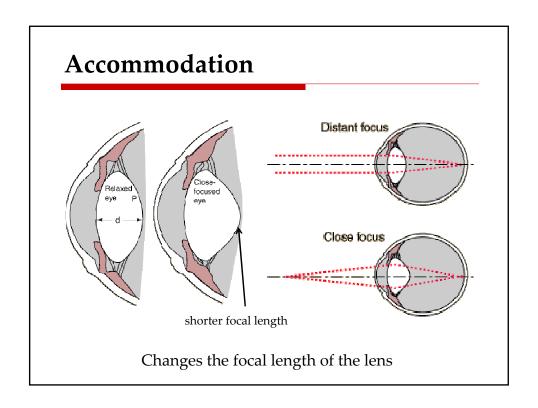
[Grossberg, Nayar]

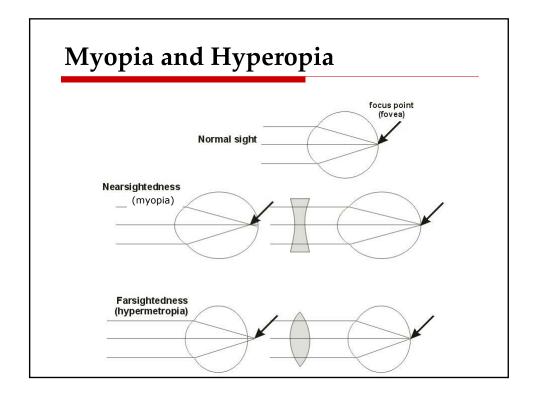
#### **Dark Current Noise Subtraction**

- Dark current noise is high for long exposure shots
- To remove (some) of it:
  - Calibrate the camera (make response linear)
  - Capture the image of the scene as usual
  - Cover the lens with the lens cap and take another picture
  - Subtract the second image from the first image

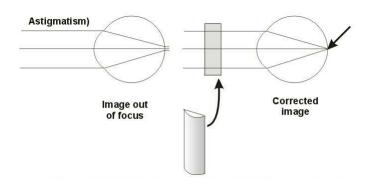








# **Astigmatism**

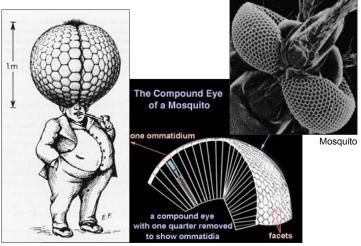


The cornea is distorted causing images to be un-focused on the retina.

# Blind Spot in Eye

Close your right eye and look directly at the "+"

# Eyes in Nature



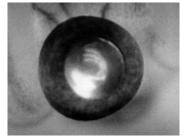
http://ebiomedia.com/gall/eyes/octopus-insect.html

Mosquitos have microscopic vision, but to focus at large distances their would need to be 1 m!

# Curved Mirrors in Scallop Eyes



Telescopic Eye



(by Mike Land, Sussex)

... More in the last part of the course

# Next Class

- Binary Image Processing
- Horn, Chapter 3